

AMERICAN
MUSEUM
OF
NATURAL
HISTORY



MORDECAI-MARK MAC LOW

PLANET & STAR FORMATION

PLANET FORMATION

MINERAL FORMATION

Alexander Hubbard (AMNH)

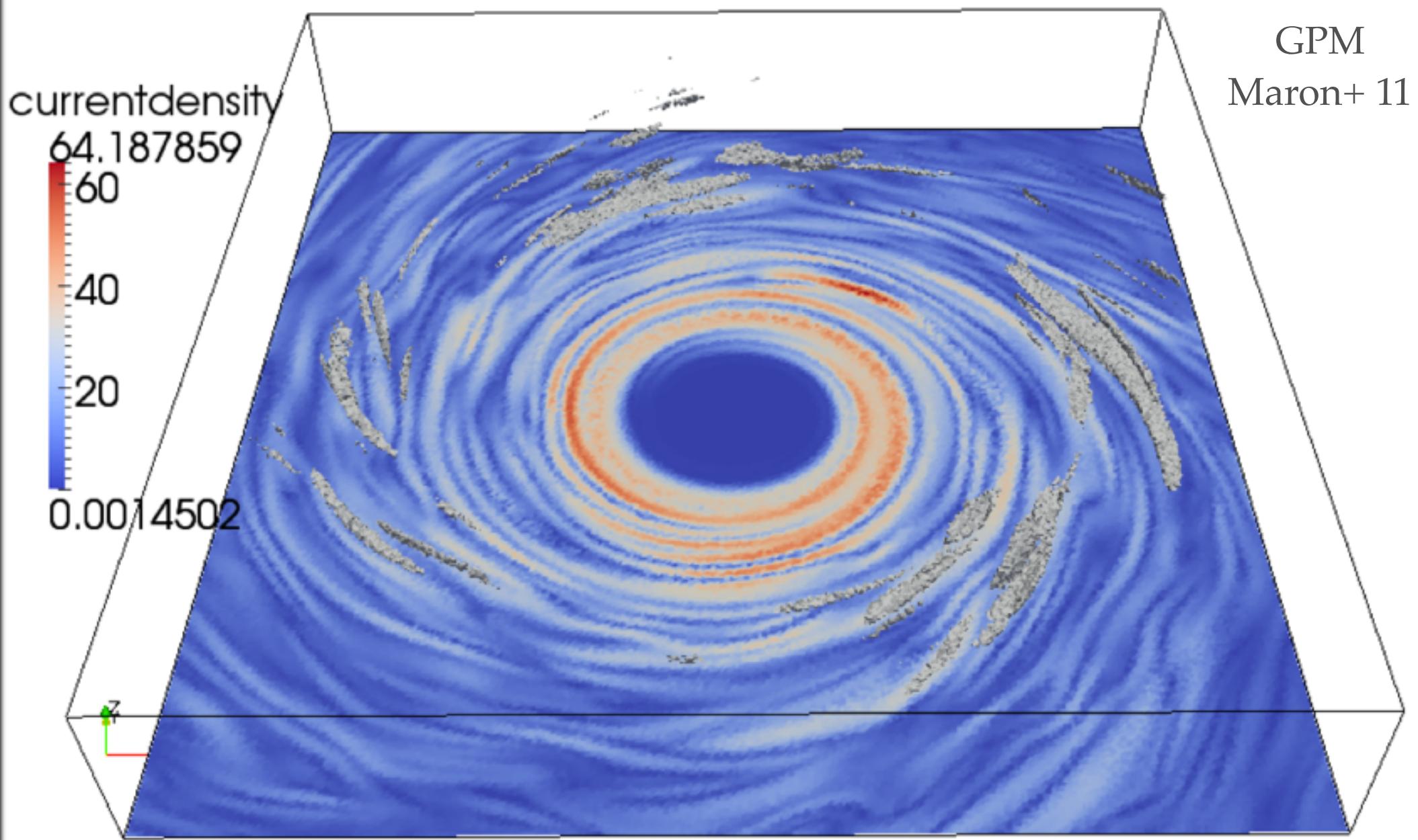
Colin McNally (NBIA, Copenhagen)

Denton Ebel (AMNH)

O. Richard Norton, <http://www.geokem.com/meteorites.html>,

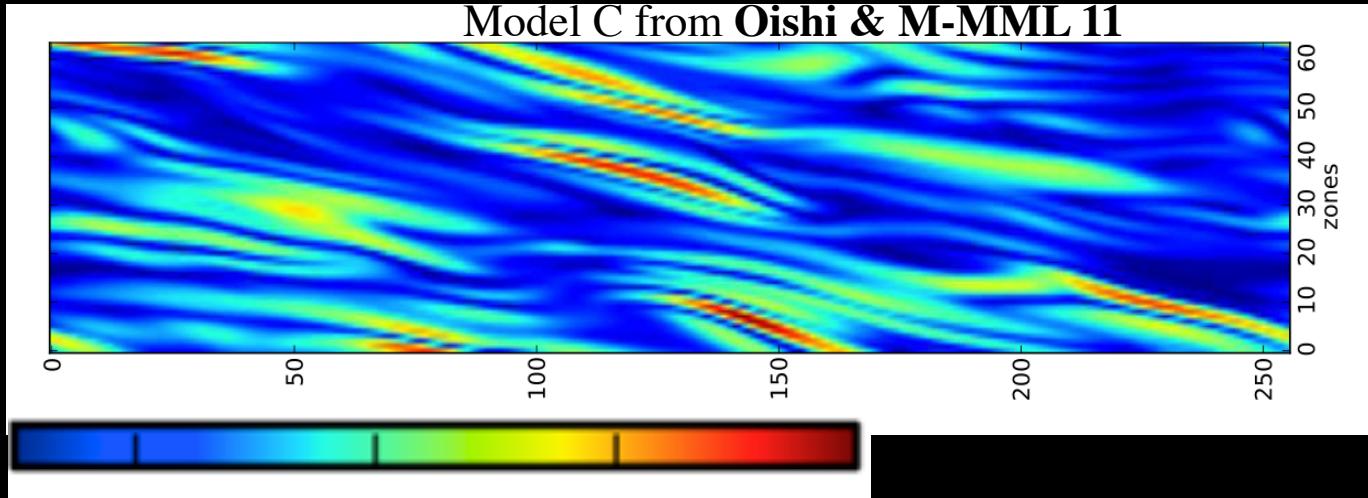
Magnetized turbulence forms ubiquitous current sheets

(Parker 1972, ApJ, 174, 499)



McNally 12 PhD Thesis, McNally, Hubbard, M-MML, Ebel, D'Alessio 13

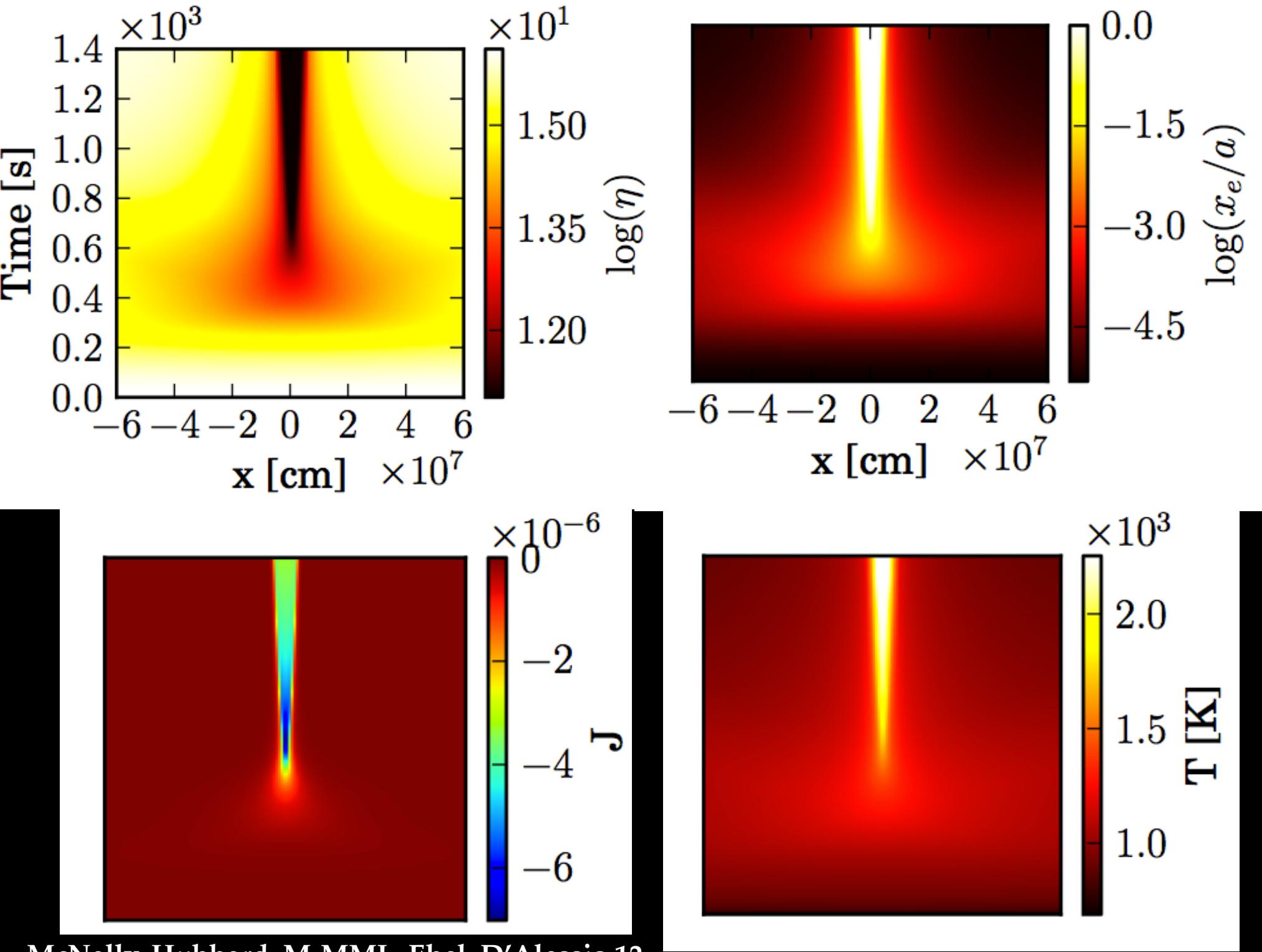
Model C from Oishi & M-MML 11



$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B} - \eta \mathbf{J})$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B}) + \eta \nabla^2 \mathbf{B} - (\nabla \eta) \times \mathbf{J}$$

