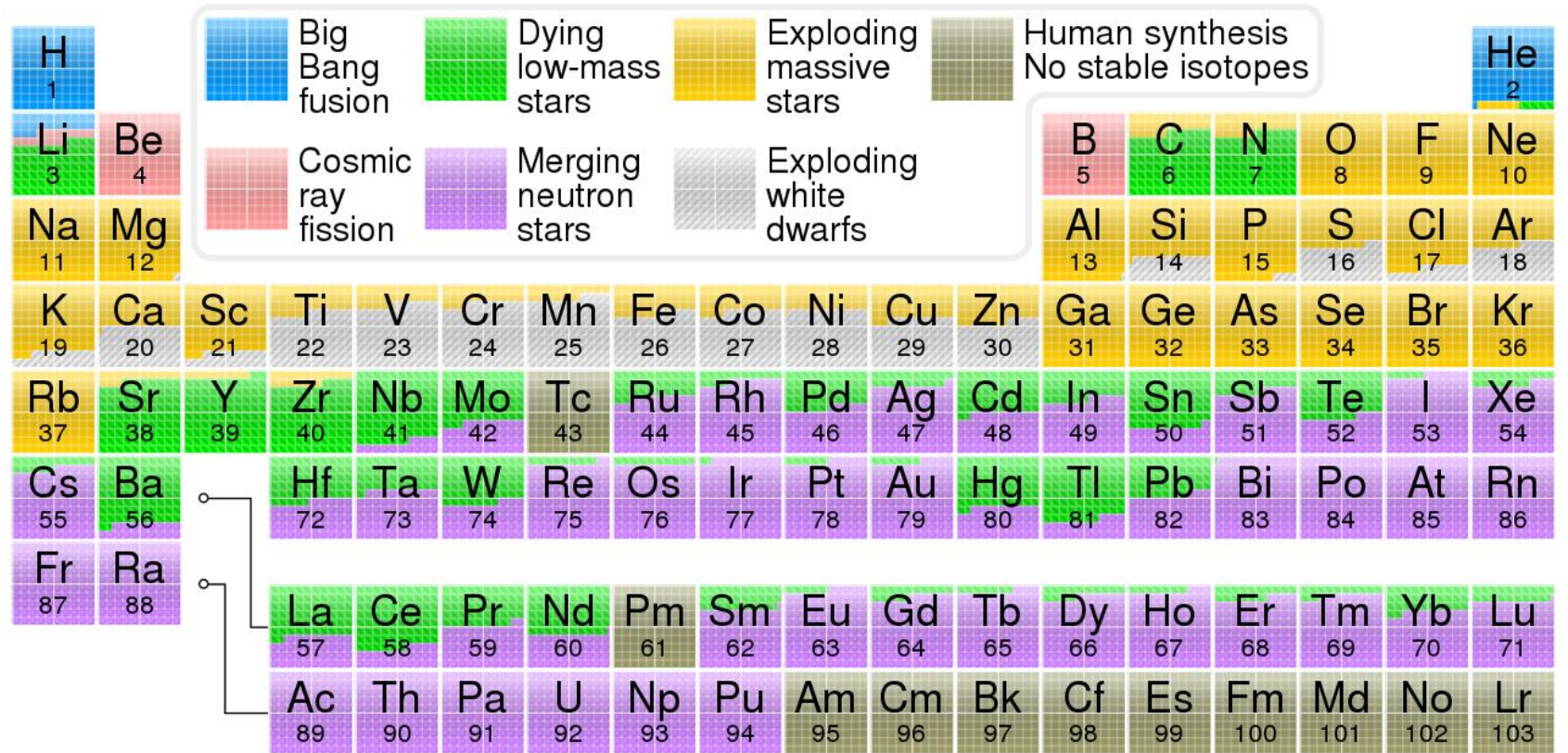
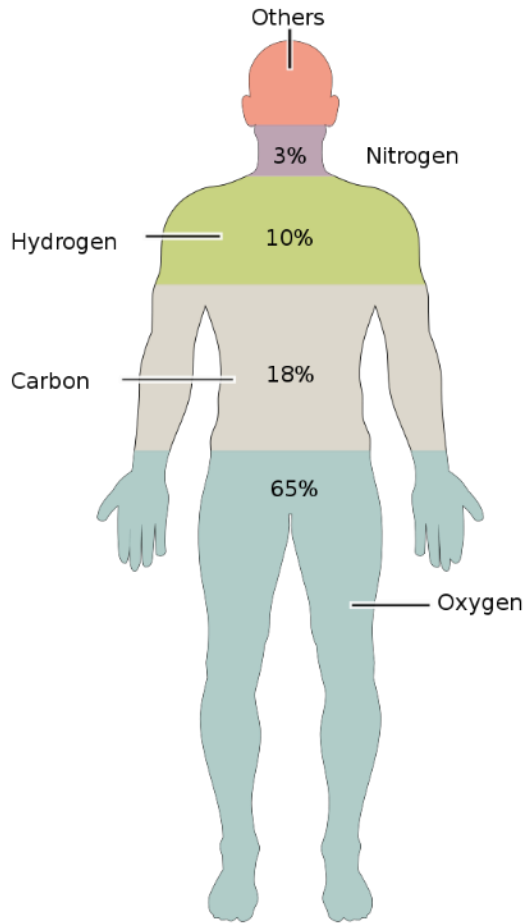


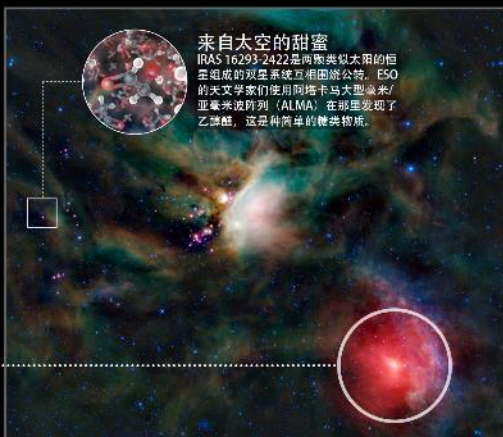
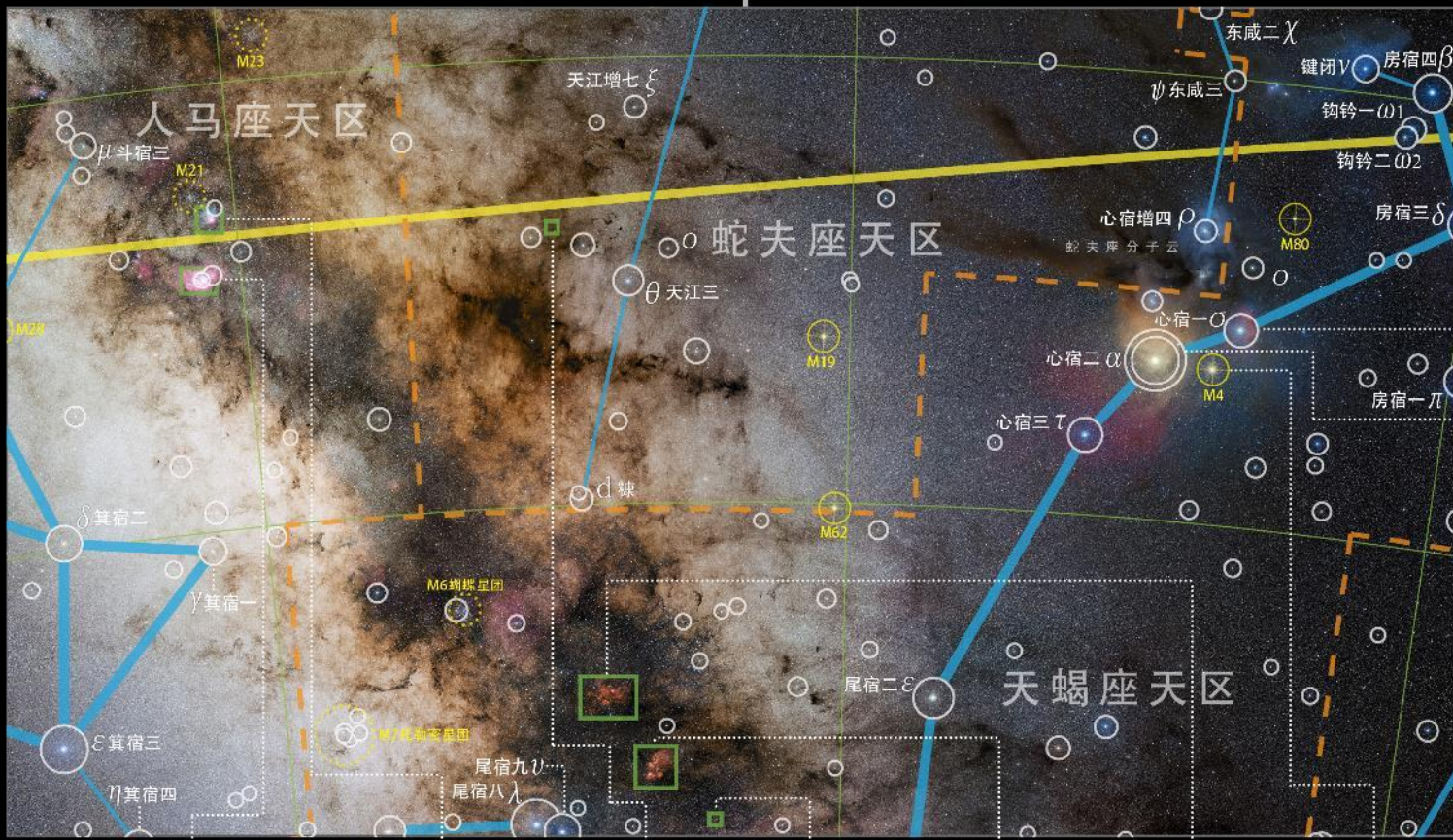
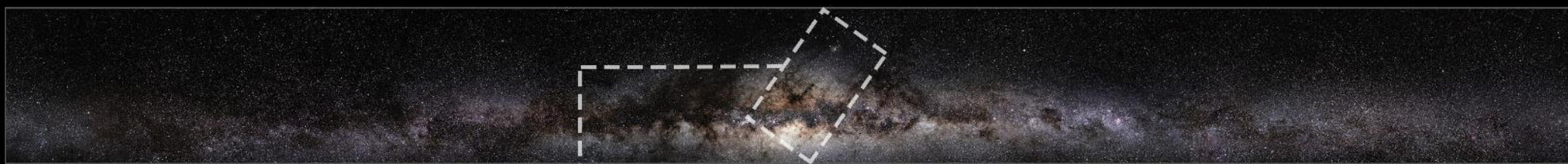
706生活实验室迈向星空系列活动
非正式演讲报告夜谈









