

# Classical Electrodynamics II

Assignment 1

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Code link: <https://github.com/Newtech66/TIFR-ED2>

**P1**

(i)

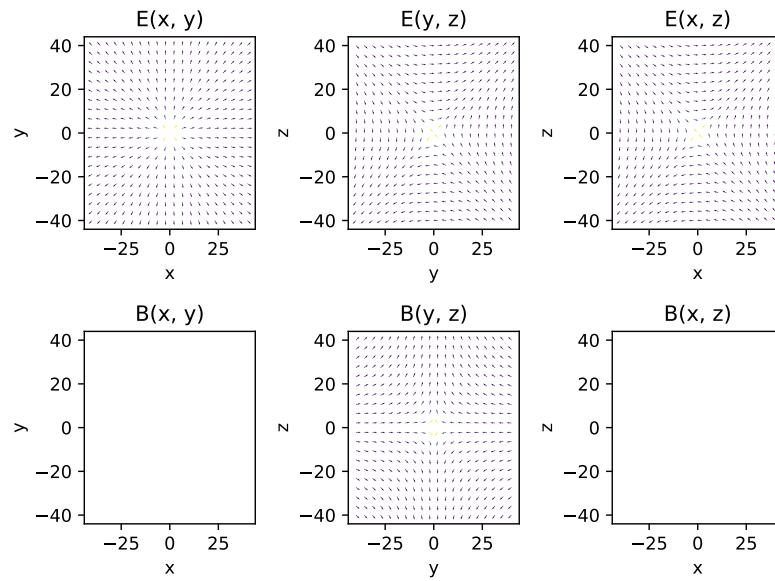


Figure 1: Field line projections for a particle moving with a constant velocity  $v = 0.1c\hat{x}$ .

(ii)

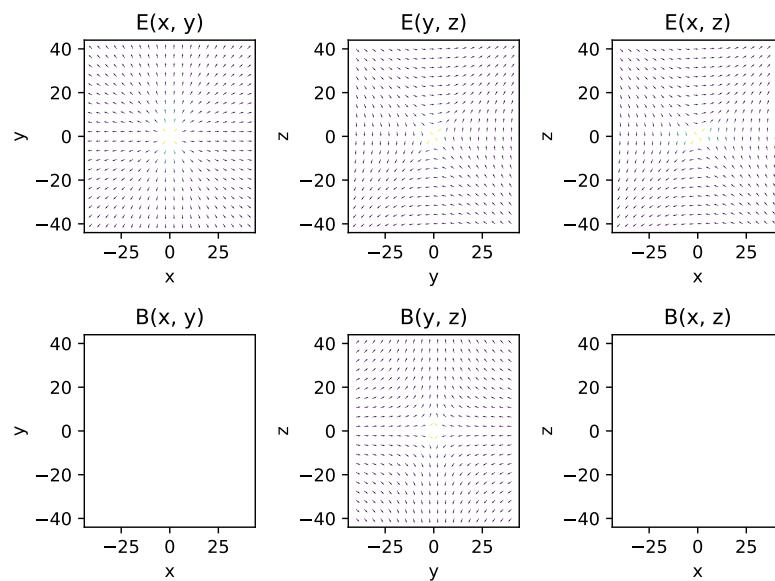


Figure 2: Field line projections for a particle moving with a constant velocity  $v = 0.9c\hat{x}$ .

(iii)