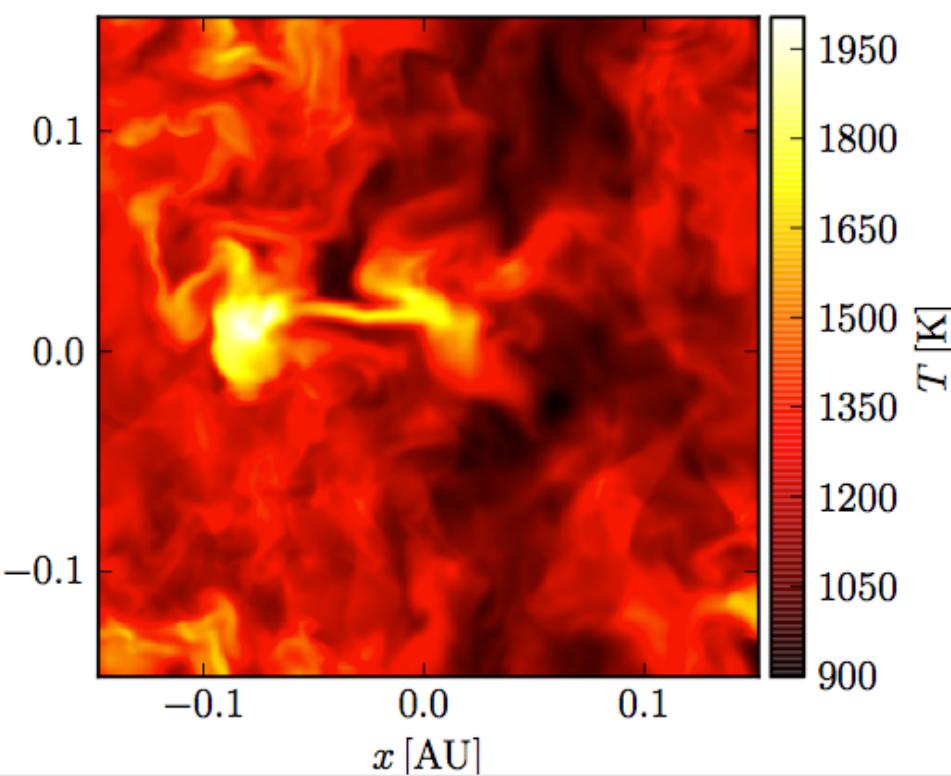
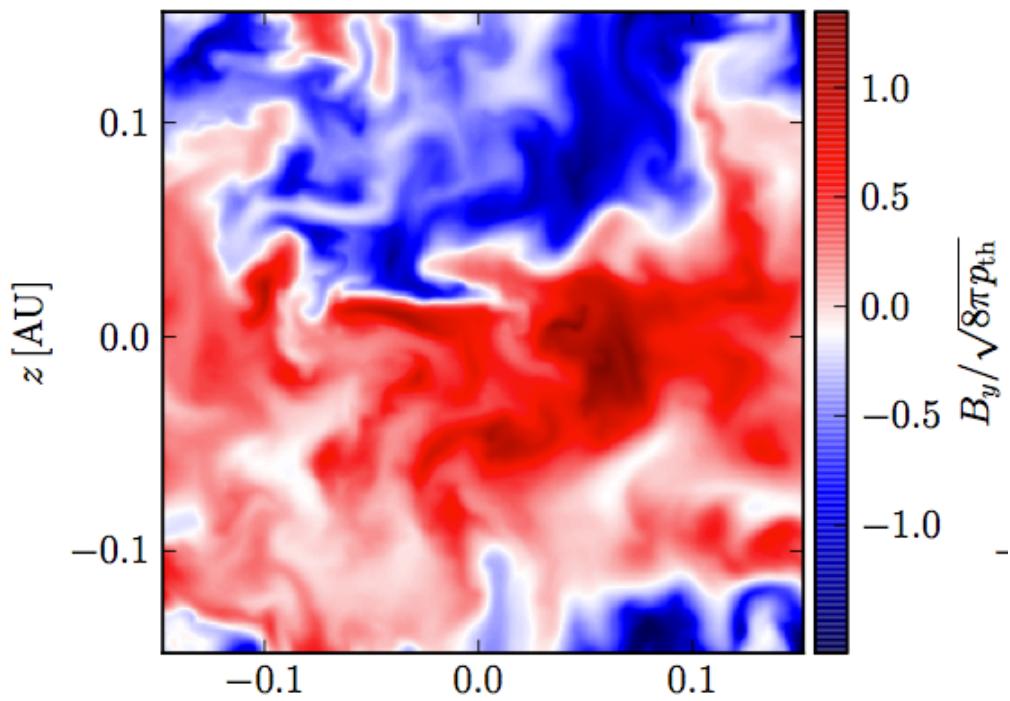
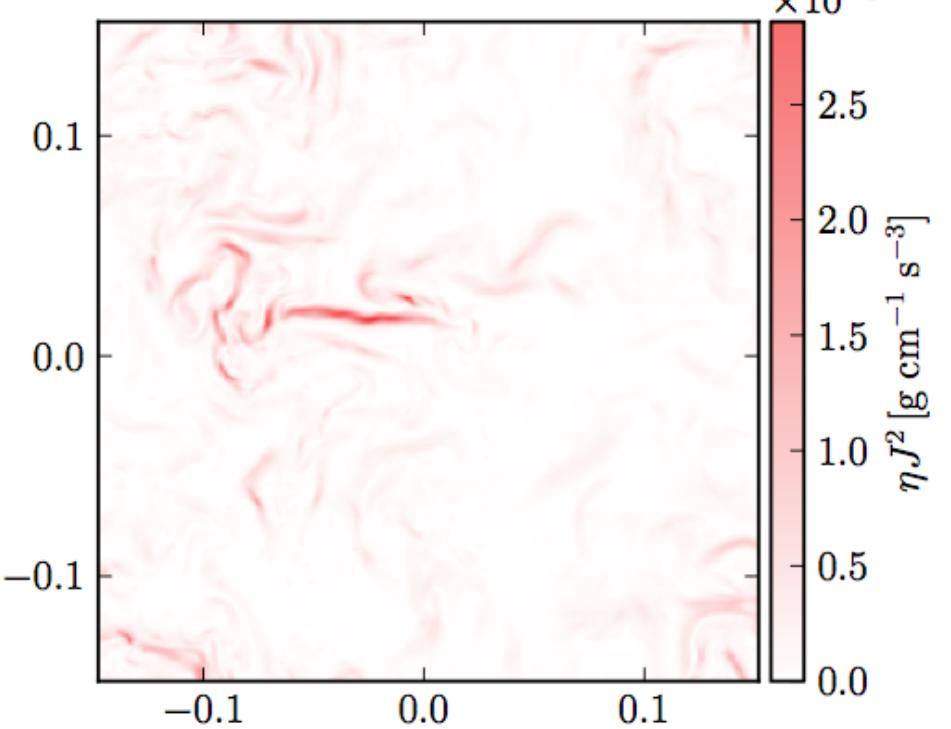
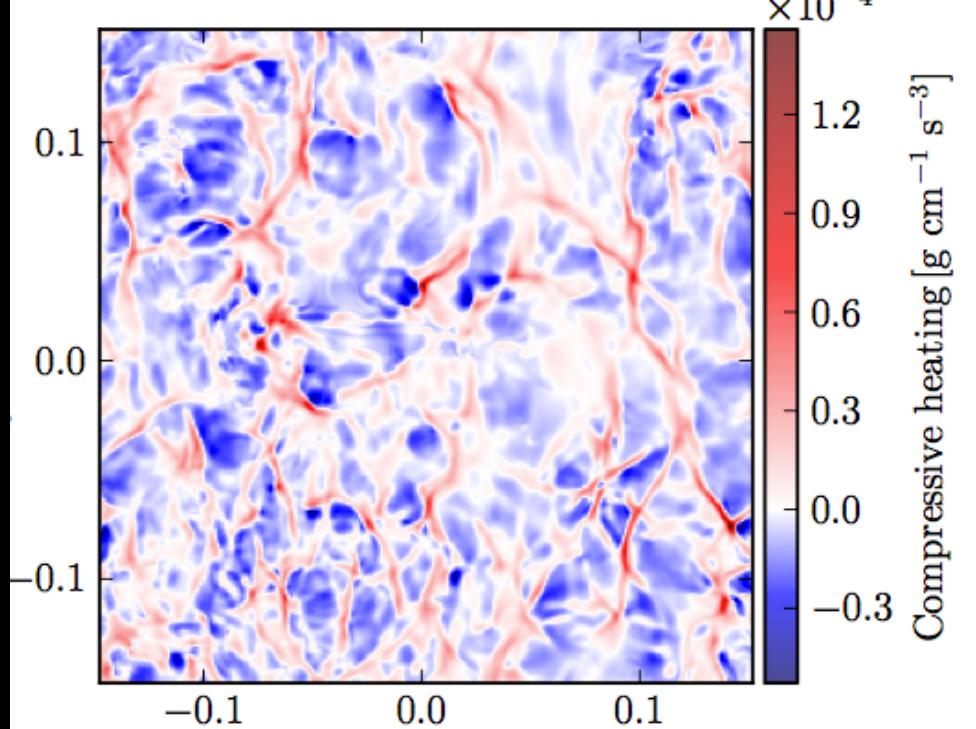
If temperatures rise high enough to ionize alkali metals, resistivity starts to drop exponentially (Hubbard, McNally, M-MML 12).

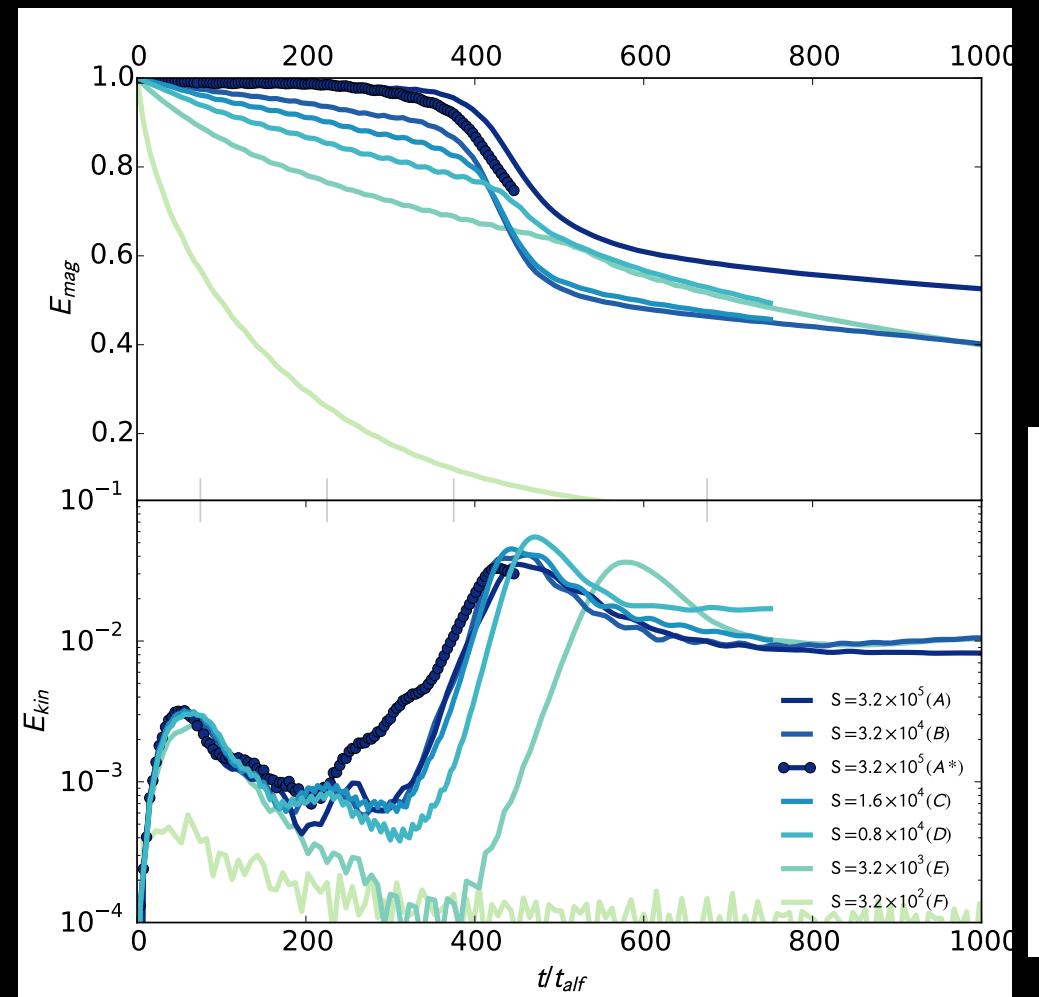
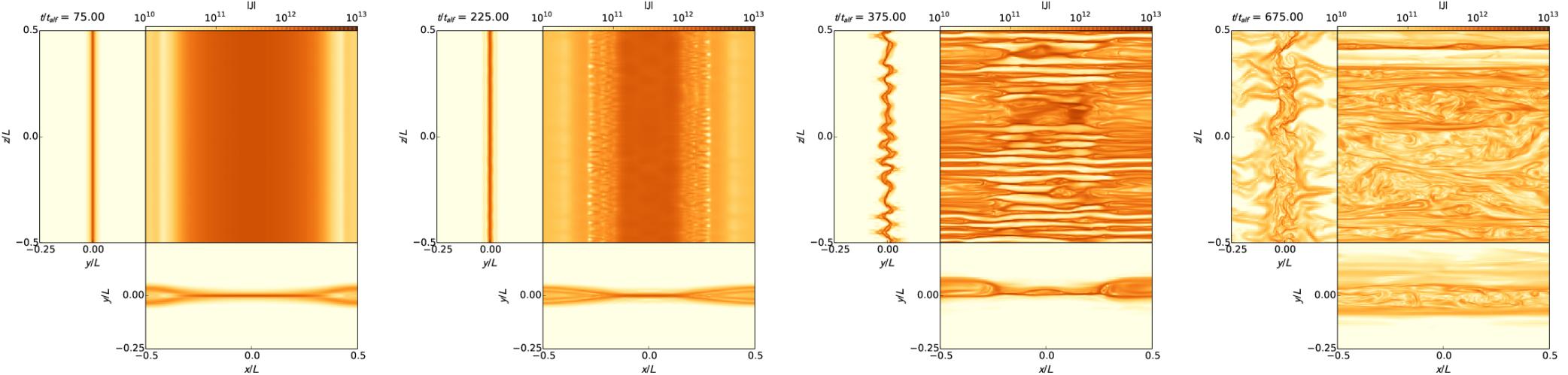


McNally, Hubbard, M-MML, Ebel, D'Alessio 12

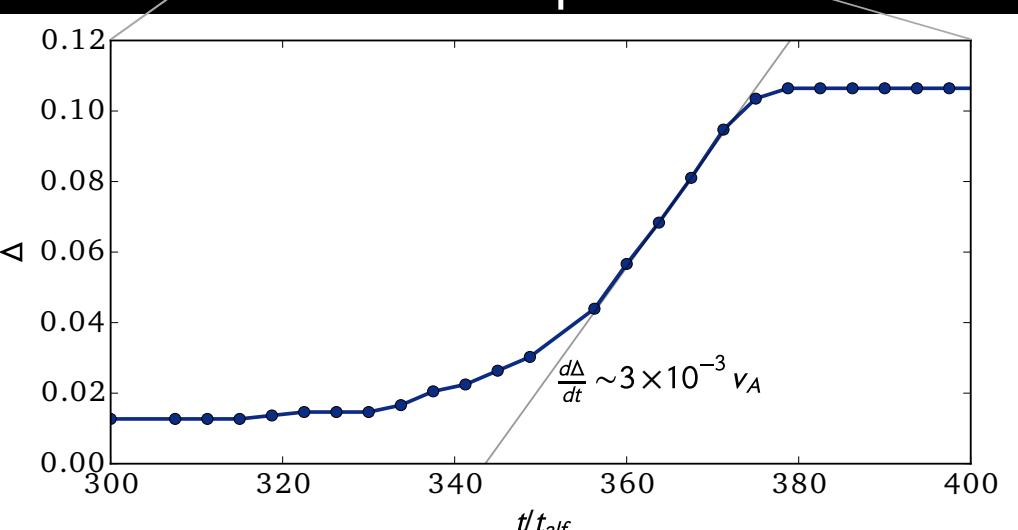
SHEARING BOX MODEL

- Pencil Code (Brandenburg & Dobler 2002)
 - www.nordita.org/software/pencil-code/
- Finite-difference MHD code
- **Explicit resistivity to resolve current sheets**
- Sixth-order spatial, third-order time
- Hyperdiffusion for time-centered scheme
- $\text{Div } \mathbf{B} = 0$ maintained using vector potential





