

AMATH 481/581 - Autumn 2022 Homework #4

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1. Presentation Skills: 2D plotting

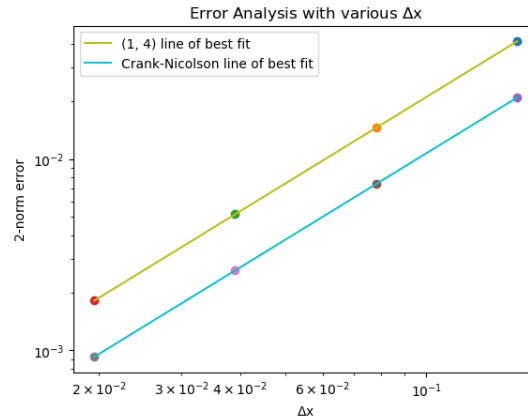


Figure 1: log-log plot of different Δx values against the two norm error for the (1, 4) accurate method against the (2, 2) accurate method (Crank-Nicolson). At every Δx value, the two norm error for the Crank-Nicolson is lower than the two norm error for the (1, 4) accurate method. This shows that the trade off between the temporal (1st order accurate vs 2nd order accurate) and spatial (2nd order accurate vs 4th order accurate) is a fine balance as they both have almost the same slope in the log-log plot (1.5).

