

Hyperbolic Plane Metrics

<https://github.com/Nazgand/nazgandMathBook>

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Abstract

The goal of this paper is to introduce the Polar Hyperbolic Plane Metric and the Cartesian Hyperbolic Plane Metric. These metrics are useful because they both map the entire Euclidean plane to the entire Hyperbolic plane.

1 Circumference as a Function of Radius and Curvature

Let $R = \frac{1}{\sqrt{-K}}$, where K is the Gaussian curvature of the Hyperbolic Plane. Then the circumference of a circle with radius r is

$$C(R, r) = 2\pi R \sinh\left(\frac{r}{R}\right) \quad (1.1)$$

However, looking ahead, the models we obtain in the general case are essentially scaled versions of the case $K = -1 \wedge R = 1$. E.g. distance d in a model where $R \in \mathbb{R}^+$ is isomorphic to distance $\frac{d}{R}$ in a model where $R = 1$.

2 Polar Hyperbolic Plane Metric

In this model, we desire the distance from the origin to be r , both in [the Euclidean sub-space the model is built on] and in [the Hyperbolic space of the model]. This model requires changing the density of space along the circles surrounding the origin by a factor of $\frac{C(R, r)}{2\pi r}$. I.e. The density perpendicular to every radius is multiplied by the Hyperbolic circumference and divided by the Euclidean circumference.

Let $(r, \theta) \in \mathbb{R}^2$ be the polar Euclidean coordinates, and let s be the Hyperbolic length along a Hyperbolic geodesic.

Thus, we arrive at the metric:

$$(ds)^2 = (1 * dr)^2 + \left(\frac{C(R, r)}{2\pi r} * r d\theta\right)^2 \quad (2.1)$$

Simplify, and let $R = 1$:

$$(ds)^2 = (dr)^2 + (\sinh(r) * d\theta)^2 \quad (2.2)$$

Note that, starting at the origin of the Euclidean subspace, turning angle θ , and moving forward distance r corresponds exactly with the same action in the Hyperbolic space.

3 Cartesian Hyperbolic Plane Metric

We play a similar game when creating this metric: Starting at the origin of the Euclidean subspace, moving 'east' distance x , and moving 'north' distance y corresponds exactly with the same action in the Hyperbolic space.

We can see that angles are preserved along the line $y = 0$ because the spatial density is the same in every direction. This fact could lead to interesting math.

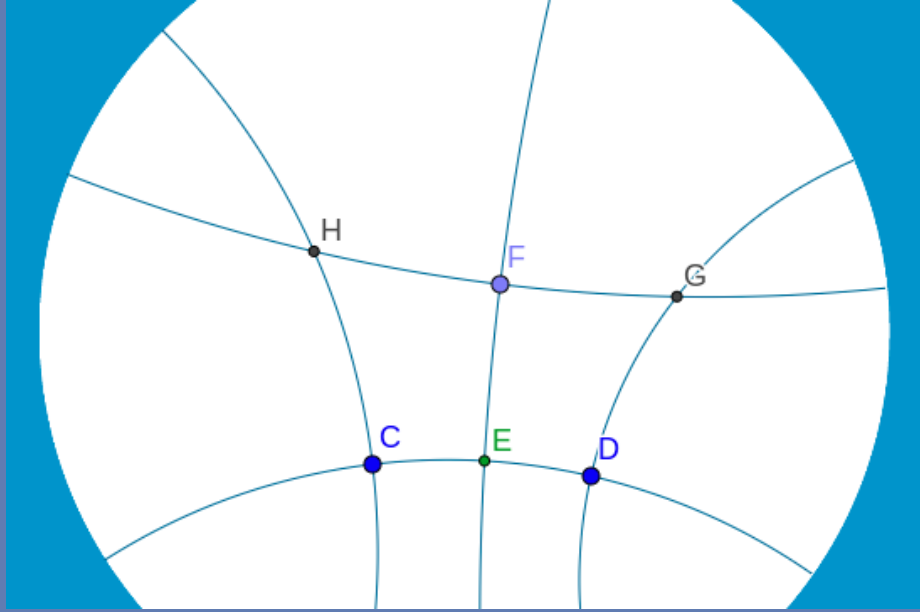


Figure 1: The Cartesian Hyperbolic Problem.

It becomes intuitively obvious that the spatial density in the horizontal direction is something like $\cosh(y)$, but to prove it, we have this problem where $C = (0, 0)$, $D = (0, \epsilon)$, $E = (0, \frac{\epsilon}{2})$, $H = (0, y)$, $G = (\epsilon, y)$ and F is the midpoint of H and G .

Noting $\angle GDE = \frac{\pi}{2}$, we have $\sin(\angle DEG) = \frac{\sinh(y)}{\sinh(\overline{EG})}$,

and $\tan(\angle DEG) = \frac{\tanh(y)}{\sinh(\frac{\epsilon}{2})}$. Thus

$$\overline{EG} = \operatorname{arsinh} \left(\frac{\sinh(y)}{\sin \left(\arctan \left(\frac{\tanh(y)}{\sinh(\frac{\epsilon}{2})} \right) \right)} \right) = \operatorname{arsinh} \left(\cosh(y) \sqrt{\sinh\left(\frac{\epsilon}{2}\right)^2 + \tanh(y)^2} \right)$$

Noting $\angle EFG = \frac{\pi}{2}$, we have $\sin(\angle GEF) = \frac{\sinh(\overline{FG})}{\sinh(\overline{EG})}$.