Enzo Nested Grid
explicit resistivity
Lundquist-number indpt.
reconnection speed



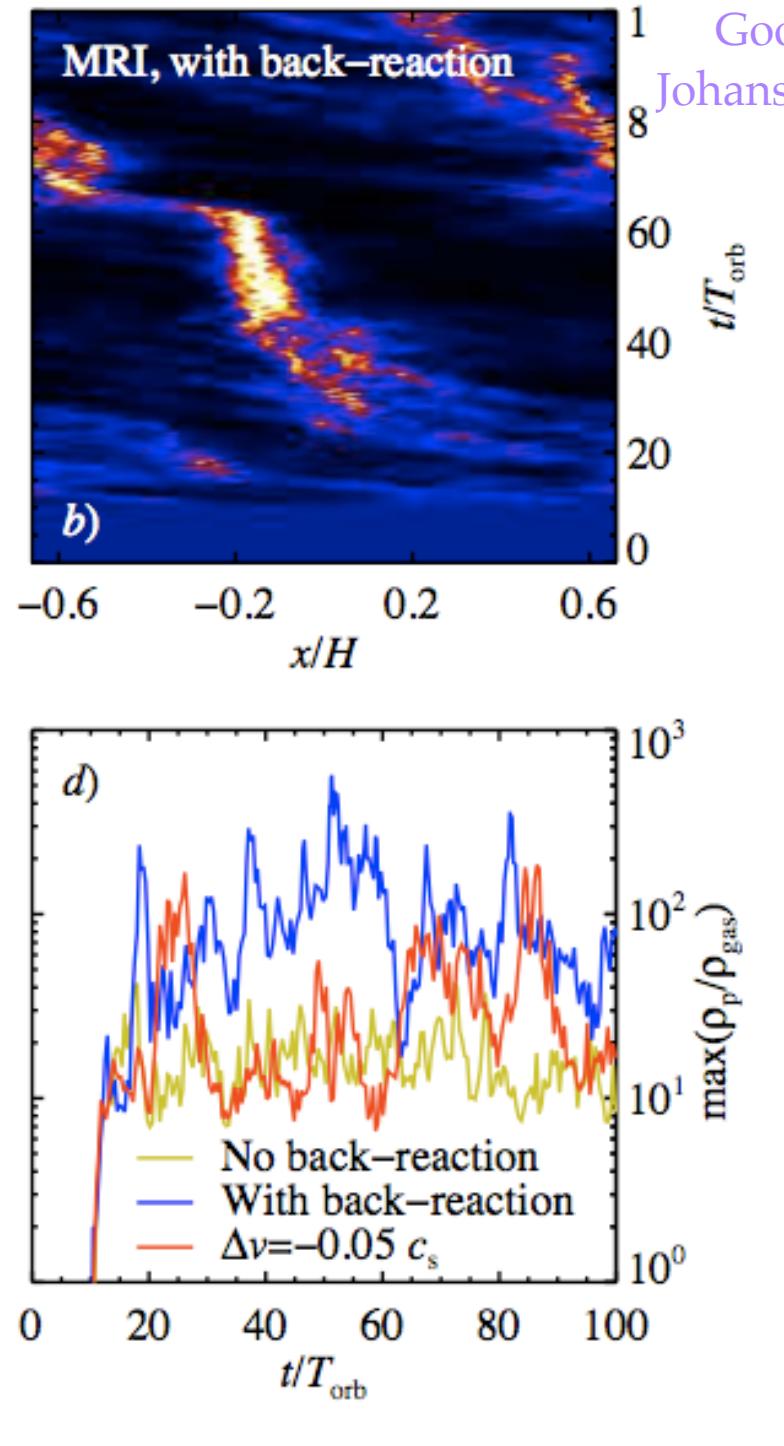
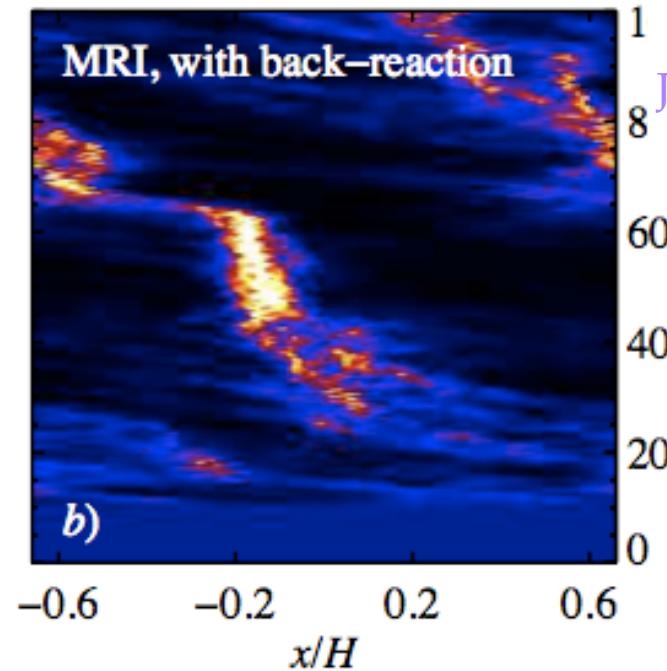
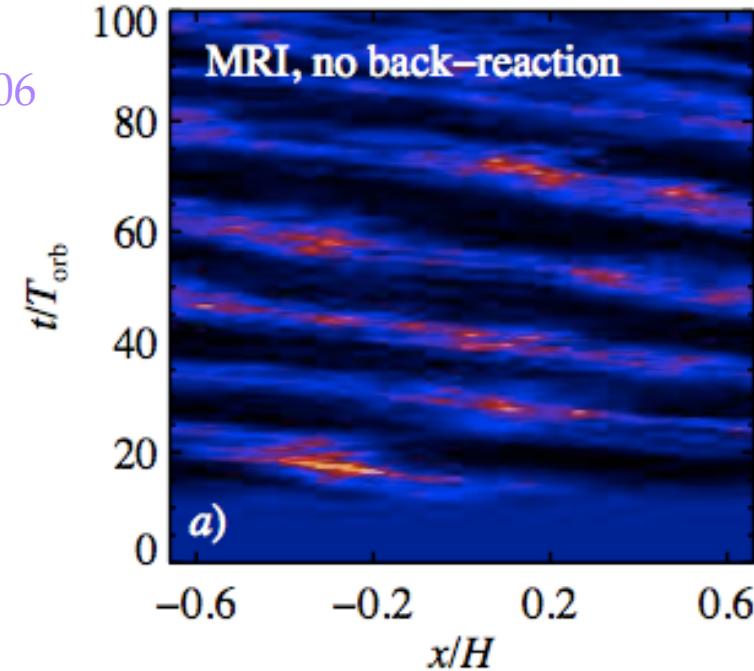
Oishi + 15

PLANET FORMATION

PLANETESIMALS

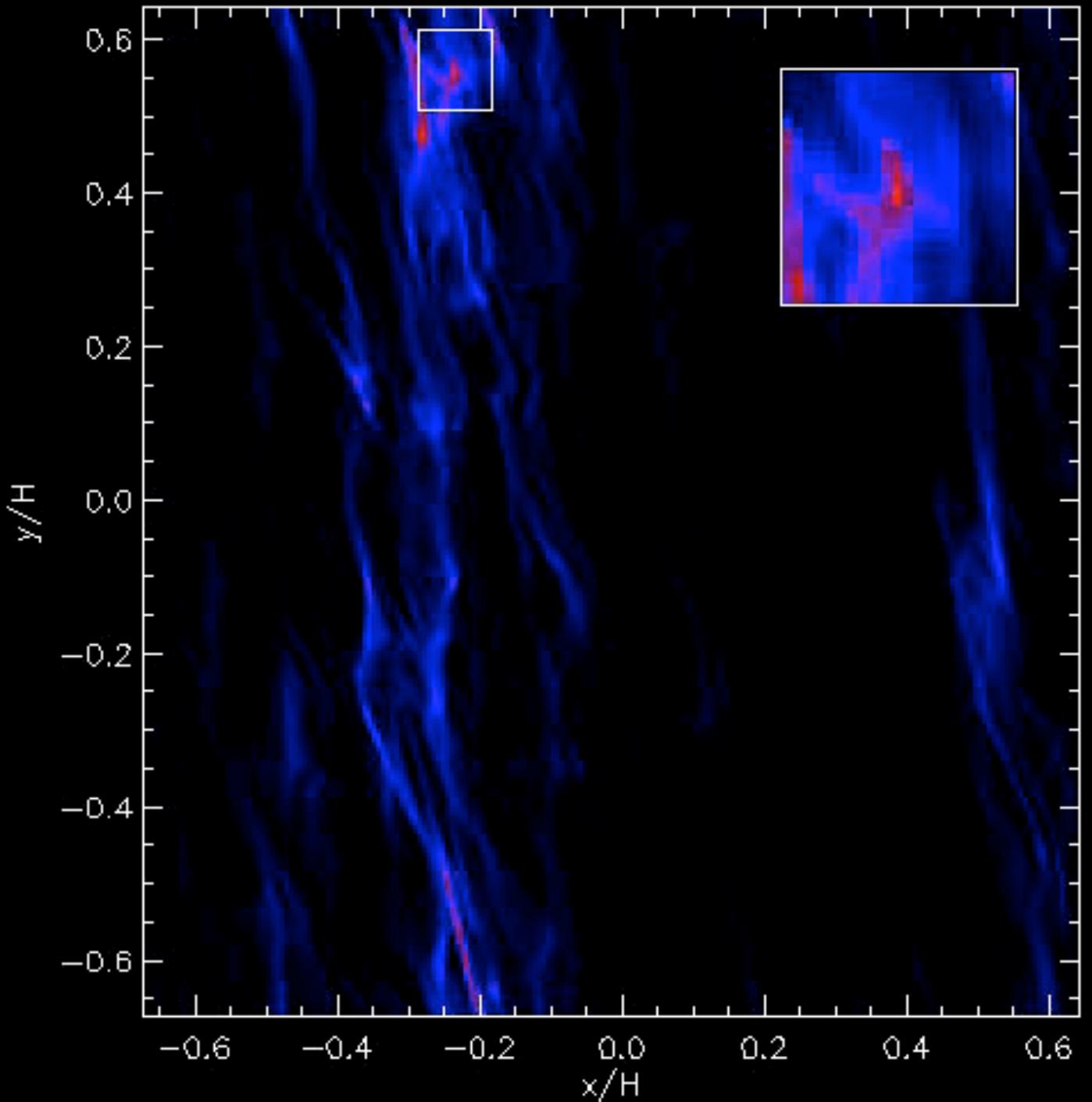
Anders Johansen (U. Lund)

O. Richard Norton, <http://www.geokem.com/meteorites.html>



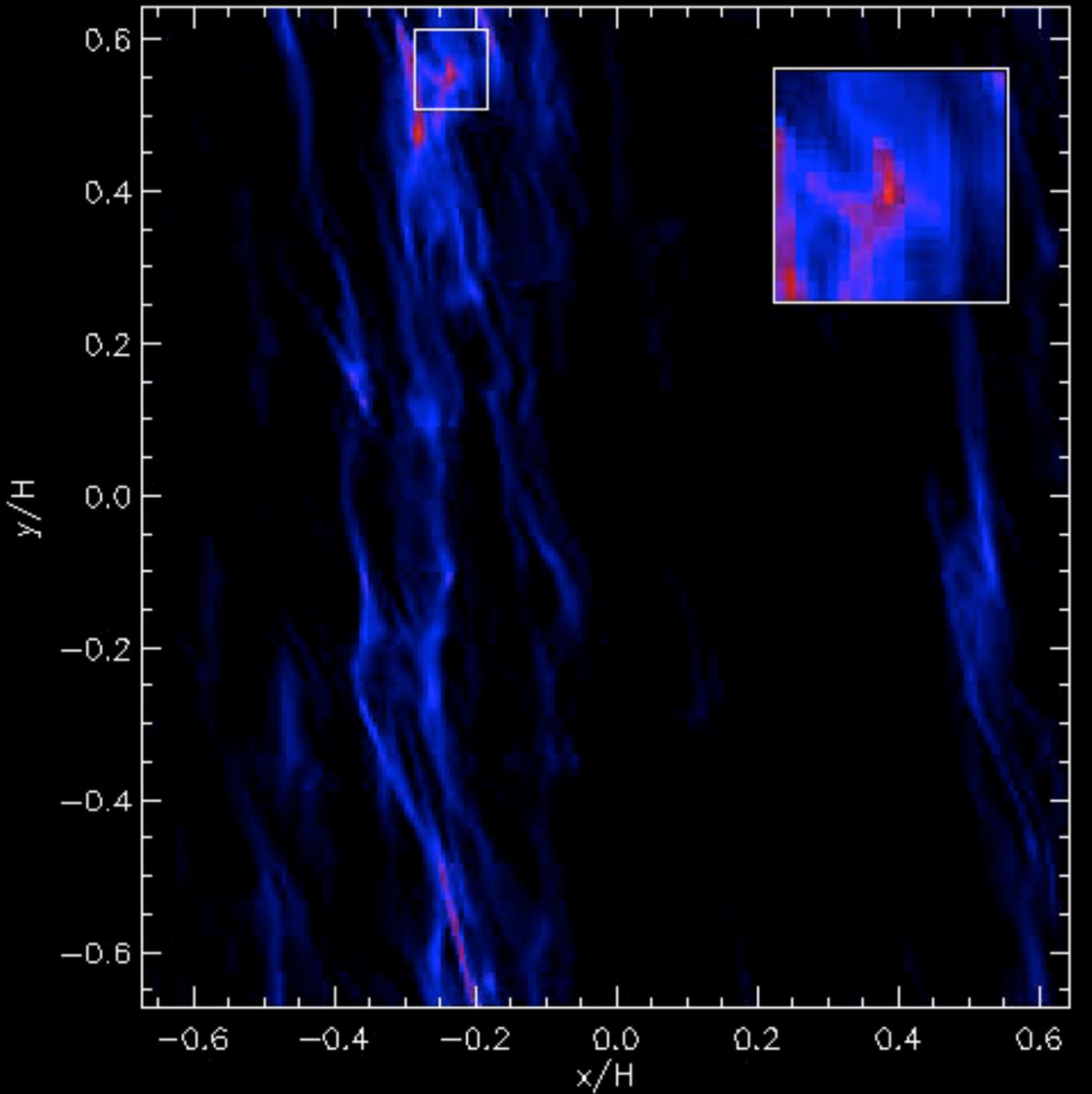
Turbulence and
streaming
instability interfere
constructively!

including self-gravity of boulders allows tracking of non-linear gravitational instability



