

Quantum Statistical Mechanics

Breakdown of classical statistical mechanics: we will look at predictions of the heat capacity (an experimentally measurable quantity) for the following three systems:

- ① Diatomic ideal gas
- ② Vibrations in a solid (phonons)
- ③ Blackbody radiation (photons)

Historically the last development among the three since the correction prediction requires knowing how to quantize the angular momentum $\hbar^2 l(l+1)$

Classical SM predicts a finite heat capacity but doesn't take into account the "freezing" of d.o.f.s at low temp.

→ Classical SM predicts an infinite energy (ultraviolet catastrophe) which led to the development of QM.

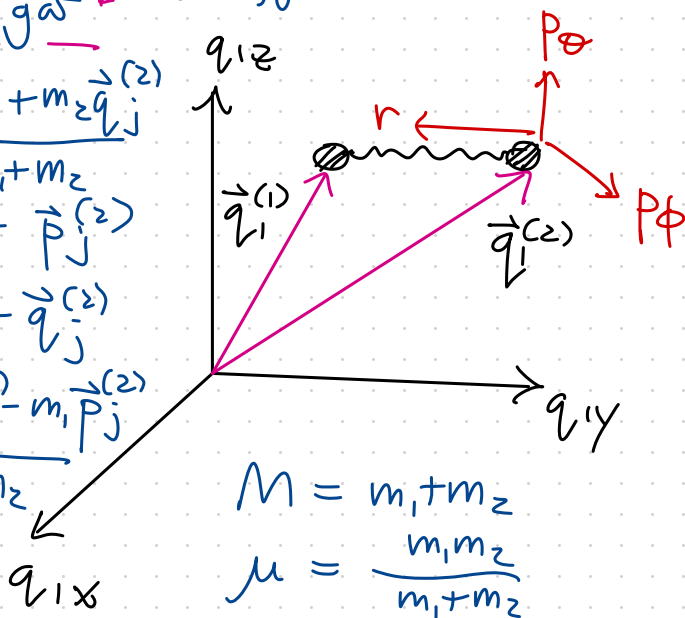
Non-interacting diatomic gas ↖ e.g. oxygen

(C.O.M. coordinates) $\vec{Q}_j = \frac{m_1 \vec{q}_j^{(1)} + m_2 \vec{q}_j^{(2)}}{m_1 + m_2}$

$\vec{P}_j = \vec{p}_j^{(1)} + \vec{p}_j^{(2)}$

(Relative coordinates) $\vec{r}_j = \vec{q}_j^{(1)} - \vec{q}_j^{(2)}$

$\vec{\pi}_j = \frac{m_2 \vec{p}_j^{(1)} - m_1 \vec{p}_j^{(2)}}{m_1 + m_2}$



$M = m_1 + m_2$

$\mu = \frac{m_1 m_2}{m_1 + m_2}$

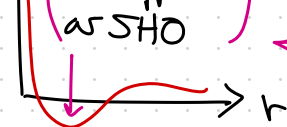
Internal d.o.f.

Relative to the second atom in the molecule, the first atom can move in three directions: along the axis connecting the two atoms, (vibrational modes) or in the other two perpendicular directions denoted θ and ϕ (associated with their own canonical momenta p_θ and p_ϕ .) (rotational modes)

$Z_N = \frac{Z_1^N}{N!}, Z_1 = \int \frac{d^3 Q_j d^3 P_j}{h^3} e^{-\beta P_j^2 / 2M}$

$V(r)$ $\times \int \frac{dr dp_r d\theta dp_\theta d\phi dp_\phi}{h^3} e^{-\beta \pi^2 / 2\mu} e^{-\beta V(r)}$

(Can approx. as SHO)

