







Global Alignment of the SciFi and the VELO (DPG titel einfuegen)

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DPG Conference, Karlsruhe

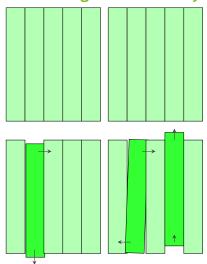








What is Alignment and why do we need it?



- Top: ideal detector, bottom: physical detector
- Surveys are used to find the rotation and position of each detector component
- →Starting positions for software alignment
- Building tracks accurately requires positions in reconstruction to be as similar as possible to real positions

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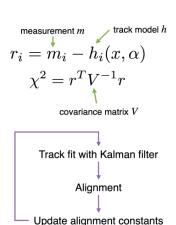








Alignment: track fits with the Kalman Filter (maybe backup)



- Starting positions: positions from laser scans of detector objects (survey)
- Alignment: χ^2 minimization of track residuals
- Add measurements one-by-one to fit
- Prediction of next measurement →minimize residuals →redo until track complete
- Why Kalman Filter?
 - easily models material interactions as well as multiple scattering

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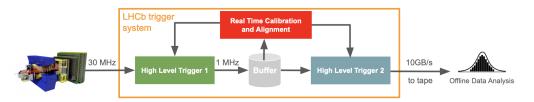






Importance of alignments

- Alignment is part of the LHCb trigger system
- Physics performance tied to alignment performance
- with optimal alignment:
 - → remove systematic biases for asymmetry measurements
 - Best possible mass resolution



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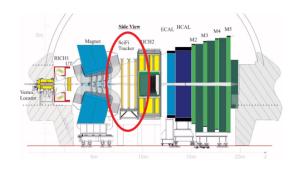




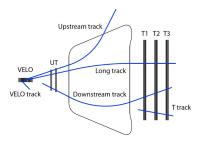




LHCb upgraded with the SciFi



- 3 stations: T1, T2, T3
- 4 layers per station: X1, U, V, X2
- Replaces former IT and OT to cope with the increased instantaneous luminosity



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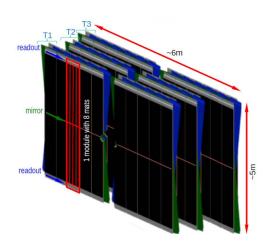








The Scintillating Fibre Tracker



- Front two stations have 5 modules per side
- Back station has 6 modules on each side
- U, V layers have a **∓5 deg** stereo angle respectively
- → Used for determining y-position of tracks by comparing hitposition at different angles

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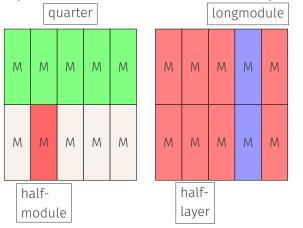






SciFi terminology (confusing for some people)

Layers are divided into two halves commonly labeled as A-side and C-side



Q2	Q3
Q0	Q1
C-side	A-side

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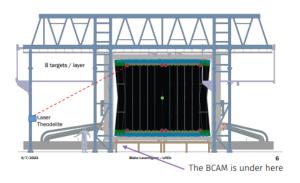






The survey: what is it and the different types (maybe backup material)

• Measure distance of some points on the detector with a laser



- Layer survey: find corners of layers
- Module survey: reflective stickers, calculate module plane
- Compare survey to simulation

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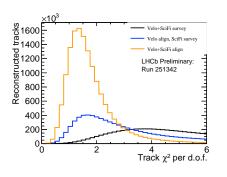








Alignment versions development in 2022 (maybe exchange with something more recent)



- V1: First ever alignments after upgrade
- Using first data sets
- Starting configuration: long modules, Tx Rz

- V2: Imroved configuaration from V1
- Half modules, Tx Rz, new time alignment

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Global Alignment and motivation

- what is global alignment?
 - alignment of the velo and the scifi together to understand the interplay in comparison to only SciFi
- why is it a good idea to do it?
 - open VELO might be a problem for other tracking stations like UT, SciFi and Muon
 - VELO is twisted between fron and back -> introduces potential biases

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problems miguel told me

• problem A

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plots about layer positions

- zig zag pattern
- dip in T2 is weird
- look at my plots and deduce the differences between configs (what happens when introduce Frames or Long layers or lagrange constraints or different dofs)

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maybe also show module positions

• make it easier to understand when looking at the plot

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Summary

- Source of complications: SciFi parts too far out of alignment to be correctly updated, corrected now with new alignment version and photogrammetry
- Correct module positions by hand: working procedure for singular low-efficiency modules
- stability tests show no substantial difference in alignment quality from magnet polarity

• sufficient statistics: 150µm movement; rerun manually if this is exceeded

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Backup

Loose Tracking config and DO selection

```
2 # (c) Copyright 2022 CERN for the benefit of the LHCb Collaboration
4 # This software is distributed under the terms of the GNU General Public
5 # Licence version 3 (GPL Version 3), copied verbatim in the file "COPYING".
6 #
7 # In applying this licence, CERN does not waive the privileges and immunities #
8 # granted to it by virtue of its status as an Intergovernmental Organization #
9 # or submit itself to any jurisdiction.
   from PvConf.Algorithms import (PrForwardTrackingVelo, PrMatchNN)
   from Moore import options, run_reconstruction
   from RecoConf.standalone import standalone hlt2 light reco without UT
14 from RecoConf.hlt2 tracking import (
       make_PrKalmanFilter_noUT_tracks,
16
       make_TrackBestTrackCreator_tracks,
18 from RecoConf.early_data import (
19
       get_loose_PrForwardTrackingVelo_params,
28
       get_loose_PrMatchNN_params,
21
       get_loose_PrKalmanFilter_params,
       get_loose_TrackBestTrackCreator_params,
23 )
24
   with PrForwardTrackingVelo.bind(**get_loose_PrForwardTrackingVelo_params()), \
        PrMatchNN.bind(**qet_loose_PrMatchNN_params()), \
26
        make_PrKalmanFilter_noUT_tracks.bind(**qet_loose_PrKalmanFilter_params()), \
        make_TrackBestTrackCreator_tracks.bind(**qet_loose_TrackBestTrackCreator_params()):
28
       run_reconstruction(options, standalone_hlt2_light_reco_without_UT)
```

```
pt_{min} = 800 MeV
pion, kaon required to have min pt and
IPCut = 60 \mu m
mass hypothesis [1760 MeV, 1960 MeV]
vertex \chi^2 < 10
```

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Backup: stability dataset

- Dataset contains magnet-up and magnet-down samples from 2022 labeled as "good" from EMTF
- Good: > 90% of datalinks are good
- Includes runs from fills: 8489, 8491, 8496 List of randomly chosen runs: 255949, 256030, 256145, 256159, 256163, 256272, 256278, 256290
- V3 Alignment from tag (loose tracking, half module alignment TxTz + Mat alignment, back layer fixed) from conditions database

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