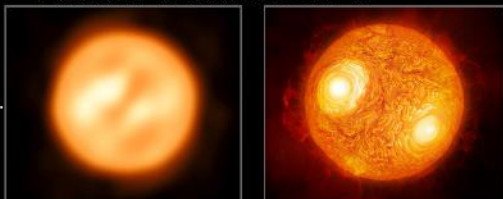


来自太空的甜蜜

IRAS 16293-2422是两颗类似太阳的恒星组成的双星系统互相围绕公转。ESO的天文学家们使用阿塔卡马大型毫米/亚毫米波阵列 (ALMA) 在那里发现了乙醛。这是一种简单的糖类物质。

蛇夫座分子云红外图像

蛇夫座分子云又名蛇夫座Rho星云复合体，是距离地球460光年由气体和尘埃构成的恒星形成区。科学家在这里发现了很多有趣的事物，例如复杂化学物和未来可能诞生行星的原行星尘埃气体盘。红外能有效穿透尘埃和气体的遮挡，洞悉宇宙背后的秘密，还有很多天体在可见光下不起眼，在红外波段下却很明亮，也就是为什么新的韦伯太空望远镜以红外观测为主。而这张照片是WISE(Wide-field Infrared Survey Explorer)红外太空望远镜拍摄的。它是NASA于2009年12月发射的，口径0.4米，主要用来巡天观测，至今仍在工作中。

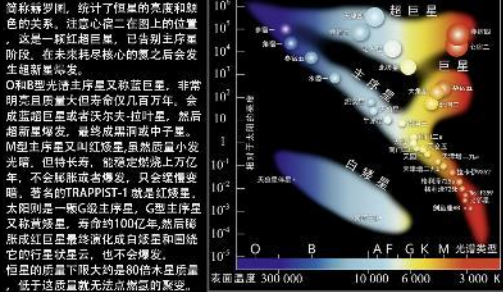


心宿二的光学干涉仪重建图像

心宿二的天文艺术画

心宿二 (Antares) 是最亮的恒星之一，位于天蝎座中心位置，距离地球约550光年。质量是太阳的12倍，也是已知最大的恒星之一。如果把心宿二取代太阳，那至少能把火星吞掉，其亮度更是太阳的上百倍。心宿二发光强度还会毫无规律的变化，是一种变星。此外心宿二有颗相对暗的蓝色伴星，质量是太阳的7.2倍，但因心宿二本身太亮而难以观测这颗伴星。

赫茨-罗素图



简称赫罗图，统计了恒星的亮度和颜色的关系，注意心宿二在图上的位置，这是一颗红超巨星，已告别主序星阶段。在未来耗尽核心的氢之后会发生超新星爆发。O和B型光谱主序星又称蓝巨星，非常明亮且质量大但寿命仅几百万年，会变成蓝超巨星或者沃尔夫-拉叶星，然后超新星爆发，最终成黑洞或中子星。M型主序星又叫红矮星，虽然质量小发光弱，但特长寿，能稳定燃烧上万亿年，不会膨胀或者爆发，只会慢慢变暗，著名的TRAPPIST-1就是红矮星，太阳则是一颗G级主序星，G型主序星又称黄矮星，寿命约100亿年，最后塌缩成红巨星最终演化成白矮星和围绕它的行星状星云，也不会爆发，恒星的质量下限大约是80倍木星质量，低于这质量就无法点燃氢的聚变。



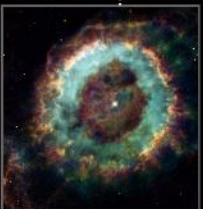
M8礁湖星云

礁湖星云 (Lagoon Nebula) 是射手座里巨大的发射星云，距离地球约4100光年，星云大约横跨100光年。因为星云里的炽热恒星的紫外辐射电离而发光，发此颜色和辐射能量强弱有关。这是活跃的恒星形成区，图中黑暗区域为气体和尘埃，是形成新恒星与行星的原料。