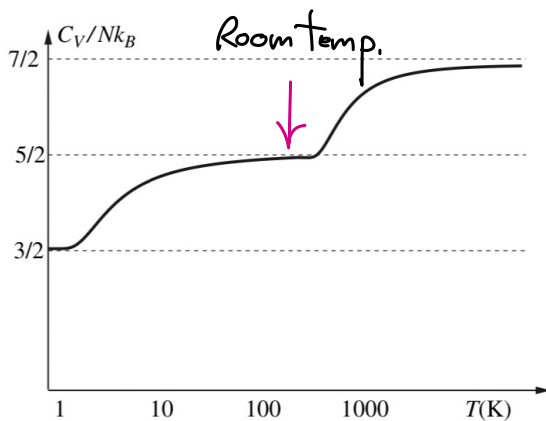


Vibrational energy is a function of the distance $\vec{q}_j^{(1)} - \vec{q}_j^{(2)}$

$$\begin{aligned}
 & \int \frac{d^3 Q_j d^3 P_j}{h^3} \overbrace{e^{-\beta P_j^2 / 2M}}^{3 \text{ translational}} \times \underbrace{\int \frac{d\theta d\phi d p_\theta d p_\phi}{h^2} e^{-\beta (p_\theta^2 + p_\phi^2) / 2\mu}}_{2 \text{ rotational}} \\
 & \times \int \frac{dr dp_r}{h} \underbrace{e^{-\beta P_r^2 / 2\mu} e^{-\beta V(r)}}_{2 \text{ vibrational}}
 \end{aligned}$$

According to the equipartition thm, each harmonic d.o.f. should give $\frac{1}{2} kT$ of energy.

$$U = NkT \left(\frac{3}{2} + 1 + 1 \right) = \frac{7}{2} NkT \Rightarrow C_v = \left. \frac{\partial U}{\partial T} \right|_v = \frac{7}{2} Nk$$



Doesn't agree with experiments!

Log plot \rightarrow

Vibrational modes

Quantum analysis

$$H_{\text{vib}} = \hbar\omega\left(n + \frac{1}{2}\right)$$

$$\sum_{n=0}^{\infty} ar^n = \frac{a}{1-r}$$

$$Z_{\text{vib}} = \sum_{n=0}^{\infty} e^{-\beta\hbar\omega\left(n + \frac{1}{2}\right)}$$

$$= \underbrace{e^{-\beta\hbar\omega/2}}_{\text{zero-point energy}} \sum_{n=0}^{\infty} (e^{-\beta\hbar\omega})^n =$$

$$\frac{e^{-\beta\hbar\omega/2}}{1 - e^{-\beta\hbar\omega}}$$

$$\ln Z_{\text{vib}} = -\frac{\beta\hbar\omega}{2} - \ln(1 - e^{-\beta\hbar\omega})$$

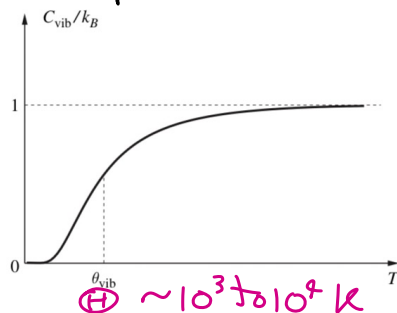
$$\langle E_{\text{vib}} \rangle = -\frac{\partial \ln Z_{\text{vib}}}{\partial \beta} = \frac{\hbar\omega}{2} + \hbar\omega \frac{e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}}$$

$$C_v = \frac{\partial \langle E_{\text{vib}} \rangle}{\partial T} = k_B \left(\frac{\hbar\omega}{k_B T} \right)^2 \frac{e^{-\beta\hbar\omega}}{(1 - e^{-\beta\hbar\omega})^2}$$

Characteristic temperature $\Theta = \hbar\omega/k_B$

$$\Rightarrow C_v = k_B \left(\frac{\Theta}{T} \right)^2 \frac{e^{-\Theta/T}}{(1 - e^{-\Theta/T})^2}$$

High temp: $k_B e^{-\Theta/T} \sim k_B$



(To ensure that quantum stat. mech. is consistent with classical stat. mech., we assert that the Planck constant "h" here is the same as the "h" in $\int \frac{dx dp}{h}$)