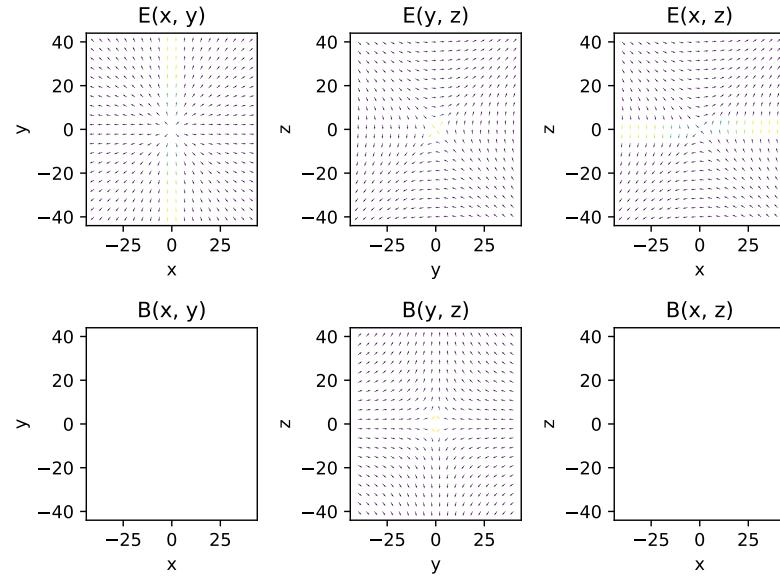


Figure 3: Field line projections for a particle moving with a constant velocity  $v = 0.999c\hat{x}$ .

**P2**

(A)

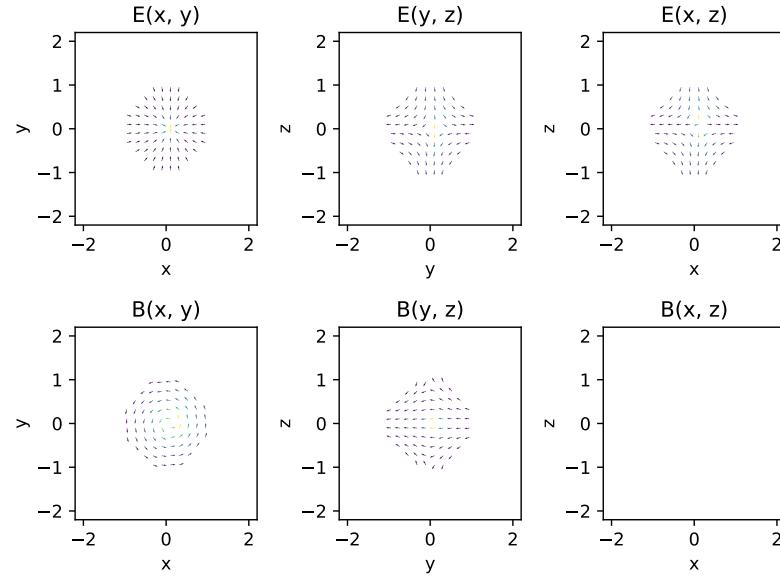


Figure 4: Radiated field line projections for a particle moving with initial velocity  $v_0 = 0.1c\hat{x}$  and acceleration  $a = 0.3c\hat{z}$  at time  $t = 1$ .

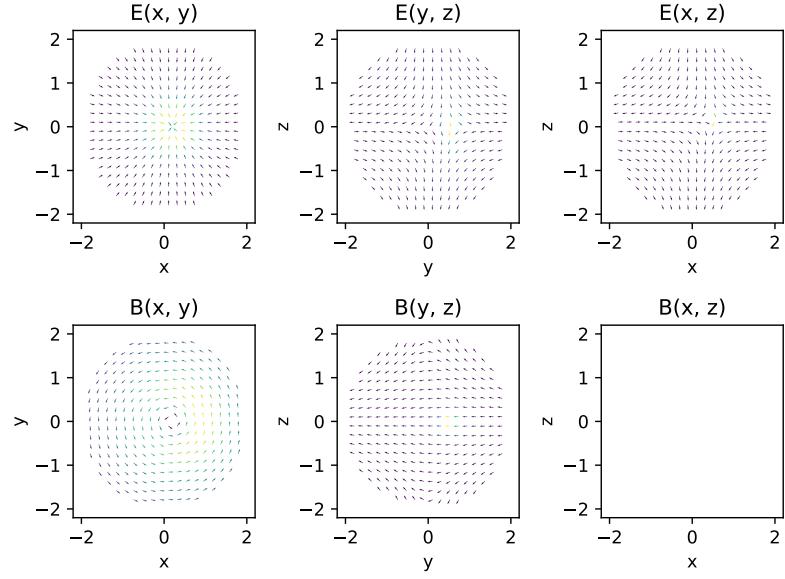


Figure 5: Radiated field line projections for a particle moving with initial velocity  $v_0 = 0.1c\hat{x}$  and acceleration  $a = 0.3c\hat{z}$  at time  $t = 2$ .

