



Figure 1 Transfer Between Earth and Mars

```
% This is the code for HW3 - Problem 3
% By Nicholas Luis (PSU ID 930841391)
```

```
clear; clc; close all;
```

```
% Constants
```

```
MU = 1.327*(10^11); % km^3 / s^2
```

```
PI = 3.141592654;
```

```
AU = 149.6*(10^6); % km
```

```
% Earth Orbit
```

```
rE = 1*AU;
```

```
aE = rE; % True for circular orbits
```

```
vE = getVelo(rE, aE);
```

```
% Mars Orbit
```

```
rM = 1.524*AU;
```

```

aM = rM; % True for circular orbits
vM = getVelo(rM, aM);

% Transfer Orbit
TOF = 1.58112*(10^6); % Time Of Flight in seconds
delta_f = 135.8624; % Degrees
a_T = get_aMin(rE, rM, delta_f);
p_T = getPT(a_T, rE, rM, delta_f);
e_T = (p_T / rE) - 1;
vET = getVelo(rE, a_T); % Velocity on the TO at intersection w/ Earth orbit
vMT = getVelo(rM, a_T); % Velocity on the TO at intersection w/ Mars orbit

% Delta V
delta_VE = getDeltaV(vE, vET, 0);
delta_VM = getDeltaV(vM, vMT, 0);

%% Plotting
fvec = 0:0.1:360; % Creates a list of true anomaly values to iterate through (in
degrees)

% X Y values. Note that (0,0) will be located on the Sun

xvec1 = rE.*cosd(fvec); % List of the x coordinates of the departure orbit
yvec1 = rE.*sind(fvec); % List of the y coordinates of the departure orbit

rTvec = (a_T*(1-e_T^2)) ./ (1+e_T*cosd(fvec)); % List of the radii of the transfer
orbit
xvecT = rTvec.*cosd(fvec); % List of the x coordinates of the transfer orbit
yvecT = rTvec.*sind(fvec); % List of the y coordinates of the transfer orbit

xvec2 = rM.*cosd(fvec-30); % List of the x coordinates of the arrival orbit
yvec2 = rM.*sind(fvec-30); % List of the y coordinates of the arrival orbit

figure(1)
hold on
plot(xvec1, yvec1, LineWidth=2)
plot(xvec2, yvec2, LineWidth=2)
plot(xvecT(1:1801), yvecT(1:1801), ':', LineWidth=2) % Only plotting half of the
transfer orbit
title('Earth-Mars Transfer')
xlabel("x (km)")
ylabel("y (km)")
legend('Earth Orbit', 'Mars Orbit', 'Transfer Orbit')
hold off

exportgraphics(gca, "HW2_Problem3_Figure1.jpg");

%% Functions

function P_output = getPT(amin, r1_in, r2_in, df)
    % This function gets the semilatus rectum of a transfer orbit given r1,
    % r2, and the change in f
    k = r1_in*r2_in*(1-cosd(df));
    m = r1_in*r2_in*(1+cosd(df));
    l = r1_in + r2_in;

```

```

P_output = (k*m - 2*amin*k*1) / (4*amin*m - 2*amin*1*1);

end

function velo = getVelo(r_input, a_input)
    % This function gets the velocity using energy given MU, r, and a
    MU = 1.327*(10^11); % km^3 / s^2

    velo = sqrt(2 * MU * ( (1/r_input) - (1 / (2*a_input)) ) );
end

function dvelo = getDeltaV(vi, vf, gamma)
    dvelo = sqrt( vf^2 + vi^2 - 2*vi*vf*cosd(gamma) );
end

function aMinT = get_aMin(r1_in, r2_in, df)
    % This function gets the semimajor axis of minimum energy transfer
    % orbit given r1, r2, and the change in f
    sqrtTerm = sqrt( r1_in^2 + r2_in^2 - 2*r1_in*r2_in*cosd(df) );
    aMinT = 0.25 * (r1_in + r2_in + sqrtTerm);
end

function fOut = getF(e, p, r)
    fOut = acosd((1/e)*((p/r)-1));
end

function gammaOut = getGamma(e, f, df)
    gammaOut = atand((e*sind(f)) / (1+e*cosd(f)));

    if (df > 180)
        gammaOut = abs(gammaOut);
    else
        gammaOut = - abs(gammaOut);
    end
end

```