Appendix: Code:

```
[language=Python]
import numpy as np
import scipy.sparse
import scipy.optimize
import matplotlib.pyplot as plt
alpha = 2
L = 10
Time = 2
n = 128
xspan, dx = np.linspace(-L, L, n, endpoint=False, retstep=True)
tspan, dt = np.linspace(0, Time, 501, retstep=True)
lambda_star = (alpha * dt) / (dx**2)
# %%
e1 = -30 * np.ones(n)
e2 = 16 * np.ones(n)
e3 = -1 * np.ones(n)
A = scipy.sparse.spdiags([e3, e2, e1, e2, e3],
                          [-2, -1, 0, 1, 2], n, n, format='csc')
A[0, -1] = 16
A[0, -2] = -1
A[1, -1] = -1
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A[-1, 0] = 16
A[-1, 1] = -1
A[-2, 0] = -1

A = 1/12 * A
A3 = A.toarray()

# Solve
sol1 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol1[:, 0] = u0
for i in range(len(tspan) - 1):
    u1 = u0 + lambda_star * (A@u0)
    sol1[:, i + 1] = u1
    u0 = u1
A5 = sol1[:, -1]
#reshape A5 to a 128x1 matrix
A5 = A5.reshape(128,1)

# %%
# Question 2
e1 = -2 * np.ones(n)
e2 = np.ones(n)
mid = 1/2 * lambda_star * scipy.sparse.spdiags([e2, e1, e2],
                                                [-1, 0, 1], n, n, format='csc')

mid[0, -1] = 1/2 * lambda_star
mid[-1, 0] = 1/2 * lambda_star
B = scipy.sparse.eye(n) - mid
C = scipy.sparse.eye(n) + mid
A7 = B.toarray()
A8 = C.toarray()
print(A8)
# %%
# Solve suing LU decomposition
sol2 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol2[:, 0] = u0
PLU = scipy.sparse.linalg.splu(B)
test = PLU.solve(C@u0)
for i in range(len(tspan) - 1):
    u1 = PLU.solve(C@u0)
    sol2[:, i + 1] = u1
    u0 = u1
A9 = sol2[:, -1]
# reshape A9 to a 128x1 matrix
A9 = A9.reshape(128,1)
print(A9)

A10 = 0

# %%
# Question 3
#Load the exact solution from the file exact_128.csv
exact = np.loadtxt('exact_128.csv', delimiter=',')
#reshape the exact solution to a 128x1 matrix
exact = exact.reshape(128,1)
#find the two norm difference between A5 and exact, save as A11
A11 = np.linalg.norm(A5 - exact)
#find the two norm difference between A9 and exact, save as A12
A12 = np.linalg.norm(A9 - exact)

# %%
# redo both problems with n = 256, load in the exact solution from the file exact_256.csv
n = 256
xspan, dx2 = np.linspace(-L, L, n, endpoint=False, retstep=True)

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tspan, dt = np.linspace(0, Time, (4 * 500) + 1, retstep=True)
e1 = -30 * np.ones(n)
e2 = 16 * np.ones(n)
e3 = -1 * np.ones(n)
A = scipy.sparse.spdiags([e3, e2, e1, e2, e3],
                          [-2, -1, 0, 1, 2], n, n, format='csc')

A[0, -1] = 16
A[0, -2] = -1
A[1, -1] = -1
A[-1, 0] = 16
A[-1, 1] = -1
A[-2, 0] = -1

A = 1/12 * A
print(A.toarray())
sol1 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol1[:, 0] = u0
for i in range(len(tspan) - 1):
    u1 = u0 + lambda_star * (A@u0)
    sol1[:, i + 1] = u1
    u0 = u1
first = sol1[:, -1]
# %%
e1 = -2 * np.ones(n)
e2 = np.ones(n)
mid = 1/2 * lambda_star * scipy.sparse.spdiags([e2, e1, e2],
                                                  [-1, 0, 1], n, n, format='csc')
#mid = scipy.sparse.lil_matrix(mid)
mid[0, -1] = 1/2 * lambda_star
mid[-1, 0] = 1/2 * lambda_star
B = scipy.sparse.eye(n) - mid
C = scipy.sparse.eye(n) + mid
# Solve suing LU decomposition
sol2 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol2[:, 0] = u0
PLU = scipy.sparse.linalg.splu(B)
for i in range(len(tspan) - 1):
    u1 = PLU.solve(C@u0)
    sol2[:, i + 1] = u1
    u0 = u1
second = sol2[:, -1]
#Load the exact solution from the file exact_256.csv
exact = np.loadtxt('exact_256.csv', delimiter=',')
#find the two norm difference between A5 and exact, save as A11
A13 = np.linalg.norm(first - exact)
#find the two norm difference between second and exact, save as A12
A14 = np.linalg.norm(second - exact)

# %%
# redo both problems with n = 512, load in the exact solution from the file exact_512.csv
n = 512
xspan, dx3 = np.linspace(-L, L, n, endpoint=False, retstep=True)
tspan, dt = np.linspace(0, Time, (16 * 500) + 1, retstep=True)
e1 = -30 * np.ones(n)
e2 = 16 * np.ones(n)
e3 = -1 * np.ones(n)
A = scipy.sparse.spdiags([e3, e2, e1, e2, e3],