(B)

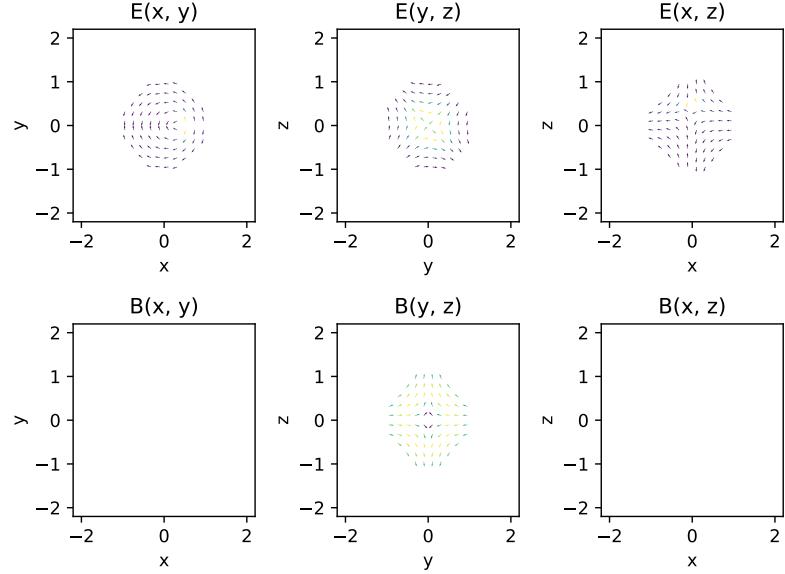


Figure 6: Radiated field line projections for a particle moving with initial velocity  $v_0 = 0.1c\hat{x}$  and acceleration  $a = 0.9c\hat{x}$  at time  $t = 1$ .

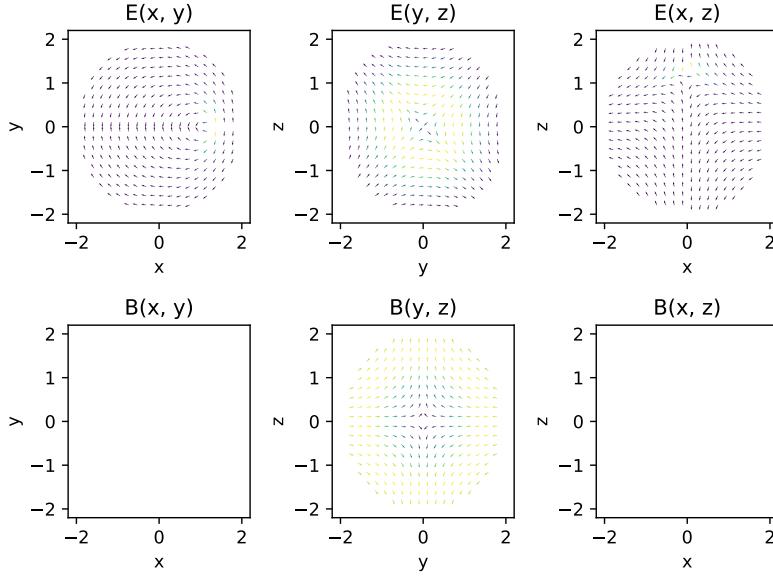


Figure 7: Radiated field line projections for a particle moving with initial velocity  $v_0 = 0.1c\hat{x}$  and acceleration  $a = 0.9c\hat{x}$  at time  $t = 2$ .

