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NWU Pulsar Timing Workshop
25 Sept 2023



**Pulsars: What they are and
how to find them**

Part 1: WHAT ARE PULSARS?

Default (GPT-3.5)



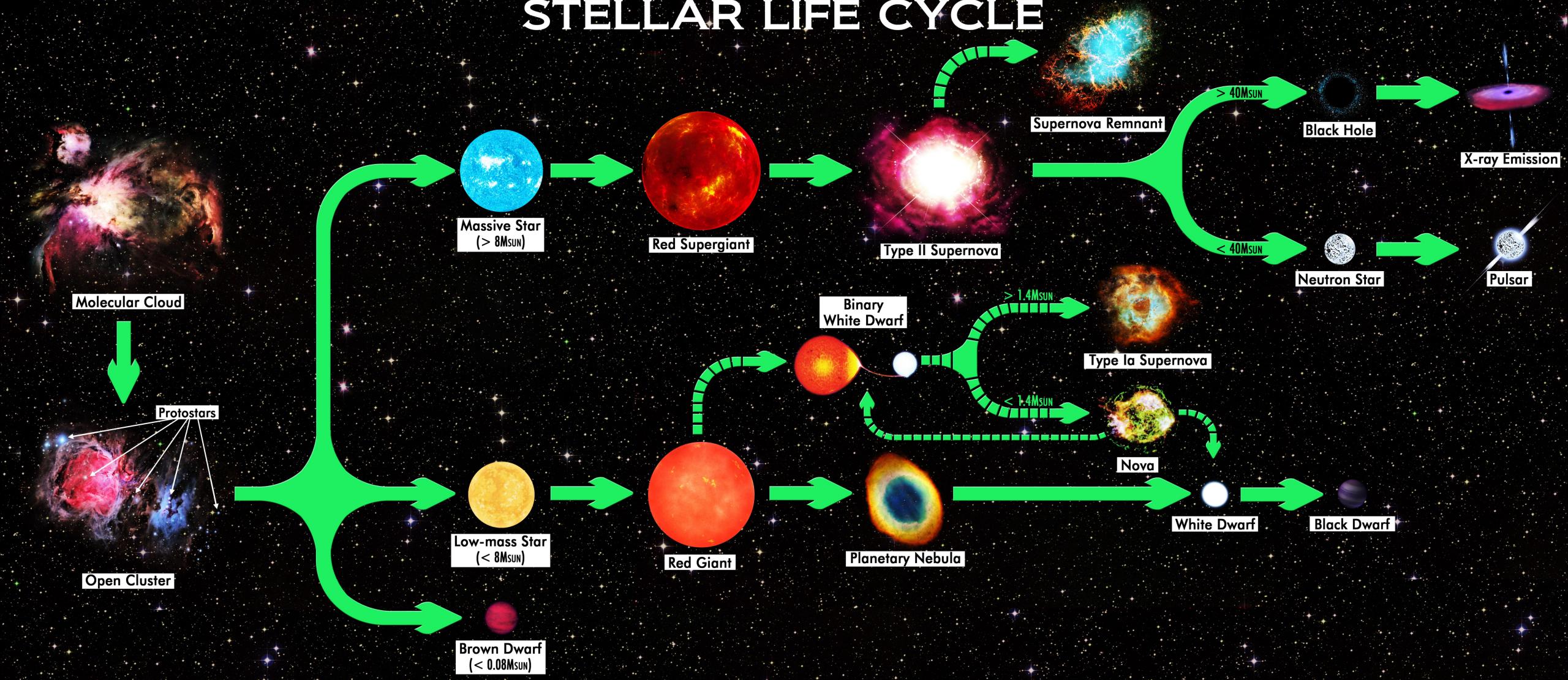
Explain to me what a pulsar is in a single sentence.



A pulsar is a highly magnetized, rapidly rotating neutron star that emits beams of electromagnetic radiation, including radio waves, with a regular and precise pulsing pattern.

**“Highly magnetised, rapidly
rotating neutron stars...”**

STELLAR LIFE CYCLE



Birth

Main Sequence

Old Age

Death

Remnant

Neutron Star matter

$1.4 M_{\text{sun}}$

20 km diameter

$\therefore \sim 10^{17} \text{ kg/m}^3$



1 tbsp. of Earth = 0.02 kg

1 tbsp. of the Sun = 2 kg

1 tbsp. of a NS = 900 billion kg

**“Highly magnetised, rapidly
rotating neutron stars...”**

Rapidly rotating

$$\Omega_f = \Omega_i \left(\frac{R_i}{R_f} \right)^2$$

$R_i \sim 10^6$ km

$R_f \sim 10$ km

\therefore factor 10^{10} speed-up

Spin periods = 1 ms - 76 s

Highly magnetic

$R_i \sim 10^6$ km

$R_f \sim 10$ km

\therefore factor 10^{10} amplification

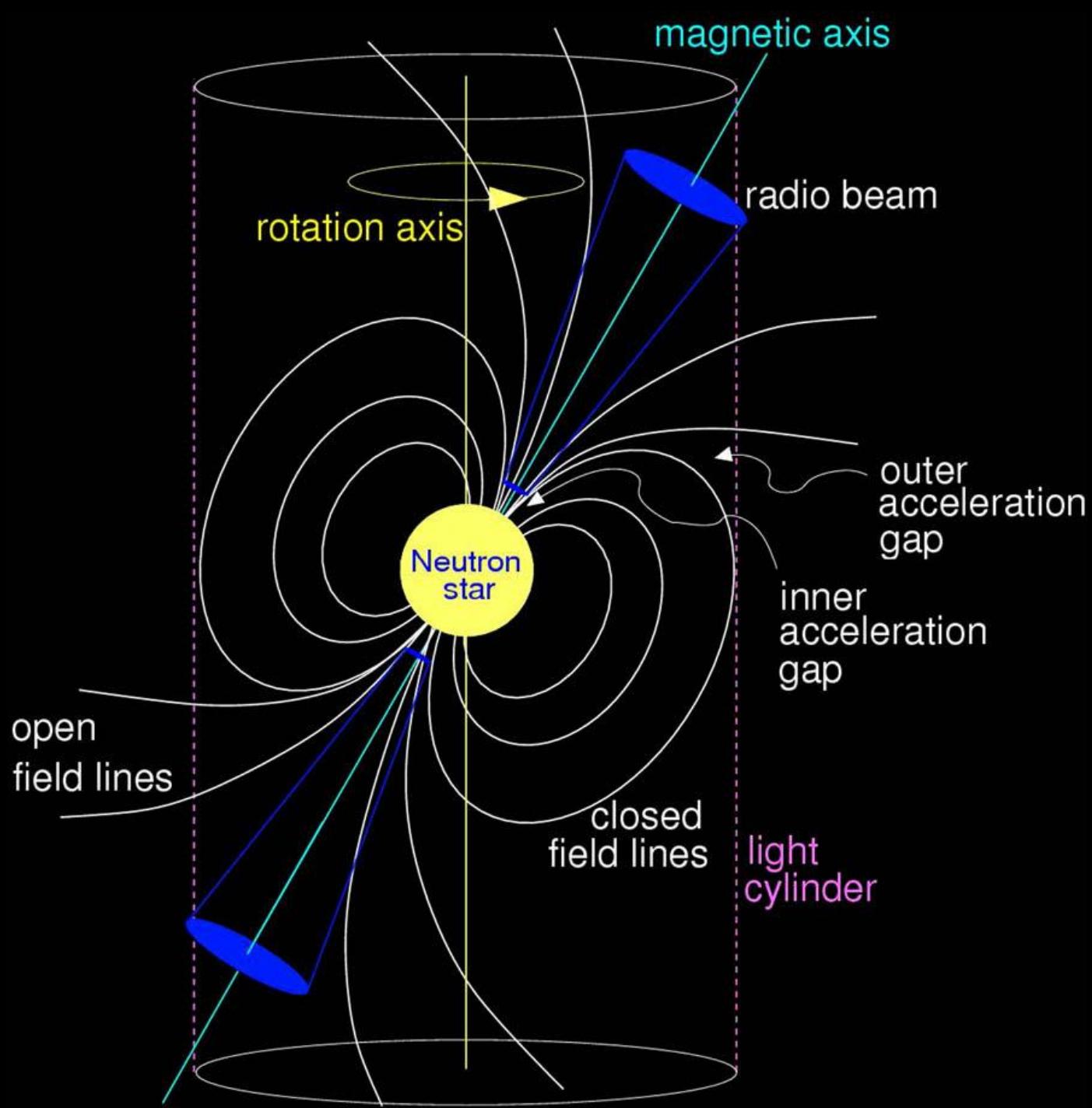
$$B_f = B_i \left(\frac{R_i}{R_f} \right)^2$$

