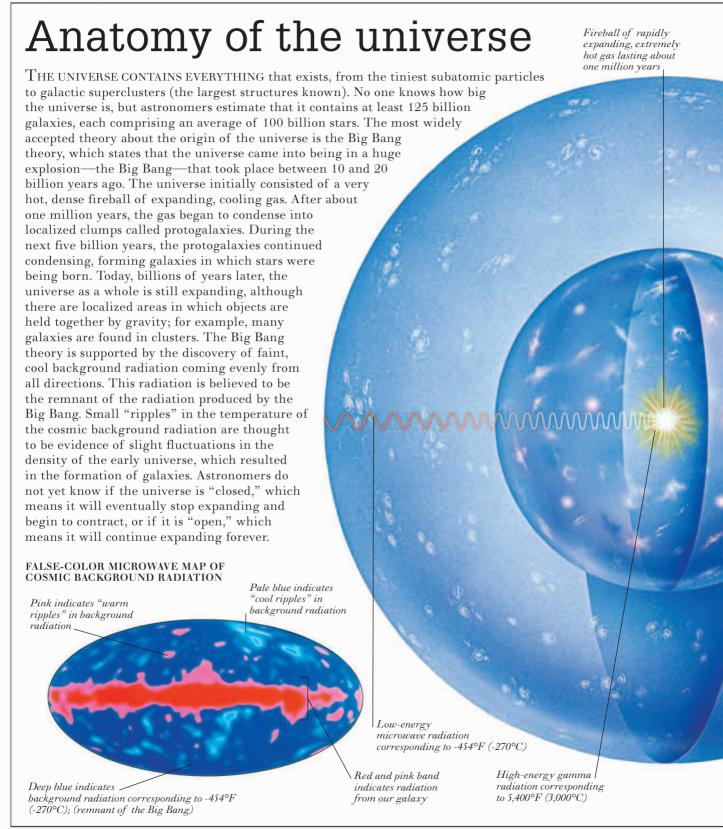


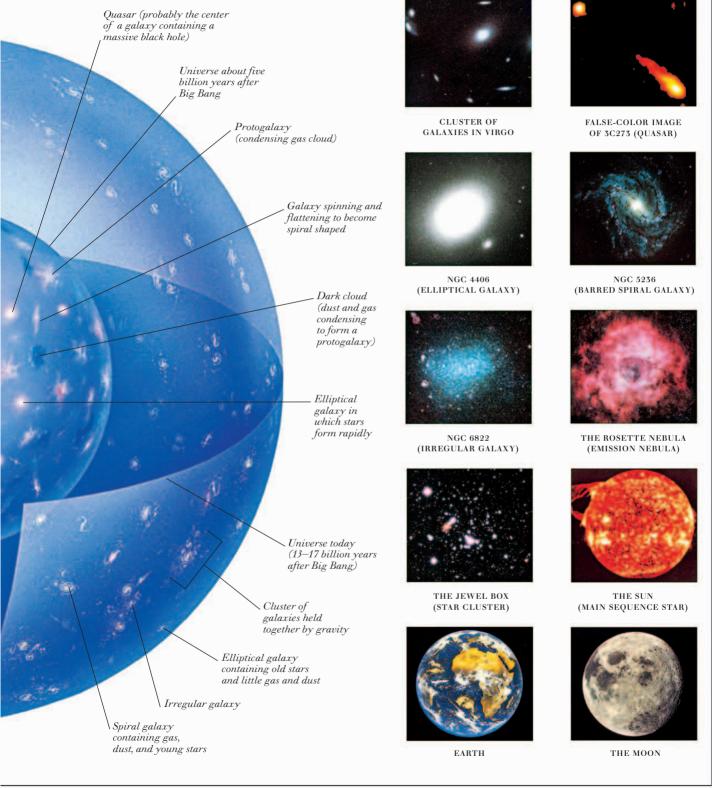
THE UNIVERSE

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ORIGIN AND EXPANSION OF THE UNIVERSE

OBJECTS IN THE UNIVERSE



Galaxies



SOMBRERO. A SPIRAL GALAXY

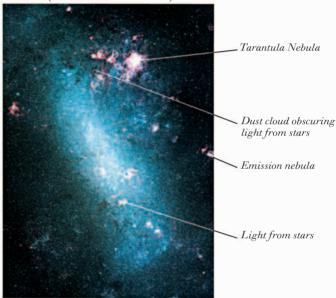
A GALAXY IS A HUGE MASS OF STARS, nebulae, and interstellar material. The smallest galaxies contain about 100,000 stars, while the largest contain up to 3 trillion stars. There are three main types of galaxy, classified according to their shape: elliptical, which are oval shaped; spiral,

which have arms spiraling

outward from a central bulge (those whose arms spiral from a bar-shaped bulge are called spirals); and irregular, which have no obvious shape. Sometimes, the shape of a galaxy is distorted by a collision with another galaxy. Quasars (quasistellar objects) are thought to be galactic nuclei but are so far away that their exact nature is still uncertain. They are compact, highly luminous objects in the outer reaches of the known universe: while the farthest known "ordinary" galaxies are about 12 billion light-years away, the farthest known quasar is about 13 billion lightyears away. Active galaxies, such as Seyfert galaxies and radio galaxies, emit intense radiation. In a Seyfert galaxy, this radiation comes from the galactic nucleus; in a radio galaxy, it also comes from huge lobes on either side of the galaxy. The radiation from active galaxies and quasars is thought to be caused by material falling into central black holes (see pp. 28-29).

