Lab Assignment 6: The Deformable Mirror and the Gerchberg-Saxton Algorithm

This lab is inspired by NASA's Nancy Grace Roman Space Telescope, whose objectives are "to unravel the secrets of dark energy and dark matter, search for and image exoplanets, and explore many topics in infrared astrophysics". A student completes a Python program, in versions, to simulate the Gerchberg-Saxton (GS) algorithm, as inspired by the control of a deformable mirror in one instrument of the space telescope. NASA has produced videos on the telescope, which include a two-minute explanation of the coronagraph instrument. A related video explains a simplified coronagraph at the 0:20 mark.

Version 0: Getting Started

We will explore a model of the GS algorithm and its application to a simplified instrument. In "<u>A Tail of Two Cats</u>" by Kevin Cowtan, the GS algorithm is explained as a way to reconstruct a missing piece of a two-dimensional (2D) image. This involves the manipulation of *magnitude* and *phase* information of the 2D *Discrete Fourier Transform* (DFT). For our purposes, the necessary functionality of the 2D DFT and its inverse are encapsulated in two functions, dft2 and idft2, of the coronaSimulate program.

Unzip V0GettingStarted.zip. In it, there are four files, including <code>coronagraph_v0.avi</code> and a copy of <code>an image</code> taken from a NASA website. In the image, i.e., 300_26a_big-vlt-s.jpg, an exoplanet called 2M1207b orbits a dwarf star called 2M1207. Next, run the <code>coronaSimulate.py</code> file. In addition to outputting text and figures to the Console and Plots panes, it outputs several <code>.png</code> files. Each <code>.png</code> file specifies one frame of a video given in the <code>.avi</code> file. Please compare the images and video.

Run the coronaAnimate.py file. Assuming a Spyder installation with OpenCV support, it produces a video, coronagraph.avi, from a sequence of .png files, coronagraph0, coronagraph1, etc., after you enter the end frame number at the prompt. Compare coronagraph_v0.avi to this video.

Version 1: Simple Occultation

Unzip the V1SimpleOccultation.zip file. It has a coronagraph_v1.avi video and a subfolder of .png frames. Watch the video. When you complete this version, modifying coronaSimulate.py only, you should be able to reproduce the video after running coronaAnimate.py on your .png frames.

Complete the occultSquare, gerchbergSaxton, and saveFrames functions, writing their comment headers also. First, modify saveFrames so that each pixel shown has equal red, green, and blue parts. Edit the title to match the given result. Use a statement or statements to hide axes ticks and labels.

Next, modify the occultSquare function so that the central part of the image, im, it returns is black (zero). Implement a square shape that is exactly width pixels high and wide (make width the second input argument). Check it with the Variable Explorer. Do not change code outside occultSquare.

The opticalSystem function returns a second argument, Dphi, which is the true phase aberration in the pupil plane of the coronagraph. A complete GS algorithm would estimate this value independently. It would start from an initial guess, at iteration zero, and approximate the true value after a maximum number of iterations. For simplicity, our simulated algorithm assumes the aberration is known.

Modify the gerchbergSaxton function so that, when it invokes idft2, it "corrects" the aberration of phase. When idft2 is invoked, the second argument is phase. Rewrite it so that the argument equals IMp, at iteration 0, and IMp+Dphi, at iteration maxIters. Use <u>linear interpolation</u> at other iterations. Do not modify the main function or, for the purposes of this requirement, any other function.

To follow an iterative and incremental approach, submit your Version 1 coronaSimulate.py file by the Version 1 deadline. Before submission, test it when other files, from Version 0, are unchanged.

Version 2: Occultation and Plot

Unzip the V2Occultation&Plot.zip file. When your Version 2 program is sufficiently correct, it will output .png frames like those in the coronagraph_v2a subfolder. Ideally, your coronaSimulate will output ones like those in the _v2b subfolder. You may produce videos, like coronagraph_v2a.avi or _v2b.avi, by also running the given coronaAnimate program, assuming an OpenCV installation.

In the saveFrames function, insert plt.subplot(1,2,1) and plt.subplot(1,2,2) statements so that images are shown on the right side only. Rename the occultSquare function to occultCircle, both where it is defined and also where it is invoked. Revise its definition so that instead of blacking out, or occulting, a square of the specified width, the function occults a circle with that diameter, or width. The occulted group of pixels must be at the centre of the image. Verify that your program runs.

Modify the occultCircle function to return a second argument, mask, that is a 2D NumPy array of boolean entries. Initialize it, using np.full, to have the same shape as the first input argument, im, but with the value False. For every entry of im you occult, assign the corresponding entry of mask to True. Modify the invocation in opticalSystem of occultCircle so that its output arguments are assigned to two variables, the first of which, im, is the same as before. Verify that your program runs.

Modify opticalSystem to return a third argument, mask, the mask obtained from occultCircle. Modify main to accept the third output argument from opticalSystem and to pass it as a fourth input argument to gerchbergSaxton. Modify gerchbergSaxton to accept a fourth input argument, mask, and to return a second output argument, errors, initialized to an empty list. Modify saveFrames to accept a second input argument, errors. So that you can verify your program runs, modify main to accept the errors argument from gerchbergSaxton and to pass it on to saveFrames.

In gerchbergSaxton, compute a float, error, from the result, im, of each idft2 invocation and the fourth input argument, mask. Append each such float, error, to the list, errors, that the function returns. The error is the sum total of the squared values of imentries where corresponding mask entries are True. Outsource computation to a function, occultError, with two input arguments and one output argument, that you define and invoke appropriately. Verify that your program runs.

Examine the computed list, errors, with the Variable Explorer using a breakpoint inside saveFrames. Expected values can be read approximately from the coronagraph_v2a*.png files. Compute the maximum, maxErrors, of errors within saveFrames. Without altering the right-side image, make and annotate the left-side plot as indicated in the given .png files. Verify your program works.

Examine the coronagraph_v2b*.png files. To meet the implied requirements, modify saveFrames. Remove all plt.subplot(...) and any plt.axis('off') statements. Plot the graph first with correct limits. Using a plt.imshow statement with an extent=(...) argument, where you determine the (...) expression, add the image to the graph (plotted lines will show). Use plt.gca().set_aspect(?), where you must compute? from maxIters and maxErrors, so that the image looks correct.

In your own words, write suitable comment headers for all functions, main aside, without comment headers. A comment header should summarize the function's purpose, input and output arguments, and side effects, if any, such as Console input or output, Plots output, and file input or output.

To follow an iterative and incremental approach, submit your Version 2 coronaSimulate.py file by the Version 2 deadline. Before submission, test it when other files, from Version 0, are unchanged.

Revision History

This document and associated files were authored in 2020 by <u>Dileepan Joseph</u> and edited by <u>Wing Hoy</u>. The assignment was revised in 2021 by Joseph and reviewed by <u>Edward Tiong</u> and Jason Myatt. Due to a programming language change, the assignment was revised in 2022 by Joseph with Hoy's help. Joseph acknowledges the support of <u>Dan Sirbu</u>, an imaging scientist at NASA's <u>Ames Research Center</u>.