M20三裂星云

三裂星云 (Trifid Nebula) 距离地球5200光年，也是恒星诞生区。混合了发射星云、反射星云 (蓝) 以及暗星云。



NGC6369小鬼星云

小鬼星云 (Little Ghost Nebula) 距离地球2000-5000光年，是行星状星云。是类似太阳的恒星死亡后抛出的物质形成的，中心有一颗白矮星。



NGC6302蝴蝶星云

NGC6302也是一颗行星状星云，距离约3800光年。“蝴蝶”的两翼是恒星两极抛出的气体形成的，红色边缘是氢发出的光。



NGC6334猫掌星云

猫掌星云 (Cat's Paw Nebula) 距离地球约5500光年，是发射星云。这片星云孕育了不少非常明亮的年轻红巨星，就形成了一颗引人注目的开放星团。



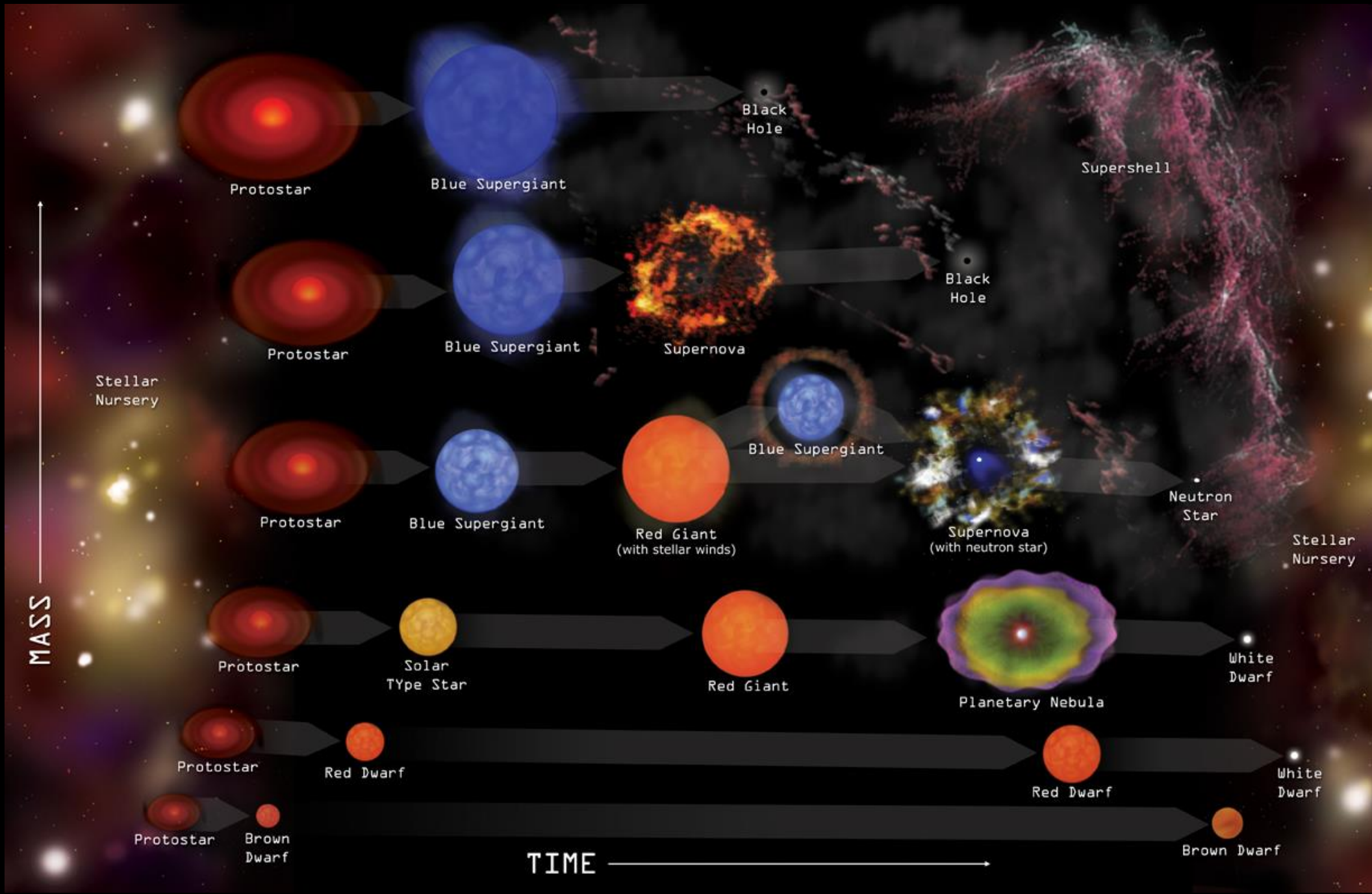
NGC6357龙虾星云

龙虾星云 (Lobster Nebula) 距离约8000光年，也是发射星云。这片星云孕育了不少非常明亮的年轻红巨星，就形成了一颗引人注目的开放星团。



M4球状星团

M4星团距离约7200光年。这个球状星团包含了约1万颗古老恒星，半径却只有35光年，使得恒星之间距离很近。



MASS ↑

TIME →

Stellar Nursery

Stellar Nursery

Protostar

Blue Supergiant

Black Hole

Supershell

Protostar

Blue Supergiant

Black Hole

Supernova

Stellar Nursery

Protostar

Blue Supergiant

Red Giant
(with stellar winds)

Blue Supergiant

Supernova
(with neutron star)