Earth: $\sim 10^{-4}$ T

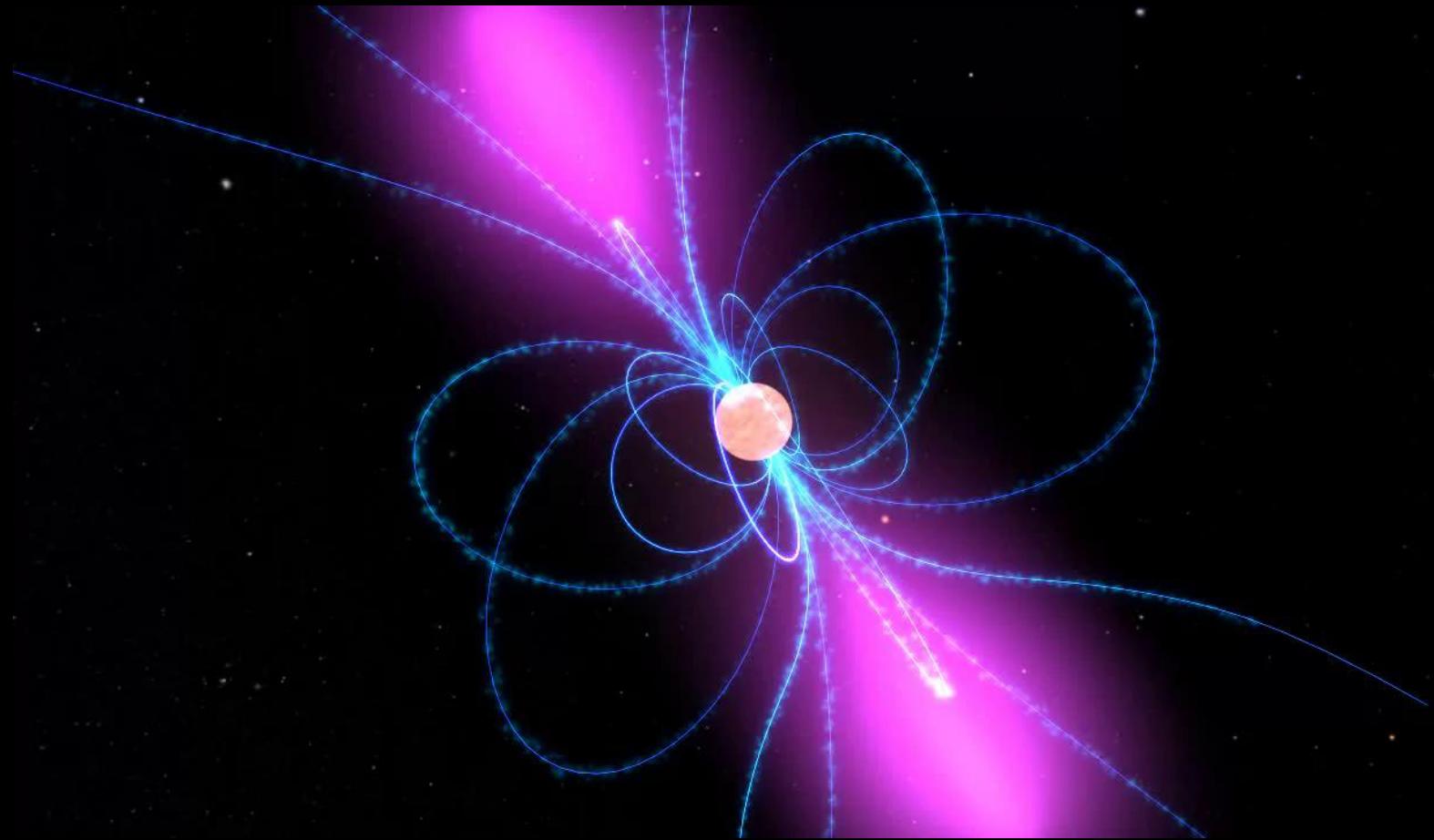
Strongest manmade: 45 T

Strongest NS: 1.6×10^9 T

**“...neutron stars that emit
beams of electromagnetic
radiation...”**

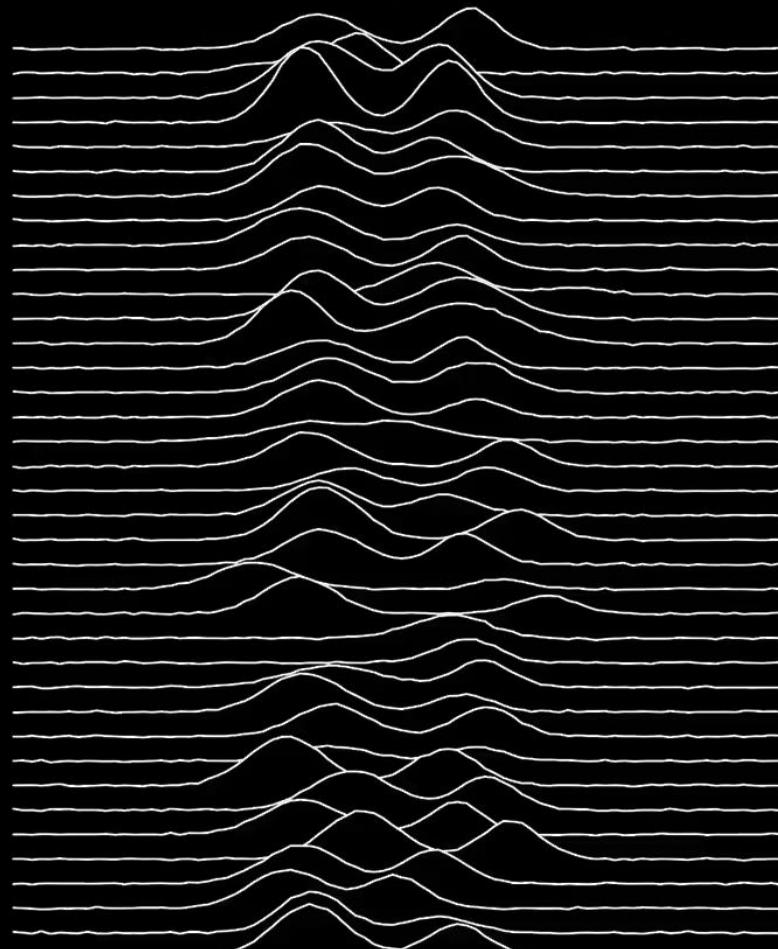


“...with a regular and precise pulsing pattern.”



Single Pulses

Pulsar B1919+21

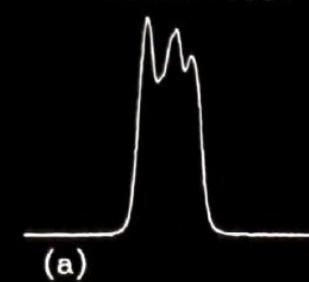


15th June 2019

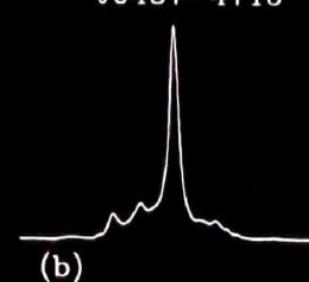
Jodrell Bank Observatory
The University of Manchester

Integrated Pulses

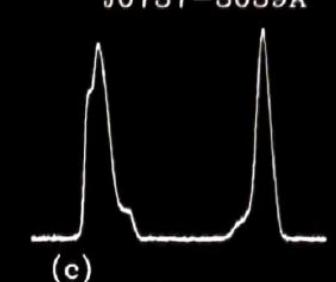
J0407+1607



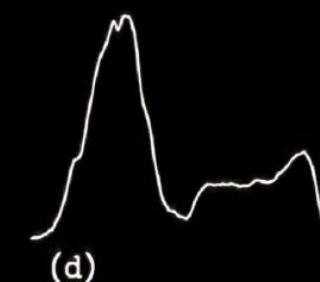
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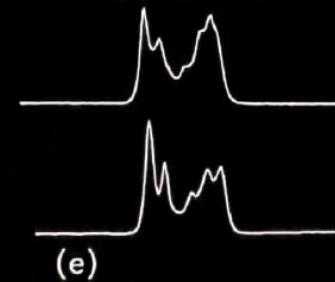
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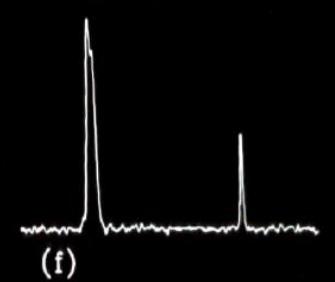
B0826-34



