

Normalization relations (barred quantities are non-normalized, starred are value at sonic radius)

$$\rho_o = \tilde{\rho}_o / \rho_*$$

$$T_o = \tilde{T}_o / T_*$$

$$v_o = \tilde{v}_o / v_*$$

$$q_o = \tilde{q}_o / q_*$$

$$\beta_o = \tilde{\beta}_o / \beta_*$$

$$E_o = \tilde{E}_o / E_*$$

$$r = \tilde{r} / r_*$$

$$\Lambda(\tilde{E}_o) = \tilde{\Lambda}(\tilde{E}_o) / \Lambda(E_*)$$

$$L(\tilde{E}_o) = \tilde{L}(\tilde{E}_o) / [E_* \tilde{\Lambda}(E_*)]$$

Relations at the sonic radius

$$\rho_* = m \frac{P_*}{T_*}$$

$$r_*^2 \frac{P_*}{T_*} V_* = \frac{G}{4\pi}$$

$$V_* = \sqrt{\frac{G T_*}{m}}$$

$$\lambda_* = \frac{P_*}{T_*} \lambda_* r_*$$

$$(\gamma - 1) \frac{4\pi}{G} \frac{m}{\gamma} \lambda_* \left(\frac{P_*}{T_*} \right)^{1/2} = 2$$

relations to get pressure and density (normalized)

$$\rho_o = \frac{\rho_*}{\tilde{\rho}_o}$$

$$r^2 \frac{\rho_o}{\tilde{\rho}_o} v_o = 1$$

Normalized system of differential equations

$$\frac{dV_o^2}{dr} = \frac{4V_o^2}{(V_o^2 - T_o)} \left[\frac{T_o}{r} - \frac{q_o \Lambda(E_o)}{V_o} \right] \quad (I_o)$$

$$\frac{dT_o}{dr} = \frac{2q_o \Lambda(E_o)}{V_o} - \frac{(\gamma - 1)}{2} \frac{dV_o^2}{dr} \quad (II_o)$$

$$\frac{dq_o}{dr} = \lambda_* \frac{q_o \Lambda(E_o)}{r^2 V_o} \quad (III_o)$$

$$\frac{\partial E_o}{\partial r} = 2 \lambda_* \frac{L(E_o)}{r^2 v_o} \quad (\underline{IV}_o)$$

singularity for $\frac{\partial v_o^2}{\partial r}|_{r=1}$:

$$\frac{\partial v_o^2}{\partial r}|_{r=1} = \frac{2}{\gamma+1} \left[-(\gamma-3) + \sqrt{(\gamma-3)^2 - 2(\lambda_* + \Psi_* - 1)(\gamma+1)} \right]$$

$$\text{with } \Psi_* = 2 \lambda_* L(E_*) \frac{\partial \Lambda}{\partial E}|_{E=E_*}$$

Empirical energy attenuation cross sections

$$\tilde{\Lambda}(E) = \hat{\sigma}_T(E) + \frac{2 \tilde{L}(E)}{E}$$

$$\tilde{L}(E) = 8.62 \cdot 10^{-15} \left[\left(\frac{E}{100} \right)^{0.823} + \left(\frac{E}{60} \right)^{-0.125} + \left(\frac{E}{98} \right)^{-1.94} \right]^{-1} \text{ eV} \cdot \text{cm}^2$$

$$\hat{\sigma}_T(E > 100 \text{ eV}) = \frac{8.8 \cdot 10^{-13}}{E^{1.71}} - \frac{1.62 \cdot 10^{-12}}{E^{1.932}} \text{ cm}^2$$

$$\hat{\sigma}_T(E < 100 \text{ eV}) = \frac{1.1 \cdot 10^{-14}}{E} \text{ cm}^2$$

boundary conditions (normalized quantities)

at sonic radius:

$$v_o^2(1) = 1 \quad T_o(1) = 1 \quad q_o(1) = 1 \quad E_o(1) = 1$$

$$\Lambda(E_o(1)) = 1 \quad L(\tilde{E}_o(1)) = \tilde{L}(E_*) / (E_* \tilde{\Lambda}(E_*))$$

at pellet radius:

$$v_o^2(r_p) = ? \quad T_o(r_p) = 0 \quad q_o(r_p) = 0 \quad E_o(r_p) = ?$$

$$r_p = \tilde{r}_p / r_*$$

at cloud boundary ($r \rightarrow \infty$):

$$p(\infty) = 0$$

$$v_o^2(\infty) = ? \quad T_o(\infty) = ? \quad q(\infty) = q_{bg} \quad E(\infty) = E_{bg}$$

make boundary value problem for $0 \leq \hat{r} \leq 1$

we want to solve eq. system (I_0) to (IV_0)

for a given pellet radius r_p , plasma temperature T_{bg} , plasma electron density $n_{e,bg}$

this eq. system is in terms of the variable $r = \frac{\tilde{r}}{r_*}$

which is $r=1$ at the sonic radius but r = unknown at pellet radius

transform the eq. system into one in terms of the

variable \hat{r} so that $\hat{r}=1$ at sonic radius and $\hat{r}=0$ at pellet radius