Neutron Star

Stellar Nursery

Protostar

Solar Type Star

Red Giant

Planetary Nebula

White Dwarf

Protostar

Red Dwarf

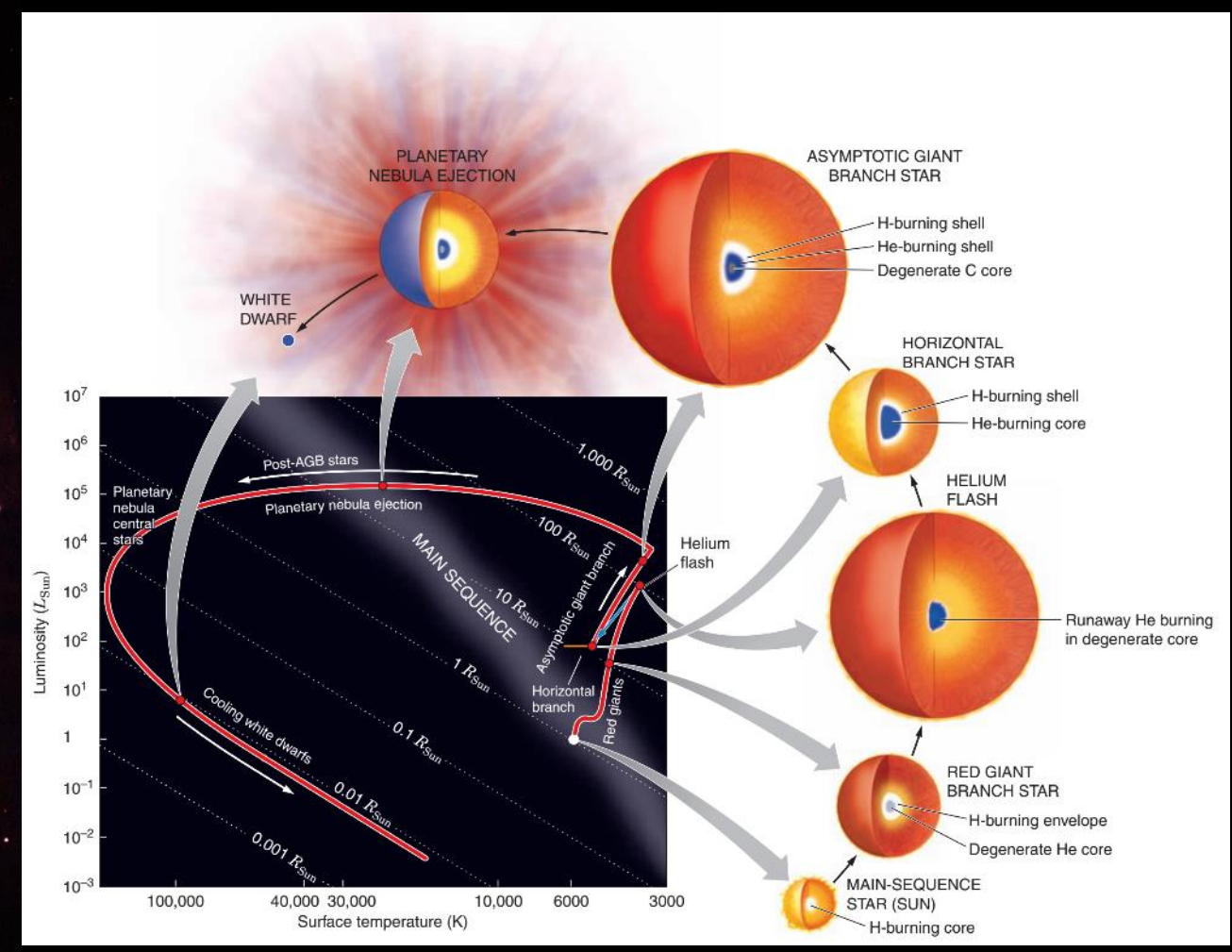
Red Dwarf

White Dwarf

Protostar

Brown Dwarf

Brown Dwarf





Before Explosion

—| 0.1 Astronomical Units

20 Days After Explosion

—| 50 Astronomical Units

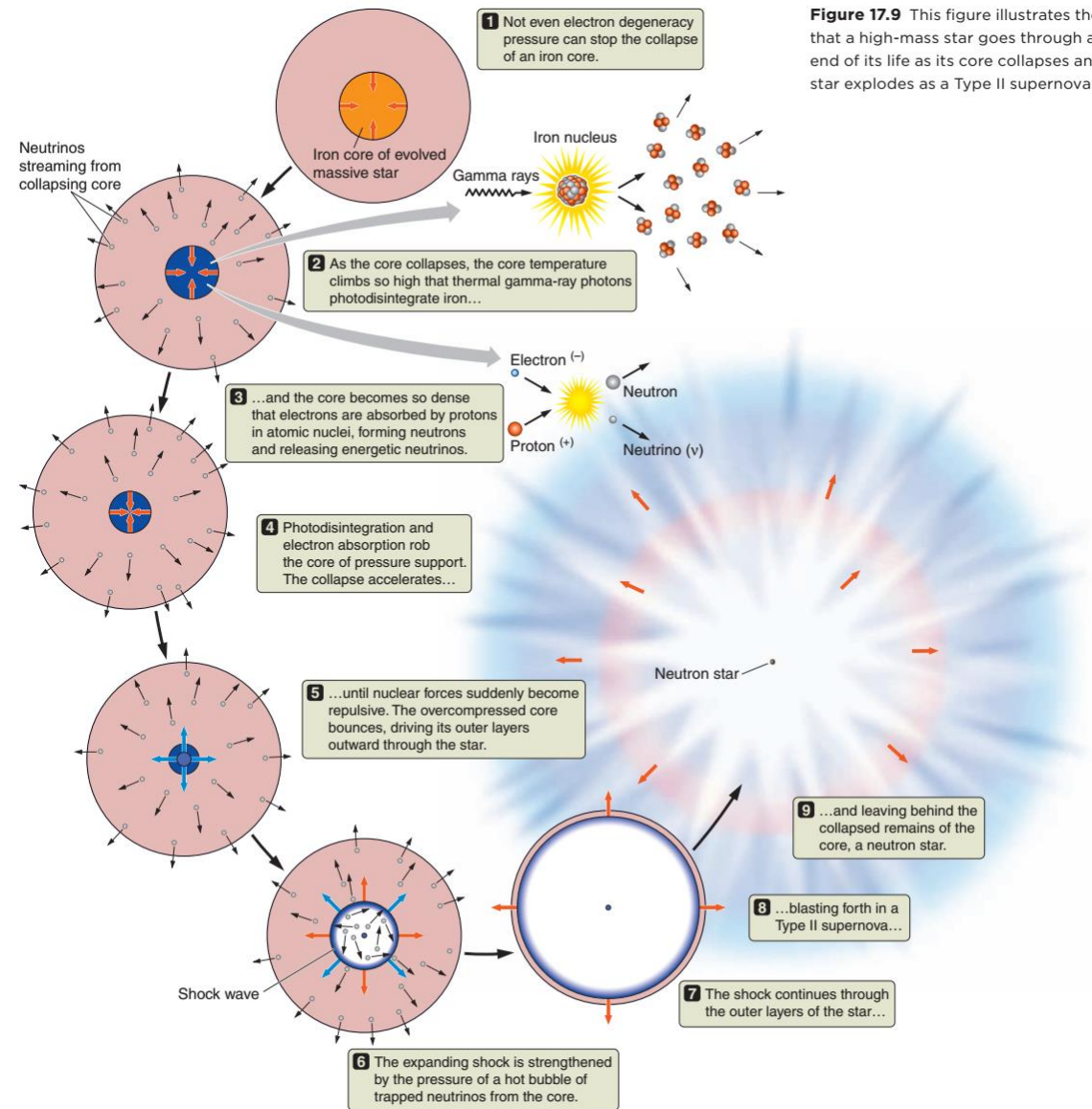
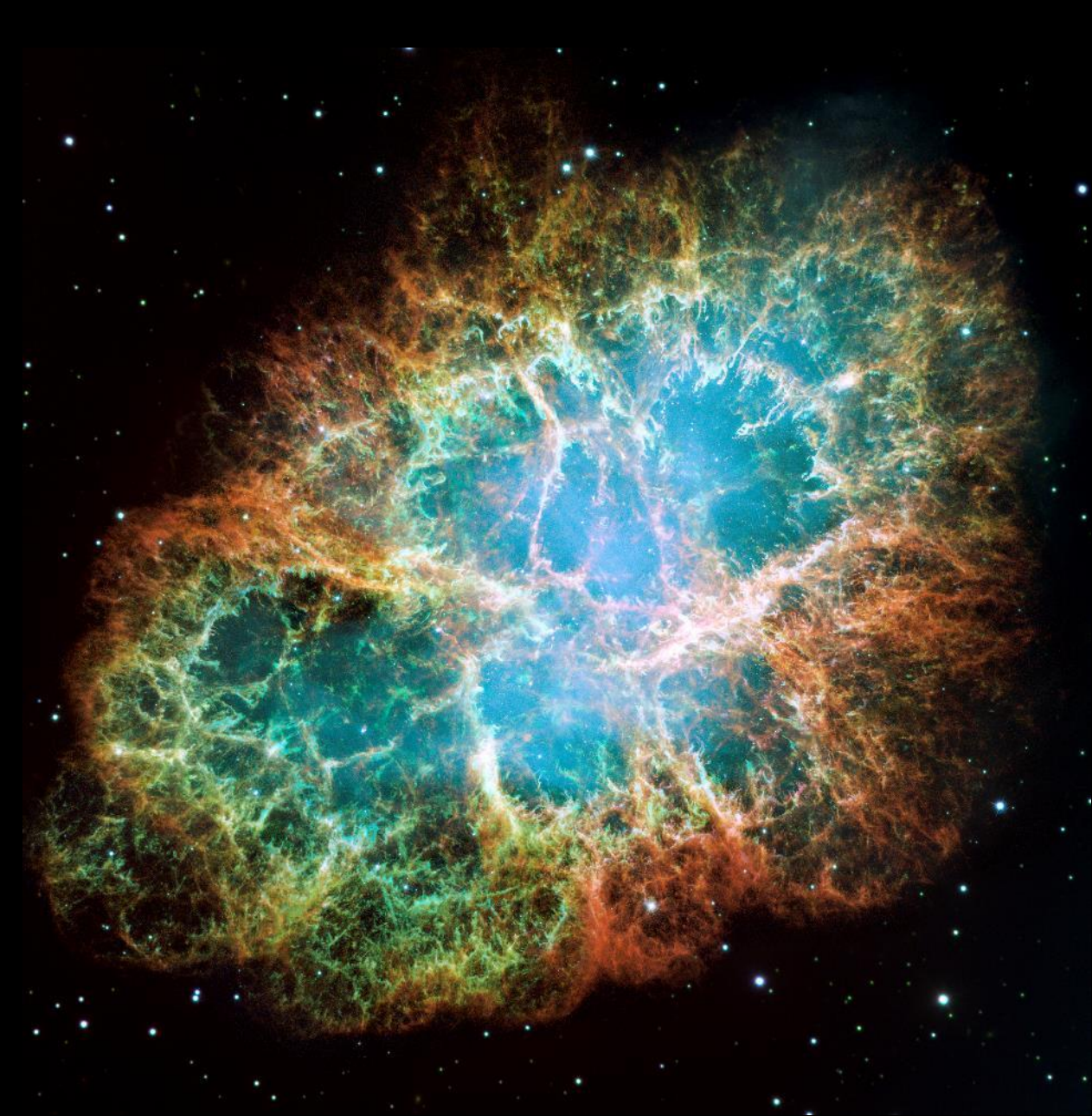
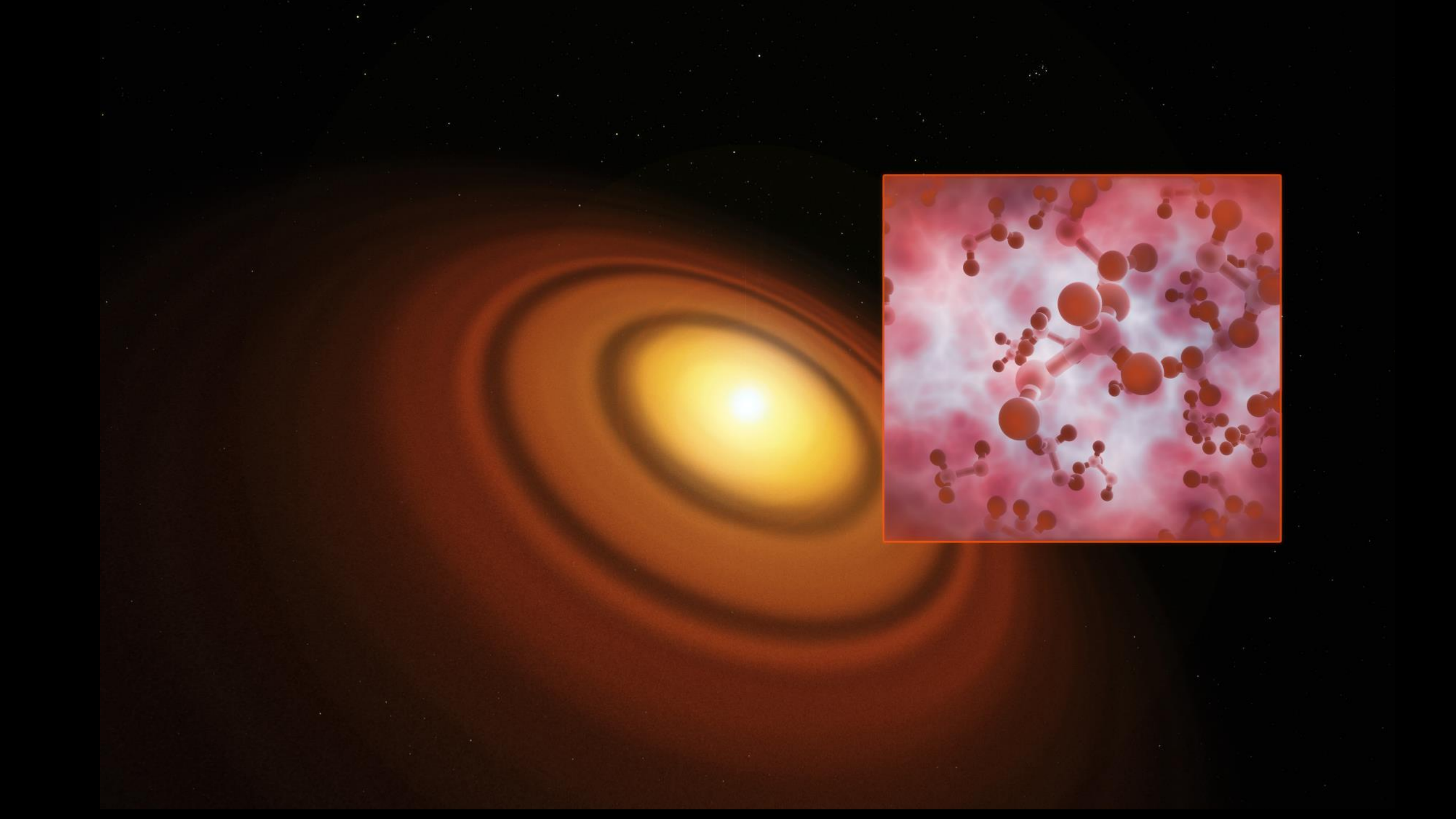