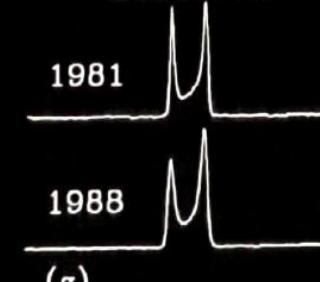
B1237+25



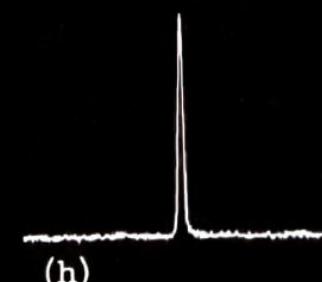
B1702-19



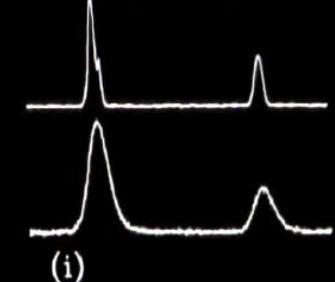
B1913+16



B1933+16



B1937+21

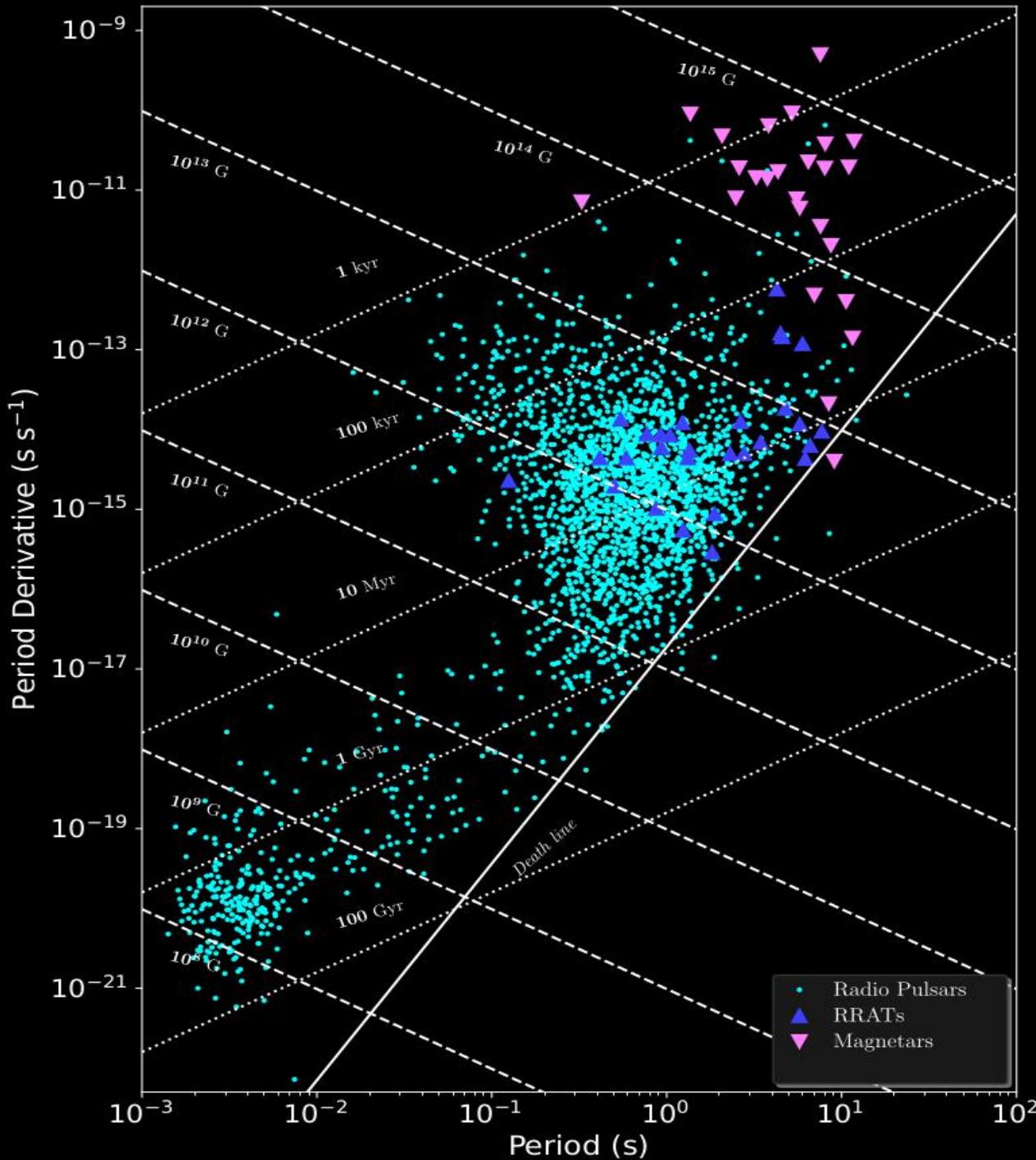


$$E_{\text{dot}} = -\frac{4\pi^2}{P^2} I \dot{P}$$

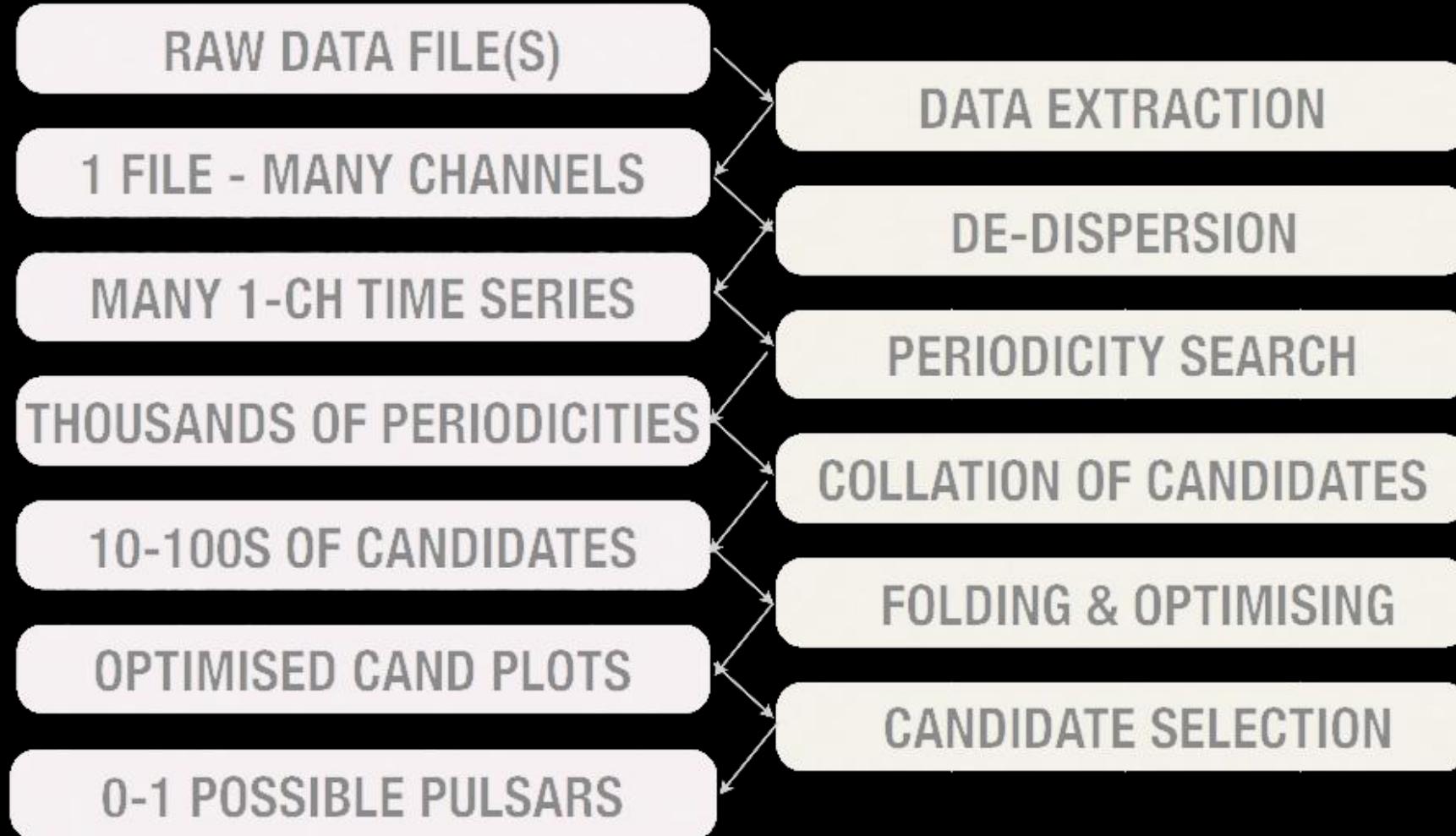