$$\Rightarrow \hat{r}(\tilde{r}=r_*)=1 \text{ and } \hat{r}(\tilde{r}=r_p)=0 \Rightarrow \hat{r} = \frac{\tilde{r} - r_p}{r_* - r_p}$$

but we need $r=r(\hat{r})$ to transform the system $(\hat{r} \in [0,1], r \in [\frac{r_p}{r_*}, 1], \tilde{r} \in [r_p, r_*])$

$$r = \frac{\tilde{r}}{r_*} \Leftrightarrow \tilde{r} = r r_*$$

$$\Rightarrow \hat{r} = \frac{r r_* - r_p}{r_* - r_p} \Leftrightarrow \frac{1}{r_*} \left[(r_* - r_p) \hat{r} + r_p \right] = r$$

$$\Leftrightarrow r = \left(1 - \frac{r_p}{r_*}\right) \hat{r} + \frac{r_p}{r_*} \quad \text{define } \eta_* = \frac{r_p}{r_*} \quad (\text{dimensionless})$$

$$r = (1 - \eta_*) \hat{r} + \eta_*$$

$$\Rightarrow \frac{dr}{d\hat{r}} = 1 - \eta_* \Leftrightarrow dr = (1 - \eta_*) d\hat{r}$$

Transform eq. sys.

$$(I_0) \Leftrightarrow \frac{dv_o^2}{d\hat{r}} = (1 - \eta_*) \frac{4 v_o^2}{(v_o^2 - T_0)} \left[\frac{T_0}{(1 - \eta_*) \hat{r} + \eta_*} - \frac{q_0 \Lambda}{V_0} \right]$$

$$(II_0) \Leftrightarrow \frac{dT_0}{d\hat{r}} = (1 - \eta_*) \frac{2 q_0 \Lambda}{V_0} - \frac{\gamma - 1}{2} \frac{dv_o^2}{d\hat{r}}$$

$$(III_0) \Leftrightarrow \frac{dq_0}{d\hat{r}} = (1 - \eta_*) \lambda_* \frac{q_0 \Lambda}{V_0} ((1 - \eta_*) \hat{r} - \eta_*)^{-2}$$

$$(\text{IV}_0) \Leftrightarrow \frac{dE_0}{dr} = (1 - \eta_*) 2 \lambda_* \frac{L}{v_0} \left((1 - \eta_*) \hat{r} - \eta_* \right)^{-2}$$

$$\frac{dv_0^2}{d\hat{r}} \Big|_{\hat{r}=1} = (1 - \eta_*) \frac{dv_0^2}{dr} \Big|_{r=1}$$

now there are 4 eqs with 2 unknown parameters η_*, λ_*

boundary conditions:

$$V_0^2(\hat{r}=1) = 1, T_0(\hat{r}=1) = 1, q_0(\hat{r}=1) = 1, E_0(\hat{r}=1) = 1$$

$$T_0(\hat{r}=0) = 0, q_0(\hat{r}=0) = 0$$

to be able to calculate λ and L we need E_*

so instead of setting $\tilde{E}(\tilde{r} \rightarrow \infty) = E_{bg}$ we set E_*

and in the second step find out which E_{bg} this corresponds to

now we have enough to solve the boundary value problem $\hat{r} \in [0, 1]$

numerical solution via `scipy.integrate.solve_bvp` yields

values for η_*, λ_*

and the values of V_0^2, T_0, q_0, E_0 on a chosen grid on $\hat{r} \in [0, 1]$

To find E_{bg} and q_{bg} , we then solve the system with known η_*, λ_*, E_* as an initial value problem $\hat{r} \in [1, \infty)$

where we stop at some \hat{r} where all parameters have converged to be constant

Then we know $V_0^2(\hat{r} \rightarrow \infty), T_0(\hat{r} \rightarrow \infty), E_0(\hat{r} \rightarrow \infty), q_0(\hat{r} \rightarrow \infty)$

Then we also know $E_{bg} = E_0(\hat{r} \rightarrow \infty) E_*$ and through $q_{bg} = \frac{n_{e,bg} v_{e,bg}}{4} E_{bg}$ (Park eq. 1)
we can find q_*

unknown still: v_*, T_*, P_*, G

$$\lambda_* = \frac{P_*}{T_*} \lambda_* r_*$$

(I_{*})

$$\lambda_* = \frac{P_*}{T_*} \lambda_* r_*$$

$$r_*^2 \frac{P_*}{T_*} v_* = \frac{G}{4\pi}$$

$$v_* = \sqrt{\frac{8T_*}{m}}$$

(I_{*})

(II_{*})

(III_{*})

$$(J-1) \frac{4\pi}{G} \frac{\mu}{J} \lambda_* \left(g_* \frac{r_*^2}{T_*} \right) = 2 \quad (\text{IV}_*)$$

Solve (I_{*}) to (IV_{*}) numerically for v_* , P_* , T_* , G