Figure 17.9 This figure illustrates the stages that a high-mass star goes through at the end of its life as its core collapses and the star explodes as a Type II supernova.





Cis-Lunar Space

A deep dpace harbor for expanded human presence

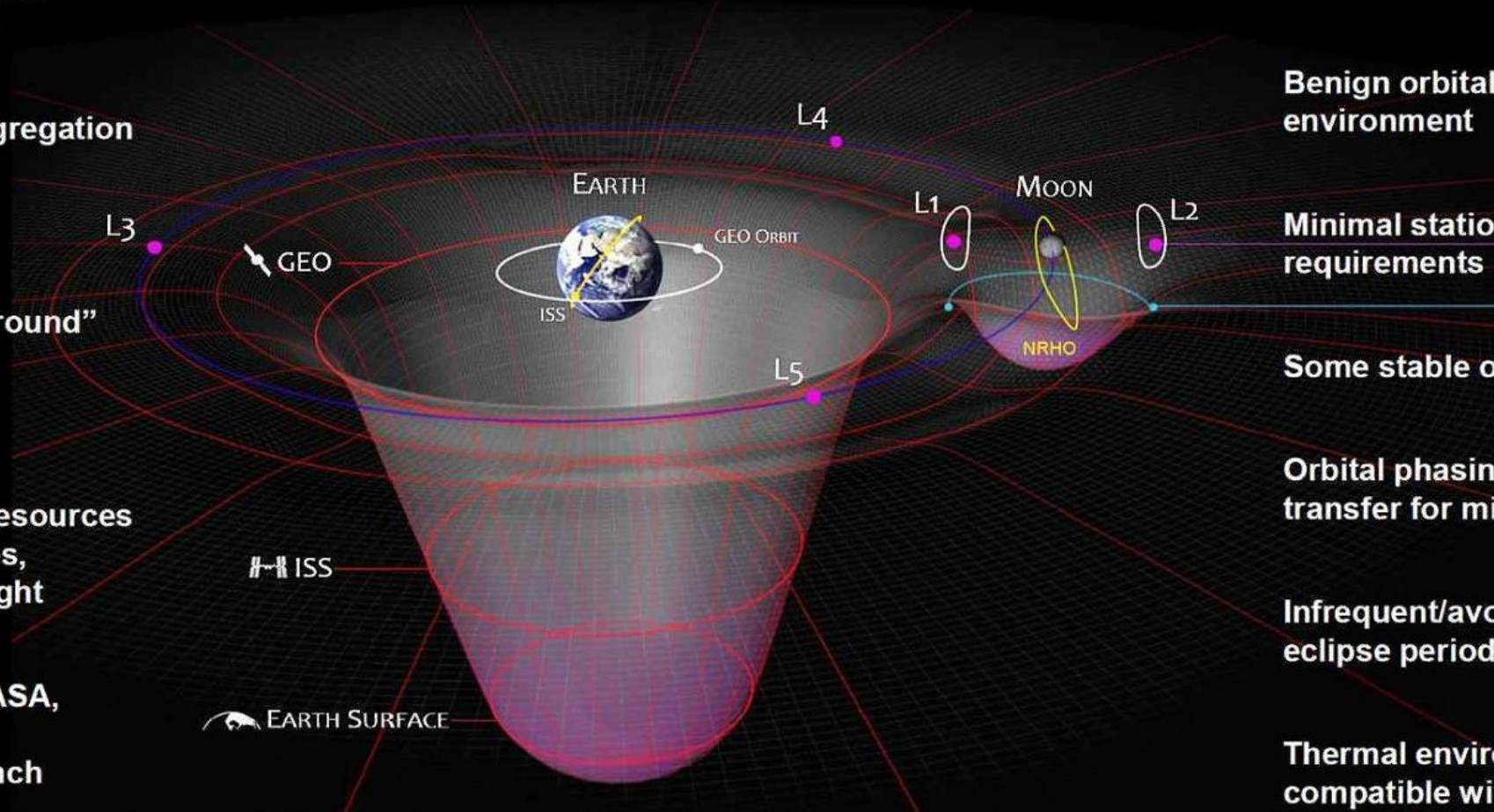
Only ~3 to 5 days away
from Earth yet farther
than Apollo went