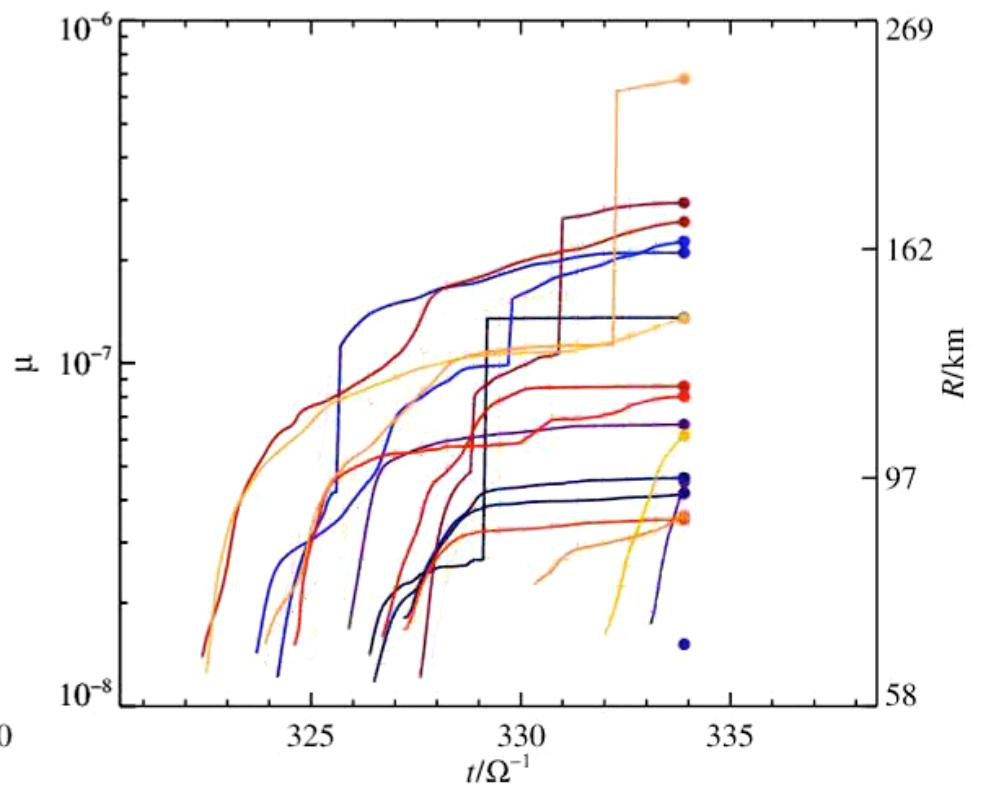
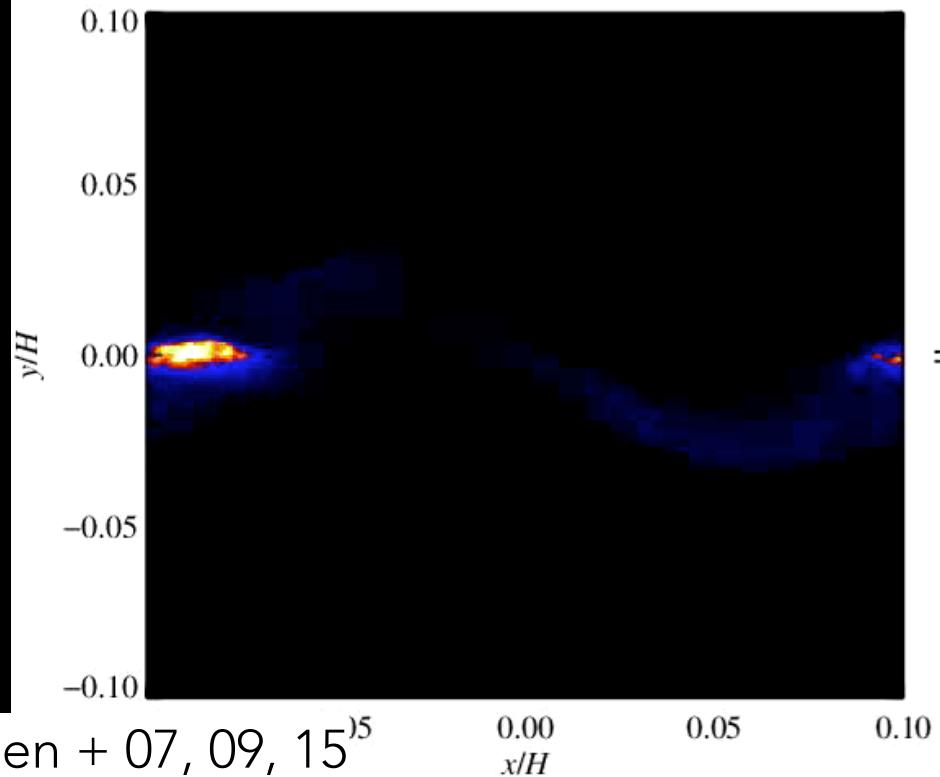
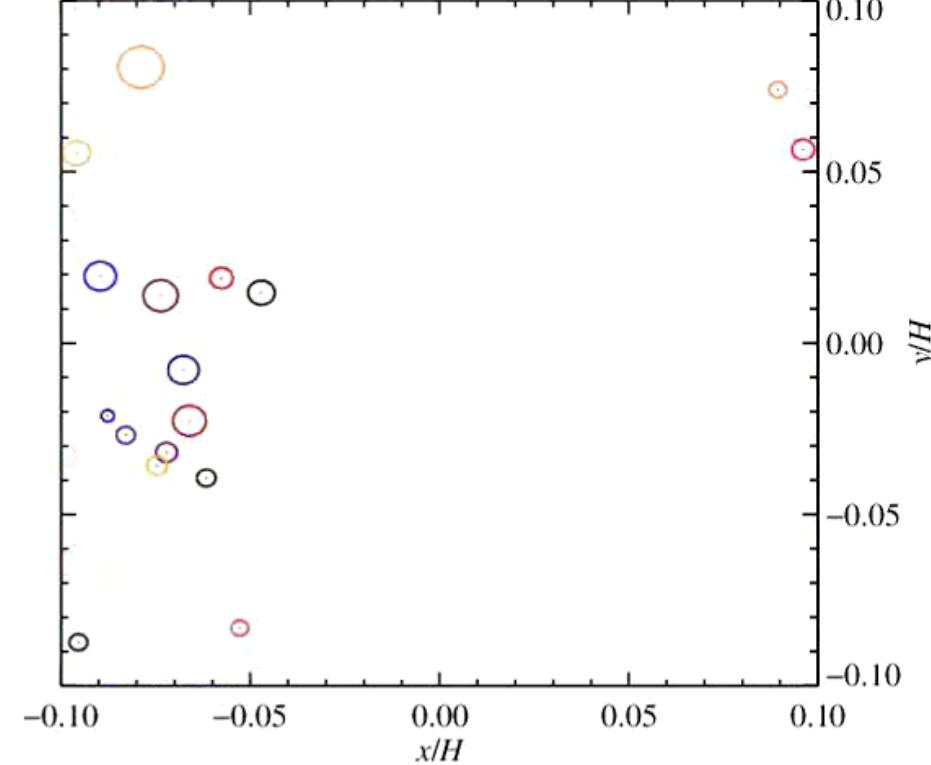
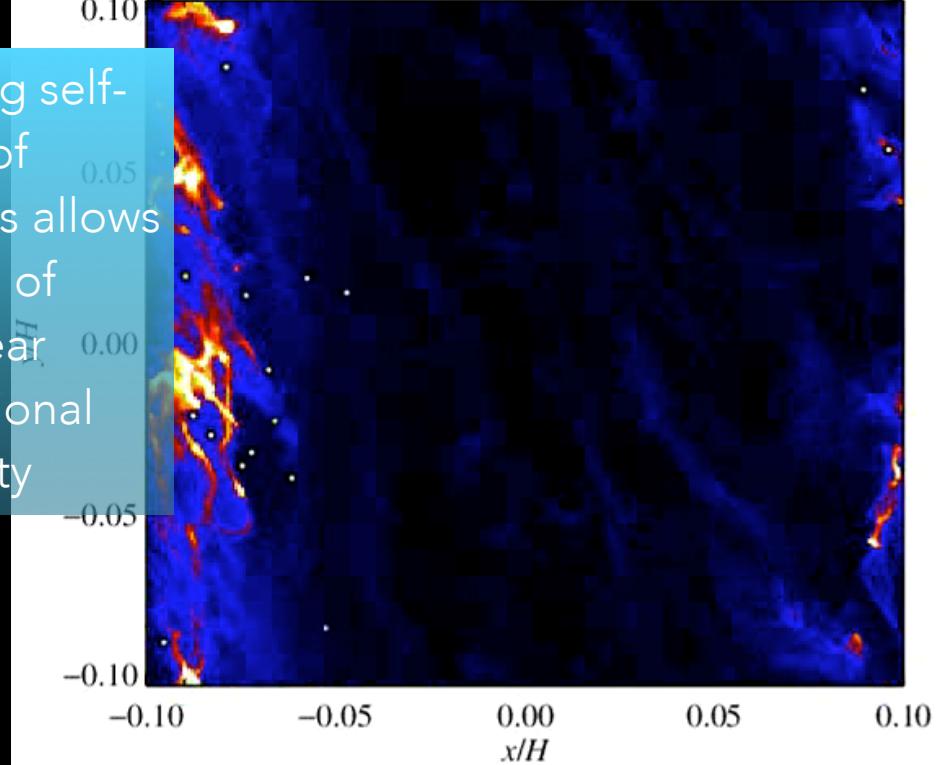
Johansen, Oishi, Mac Low, Klahr, Henning, & Youdin 07

including self-gravity of boulders allows tracking of non-linear gravitational instability

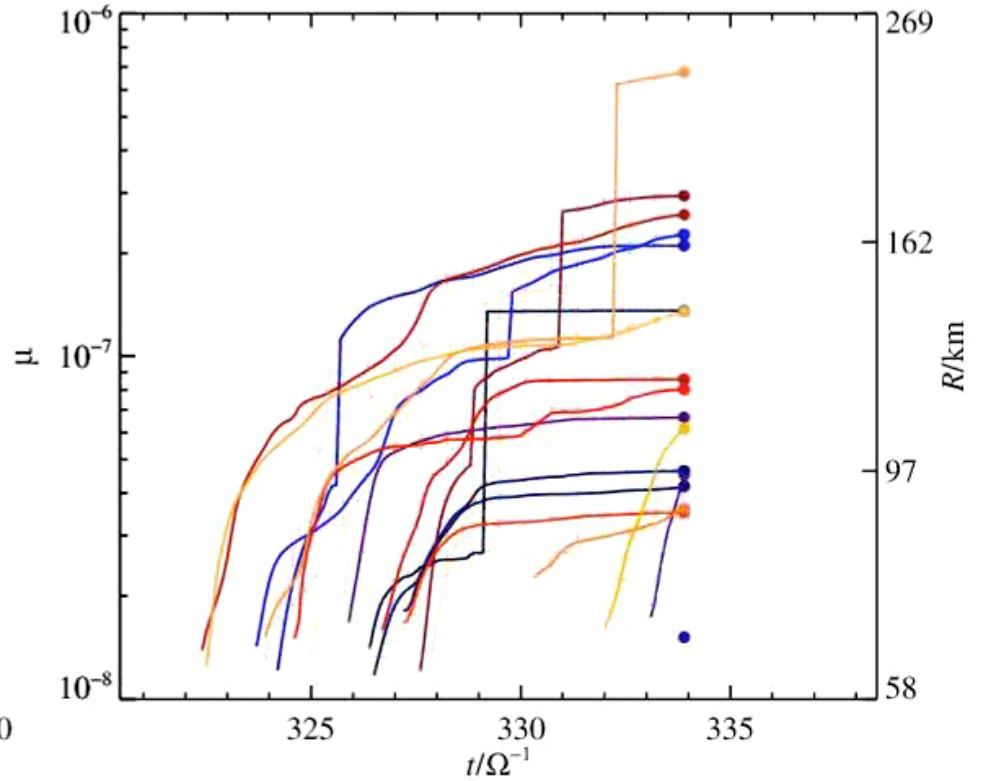
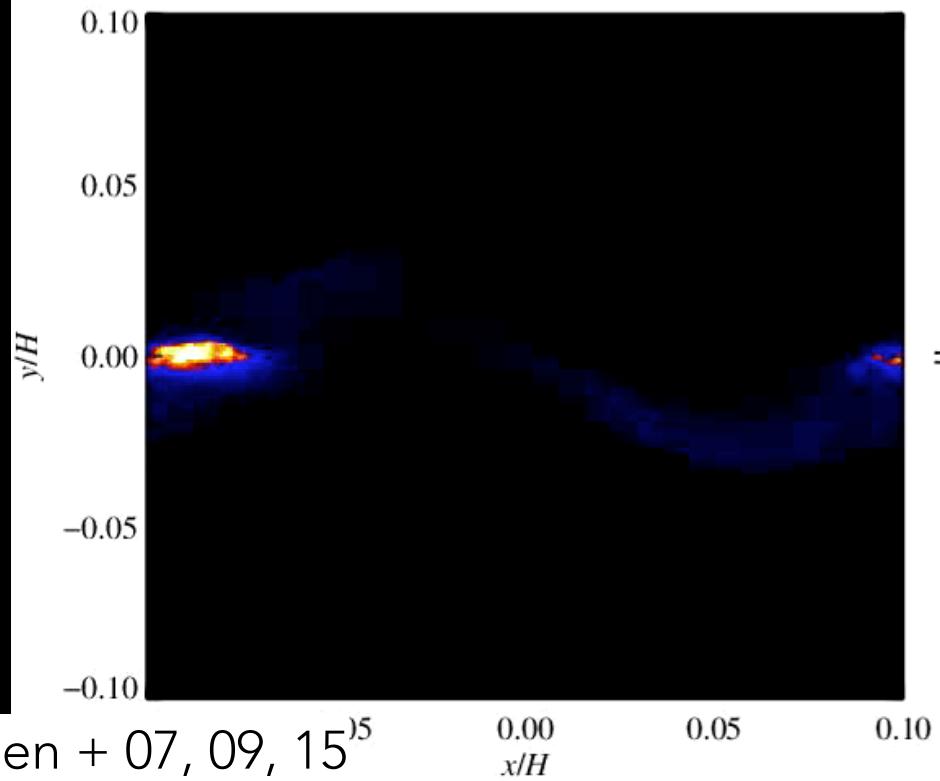
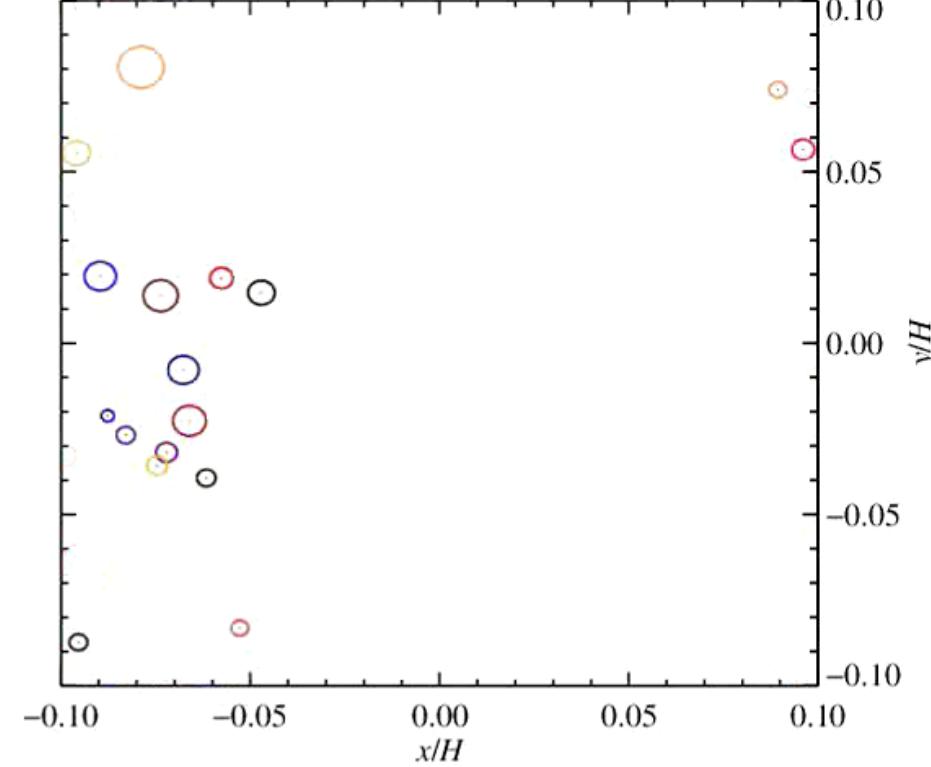
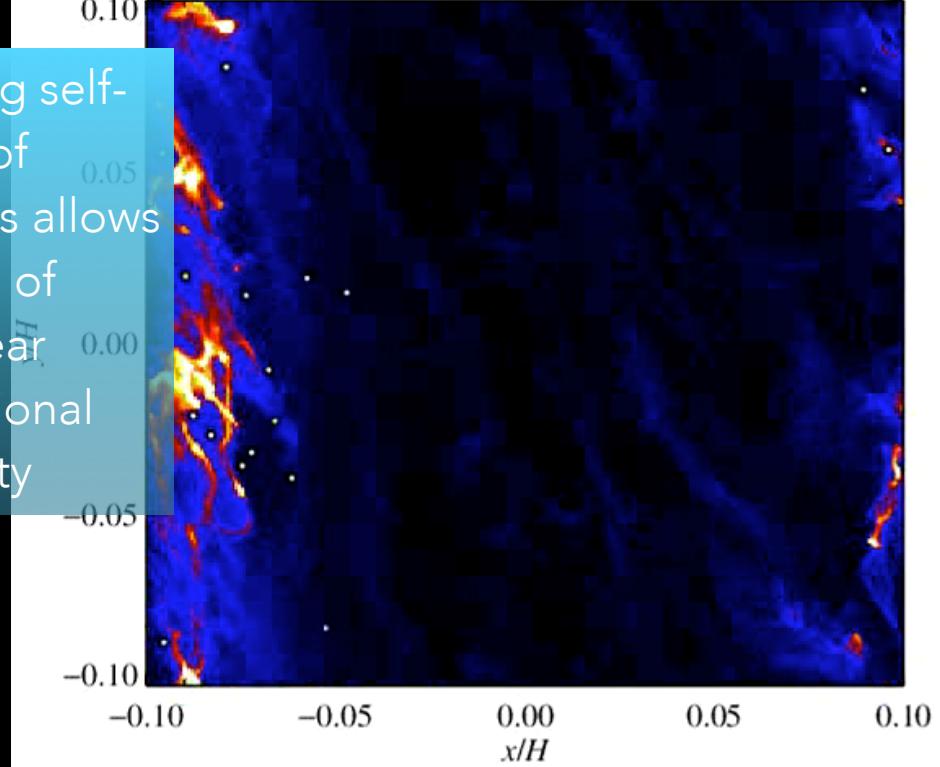