$\Theta \sim 10^3 \text{ to } 10^4 \text{ K}$

Rotational modes

Classical analysis

$$H_{\text{rot}} = \frac{1}{2I} \left(p_{\theta}^2 + \frac{p_{\phi}^2}{\sin^2 \theta} \right) = \frac{L^2}{2I}$$

$$Z_{\text{rot}} = \int d\theta dp_{\theta} d\phi dp_{\phi} e^{-\beta (p_{\theta}^2 + p_{\phi}^2 / \sin^2 \theta) / 2I}$$

$$= \underbrace{\frac{2\pi}{h^2}}_{\phi \text{ integral}} \cdot \underbrace{\sqrt{2\pi I kT}}_{p_{\theta} \text{ integral}} \cdot \int_0^{\pi} d\theta \underbrace{\sqrt{2\pi I \sin^2 \theta kT}}_{p_{\phi} \text{ integral}}$$

$$= \frac{4\pi}{h^2} \cdot 2\pi I k_B T = \frac{2I k_B T}{h^2} \leftarrow \hbar = \frac{h}{2\pi}$$

$$\langle E \rangle = -\frac{\partial \ln Z}{\partial \beta} = +\frac{\partial \ln \beta}{\partial \beta} = k_B T \Rightarrow C_v = k_B$$

Quantum analysis

$$Z_{\text{rot}} = \sum_{l=0}^{\infty} \exp \left[-\frac{\beta}{2I} \hbar^2 l(l+1) \right] \overbrace{(2l+1)}^{\text{degeneracy}}$$

$$\text{Let } \Theta = \hbar^2 / 2I k_B \Rightarrow Z_{\text{rot}} = \sum_{l=0}^{\infty} e^{-\Theta l(l+1) / T} (2l+1)$$

$T \gg \Theta$

$$\sum_{l=0}^{\infty} \rightarrow \int_0^{\infty} dx, \quad x = l(l+1), \quad dx = (2l+1) dl$$

$$Z_{\text{rot}} = \int_0^{\infty} dx e^{-(\Theta/T)x} = \boxed{\frac{T}{\Theta}} \quad \left(\begin{array}{c} \text{Classical} \\ \text{limit} \end{array} \right)$$

$T \ll \Theta$ First few terms dominate

$$Z_{\text{rot}} = 1 + 3e^{-2\Theta/T} + \mathcal{O}(e^{-6\Theta/T})$$

Using $\ln(1+x) = x - \frac{x^2}{2} + \mathcal{O}(x^3)$

$$\ln Z_{\text{rot}} \approx 3e^{-2\Theta/T} = 3e^{-2\Theta/k_B T}$$

$$\therefore \langle E \rangle = -\frac{\partial \ln Z}{\partial \beta} \approx 6k_B \Theta e^{-2\Theta/T}$$

$$\Rightarrow C_v = \frac{\partial \langle E \rangle}{\partial T} \approx 3k_B \underbrace{\left(\frac{2\Theta}{T}\right)^2 e^{-2\Theta/T}}$$

$\Theta \sim 1$ to 10 K

vanishes at lower T

due to quantum statistics

$$x^2 e^{-x} \rightarrow \infty \cdot 0 \text{ as } T \rightarrow 0$$

But by L'Hospital rule

$$\lim_{x \rightarrow 0} x^2 e^{-x} = \lim_{x \rightarrow 0} \frac{d^2 x^2}{dx^2} \frac{d^2 e^{-x}}{dx^2} = 0$$

Solids

Classically, atoms that form a solid can have small vibrations around their equilibrium in all three directions, except very long wavelength modes (small wave numbers k) that correspond to the translations of the whole solid. So in the thermodynamic limit, $\langle E \rangle \approx 3Nk_B T \Rightarrow C_v \approx 3k_B$

\uparrow
kinetic + potential

No restoring force

Quantum analysis

$$Z_1 = \frac{e^{-\beta \hbar \omega (k)/2}}{1 - e^{-\beta \hbar \omega (k)}} \quad \left(\text{Same as the vibrational modes} \right)$$

$$Z_N = e^{-\beta E_0} \prod_{j, \vec{k}} \frac{1}{1 - e^{-\beta \hbar \omega_j(\vec{k})}}$$

Include the zero-point energies of all oscillators

No $N!$ factor because the atoms are distinguishable by their locations.