Ideal mission aggregation
location

The next "high ground"
beyond GEO

Access to local resources
including volatiles,
gravity and sunlight

Accessible by NASA,
commercial, and
international launch
systems



True deep space radiation
environment

Benign orbital debris
environment

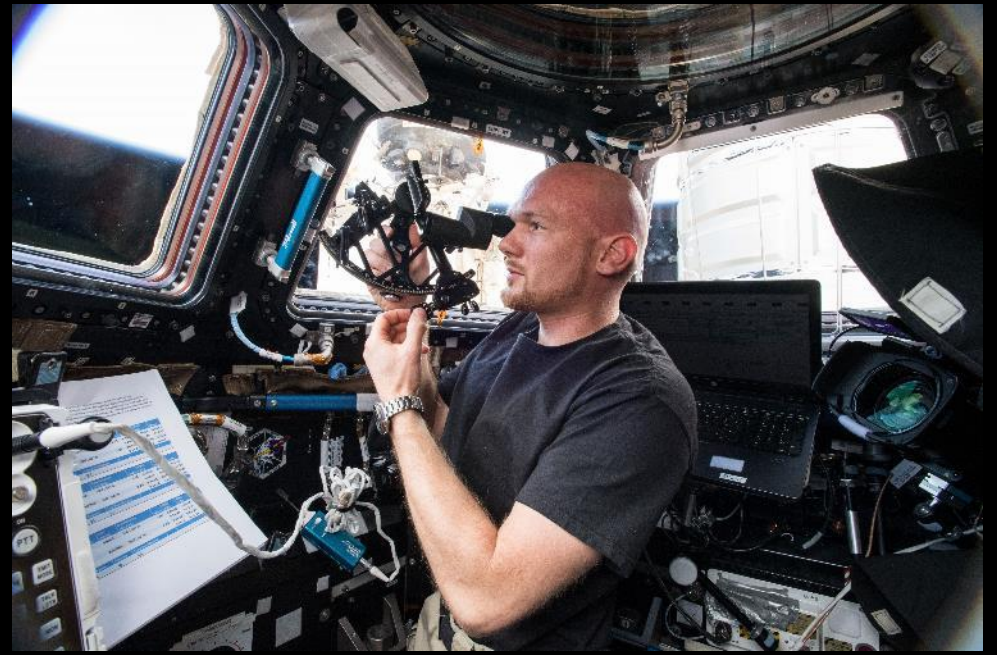
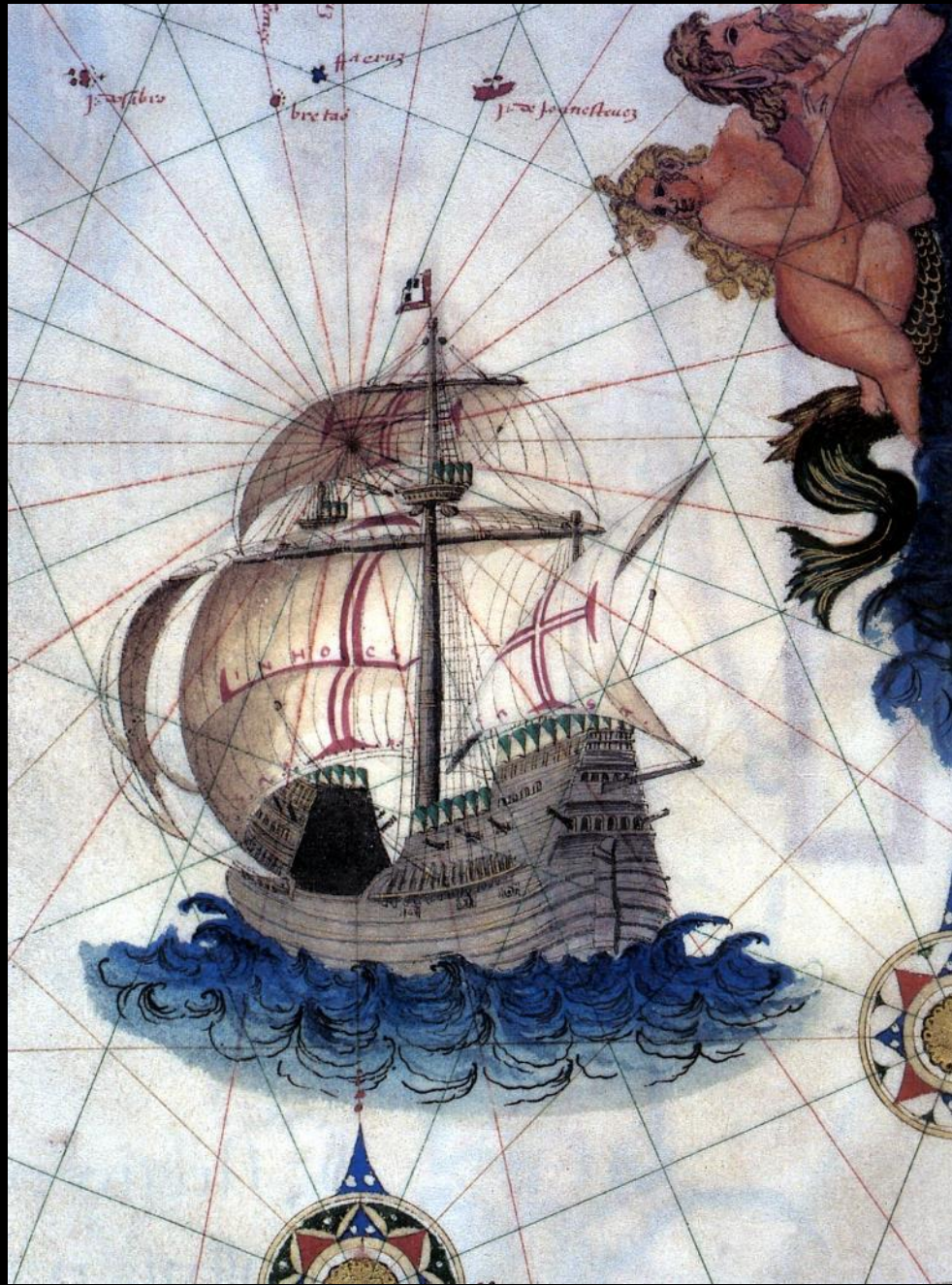
Minimal station keeping
requirements