Johansen, Oishi, Mac Low, Klahr, Henning, & Youdin 07

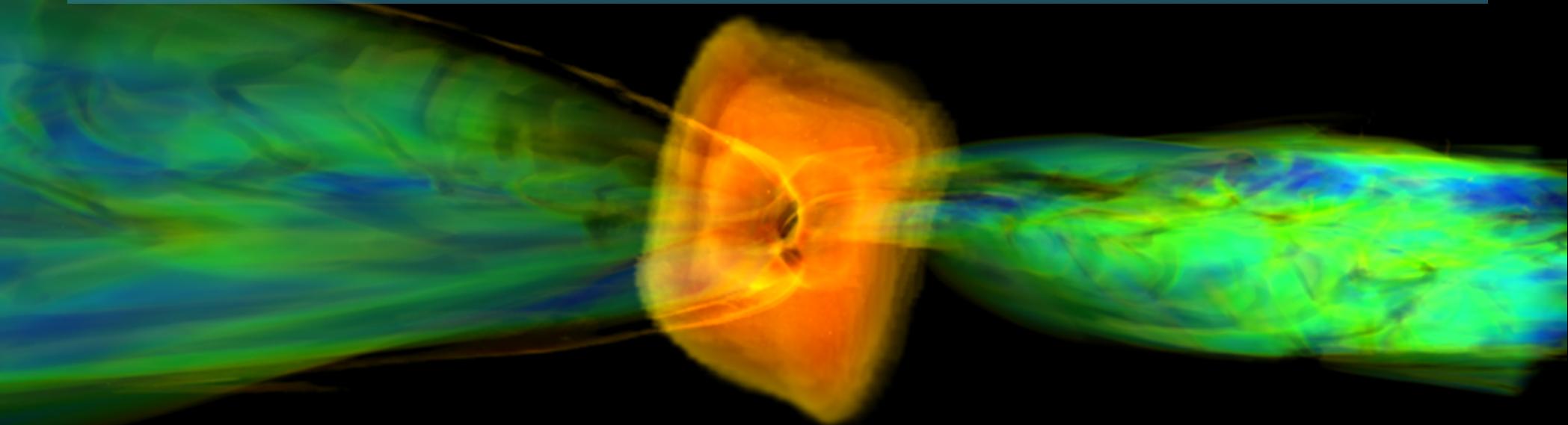
including self-gravity of boulders allows tracking of non-linear gravitational instability



including self-gravity of boulders allows tracking of non-linear gravitational instability



CLUSTER FORMATION



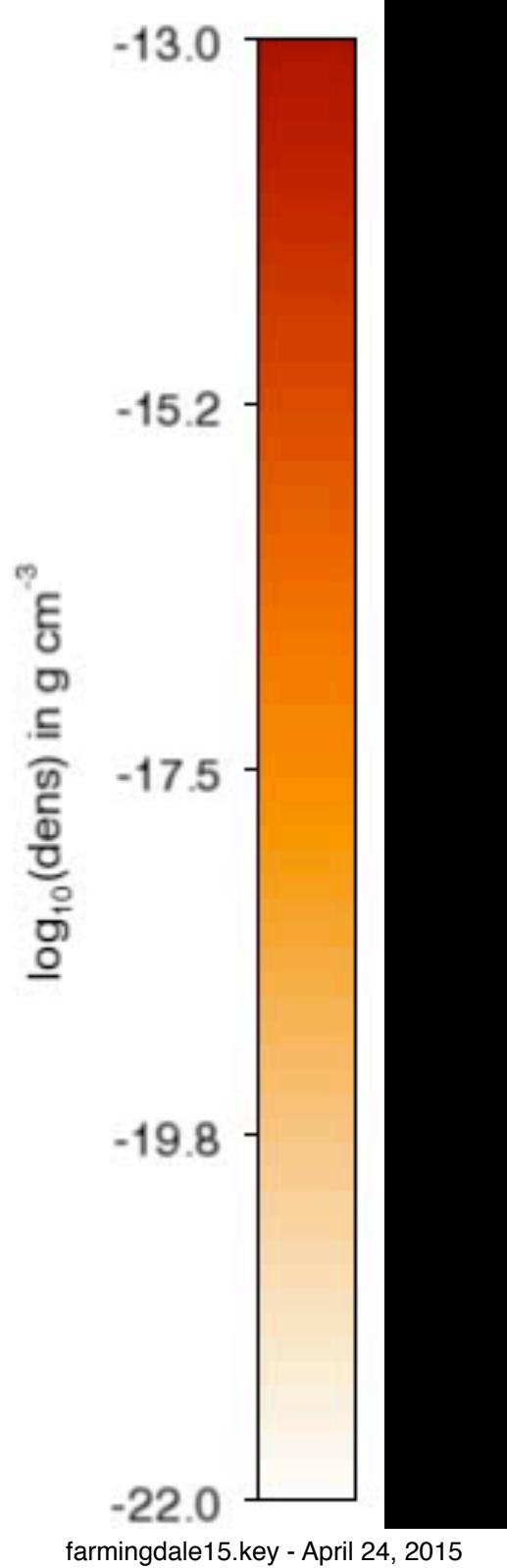
Thomas Peters (MPA)
Joshua Wall (Drexel)
Steven McMillan (Drexel)
Simon Portegies-Zwart (Leiden)
Juan Ibañez-Mejia (AMNH/ITA-ZAH)
Kevin Olson (Drexel)
Ralf Klessen (ITA-ZAH)