We assume that the particle starts accelerating at  $t = 0$ . We only plot the field line projections for  $v = 0.1c$  because it would be redundant to plot the rest. For part (B), the equations for velocity are given in [this Physics StackExchange question](#). We can follow the same procedure for part (A) and end up with the equations

$$v = c \sqrt{\frac{k^2 + a^2 t^2}{c^2 + k^2 + a^2 t^2}} \quad (1)$$

$$\dot{x} = k \sqrt{1 - \frac{v^2}{c^2}} \quad (2)$$

$$\dot{z} = \frac{at}{k} \dot{x} \quad (3)$$

The position of the particle over time is computed by numerically integrating the velocity.

## P3

### (A)

In this case, the velocity is along the x-axis and the acceleration is along the z-axis. So this is a case of the acceleration being perpendicular to the velocity. The answer is given by equation (4.103) of Rybicki and Lightman.

$$\frac{dP}{d\Omega} = \frac{q^2 a^2}{4\pi c^3} \frac{1}{(1 - \beta \cos\theta)^4} \left[ 1 - \frac{\sin^2 \theta \cos^2 \phi}{\gamma^2 (1 - \beta \cos\theta)^2} \right] \quad (4)$$

### (B)

In this case, the velocity is along the x-axis and the acceleration is along the x-axis. So this is a case of the acceleration being parallel to the velocity. The answer is given by equation (4.101) of Rybicki and Lightman.

$$\frac{dP}{d\Omega} = \frac{q^2 a^2}{4\pi c^3} \frac{\sin^2 \theta}{(1 - \beta \cos\theta)^6} \quad (5)$$

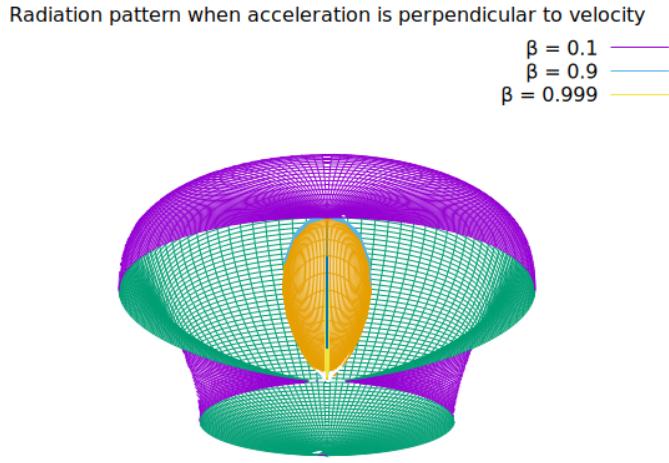


Figure 8: As  $\beta$  increases, the radiation gets more sharply peaked in the direction of motion. The plots here are normalized by the maximum value. We show a cross section so that the side lobes are visible.

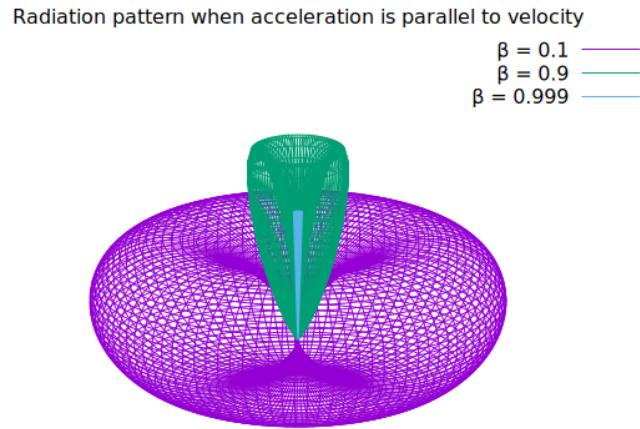


Figure 9: As  $\beta$  increases, the radiation gets more sharply peaked in the direction of motion. The plots here are normalized by the maximum value.