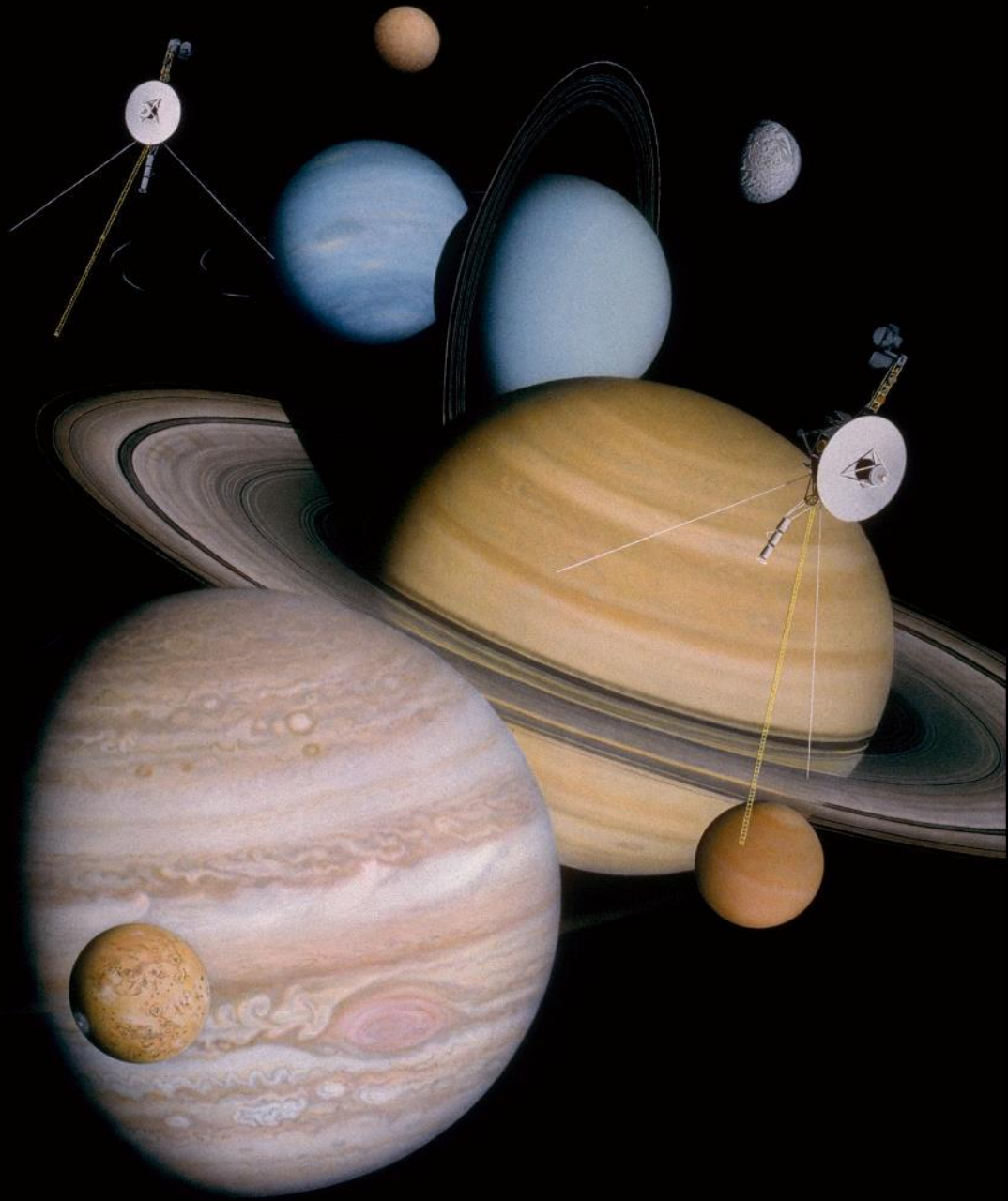
Some stable orbits

Orbital phasing and
transfer for minimal energy

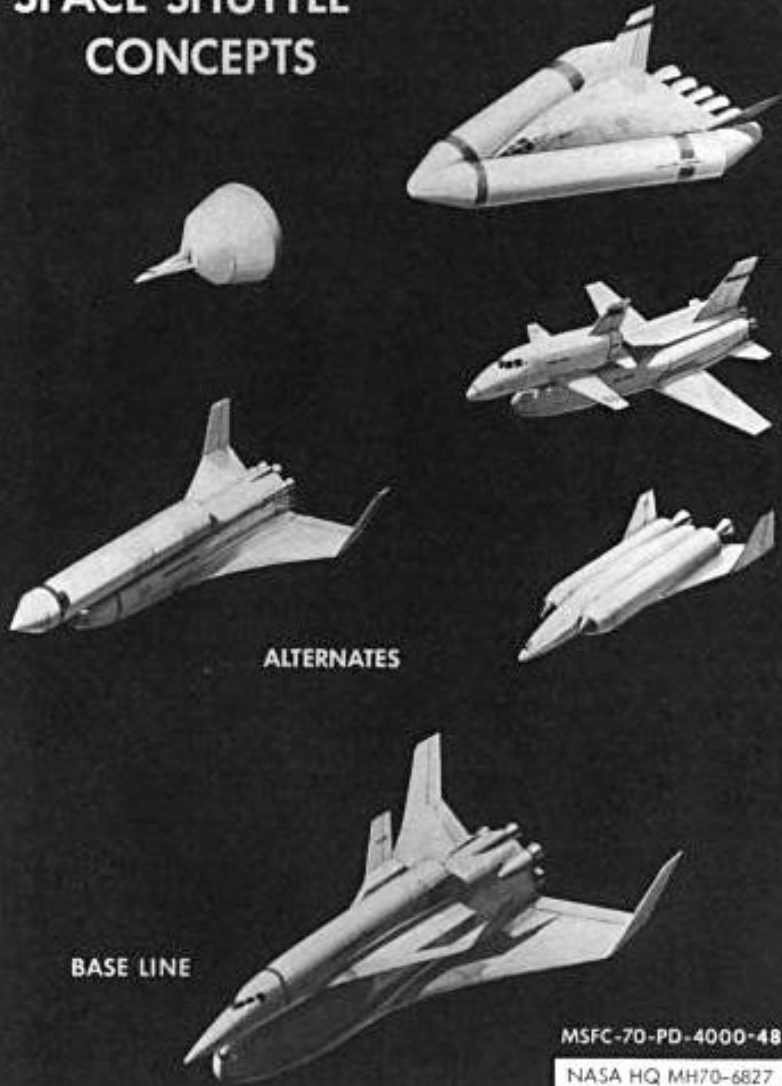
Infrequent/avoidable
eclipse periods

Thermal environment
compatible with cryogenic
oxygen and methane





SPACE SHUTTLE CONCEPTS



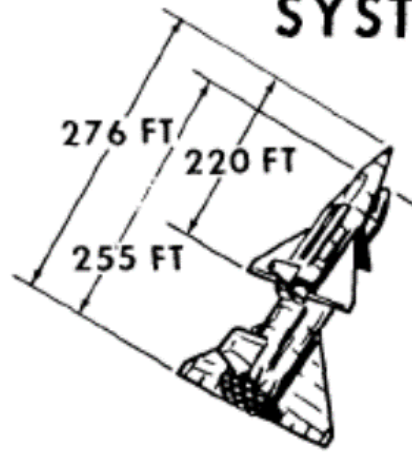
ALTERNATES

BASE LINE

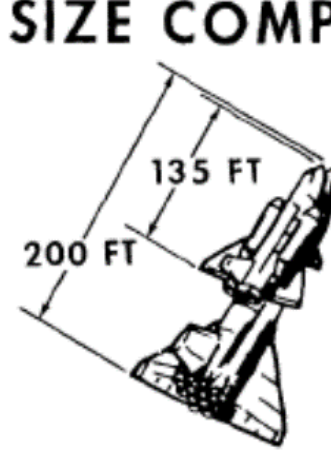
MSFC-70-PD-4000-48

NASA HQ MH70-6827

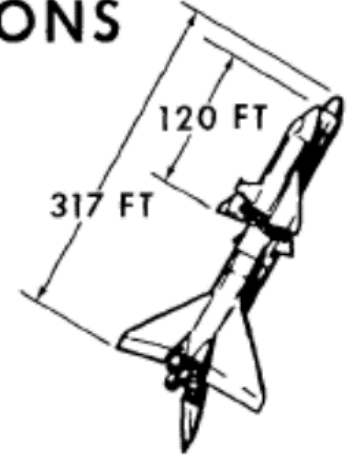