Peters + 14, Viz: Zilken, Jülich

0.594 Myr

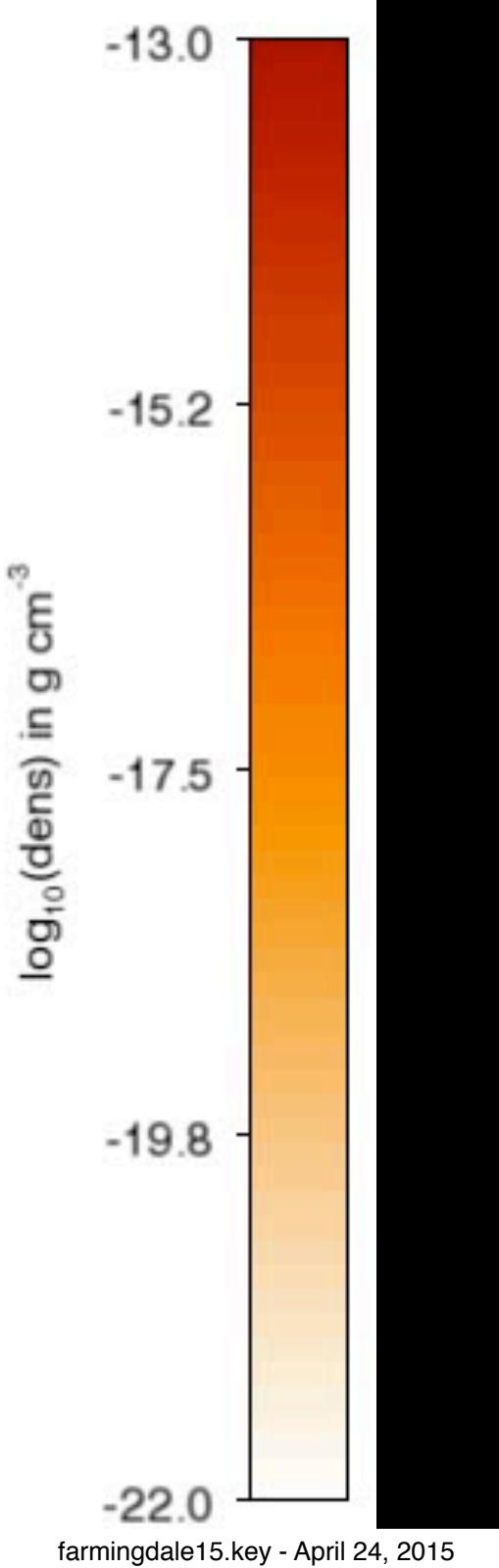
box size 0.324 pc

14-1



0.594 Myr

box size 0.324 pc



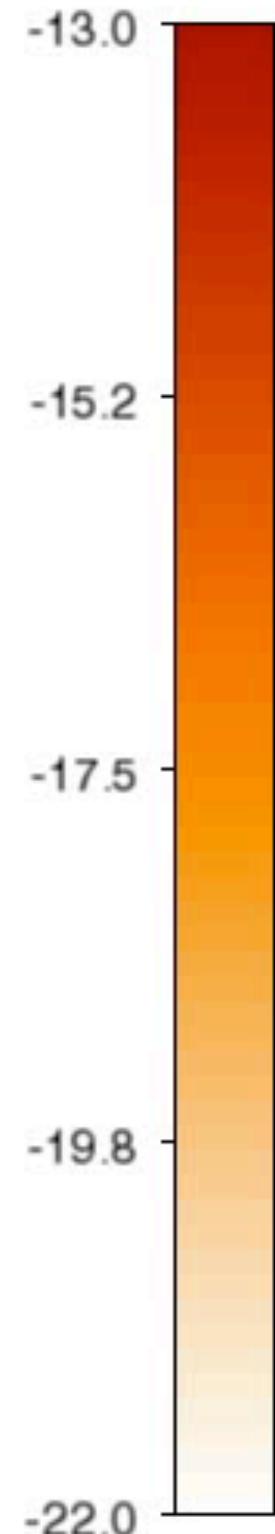
0.594 Myr

box size 0.324 pc

15-1

farmingdale15.key - April 24, 2015

$\log_{10}(\text{dens}) \text{ in g cm}^{-3}$



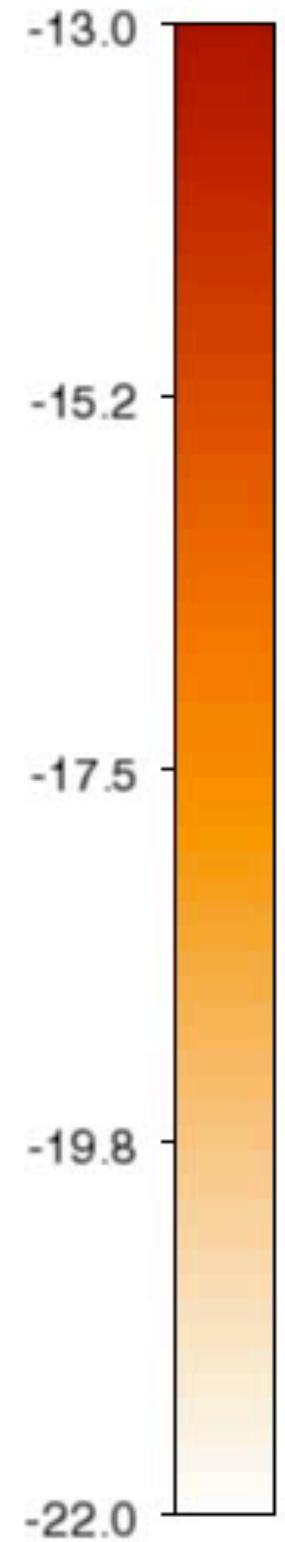
0.594 Myr

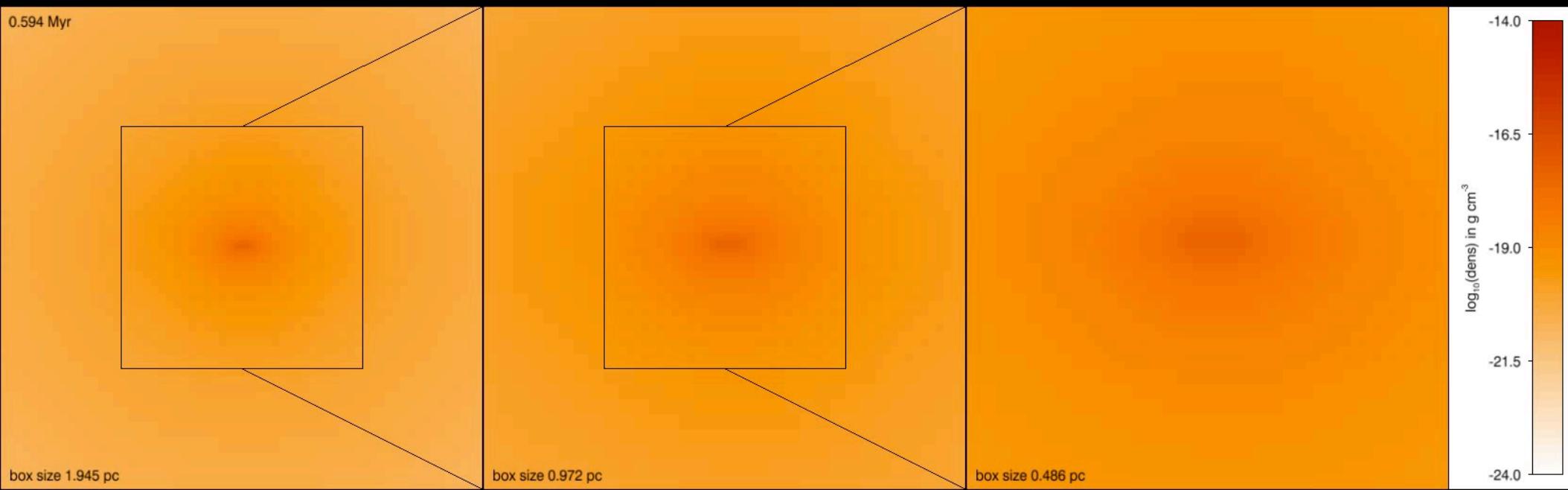
box size 0.324 pc

15-2

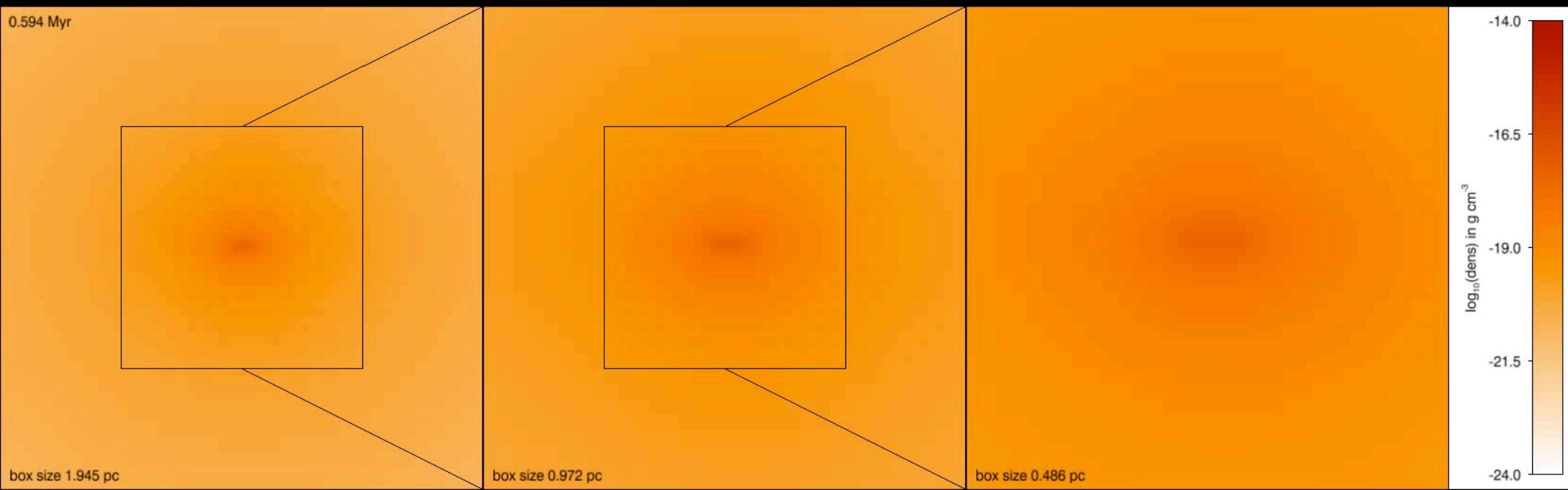
farmingdale15.key - April 24, 2015

$\log_{10}(\text{dens}) \text{ in g cm}^{-3}$





Peters + 14



Peters + 14

NEXT STEPS

- Update ionization to FERVENT (Baczynski +15), based on Wise & Abel MORAY algorithm (HEALPIX ray-tracing)
- Integrate N-body dynamics (ph4, McMillan) using the AMUSE framework (<http://amusecode.org/doc/>, Portegies-Zwart)
- Incorporate Glover chemistry module following lead of SILCC collaboration (Walch)

STAR FORMATION

GALACTIC STAR FORMATION

Juan Ibañez-Mejia (AMNH/ITA-ZAH)

Ralf Klessen (ITA-ZAH)

Andrea Gatto (MPA)

Stefanie Walch (U. Köln)

Christian Baczynski (ITA-ZAH)

Simon Glover (ITA-ZAH)

Alex Hill (Haverford)

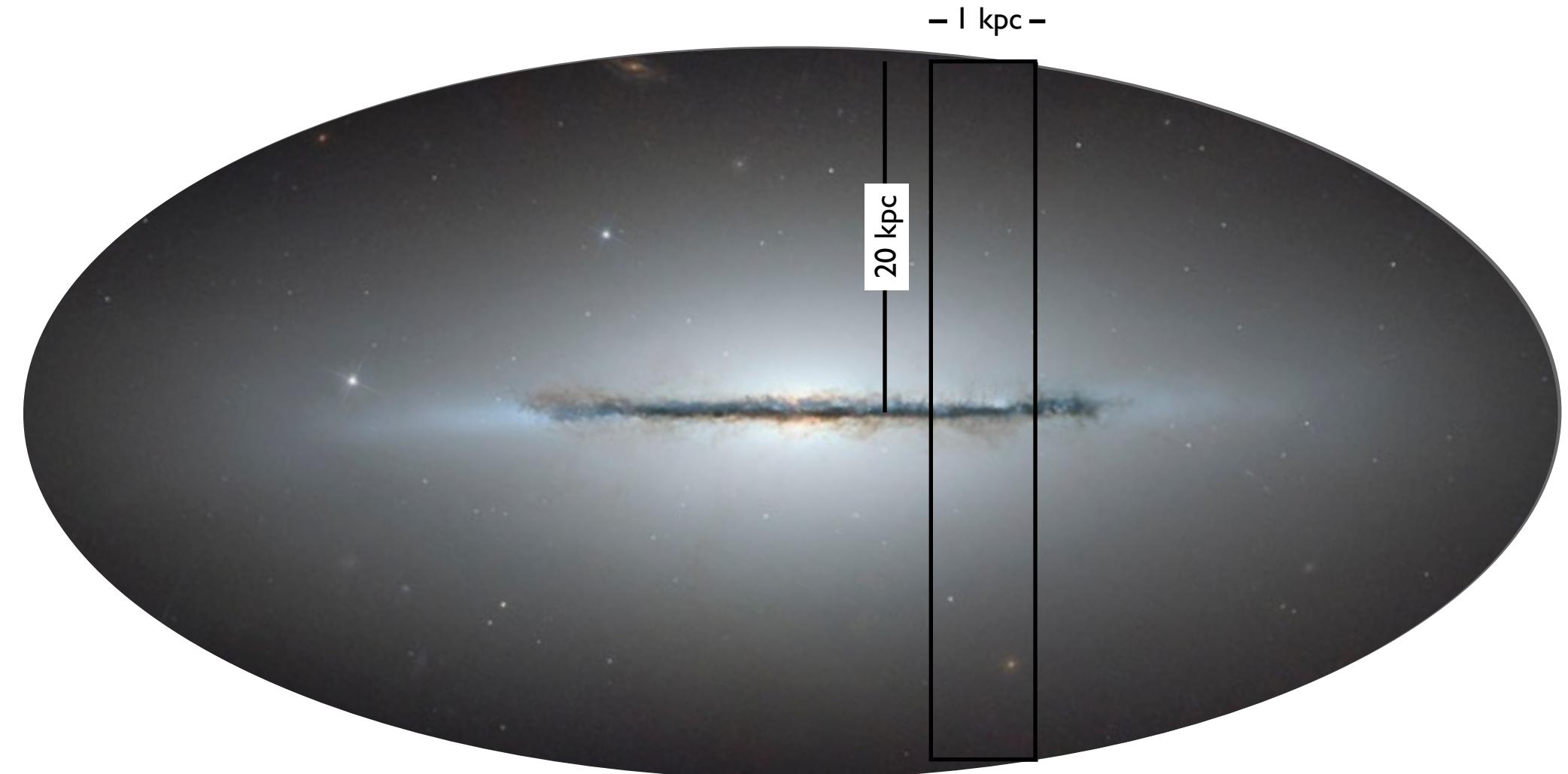
Hill + 12

250 pc

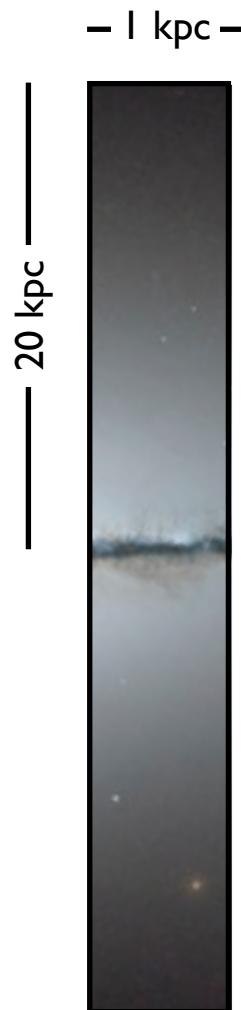
Our Model



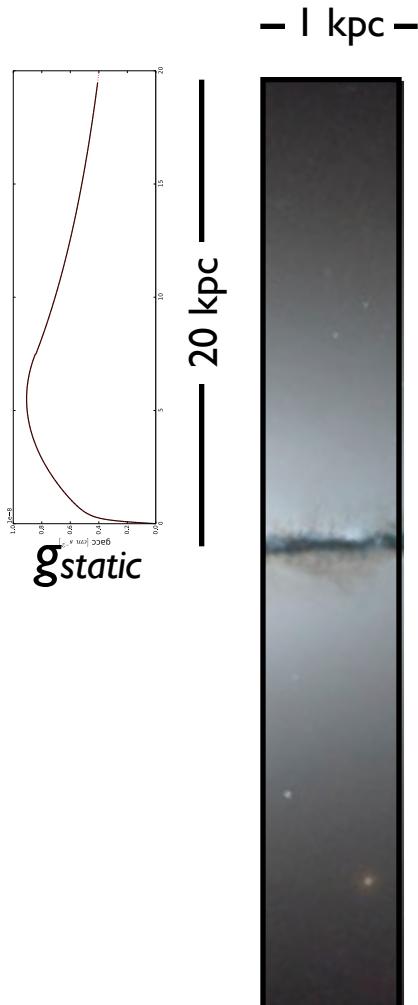
Our Model



Our Model

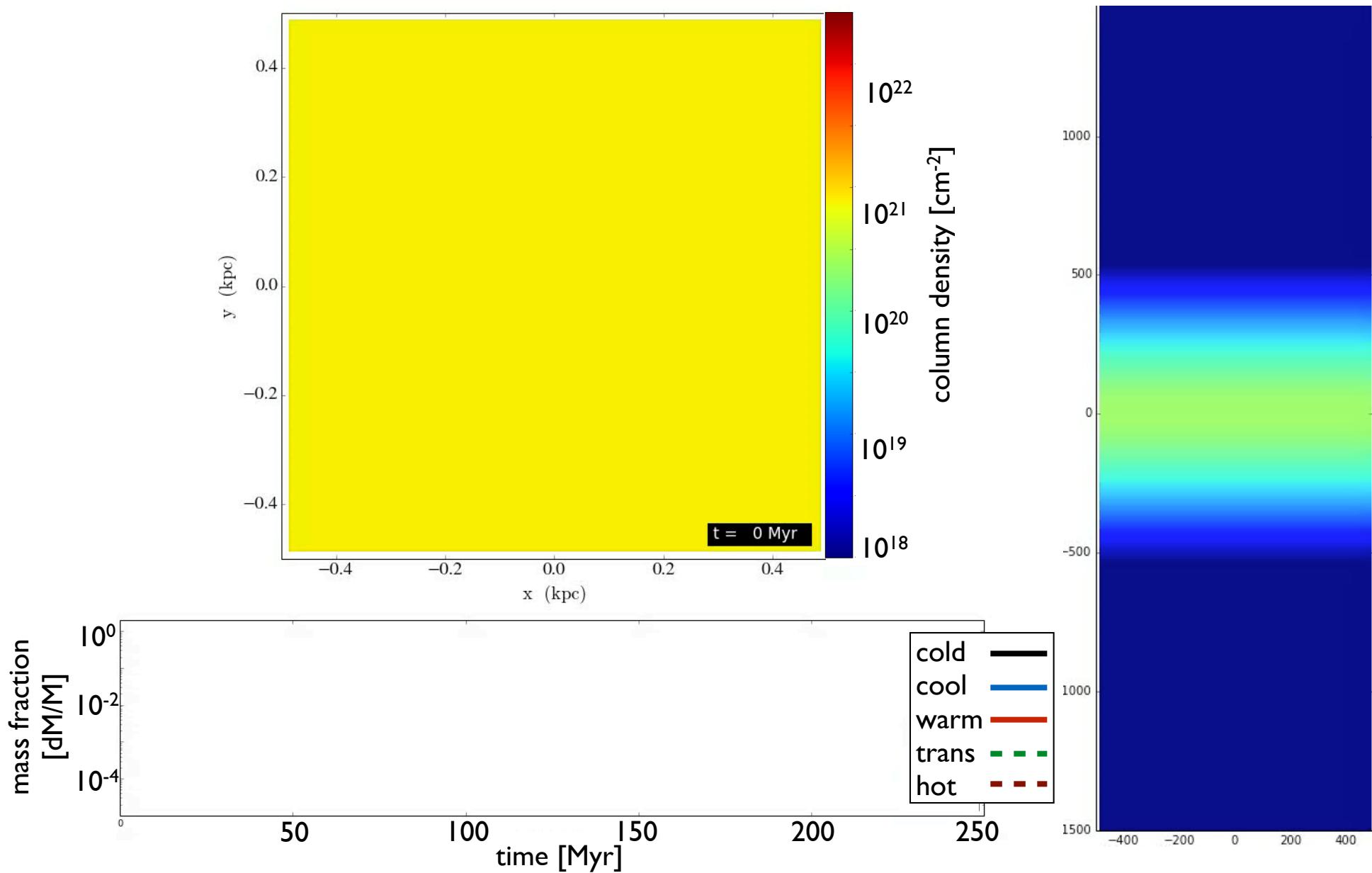


Our Model

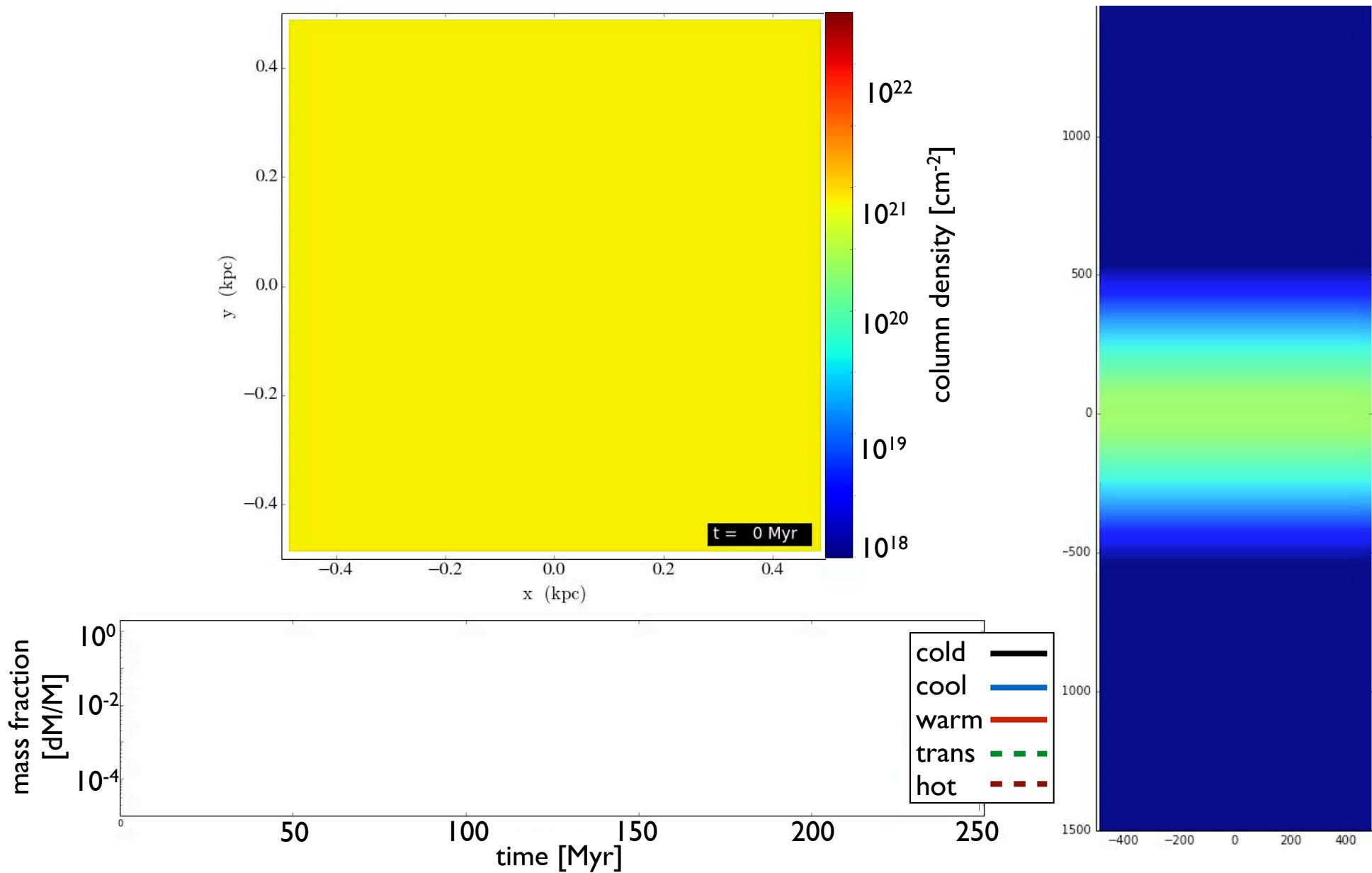


- Code: Flash 4.0
 - Eulerian grid based code
 - 3 dimensional MHD solver.
 - Adaptive mesh refinement
- Physics
 - Static gravitational potential (Gas + Stars + DM)
 - Random & correlated supernova (SN & SB)
 - Parameterized heating & cooling
 - Ideal magnetic fields
 - Self-gravity