$$\tau_c = \frac{P}{2\dot{P}}$$

$$B_s = 3.2 \times 10^{19} \sqrt{P \dot{P}} \text{ G}$$

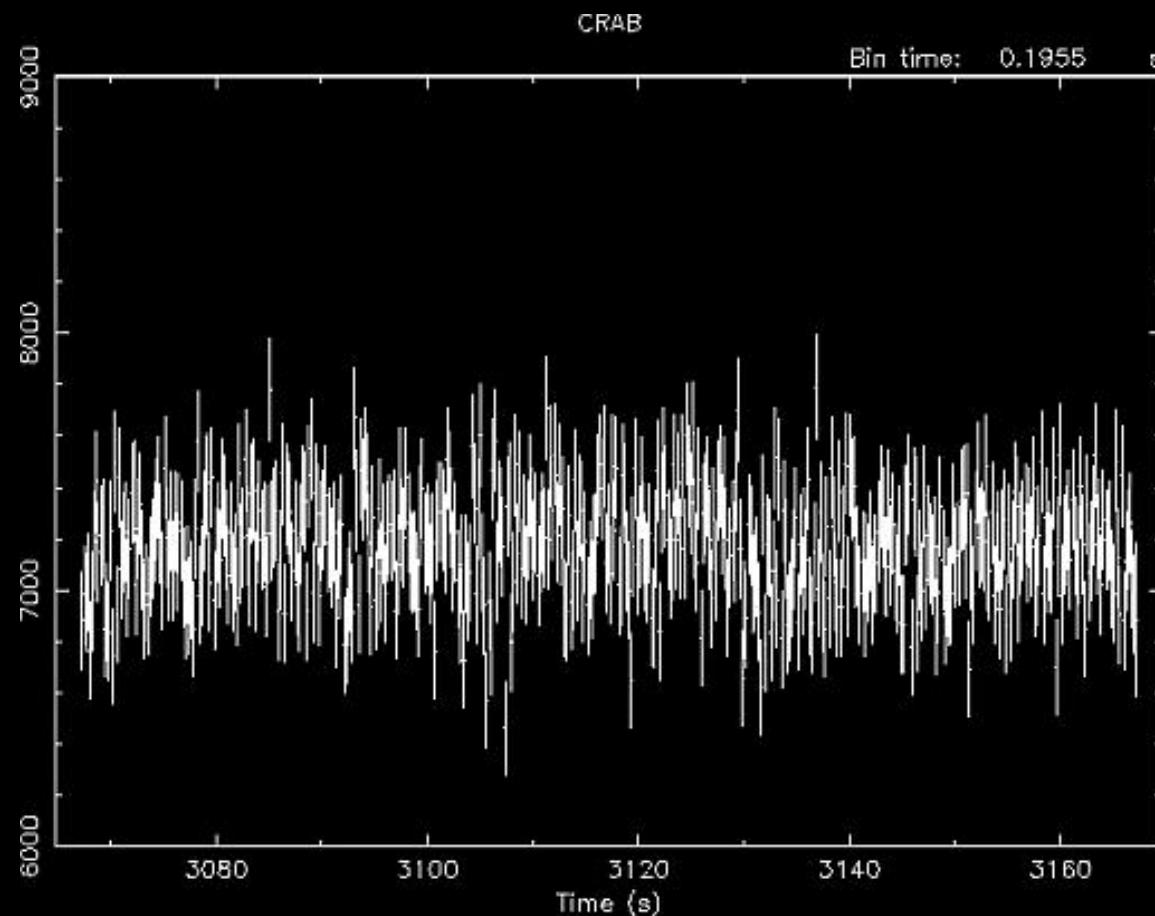
Numbers so far:
3000 radio pulsars
120 RRATs
30 magnetars



Part 2: HOW DO WE FIND THEM?



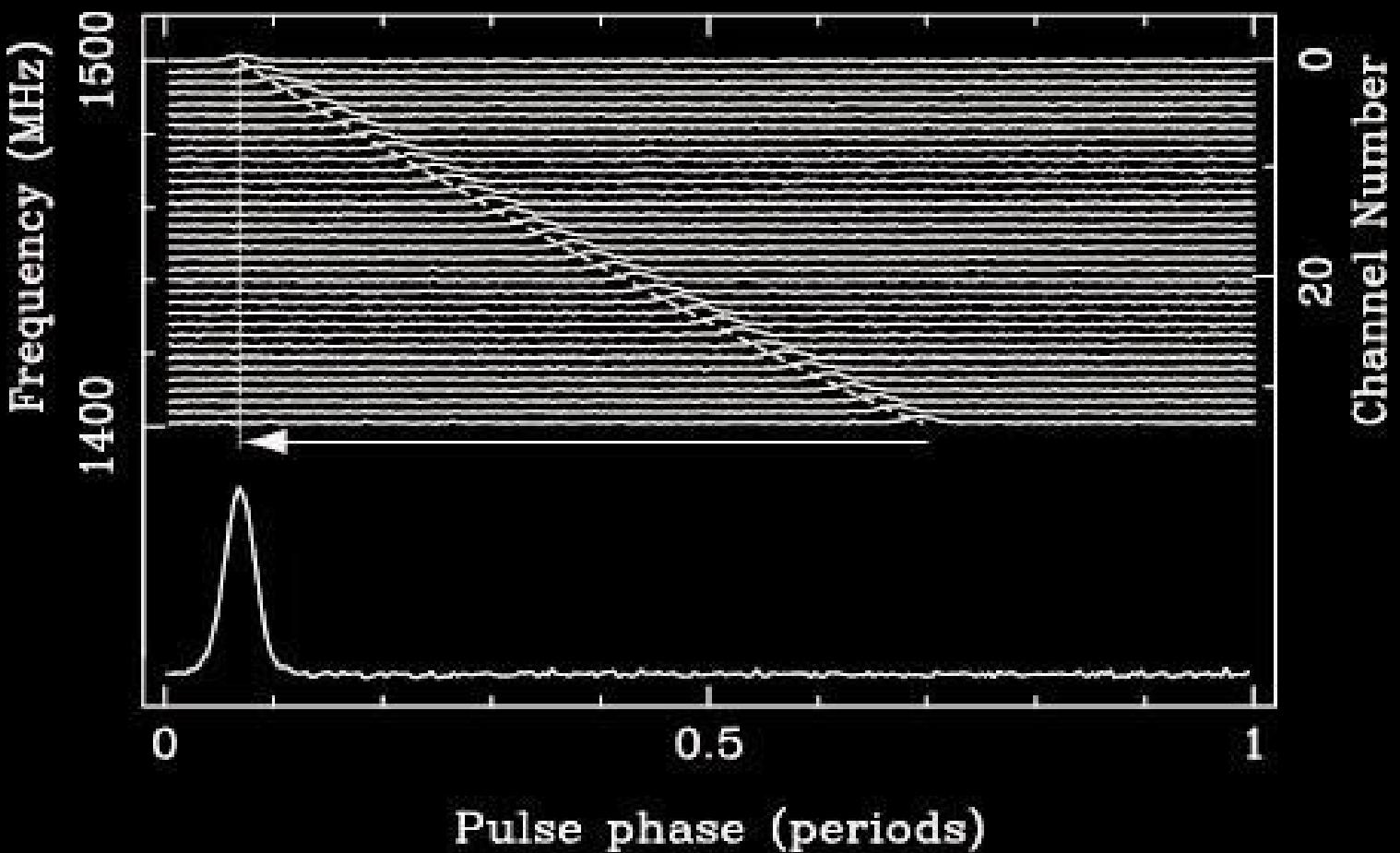
Periodicity Searches: Picking the signal out of the noise