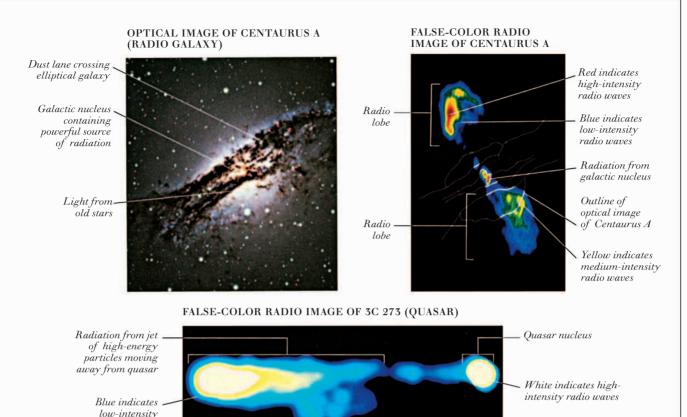
OPTICAL IMAGE OF NGC 4486 Globular cluster (ELLIPTICAL GALAXY) containing very old red giants Central region containing old red giants Less densely populated region Neighbouring galaxy

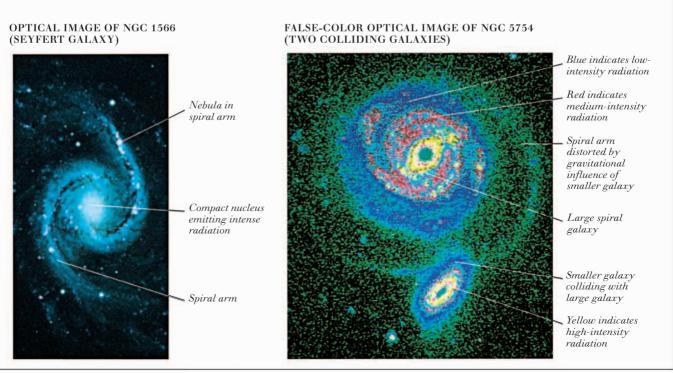
OPTICAL IMAGE OF LARGE MAGELLANIC CLOUD (IRREGULAR GALAXY)











radio waves

The Milky Way



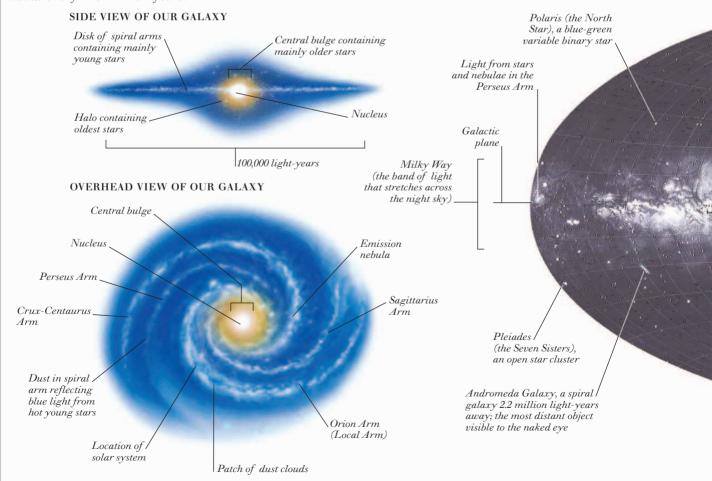
VIEW TOWARD GALACTIC CENTER

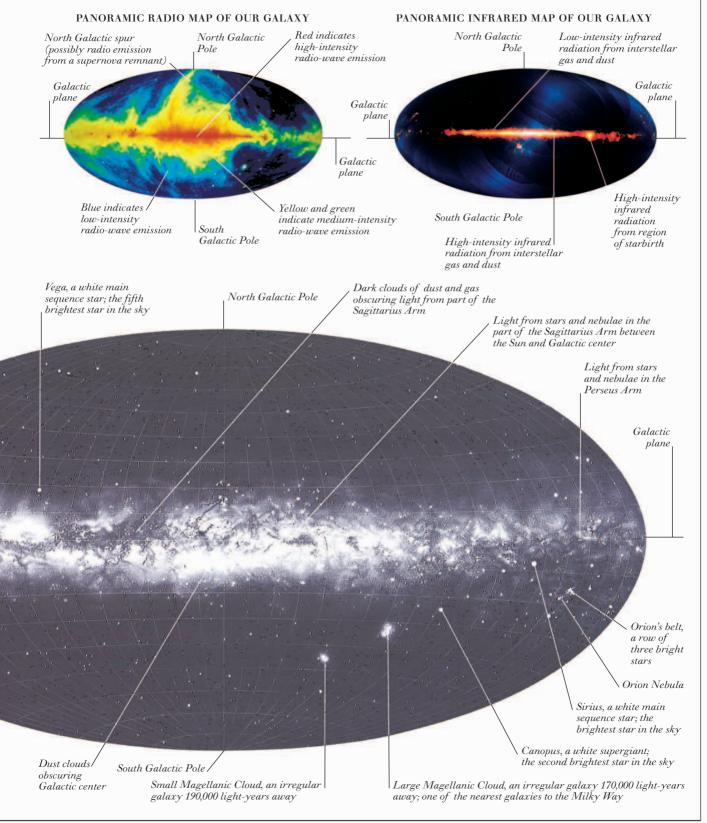
The Milky way is the name given to the faint band of light that stretches across the night sky. This light comes from stars and nebulae in our galaxy, known as the Milky Way Galaxy or simply as "the Galaxy." The Galaxy is believed to be a barred spiral, with a dense central bar of stars encircled by four arms spiraling outward and surrounded by a less dense halo. We cannot see the spiral shape because the solar system is in one of the spiral arms, the Orion Arm (also called the Local Arm). From our position, the center of the Galaxy is completely obscured by dust clouds; as a result, optical maps give only a limited view of the Galaxy. However, a more complete picture can be