                          [-2, -1, 0, 1, 2], n, n, format='csc')

A[0, -1] = 16
A[0, -2] = -1
A[1, -1] = -1
A[-1, 0] = 16

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A[-1, 1] = -1
A[-2, 0] = -1

A = 1/12 * A
print(A.toarray())
sol1 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol1[:, 0] = u0
for i in range(len(tspan) - 1):
    u1 = u0 + lambda_star * (A@u0)
    sol1[:, i + 1] = u1
    u0 = u1
first = sol1[:, -1]
# %%
e1 = -2 * np.ones(n)
e2 = np.ones(n)
mid = 1/2 * lambda_star * scipy.sparse.spdiags([e2, e1, e2],
                                                [-1, 0, 1], n, n, format='csc')

#mid = scipy.sparse.lil_matrix(mid)
mid[0, -1] = 1/2 * lambda_star
mid[-1, 0] = 1/2 * lambda_star
B = scipy.sparse.eye(n) - mid
C = scipy.sparse.eye(n) + mid
# Solve using LU decomposition
sol2 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol2[:, 0] = u0
PLU = scipy.sparse.linalg.splu(B)
for i in range(len(tspan) - 1):
    u1 = PLU.solve(C@u0)
    sol2[:, i + 1] = u1
    u0 = u1
second = sol2[:, -1]
#Load the exact solution from the file exact_512.csv
exact = np.loadtxt('exact_512.csv', delimiter=',')
#find the two norm difference between A5 and exact, save as A11
A15 = np.linalg.norm(first - exact)
#find the two norm difference between second and exact, save as A12
A16 = np.linalg.norm(second - exact)

# %%
# redo both problems with n = 1024, load in the exact solution from the file exact_1024.csv
n = 1024
xspan, dx4 = np.linspace(-L, L, n, endpoint=False, retstep=True)
tspan, dt = np.linspace(0, Time, (64 * 500) + 1, retstep=True)
e1 = -30 * np.ones(n)
e2 = 16 * np.ones(n)
e3 = -1 * np.ones(n)
A = scipy.sparse.spdiags([e3, e2, e1, e2, e3],
                        [-2, -1, 0, 1, 2], n, n, format='csc')

A[0, -1] = 16
A[0, -2] = -1
A[1, -1] = -1
A[-1, 0] = 16
A[-1, 1] = -1
A[-2, 0] = -1

A = 1/12 * A
print(A.toarray())
sol1 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol1[:, 0] = u0
for i in range(len(tspan) - 1):
    u1 = u0 + lambda_star * (A@u0)
    sol1[:, i + 1] = u1

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u0 = u1
first = sol1[:, -1]
# %%
e1 = -2 * np.ones(n)
e2 = np.ones(n)
mid = 1/2 * lambda_star * scipy.sparse.spdiags([e2, e1, e2],
                                                [-1, 0, 1], n, n, format='csc')

#mid = scipy.sparse.lil_matrix(mid)
mid[0, -1] = 1/2 * lambda_star
mid[-1, 0] = 1/2 * lambda_star
B = scipy.sparse.eye(n) - mid
C = scipy.sparse.eye(n) + mid
# Solve using LU decomposition
sol2 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol2[:, 0] = u0
PLU = scipy.sparse.linalg.splu(B)
for i in range(len(tspan) - 1):
    u1 = PLU.solve(C@u0)
    sol2[:, i + 1] = u1
    u0 = u1
second = sol2[:, -1]
#Load the exact solution from the file exact_1024.csv
exact = np.loadtxt('exact_1024.csv', delimiter=',')
#find the two norm difference between A5 and exact, save as A11
A17 = np.linalg.norm(first - exact)
#find the two norm difference between second and exact, save as A12
A18 = np.linalg.norm(second - exact)

# %%
#plot the norms of dx, dx2, dx3, and dx4 against A11, A13, A15, A17 on a loglog plot

fig, ax = plt.subplots()
ax.loglog(dx, A11, 'o')
ax.loglog(dx2, A13, 'o')
ax.loglog(dx3, A15, 'o')
ax.loglog(dx4, A17, 'o')
ax.loglog(dx, A12, 'o')
ax.loglog(dx2, A14, 'o')
ax.loglog(dx3, A16, 'o')
ax.loglog(dx4, A18, 'o')
ax.set_xlabel('$\Delta x$')
ax.set_ylabel('2-norm error')
#set the title of the plot to be '(1, 4) method error analysis with various dx'
ax.set_title('Error Analysis with various $\Delta x$')
#make a line of best fit for the data
m, b = np.polyfit(np.log([dx, dx2, dx3, dx4]), np.log([A11, A13, A15, A17]), 1)
ax.loglog([dx, dx2, dx3, dx4], np.exp(m * np.log([dx, dx2, dx3, dx4]) + b), label='(1, 4) line of best fit')
m, b = np.polyfit(np.log([dx, dx2, dx3, dx4]), np.log([A12, A14, A16, A18]), 1)
ax.loglog([dx, dx2, dx3, dx4], np.exp(m * np.log([dx, dx2, dx3, dx4]) + b), label='Crank-Nicolson line of best fit')
ax.legend()
plt.show()

plt.savefig('error_analysis.png')