## P4

### (A)

A free charge experiences a force  $\mathbf{F} = q\mathbf{E}_{\text{in}}$ . The dipole moment of a free charge is given by  $\mathbf{d} = qr$ . That means

$$\ddot{\mathbf{d}} = \frac{q^2}{m} \mathbf{E}_{\text{in}} \quad (6)$$

Unpolarised light can always be decomposed into a superposition of two linearly polarised beams with polarisations along perpendicular axes with any orientation in the plane perpendicular to the direction of propagation. What

this means is that we can always decompose the unpolarised light into one beam with a polarisation in the plane containing both the charged particle and the observation point, and one with a polarisation perpendicular to it.

Consider the beam with polarisation in the plane. The angle between the polarisation vector and the position vector is  $\theta$ .

## P5

The radiation reaction force is given by

$$F_{\text{rad}} = \frac{2q^2\ddot{v}}{3c^3} \quad (7)$$

We can find that

$$\ddot{v}(t) = V_0 \left( -900 \cos(30t) + \frac{1}{10000} \sin(30t) \exp\left(\frac{t}{100}\right) + \frac{3}{5} \cos(30t) \exp\left(\frac{t}{100}\right) \right) \quad (8)$$

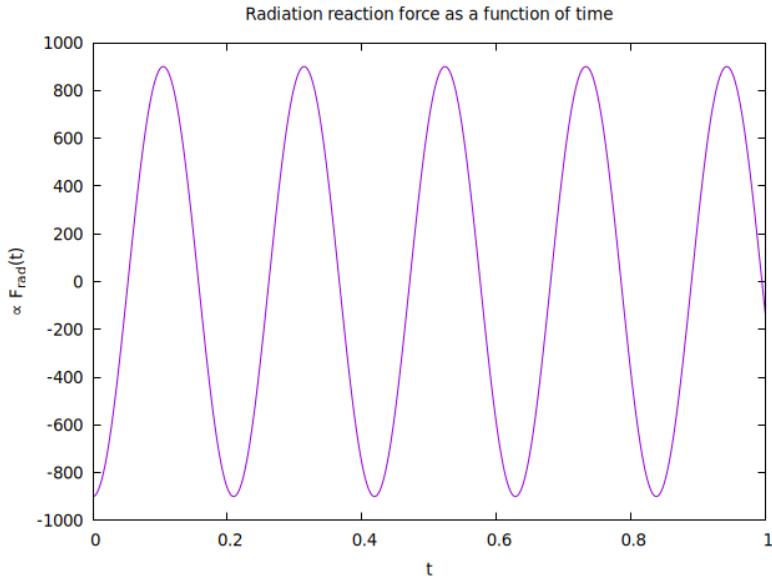


Figure 10: The radiation reaction force oscillates very quickly, therefore it is plotted only up to 1 seconds, and not 1000 seconds as asked in the problem.

## P6

The emitted power density per unit frequency is given by Rybicki and Lightman equation (5.11).

$$\frac{dW}{d\omega dV dt}(v, \omega) = \frac{16\pi e^6}{3\sqrt{3}c^3 m^2 v} n_e n_i Z^2 g_{ff}(v, \omega) \quad (9)$$

Assuming the velocity distribution is given as a volume density in  $(v_x, v_y, v_z)$  space, we need to do the integral

$$\frac{dW}{d\omega dV dt}(T, \omega) = \frac{\int_{v_{\min}}^{\infty} dv \frac{dW}{d\omega dV dt}(v, \omega) v^6 \exp\left(-\frac{mv^2}{2kT}\right) \left(1 + 0.3 \times \left(\frac{mv^2}{kT}\right)\right)}{\int_0^{\infty} dv v^6 \exp\left(-\frac{mv^2}{2kT}\right) \left(1 + 0.3 \times \left(\frac{mv^2}{kT}\right)\right)} \quad (10)$$

where  $v_{\min} = \sqrt{2h\nu/m}$  and  $\omega = 2\pi\nu$ .