obtained by studying radio, infrared, and other radiation. The central part of the Galaxy is relatively small and dense and contains mainly older red and yellow stars. The halo is a less dense region in which the oldest stars are situated; some of these stars are as old as the Galaxy itself (possibly 13 billion years). The spiral arms contain main sequence stars and hot, young, blue stars, as well as nebulae (clouds of dust and gas inside which stars are born). The Galaxy is vast, about 100,000 light-years across (a light-year is about 5,870 billion miles/9,460 billion km); in comparison, the solar system seems small, at about 12 light-hours across (about 8 billion miles/13 billion km). The entire Galaxy is rotating in space, although the inner stars travel faster than those farther out. The Sun, which is

about two-thirds out from the center, completes one lap of the Galaxy about every 220 million years.

PANORAMIC OPTICAL MAP OF OUR GALAXY AND NEARBY GALAXIES





Nebulae and star clusters

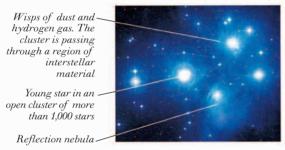


HODGE 11, A GLOBULAR CLUSTER

A NEBULA IS A CLOUD OF DUST AND GAS inside a galaxy. Nebulae become visible if the gas glows, or if the cloud reflects starlight or obscures light from more distant objects. Emission nebulae shine because their gas emits light when it is stimulated by radiation from hot young stars. Reflection nebulae shine because their dust reflects light from stars in or around the nebula. Dark nebulae appear as silhouettes because they block out light from shining nebulae or stars behind them. Two types of nebula are associated with dying stars: planetary nebulae and supernova remnants. Both consist of expanding shells of gas that were once the outer layers of a star. A planetary nebula is a gas shell drifting away

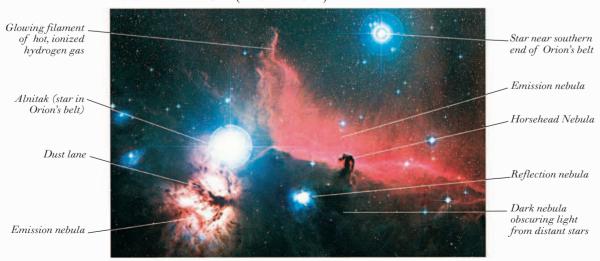
from a dying stellar core. A supernova remnant is a gas shell moving away from a stellar core at great speed following a violent explosion called a supernova (see pp. 26-27). Stars are often found in groups known as clusters. Open clusters are loose groups of a few thousand young stars that were born from the same cloud and are drifting apart. Globular clusters are densely packed, roughly spherical groups of hundreds of thousands of older stars.

PLEIADES (OPEN STAR CLUSTER) WITH A REFLECTION NEBULA



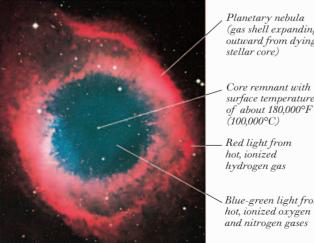
Reflection nebula Emission nebula Dust lane Starbirth region (area in which dust and gas clump together to form stars)

HORSEHEAD NEBULA (DARK NEBULA)



ORION NEBULA (DIFFUSE EMISSION NEBULA) Glowing Gas cloud cloud of dust emitting and hydrogen light due to gas forming ultravioletpart of Orion radiation from the four young Trapezium stars Nebula Dust cloud -Trapezium(group of four young stars) Green light from hot, ionized oxygen gas Glowing filament of Red light hot, ionized from hot, hydrogen gas ionized hydrogen gas





(gas shell expanding outward from dying

surface temperature of about 180,000°F

Blue-green light from

VELA SUPERNOVA REMNANT



Supernova remnant (gas shell consisting of outer layers of star thrown off in supernova explosion)

Hydrogen gas emitting red light due to being heated by supernova explosion

Glowing filament of hot, ionized hydrogen gas

