

Astronomy Formulas

Radiation

	$m_B - m_A = 2.5 \log \left(\frac{I_A}{I_B} \right)$	
Intensity Ratio	$\frac{I_A}{I_B} = 10^{0.4(m_B - m_A)}$	$m = \text{Magnitude}$ $I = \text{Intensity}$
Wien's Law	$\lambda_{peak} = \frac{b}{T}$	$b = 2.898 \cdot 10^6 \text{ nm} \cdot \text{K}$ $T = \text{Temperature [K]}$
Stefan Boltzmann Law	$F = \sigma T^4$ $L = 4\pi\sigma R^2 T^4$ $\frac{L}{L_\odot} = \left(\frac{R}{R_\odot} \right)^2 \left(\frac{T}{T_\odot} \right)^4$	$F = \text{Flux (luminosity/area) [W/m}^2]$ $\sigma = 5.6704 \cdot 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$ $T = \text{Temperature [K]}$ $L = \text{Luminosity [W]}$ $R = \text{Radius [m]}$
Redshift	$z + 1 = \frac{\lambda_O}{\lambda_E}$	$z = \text{Redshift}$ $\lambda_O = \text{Observed wavelength}$ $\lambda_E = \text{Emitted wavelength}$
Doppler Shift	$v = cz$	$v = \text{Velocity of source [m/s]}$ $c = \text{Speed of light [m/s]}$ $z = \text{Redshift}$
Inverse Square	$I = \frac{L}{4\pi d^2}$	$I = \text{Light intensity [W/m}^2]$ $L = \text{Luminosity [W]}$ $d = \text{Distance [m]}$
Distance Modulus	$m - M = -5 + 5 \log(d)$ $d = 10^{\frac{m-M+5}{5}}$	$m = \text{Apparent magnitude}$ $M = \text{Absolute magnitude}$ $d = \text{Distance [pc]}$

Telescopes and Optics

Small Angle Formula	$\theta = 206265 \frac{D}{d}$	$\theta = \text{Angular diameter [arcsec]}$ $D = \text{Diameter}$ $d = \text{Distance}$
Resolving Power	$\Delta\theta_{arcsec} = 251643 \frac{\lambda}{d}$ $\Delta\theta_{rad} = 1.22 \frac{\lambda}{d}$	$\Delta\theta = \text{Angular separation [arcsec, rad]}$ $\lambda = \text{Wavelength}$ $d = \text{Optical diameter}$
Compare LGP	$\frac{LGP_A}{LGP_B} = \left(\frac{D_A}{D_B} \right)^2$	$LGP = \text{Light gathering power}$ $D = \text{Optic diameter}$
Magnification	$M = \frac{F_O}{F_E}$	$M = \text{Magnification power}$ $F_O = \text{Objective focal length}$ $F_E = \text{Eyepiece focal length}$
F Ratio	$\frac{L_f}{D_O}$	$L_f = \text{Focal Length}$ $D_O = \text{Objective Diameter}$
Parallax	$d = \frac{1}{p}$	$d = \text{Distance [pc]}$ $p = \text{Parallax angle [arcsec]}$

Orbital Mechanics

Gravitation	$F_G = G \frac{m_1 m_2}{r^2}$	F_G = Gravitational force [N] m = Mass [kg] r = Distance [m] $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$
Vis Viva Equation	$v = \sqrt{GM \left(\frac{2}{r} - \frac{1}{a} \right)}$	v = Velocity [m/s] $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$ M = Mass [kg] r = Distance [m] a = Semimajor axis [m]
Circular Velocity	$v_{circ} = \sqrt{\frac{GM}{r}}$	v = Velocity [m/s] $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$ M = Mass [kg] r = Distance [m]
Escape Velocity	$v_{esc} = \sqrt{\frac{2GM}{r}}$	v = Velocity [m/s] $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$ M = Mass [kg] r = Distance [m]
Ellipse Equations	$r = \frac{a(1 - e^2)}{1 + e \cos(\theta)}$ $b^2 = a^2(1 - e^2)$ $\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$	a = semimajor axis b = semiminor axis e = eccentricity θ = angle from periaxis
Eccentricity	$e = \frac{c}{a}$	a = semimajor axis c = linear eccentricity = distance from foci to center
Apoapsis/Periapsis	$r_a = a(1 + e)$ $r_p = a(1 - e)$	r = distance a = semimajor axis e = eccentricity
Kepler's Second Law	$\frac{dA}{dt} = \frac{\pi ab}{T}$	A = area a = semimajor axis b = semiminor axis T = period
Kepler's Third Law	$\frac{a^3}{T^2} = \frac{G(M + m)}{4\pi^2}$	a = semimajor axis T = period M = larger mass m = smaller mass $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$
Virial Theorem	$\langle T \rangle = -\frac{1}{2} \langle U \rangle$ $\langle T \rangle = -\frac{1}{2} \sum_{k=1}^N \langle \mathbf{F}_k \cdot \mathbf{r}_k \rangle$	$\langle T \rangle$ = Time averaged kinetic energy $\langle U \rangle$ = Time averaged potential energy
Binary System	$a = r_1 + r_2$ $r_1 = a \frac{m_2}{m_1 + m_2}$ $\frac{m_1}{m_2} = \frac{\alpha_2}{\alpha_1} = \frac{r_2}{r_1} = \frac{v_2}{v_1}$	r = Distance from barycenter a = semimajor axis α = Angular distance from barycenter m = mass v = radial velocity