First let us compute the denominator.

$$\int_0^\infty dv v^6 \exp\left(-\frac{mv^2}{2kT}\right) \left(1 + 0.3 \times \left(\frac{mv^2}{kT}\right)\right) = \frac{1}{2} \left(\frac{2kT}{m}\right)^{7/2} \Gamma\left(\frac{7}{2}\right) + 0.3 \times \left(\frac{2kT}{m}\right)^{7/2} \Gamma\left(\frac{9}{2}\right) \quad (11)$$

$$= 2.9\sqrt{\pi} \left(\frac{2kT}{m}\right)^{7/2} \quad (12)$$

We straightforwardly used the properties and values of the gamma function in the above. Now we compute the numerator.

$$= \frac{16\pi e^6}{3\sqrt{3}c^3m^2} n_e n_i Z^2 \int_{v_{\min}}^\infty dv v^5 \exp\left(-\frac{mv^2}{2kT}\right) \left(1 + 0.3 \times \left(\frac{mv^2}{kT}\right)\right) g_{ff}(v, \omega) \quad (13)$$

$$= \frac{16\pi e^6}{3\sqrt{3}c^3m^2} n_e n_i Z^2 \bar{g}_{ff}(v, \omega) \frac{1}{2} \left(\frac{2kT}{m}\right)^3 \int_{h\nu/kT}^\infty dt t^2 \exp(-t) (1 + 0.6t) \quad (14)$$

$$= \frac{16\pi e^6}{3\sqrt{3}c^3m^2} n_e n_i Z^2 \bar{g}_{ff} \frac{1}{2} \left(\frac{2kT}{m}\right)^3 \times \left[0.6 \left(\frac{h\nu}{kT}\right)^3 + 2.8 \left(\frac{h\nu}{kT}\right)^2 + 5.6 \left(\frac{h\nu}{kT}\right) + 5.6\right] \exp\left(-\frac{h\nu}{kT}\right) \quad (15)$$

We extracted  $g_{ff}$  as a thermal averaged constant in going from (13) to (14). So finally we get

$$\frac{dW}{d\nu dV dt}(T, \omega) = \frac{1.06\pi e^6}{c^3 m^2} n_e n_i Z^2 \bar{g}_{ff} \left(\frac{2kT}{\pi m}\right)^{-1/2} \left[0.6 \left(\frac{h\nu}{kT}\right)^3 + 2.8 \left(\frac{h\nu}{kT}\right)^2 + 5.6 \left(\frac{h\nu}{kT}\right) + 5.6\right] \exp\left(-\frac{h\nu}{kT}\right) \quad (16)$$

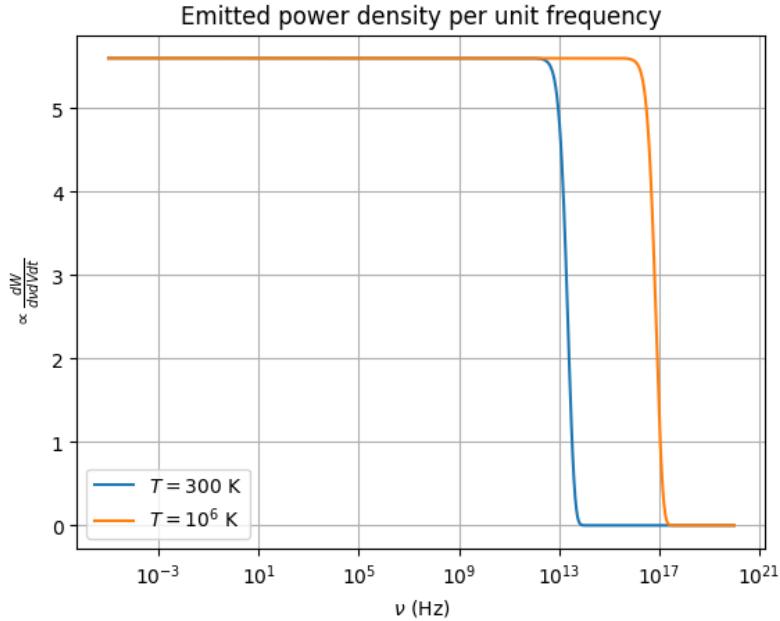


Figure 11: The emitted power density per unit frequency at (i)  $T=300$  K and (ii)  $T=10^6$  K.