(j, \vec{k}) enumerates the quantum microstates
 j labels modes and \vec{k} the wave vectors

$$\langle E \rangle = E_0 + \sum_{j, \vec{k}} \hbar \omega_j(\vec{k}) \langle n_j(\vec{k}) \rangle$$

$$\langle n_j(\vec{k}) \rangle = \frac{\sum_{n=0}^{\infty} n e^{-\beta \hbar \omega_n}}{\sum e^{-\beta \hbar \omega_n}} = -\frac{1}{\hbar \omega} \frac{\partial \ln}{\partial \beta} \frac{1}{1 - e^{-\beta \hbar \omega}} = + \frac{e^{-\beta \hbar \omega}}{1 - e^{-\beta \hbar \omega}}$$

The Einstein model

All vibrations are quantized and have the same universal frequency ω .

$$\langle E \rangle = E_0 + 3N \hbar \omega \frac{e^{-\beta \hbar \omega}}{1 - e^{-\beta \hbar \omega}}$$

$$\Theta = \hbar \omega / k_B$$

$$C_v = 3N k_B \left(\frac{\Theta}{T} \right)^2 \frac{e^{-\Theta/T}}{(1 - e^{-\Theta/T})^2}$$

However, experimentally C_v decays much more slowly as T^3 (without the exponential factor). A more sophisticated model solves this problem

The Debye model

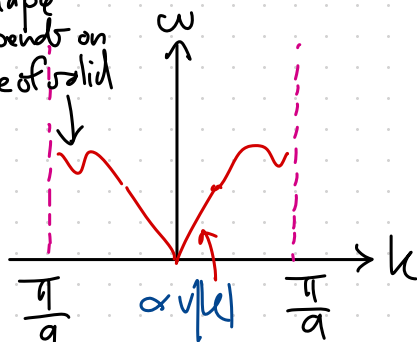
Wavelength λ

wave number $k = \frac{2\pi}{\lambda}$

frequency $\omega = \omega(k)$

(Dispersion relation)

Shape depends on type of solid



Brillouin zone

At low T , thermal energy only excites long wavelength modes. These modes are **phonons** (acoustic modes = sound)

that can propagate through the solid without much distortion with

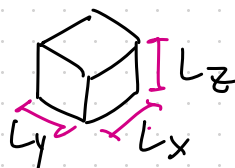
$$\omega = v k$$

Speed of sound in the solid

(In general,

$$v = \frac{\partial \omega}{\partial k}$$
)

In a box



$$\vec{k} = \left(\frac{2\pi n_x}{L_x}, \frac{2\pi n_y}{L_y}, \frac{2\pi n_z}{L_z} \right)$$

$$\Rightarrow dN = \prod_{j=1}^3 \frac{dk_j}{2\pi/L_j} = \frac{V}{(2\pi)^3} d^3k$$

$$\langle E \rangle \rightarrow E_0 + 3 \int dN \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$$

Integral: $\omega = v k \Rightarrow$

$$3 \frac{V}{(2\pi)^3} \int d^3k \frac{\hbar v k}{e^{\beta \hbar v k} - 1}$$

$T \ll \Theta$ Small \hbar 's dominate and can take the integration lim to ∞

$$x = \beta \hbar v k$$

$$|\text{Integral}| = \frac{3V}{(2\pi)^3} \int d^3k \frac{\hbar v k}{e^{\beta \hbar v k} - 1}$$

$$\approx \frac{3V}{8\pi^3} \int \left(\frac{k_B T}{\hbar v}\right)^3 d^3x \frac{k_B T x}{e^x - 1}$$

$$= \frac{3V}{8\pi^3} \left(\frac{k_B T}{\hbar v}\right)^3 k_B T \int d\Omega \int_0^\infty dx \frac{x^3}{e^x - 1}$$

$$G_\nu(z) = \frac{1}{\Gamma(\nu)} \int_0^\infty dx \frac{x^{\nu-1}}{e^x - 1}$$

$$\lim_{z \rightarrow 1} G_\nu(z) = \zeta(\nu)$$

$$\Gamma(n) = (n-1)!$$

$$3! \cdot \frac{1}{\Gamma(4)} \int_0^\infty dx \frac{x^{4-1}}{e^x - 1}$$

$$\zeta(4) = \frac{\pi^4}{90}$$

$$\rightarrow = \frac{3V}{8\pi^3} \frac{(k_B T)^4}{(\hbar v)^3} 4\pi \cdot \frac{\pi^4}{15} = \frac{\pi^2}{10} v \frac{(k_B T)^4}{(\hbar v)^3}$$

Black-body radiations

2 polarization states

$$\frac{2V}{(2\pi)^3} \int d^3k \frac{\hbar c k}{e^{\beta \hbar c k} - 1} \quad \text{Same integral}$$

$$E = V \frac{\pi^2}{15} \frac{(k_B T)^4}{(\hbar c)^3}$$