Binary Mass Function	$K = v_1 \sin i = \omega_{orb} r_1 \sin i$	i = orbital inclination = degrees from face-on
	$f = \frac{M_2^3 \sin^3 i}{(M_1 + M_2)^2} = \frac{P_{orb} K^3}{2\pi G}$	f = Mass function [kg] M = mass [kg]
	$\omega_{orb} = \frac{2\pi}{P_{orb}}$	P_{orb} = Orbital period [s] K = peak radial velocity [rad/s] = semiamplitude of rad. vel. curve

ω_{orb} = orbital frequency
 $G = 6.67 \cdot 10^{-11} \text{ m}^3 \cdot \text{kg}^{-1} \cdot \text{s}^{-2}$

Stars

Stellar Lifetime	$\tau \approx 10^{10} M^{-2.5}$	M = Mass (solar masses) τ = Stellar lifetime (years)
Mass-Luminosity Relation	$L \approx M^{3.5}$	M = Mass (solar masses) L = Luminosity (solar luminosity)
RR Lyrae P-L Rel.	$M_k = -2.34 \log(P) - 0.88$	M_K = Absolute K-band IR magnitude P = Period [days]
RV Tau P-L Rel.	$M = 7.36 \log(P) + 10.21$	M_v = Absolute magnitude P = Period [days]
Type I Ceph. P-L Rel.	$M_v = -2.43(\log_{10}(P) - 1) - 4.05$	M_v = Absolute visual magnitude P = Period [days]
Type II Ceph. P-L Rel.	$M_v = -2.43(\log_{10}(P) - 1) - 2.45$	M_v = Absolute visual magnitude P = Period [days]

Cosmology and Relativity

Relativistic Doppler Shift	$z + 1 = \sqrt{\frac{1 + (v/c)}{1 - (v/c)}}$	v = Velocity of source [m/s] c = Speed of light [m/s] z = Redshift
Hubble's Law	$v = H_0 d$	v = Recessional Velocity $H_0 \approx 70 \text{ km/s/Mpc}$ d = distance
Hubble Time	$T_U = \frac{9.778 \cdot 10^{11}}{H_0}$	T_U = Age of universe $H_0 \approx 70 \text{ km/s/Mpc}$
Schwarzschild Radius	$r_s = \frac{2GM}{c^2}$	r_s = Schwarzschild radius G = gravitational constant M = mass c = speed of light

Time Dilation	$t = \frac{t_0}{\sqrt{1 - \frac{v^2}{c^2}}}$
Length contraction	$L = L_0 \sqrt{1 - \frac{v^2}{c^2}}$

Climate and Exoplanets

Planetary Equilibrium Temperature	$T_{eq} = \left(\frac{I_0(1 - A_B)}{4\sigma} \right)^{1/4}$ $I_x = \frac{L}{4\pi a^2}$ $T_{eq} = T_{star} \sqrt{\frac{R}{2a}} (1 - A_B)^{1/4}$	A_B = bond albedo $\sigma = 5.6704 \cdot 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$ a = semimajor axis [m] L = luminosity [W/m^2]
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Radiation Balance	$ASR = (1 - A_B) \cdot 340 \approx 240 W/m^2$	$\varepsilon = \text{emissivity}$
	$OLR \approx \varepsilon \sigma T_{atm}^4 + (1 - \varepsilon) \sigma T_{surf}^4$	$ASR = \text{absorbed solar radiation}$
		$OLR = \text{outgoing longwave radiation}$
		$\sigma = 5.6704 \cdot 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$
	$\Delta F = I_0 \cdot (1 - R) \cdot \Delta \tau = 238 \cdot \Delta \tau$	
Radiative Forcing	$\Delta F = -I_0 \cdot R \cdot \Delta A_B = -102 \cdot \Delta A_B$	$\Delta F = \text{radiative forcing [W m}^2]$
	$\Delta F_{CO2} = 5.35 \cdot \ln \frac{C_0 + \Delta C}{C_0}$	$\Delta \tau = \Delta \text{ solar irradiance [W/m}^2]$
	$\Delta T_s = \lambda \Delta F$	$\lambda = \text{climate sens. param. [K m/W}^2]$
Emissivity	$\varepsilon = \frac{M_e}{M_e^\circ}$	$\varepsilon = \text{emissivity}$
		$M_e = \text{Radiant exitance [W} \cdot \text{m}^{-2}]$
		$M_e^\circ = \text{Radiant exitance of blackbody}$
Extinction	$\Delta F = \left(\frac{R_p}{R_s} \right)^2$	
Albedo		
Habitable Zone	$r = \sqrt{\frac{(1 - A)L}{16\pi\sigma T^4}}$	$T = 273 - 373 K$
	$T_s^4 = \frac{Q(1 - A)}{2\sigma(2 - \varepsilon)}$	