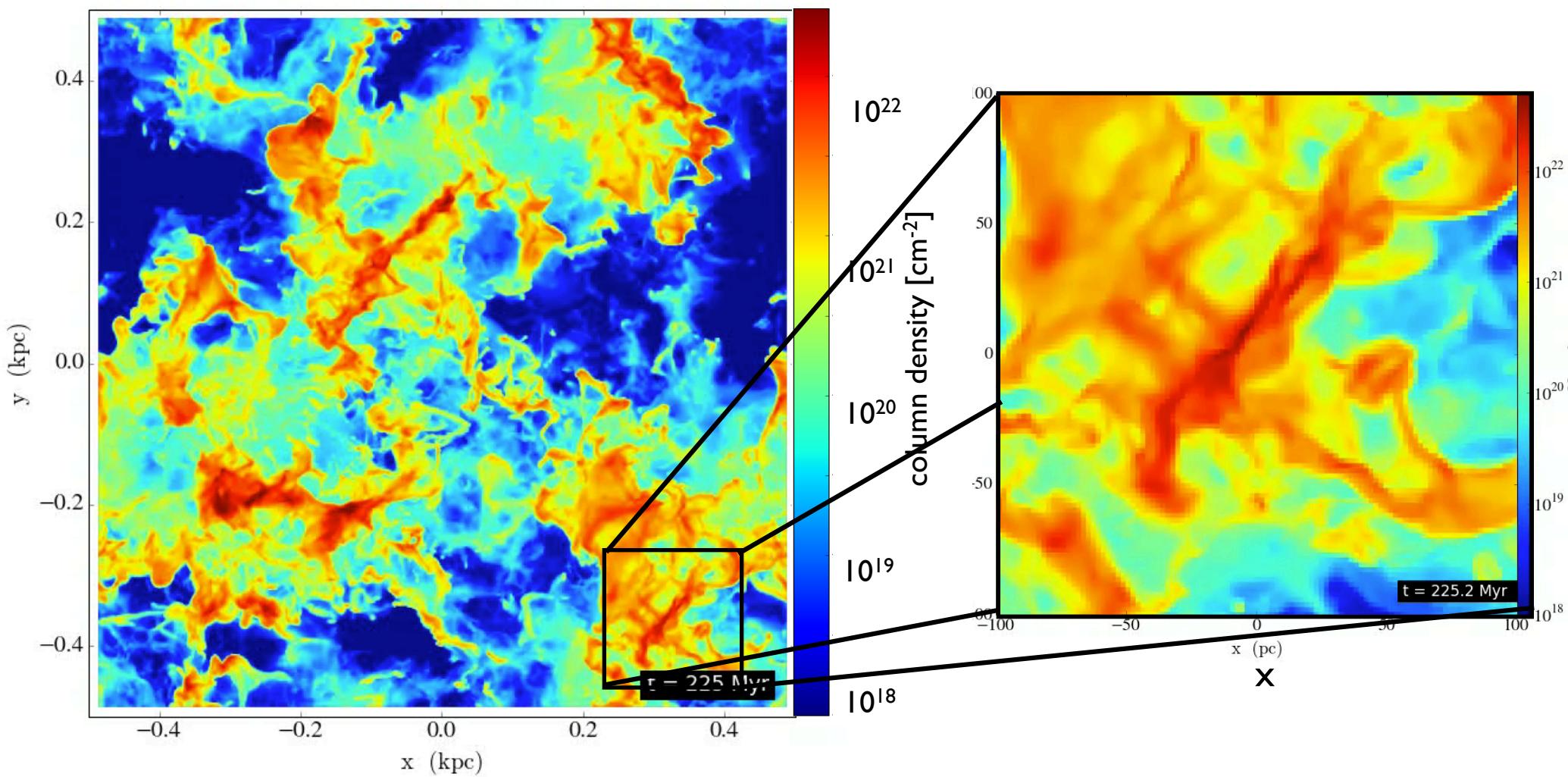
The Stratified, Magnetized, Turbulent ISM



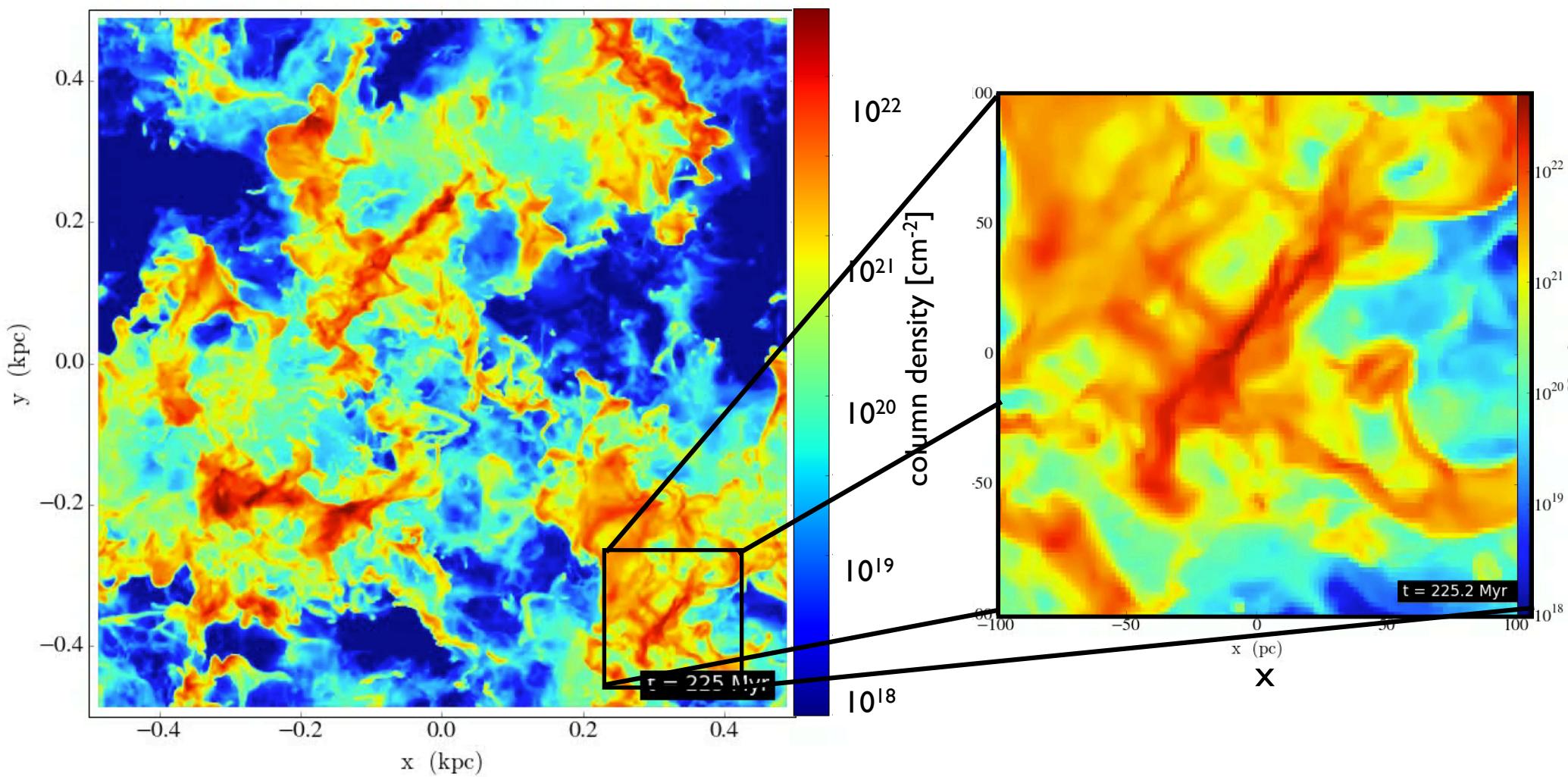
The Stratified, Magnetized, Turbulent ISM