De-dispersion

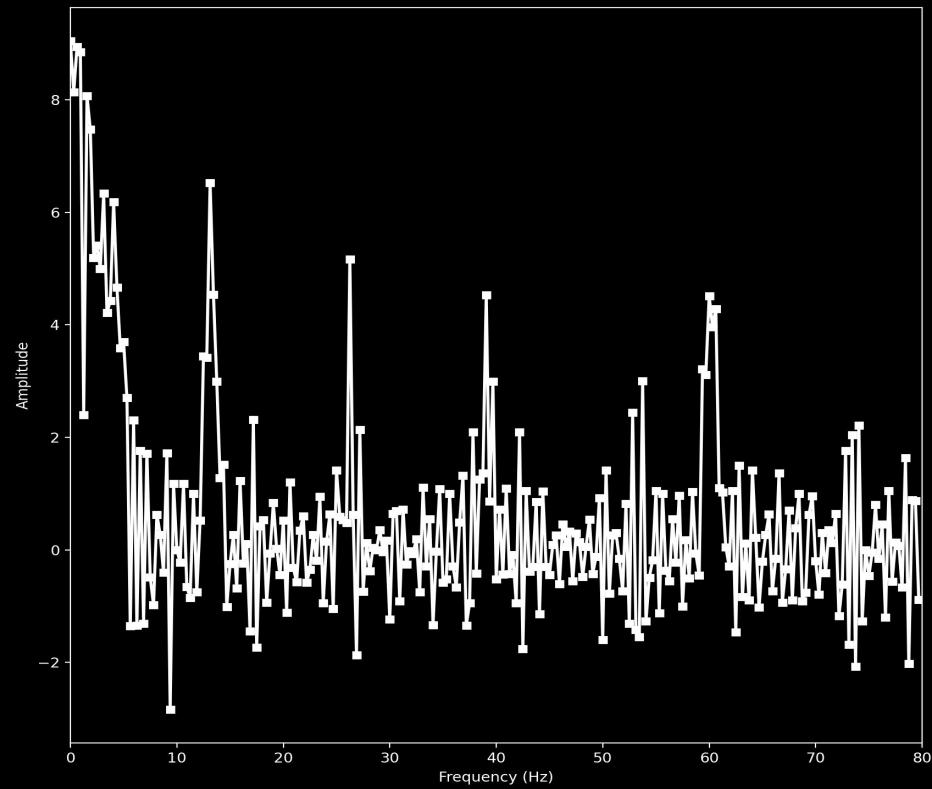
$$\Delta t_{\text{DM}} = 8.3 DM \nu^{-3} B \text{ } \mu\text{s},$$