Astronomy Constants and Conversions

Constants

Fundamental Constants

Speed of light = $c = 3 \cdot 10^8$ m/s

Gravitational constant = $G = 6.67 \cdot 10^{-11}$ m³ · kg⁻¹ · s⁻²

Mass of a proton = $m_p = 1.6726 \cdot 10^{-27}$ kg

Mass of an electron = $m_{e^-} = 9.109 \cdot 10^{-31}$ kg

Charge of an electron = $e = 1.6022 \cdot 10^{-19}$ C

Hubble's Constant = $H_0 = 70$ km/s/Mpc

Stefan-Boltzmann Constant = $\sigma = 5.67 \cdot 10^{-8}$ W/m²K⁴

Planck's Constant = $h = 6.626 \cdot 10^{-34}$ J s

Reduced Planck's Constant = $\hbar = h/(2\pi) = 1.05457 \cdot 10^{-34}$ J s

Boltzmann Constant = $k = k_B = 1.380 \cdot 10^{-23}$ m² kg/s²K

Wein's Displacement Constant = $b = 2.898 \cdot 10^{-3}$ m · K

Molar gas constant = $R = 8.315$ J/mol/K

Rydberg constant = $R_H = 1.0968 \cdot 10^7$ m⁻¹

Solar System

Mass of Earth = $M_\oplus = 5.97 \cdot 10^{24}$ kg

Radius of Earth = $R_\oplus = 6378$ km

Mass of the Moon = $M_\odot = 7.346 \cdot 10^{22}$ kg

Radius of the Moon = $R_\odot = 1738.1$ km

Mass of Jupiter

Radius of Jupiter

Mass of Sun = $M_\odot = 1.99 \cdot 10^{30}$ kg

Radius of Sun = $R_\odot = 6.96 \cdot 10^5$ km

Temperature of Sun = $T_\odot = 5778$ K

Luminosity of Sun = $L_\odot = 3.9 \cdot 10^{26}$ W

Absolute magnitude of Sun = $M_\odot = 4.83$

SGP of Sun = $\mu_\odot = GM_\odot = 1.327 \cdot 10^{20}$ m³ · s⁻²

Solar metallicity = $Z_\odot = 0.0134$

Miscellaneous

H α spectral line = 656.28 nm

Type Ia supernova absolute magnitude = -19.3

Chandrasekhar limit = $M_{\text{limit}} \approx 1.4M_\odot$

TOV limit = $M_{\text{limit}} \approx 2.1 M_\odot$

Conversions

Distance/Area

1 AU = $1.5 \cdot 10^8$ km

1 pc = $3.09 \cdot 10^{16}$ m = 3.26 ly

1 ly = $9.46 \cdot 10^{15}$ m

1 m = 3.28 ft

1 in = 2.54 cm

1 mi = 5280 ft

1 Å = 0.1 nm

1 fermi = 1 fm = 10^{-15} m

1 barn = 1 b = 10^{-28} m²

1 acre = 4046.86 m²

Angles

1 arcmin = $1' = (\frac{1}{60})^\circ$

1 arcsec = $1'' = (\frac{1}{60})' = (\frac{1}{3600})^\circ$

Time

1 year = 31557600 s

Energy

1 Megaton of TNT = $4.184 \cdot 10^{15}$ J

1 electron volt = 1 eV = $1.6 \cdot 10^{-19}$ J

Temperature

°C = K - 273.15

°F = $\frac{9}{5}^\circ\text{C} + 32$

CGS Units

1 gal = 10^{-2} m/s²

1 dyne = 10^{-5} N

1 erg = 10^{-7} J

1 barye = 10^{-1} Pa

Planck Units

Planck temperature = $T_p = \sqrt{\frac{\hbar c^5}{G k_B^2}} = 1.4168 \cdot 10^{32}$ K

Planck time = $t_p = \sqrt{\frac{\hbar G}{c^5}} = 5.391 \cdot 10^{44}$ s

Planck length = $\sqrt{\frac{\hbar G}{c^3}} = 1.61626 \cdot 10^{-35}$ m

Yotta (Y)	10^{24}	Deci (d)	10^{-1}
Zetta (Z)	10^{21}	Centi (c)	10^{-2}
Exa (E)	10^{18}	Milli (m)	10^{-3}
Peta (P)	10^{15}	Micro (μ)	10^{-6}
Tera (T)	10^{12}	Nano (n)	10^{-9}
Giga (G)	10^9	Pico (p)	10^{-12}
Mega (M)	10^6	Femto (f)	10^{-15}
Kilo (k)	10^3	Atto (a)	10^{-18}
Hecto (h)	10^2	Zepto (z)	10^{-21}
Deca (da)	10^1	Yocto (y)	10^{-24}