The Action of Self Gravity

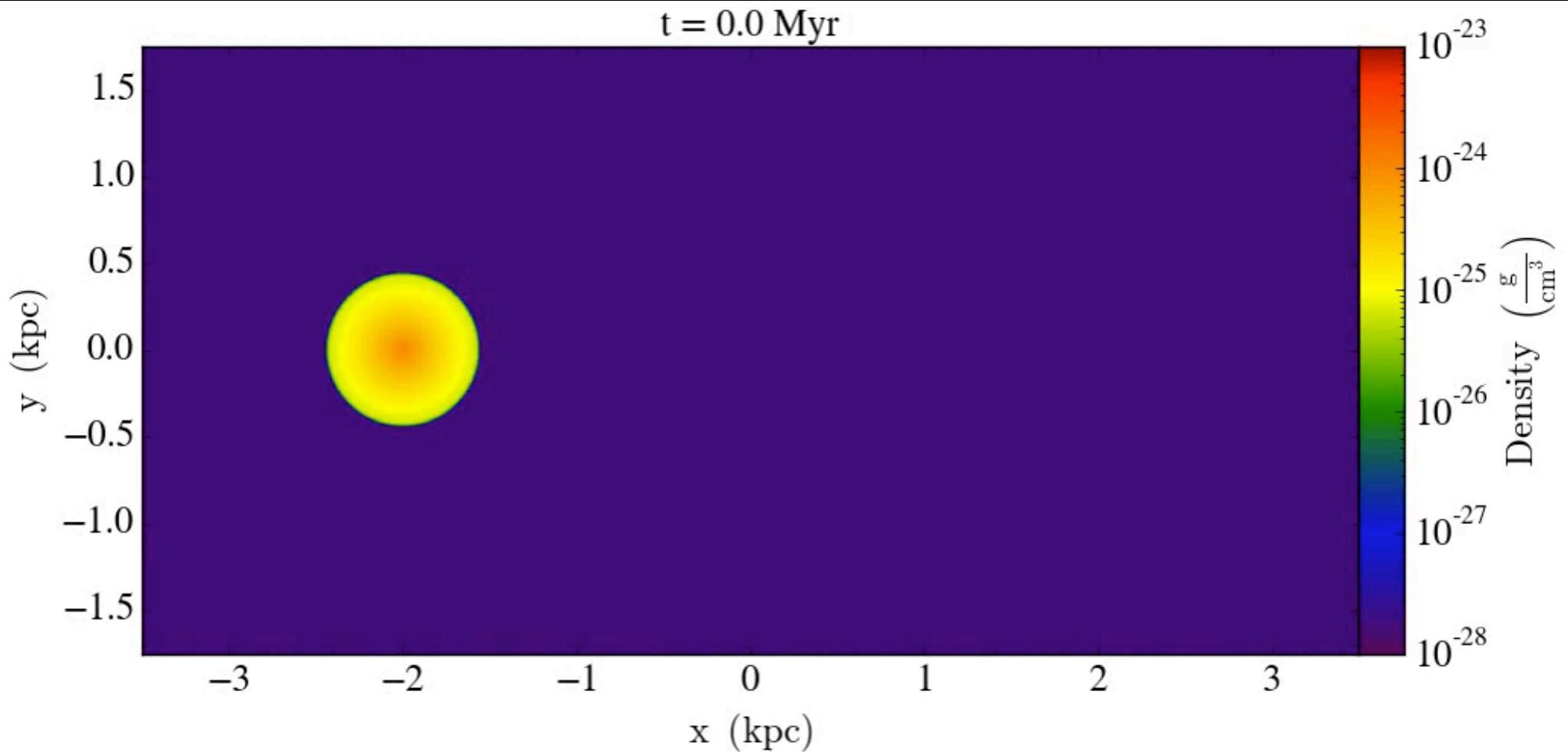


The Action of Self Gravity



STRIPPING OF DWARF GALAXIES

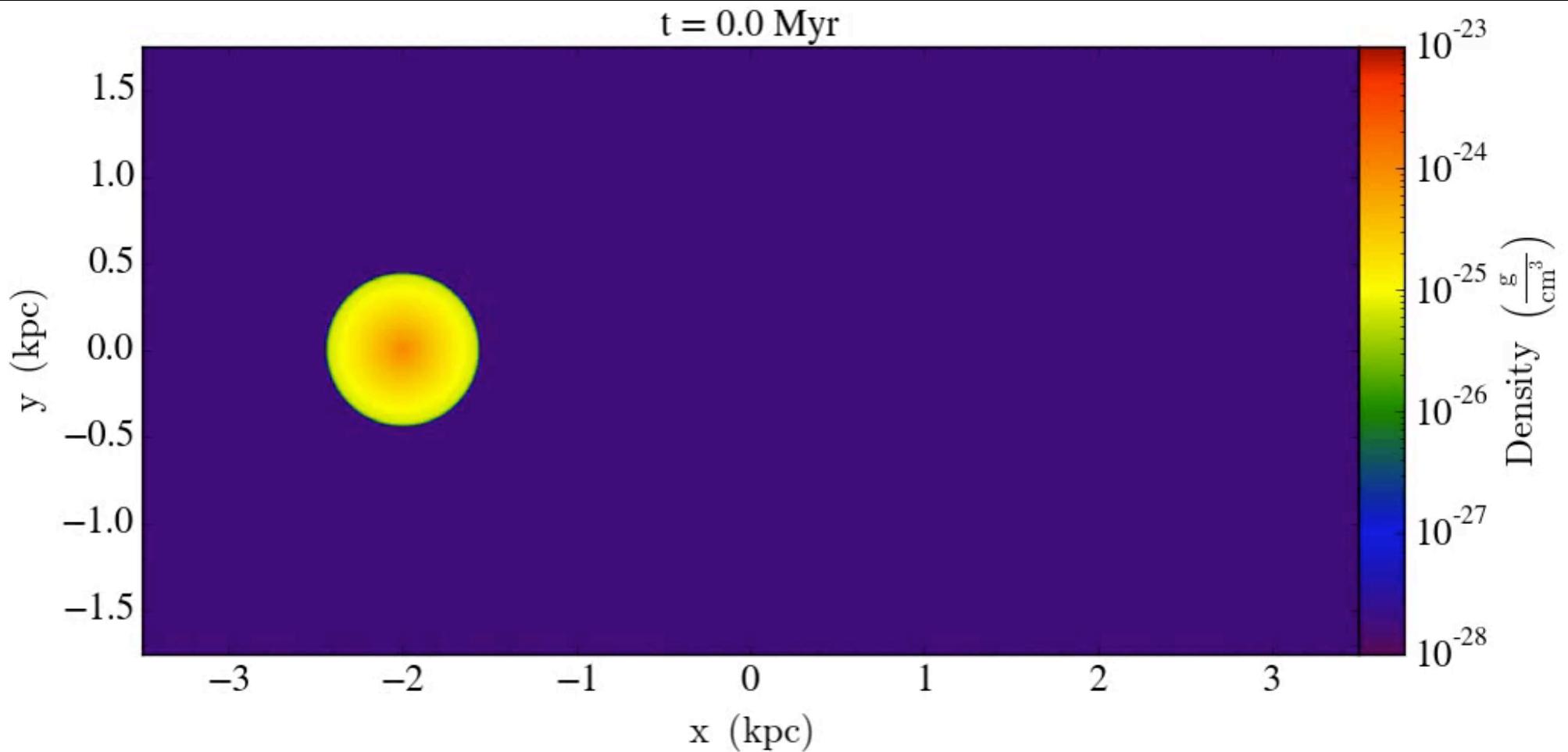
ANDREW EMERICK (COLUMBIA)
JANA GRCEVICH (AMNH)



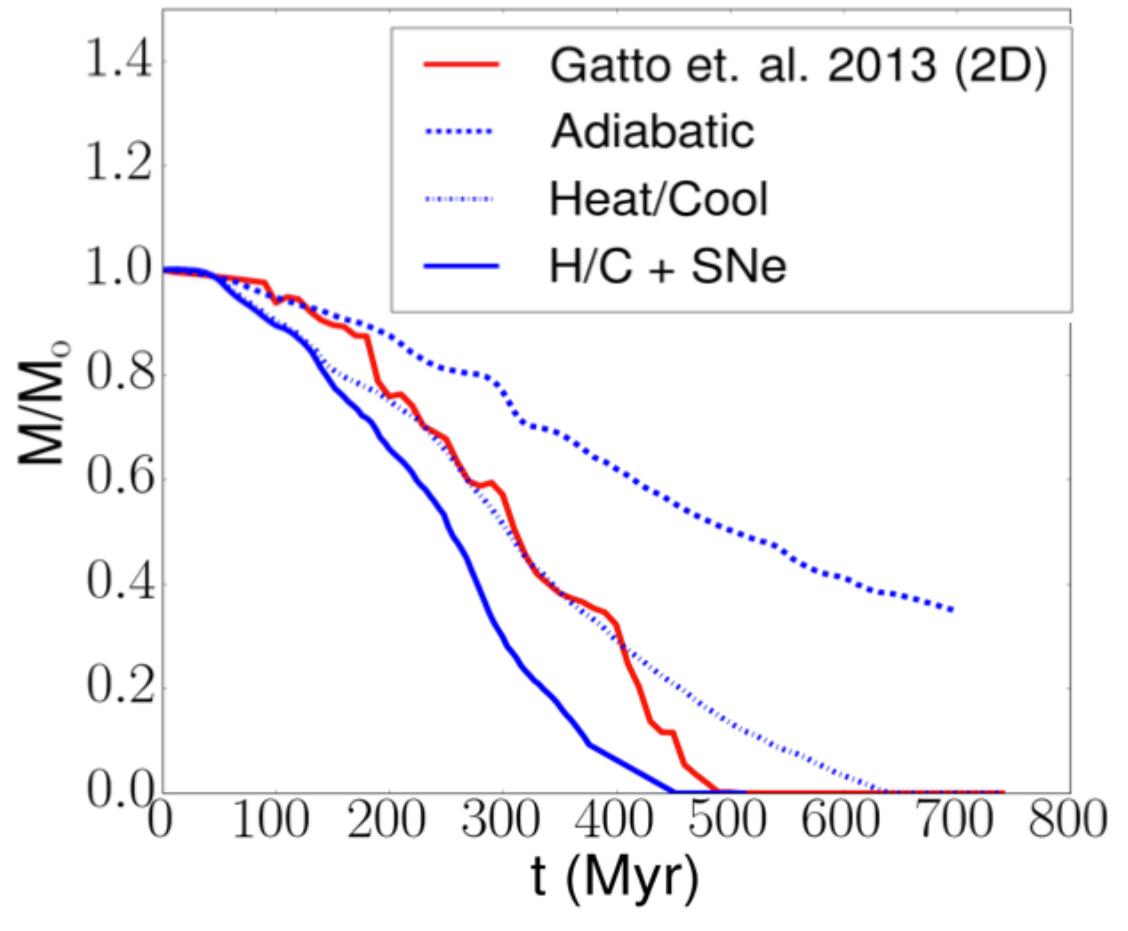
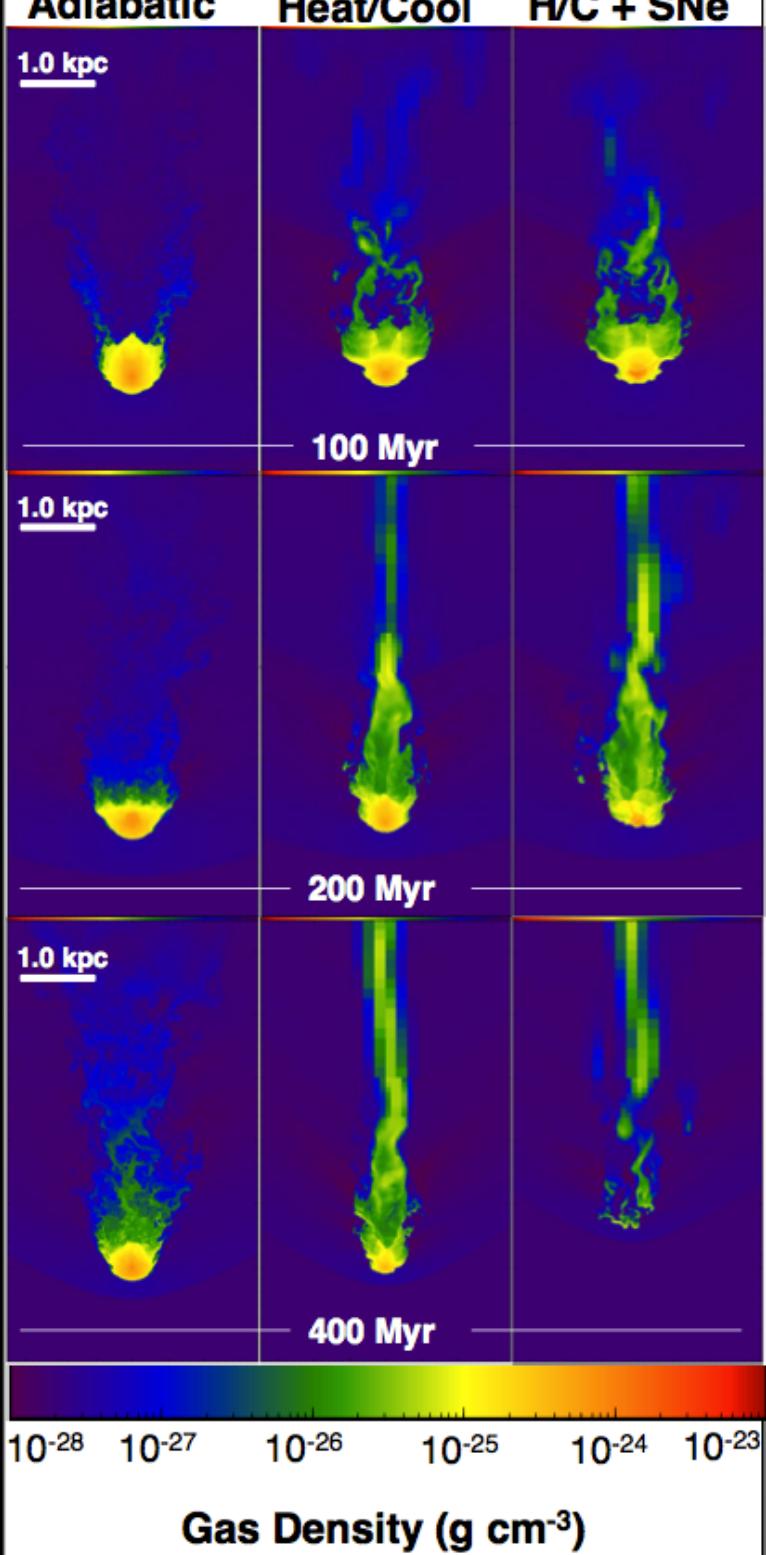
Emerick, Grcevich, Mac Low, Gatto 15

STRIPPING OF DWARF GALAXIES

ANDREW EMERICK (COLUMBIA)
JANA GRCEVICH (AMNH)

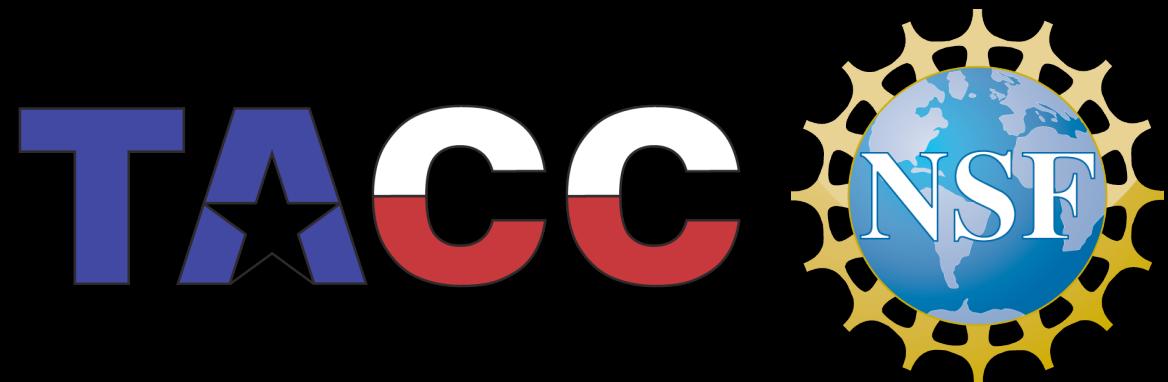


Emerick, Grcevich, Mac Low, Gatto 15



Emerick, Grcevich, Mac Low, Gatto 15

Computational Resources



Collaborators

People I have mentored in my group in orange

T. Abel (Stanford), J. Ballesteros-Paredes (CRyA, UNAM), R. Banerjee (U. Hamburg), R. A. Benjamin (UW Whitewater), E. Brinks (U. Herts.), C. Brunt (U. Exeter), G. Bryan (Columbia), P. Ciecielag (USM), P. D'Alessio (CRyA, UNAM), W. de Blok (U. Cape Town), O. De Marco (Macquarie), C. De Pree (Agnes Scott), S. Diehl (unaffiliated), L. Dursi (CITA), C. Dullemond (U. Heidelberg), D. S. Ebel (AMNH), C. Federrath (Monash), A. Ferrara (Pisa), C. L. Fryer (LANL), A. Fujita (Shinshu U.), R. Galván-Madrid (ESO), G. García-Segura (UNAM), S. Glover (ITA, U. Heidelberg), M. Gritschneider (USM), L. M. Haffner (UW Madison), Z. Haiman (Columbia), F. Heitsch (UNC), D. B. Henley (U. Ga.), P. Hennebelle (ENS Paris), F. Herwig (Victoria), M. Heyer (UMass Amherst), A. Hill (CSIRO), B. Horn (unaffiliated), A. Hubbard (AMNH), A.-K. Jappsen (unaffiliated), A. Johansen (Lund), K. V. Johnston (Columbia), M. R. Joung (Columbia), R. Kennicutt (Cambridge), E. Keto (CfA), Jongsoo Kim (KASSI), S. Kitsionas (Amer. School, Athens), H. Klahr (MPIfA), P. D. Klaassen (Leiden), R. Klessen (ZAH, Uni. Heidelberg), C. Klingenberg (U. Würzburg), M. Korpi (Helsinki), M. Krumholz (UCSC), K. Kwak (U. Ga.), R. Larson (Yale), Yuexing Li (Penn State), W. Lyra (JPL), G. Madsen (UW Madison), J. Maron (unaffiliated), C. Martin (UCSB), G. Mathews (Notre Dame), C. P. McNally (NBIA, Copenhagen), K. Menou (Toronto), M. Moe (Harvard), K. C. B. New (LANL), J. Oishi (AMNH/Farmingdale), S.-J. Paardekooper (Cambridge), J.-C. Passy (AlfA, Bonn), B. Paxton (KITP/UCSB), T. Peters (U. Zürich), R. Piontek (Astr. Inst. Potsdam), D. Price (Monash), R. Reynolds (UW Madison), G. Rockefeller (LANL), H.-W. Rix (MPIfA), J. Roman-Duval (STScI), S. Sharma (U. Sydney), Z. Sándor (U. Innsbrück), E. Sandquist (San Diego State), W. Schmidt (Göttingen), M. Shara (AMNH), R. L. Shelton (U. Ga.), C. M. Simpson (Columbia), B. D. Smith (Michigan State), D. Tamburro (unaffiliated), S. Tang (Mass. Gen. Hospital), J. Toraskar (AMNH), L. Townsley (Penn State) and the *Chandra Carina Complex Project* collaboration, J. Tumlinson (STScI), A. Ud-Doula (Penn State Worthington Scranton), E. Vázquez-Semadeni (UNAM), K. Waagan (UW Seattle), F. Walter (MPIfA), S. Walsh (Uni. Köln), Q. Wang (U. Mass. Amherst), R. Weaver (LANL), D. Wilner, K. Wood (St. Andrews), C.-C. Yang (Lund), A. Youdin (CfA).