$$DM \equiv \int_0^S n_e(l) dl$$



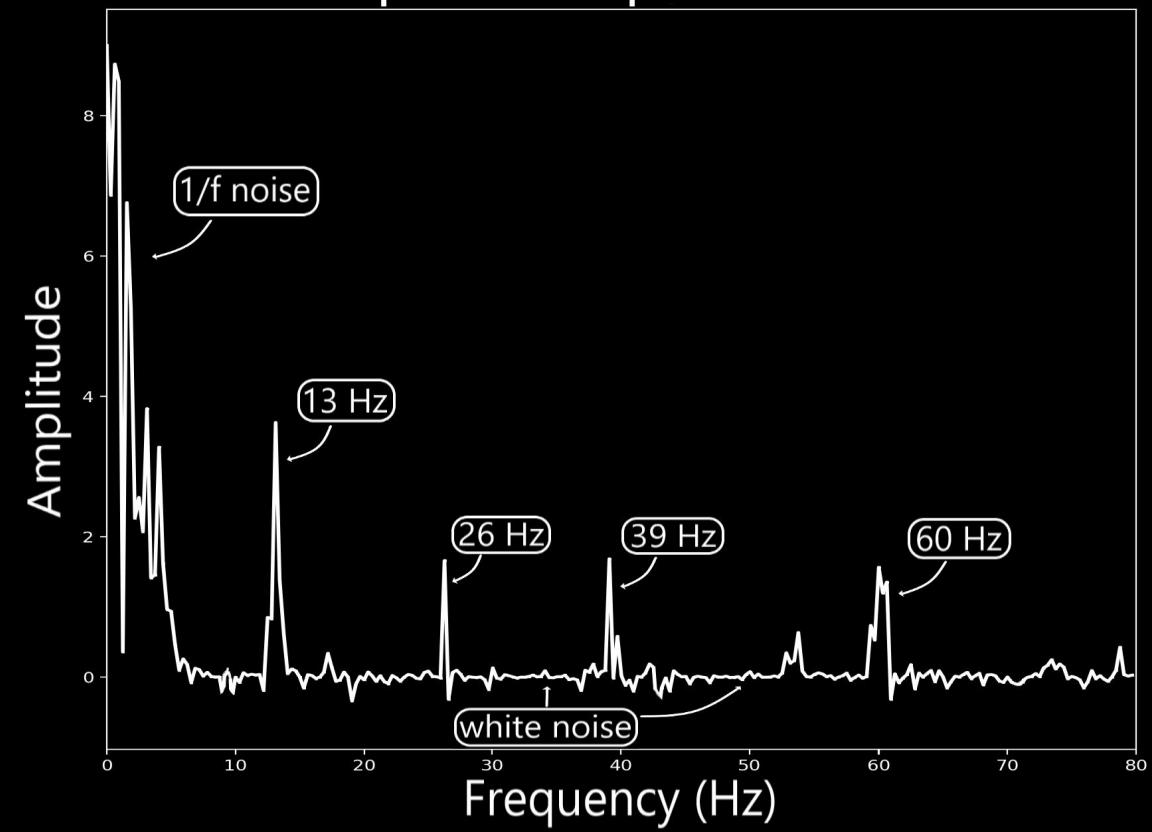
Fourier Transform

Amplitude Spectrum

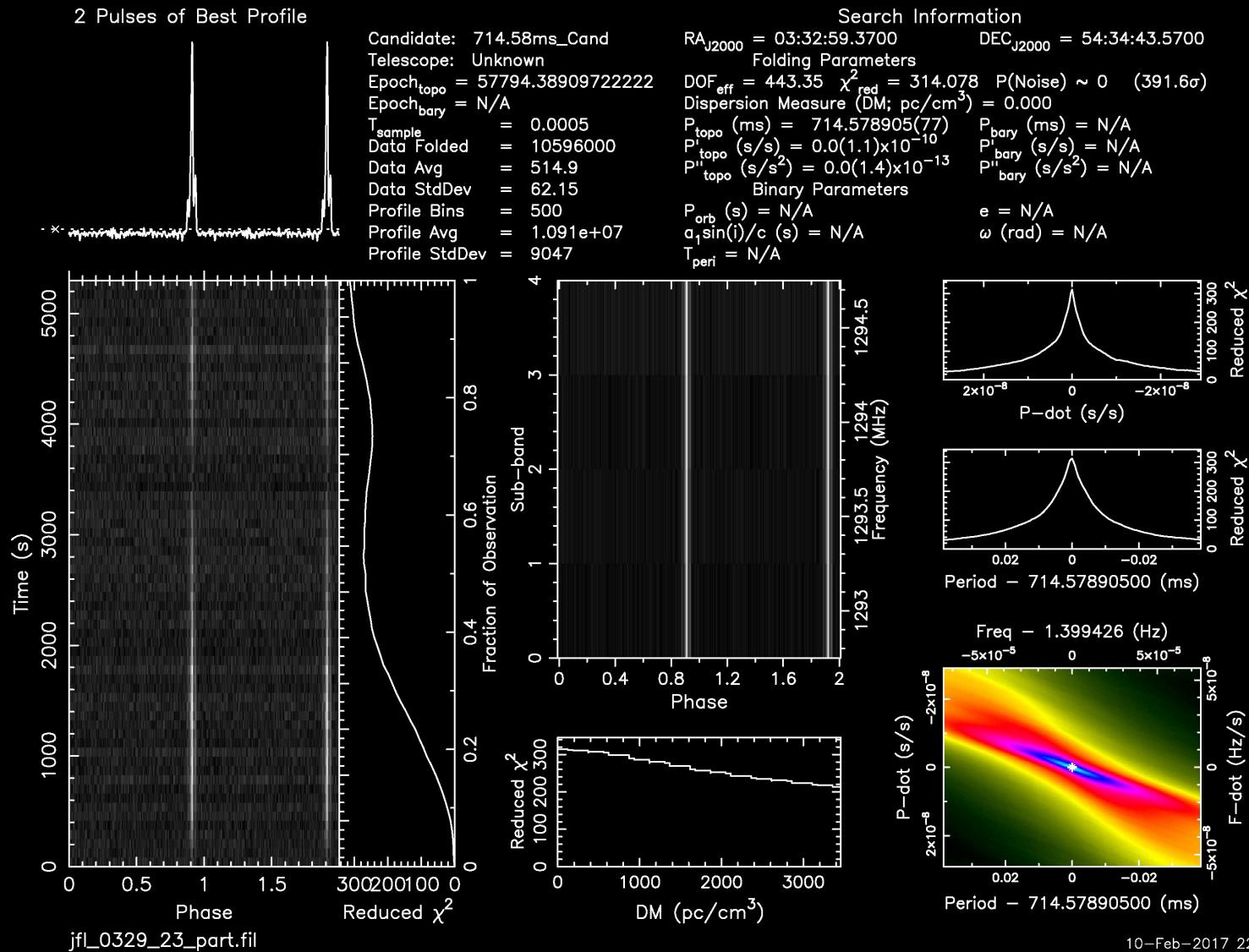


Averaging

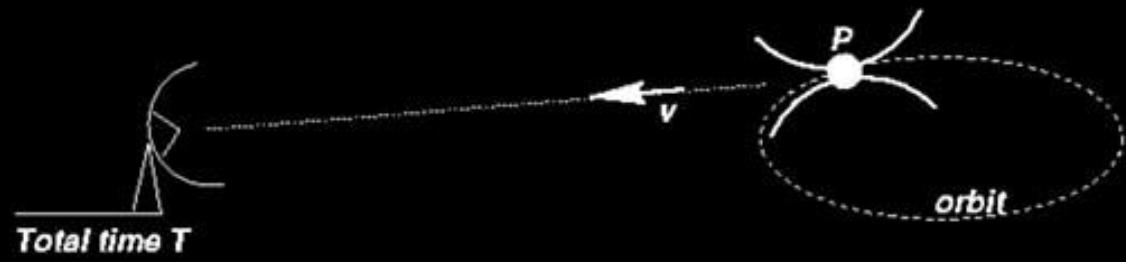
Amplitude Spectrum



Refolding & Classifying

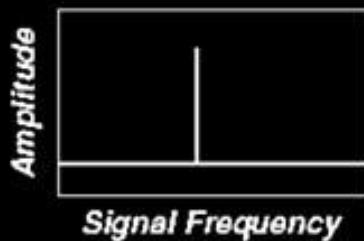


Acceleration Searches:



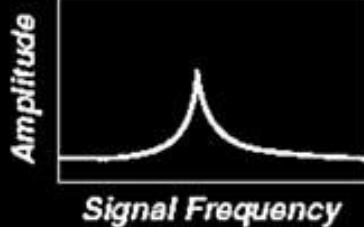
$$\text{Apparent Period} = P (1 + v/c) \sim P (1 + at/c) \quad \text{i.e. time dependent}$$

Case (a)
Isolated Pulsar



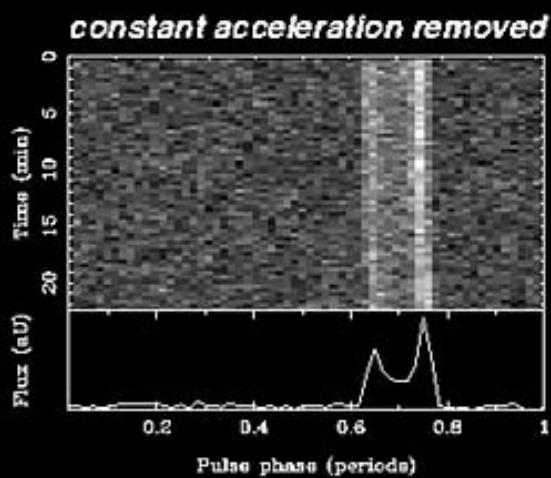
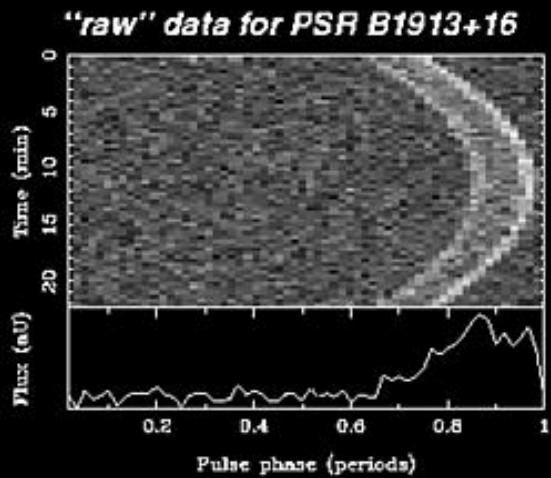
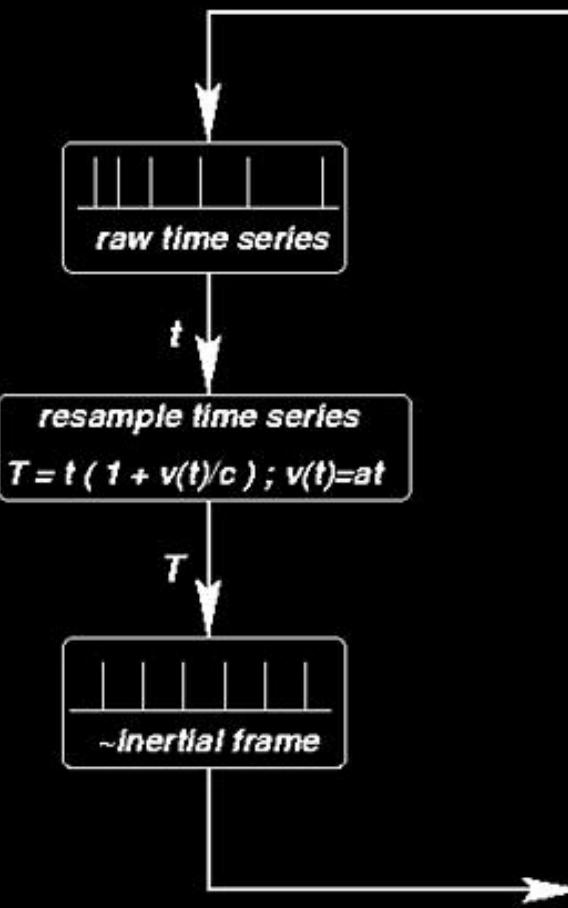
$$\Delta f \sim aT/Pc \ll 1/T$$