Noting $\angle DEG + \angle GEF = \frac{\pi}{2}$, we have...

$$\overline{FG} = \operatorname{arsinh} \left(\sin \left(\frac{\pi}{2} - \arctan \left(\frac{\tanh(y)}{\sinh(\frac{\epsilon}{2})} \right) \right) \cosh(y) \sqrt{\sinh\left(\frac{\epsilon}{2}\right)^2 + \tanh(y)^2} \right)$$

Demanding $\epsilon > 0$ allows further simplification:

$$\overline{FG} = \operatorname{arsinh} \left(\cosh(y) \sqrt{\sinh\left(\frac{\epsilon}{2}\right)^2} \right) = \operatorname{arsinh} \left(\cosh(y) \sinh\left(\frac{\epsilon}{2}\right) \right)$$

Now we can calculate the density of space in the horizontal direction!

$$\lim_{\epsilon \rightarrow 0^+} \frac{\overline{HG}}{\overline{CD}} = \lim_{\epsilon \rightarrow 0^+} \frac{\overline{FG}}{\overline{ED}} = \lim_{\epsilon \rightarrow 0^+} \frac{\operatorname{arsinh}(\cosh(y) \sinh(\frac{\epsilon}{2}))}{\frac{\epsilon}{2}} = \cosh(y)$$

Just as one might suspect, where s is the length of a hyperbolic geodesic, the Cartesian Hyperbolic Plane Metric is:

$$(ds)^2 = (\cosh(y) * dx)^2 + (1 * dy)^2$$

Alternately,

$$\frac{ds}{dx} = \sqrt{\cosh(y)^2 + \left(\frac{dy}{dx}\right)^2}$$

3.1 Cartesian Hyperbolic Plane Geodesics

Let $g(x)$ be a function of a geodesic. Then the length of a line segment is:

$$\int_a^b \sqrt{\cosh(g(x))^2 + g'(x)^2} dx$$

The Euler-Lagrange equations guarantee

$$\frac{\cosh(g(x)) \sinh(g(x))}{\sqrt{\cosh(g(x))^2 + g'(x)^2}} = \frac{\partial}{\partial x} \left(\frac{g'(x)}{\sqrt{\cosh(g(x))^2 + g'(x)^2}} \right)$$

$$\frac{\cosh(g(x)) \sinh(g(x))}{\sqrt{\cosh(g(x))^2 + g'(x)^2}} = \frac{\cosh(g(x)) \left(\cosh(g(x)) g''(x) - \sinh(g(x)) g'(x)^2 \right)}{\left(\cosh(g(x))^2 + g'(x)^2 \right)^{3/2}}$$

$$\sinh(g(x)) \left(\cosh(g(x))^2 + g'(x)^2 \right) = \cosh(g(x)) g''(x) - \sinh(g(x)) g'(x)^2$$

$$\sinh(g(x)) \left(\cosh(g(x))^2 + 2g'(x)^2 \right) = \cosh(g(x)) g''(x)$$

$$\tanh(g(x)) \left(\cosh(g(x))^2 + 2g'(x)^2 \right) = g''(x)$$

All geodesics can be divided into 4 types of geodesics:

The trivial case $y = 0$.

Those that do not cross the x-axis.

$$g_0(0) \in \mathbb{R}^+, g'_0(0) = 0, g_0(x) = g_0(-x), y = \pm g_0(x - c_0), c_0 \in \mathbb{R}$$

$$|x| < \operatorname{artanh} \left(\sqrt{1 - \tanh(g_0(0))^2} \right) = \operatorname{artanh} \left(\cosh(g_0(0))^{-1} \right)$$

Those that cross the x-axis.

$$g_1(0) = 0, g'_1(0) \in \mathbb{R}^+, g_1(x) = -g_1(-x), y = \pm g_1(x - c_1), c_1 \in \mathbb{R}$$

$$g_1(x) = \operatorname{artanh}(g'_1(0) \sinh(x)), |x| < \operatorname{arsinh}(g'_1(0)^{-1})$$

Those that touch the x-axis once infinitely far from the origin.

$$\lim_{x \rightarrow -\infty} g_2(x) = 0, \lim_{h \rightarrow \infty} g_2(-1/h) = \infty, y = \pm g_2(\pm x - c_2), c_2 \in \mathbb{R}$$

$$x < 0$$

4 Future work

Future work would use the Euler-Lagrange Equations to determine formulae for geodesics. Geodesics of the Cartesian Hyperbolic Plane Metric are particularly interesting because geodesics look like (and might exactly be!) $[y = c_0 \sec(c_2(x - c_1))$ if the geodesic does not cross the line $y = 0]$ and $[y = c_0 \tan(c_2(x - c_1))$ if the geodesic crosses the line $y = 0]$.