```

2. Presentation Skills: 3D plotting

Crank-Nicolson method solution

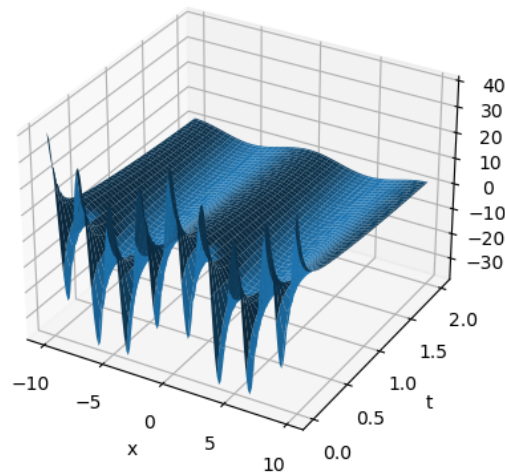


Figure 2: 3d plot of the entire final solution with $n = 128$ for the most accurate method (Crank-Nicolson). The plot shows that near $t = 0$, the solution is much more chaotic and fluctuates wildly. In contrast, as time progresses, the solution becomes much more smooth and the heat becomes more normalized. This signifies the heat is spreading out and becoming more uniform through the rod.

Appendix: Code:

```
[language=Python]
# %%
import numpy as np
import scipy.sparse
import scipy.optimize
import matplotlib.pyplot as plt
# %%
# Question 1
alpha = 2
L = 10
Time = 2
n = 128
xspan, dx = np.linspace(-L, L, n, endpoint=False, retstep=True)
tspan, dt = np.linspace(0, Time, 501, retstep=True)
lambda_star = (alpha * dt) / (dx**2)
# %%
e1 = -30 * np.ones(n)
e2 = 16 * np.ones(n)
e3 = -1 * np.ones(n)
A = scipy.sparse.spdiags([e3, e2, e1, e2, e3],
                          [-2, -1, 0, 1, 2], n, n, format='csc')

A[0, -1] = 16
A[0, -2] = -1
A[1, -1] = -1
A[-1, 0] = 16
A[-1, 1] = -1
A[-2, 0] = -1

A = 1/12 * A
```

```

A3 = A.toarray()

# Solve
sol1 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol1[:, 0] = u0
for i in range(len(tspan) - 1):
    u1 = u0 + lambda_star * (A@u0)
    sol1[:, i + 1] = u1
    u0 = u1
A5 = sol1[:, -1]
#reshape A5 to a 128x1 matrix
A5 = A5.reshape(128,1)
# %%
# stability analysis
g = lambda z: 1 + 1/12 * lambda_star * (-30 + 32 * np.cos(z) - 2 * np.cos(2 * z))
A1 = abs(g(1))
minimum_index = scipy.optimize.fminbound(lambda z: -abs(g(z)), -np.pi, np.pi)
A2 = g(minimum_index)

# %%
# Question 2
e1 = -2 * np.ones(n)
e2 = np.ones(n)
mid = 1/2 * lambda_star * scipy.sparse.spdiags([e2, e1, e2],
                                                [-1, 0, 1], n, n, format='csc')

mid[0, -1] = 1/2 * lambda_star
mid[-1, 0] = 1/2 * lambda_star
B = scipy.sparse.eye(n) - mid
C = scipy.sparse.eye(n) + mid
A7 = B.toarray()
A8 = C.toarray()
print(A8)
# %%
# Solve suing LU decomposition
sol2 = np.zeros((len(xspan), len(tspan)))
u0 = 10 * np.cos(2 * np.pi * xspan / L) + 30 * np.cos(8 * np.pi * xspan / L)
sol2[:, 0] = u0
PLU = scipy.sparse.linalg.splu(B)
test = PLU.solve(C@u0)
for i in range(len(tspan) - 1):
    u1 = PLU.solve(C@u0)
    sol2[:, i + 1] = u1
    u0 = u1
A9 = sol2[:, -1]
# reshape A9 to a 128x1 matrix
A9 = A9.reshape(128,1)
print(A9)

A10 = 0
# %%
# make a 3d plot of the solution of sol2
fig = plt.figure()
ax = fig.add_subplot(111, projection='3d')
X, T = np.meshgrid(xspan, tspan)
ax.plot_surface(X, T, sol2.T)
ax.set_xlabel('x')
ax.set_ylabel('t')
ax.set_zlabel('u')
ax.set_title('Crank-Nicolson method solution')
plt.show()

```