Case (b)
Binary Pulsar



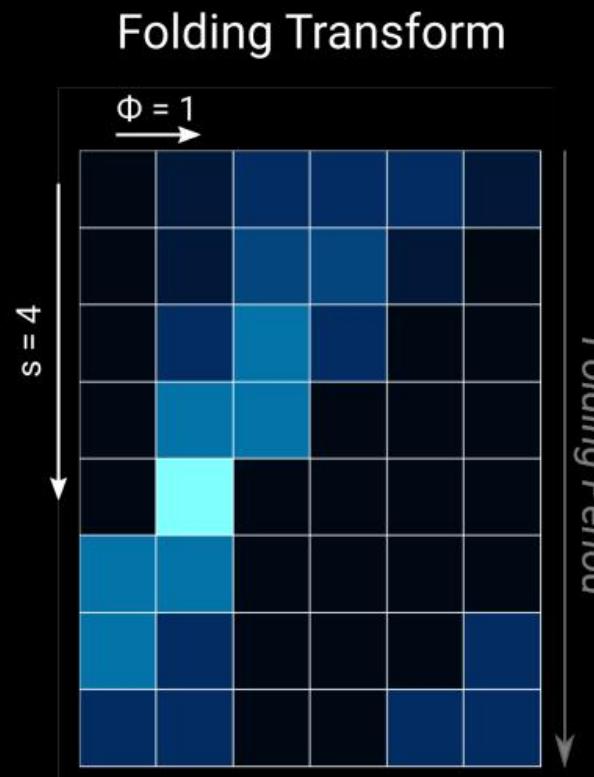
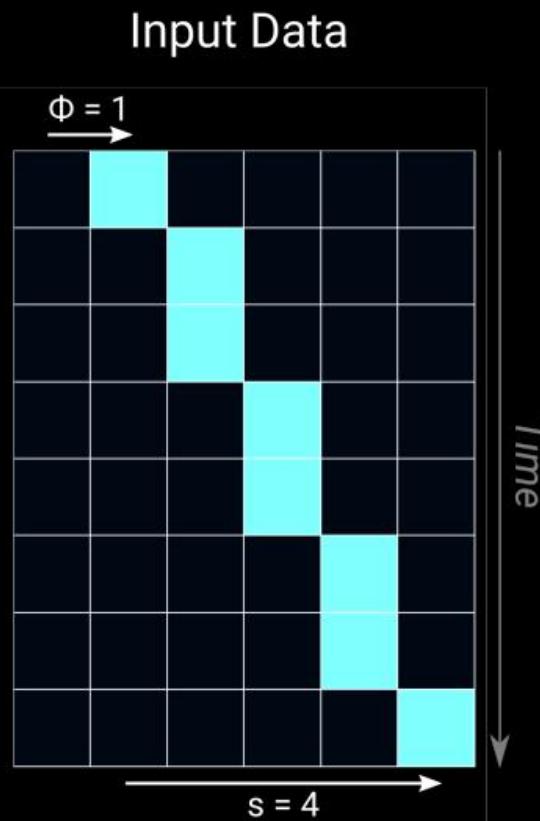
$$\Delta f \gg aT/Pc$$

$$aT^2/Pc \gg 1$$

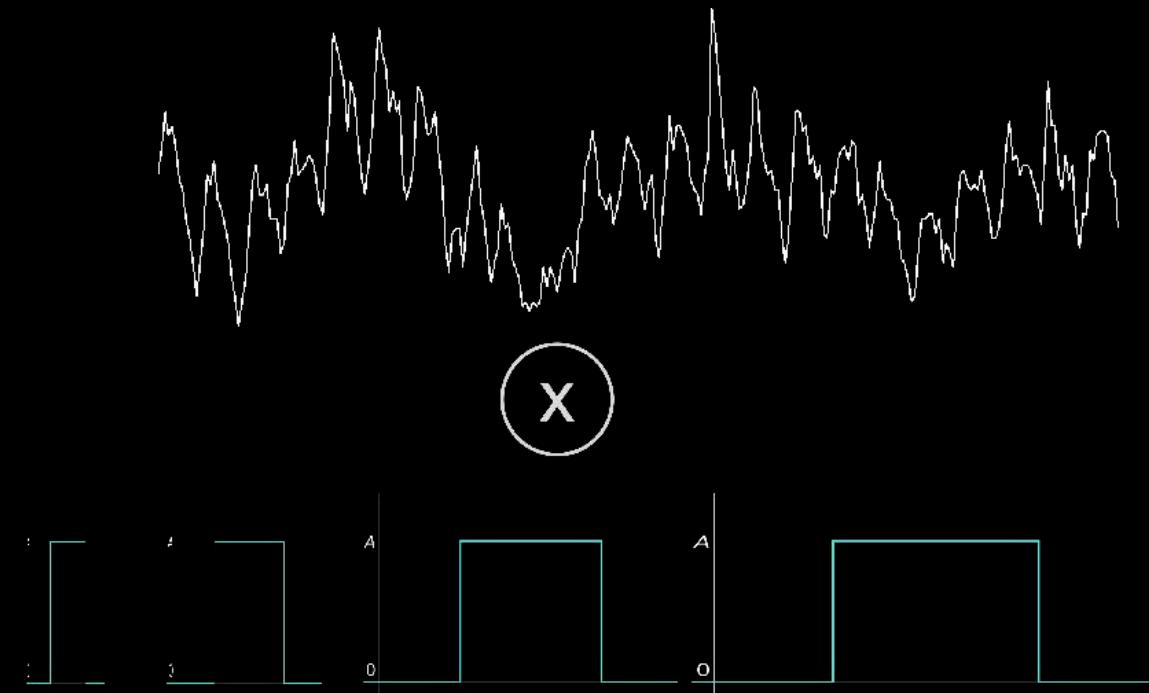


OR:

Fast Folding Algorithms



Single Pulse Searches



But what about the hardware?

Pulsar discovery tally:

1441 | Murriyang



392 | Arecibo

362 | Green bank

262 | FAST

198 | MeerkAT



What does the future hold?

SKA 1-Mid: ~7 000 canonical pulsar
 + ~900 MSPs

SKA 2-Mid: ~24 000 - 30 000 pulsars

(Keane et al; 2015)