SOME PHASE B SPACE SHUTTLE SYSTEM SIZE COMPARISONS



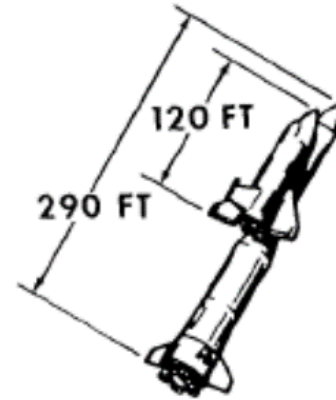
FULLY REUSABLE



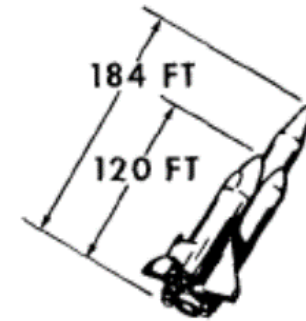
EXTERNAL LH2 TANKS



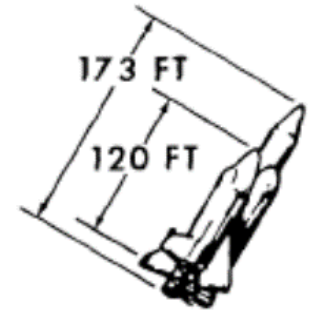
SERIES LIQUID



F-1 FLYBACK



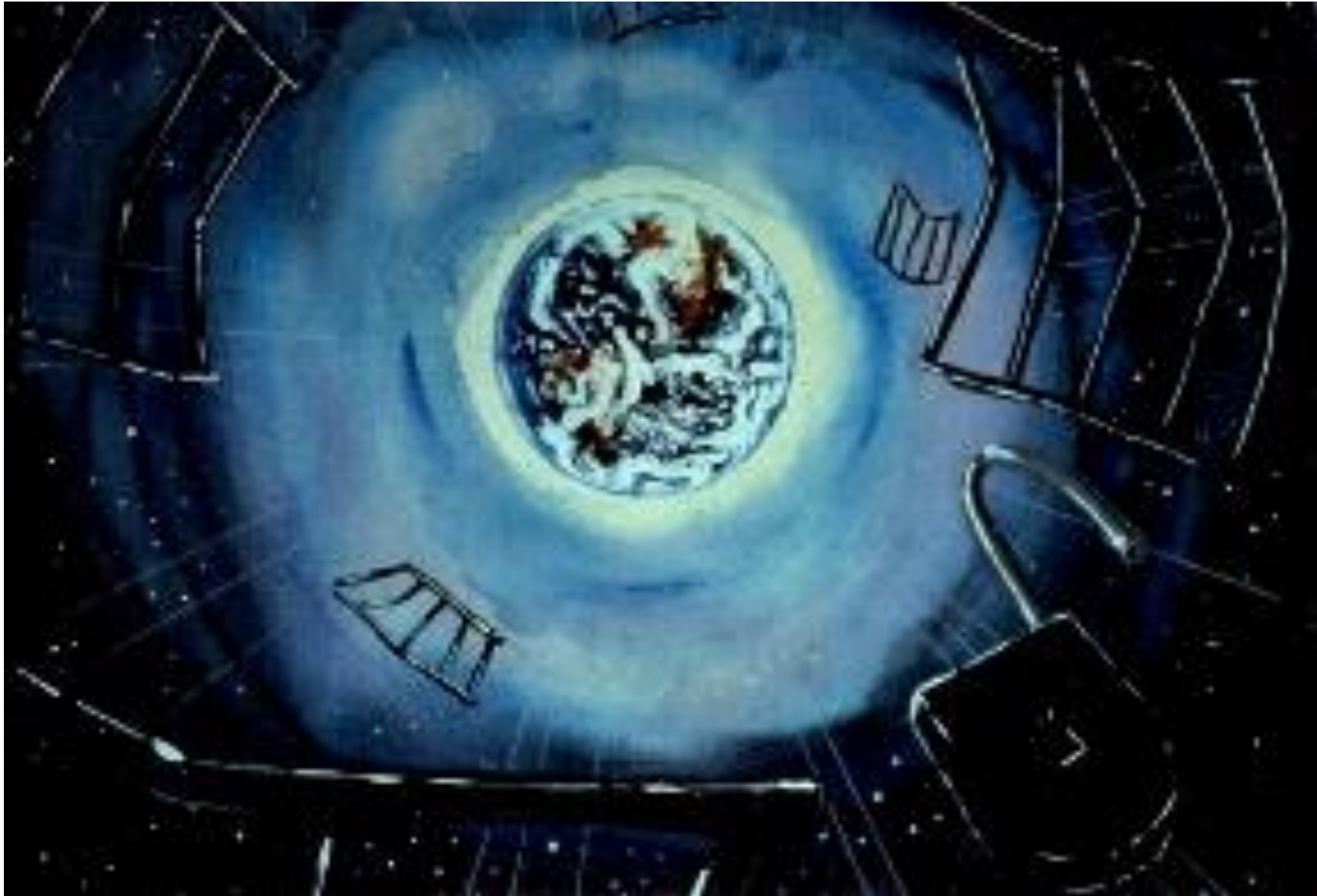
PARALLEL LIQUID



PARALLEL SOLID ROCKET MOTOR







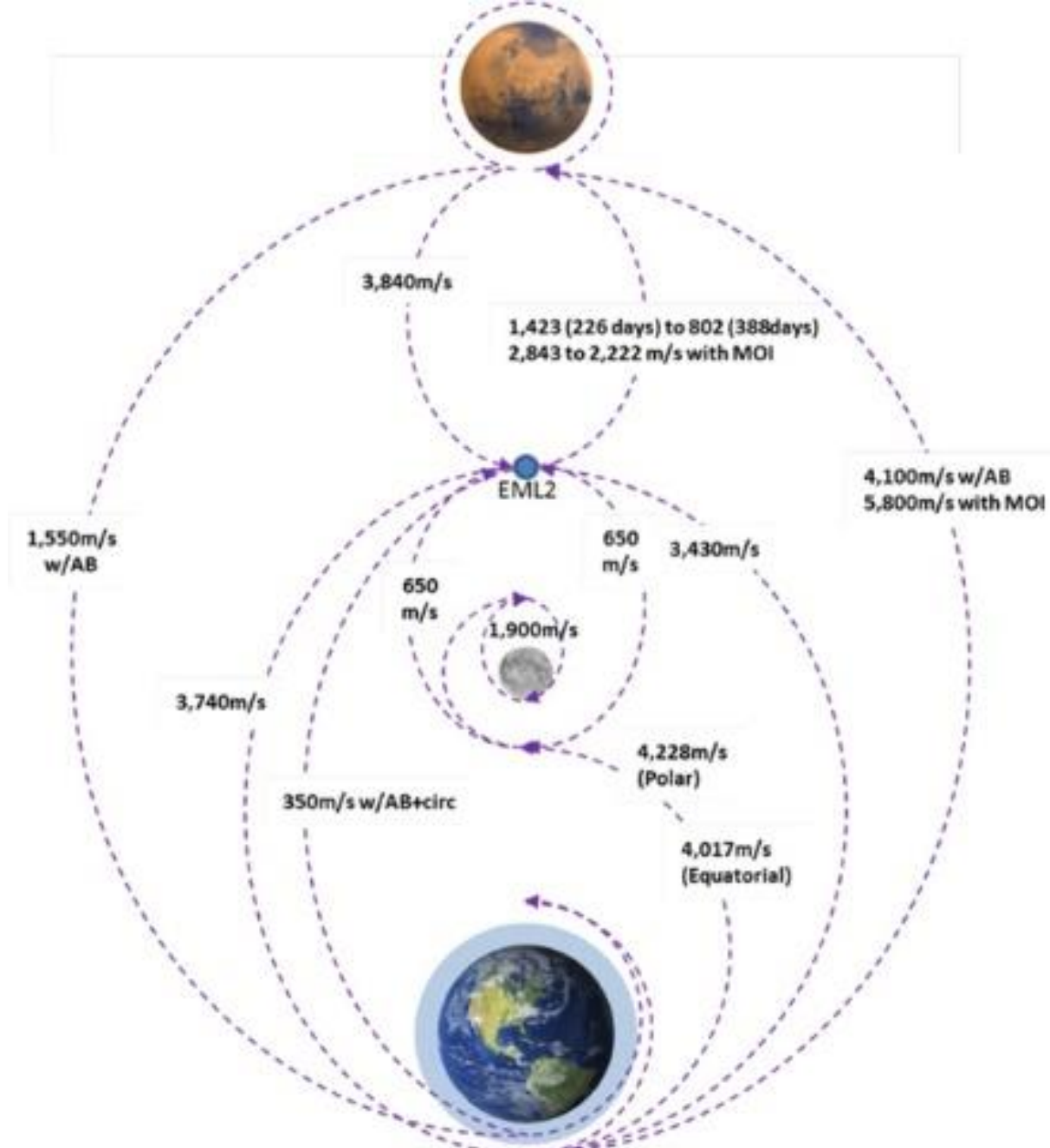
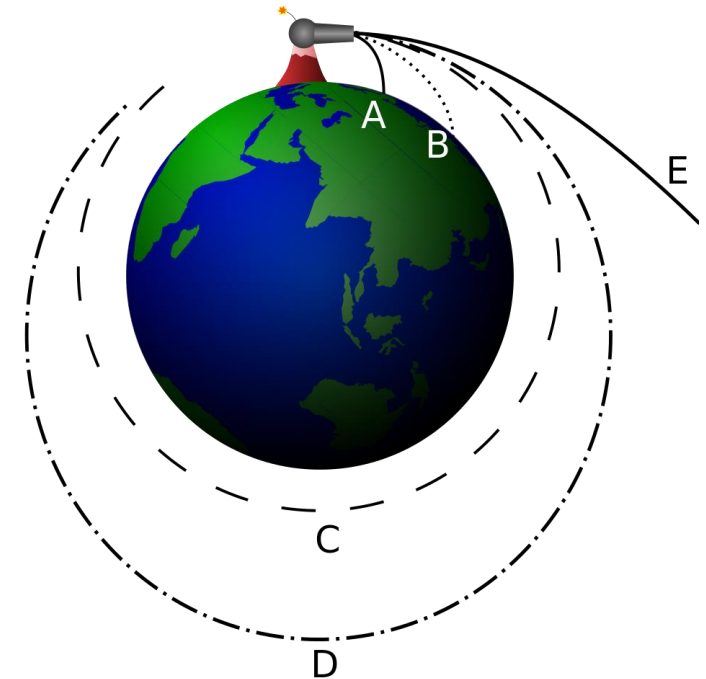
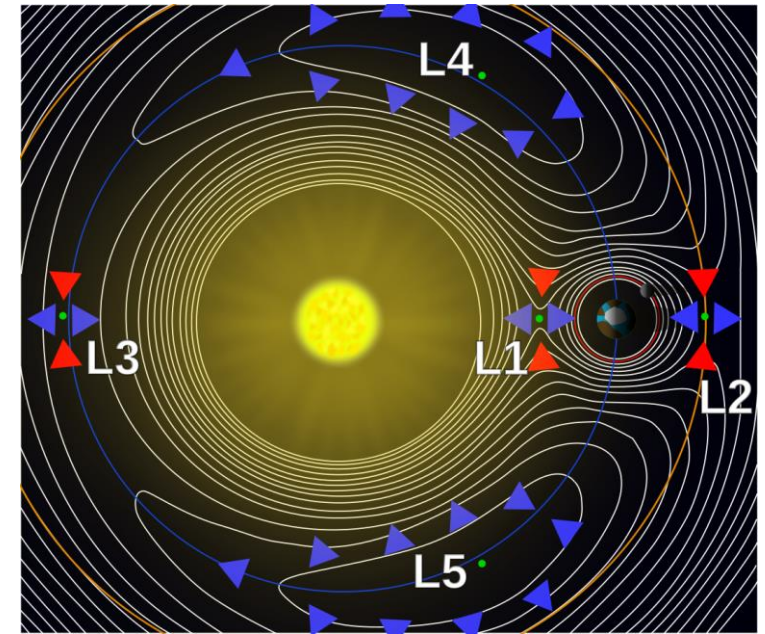
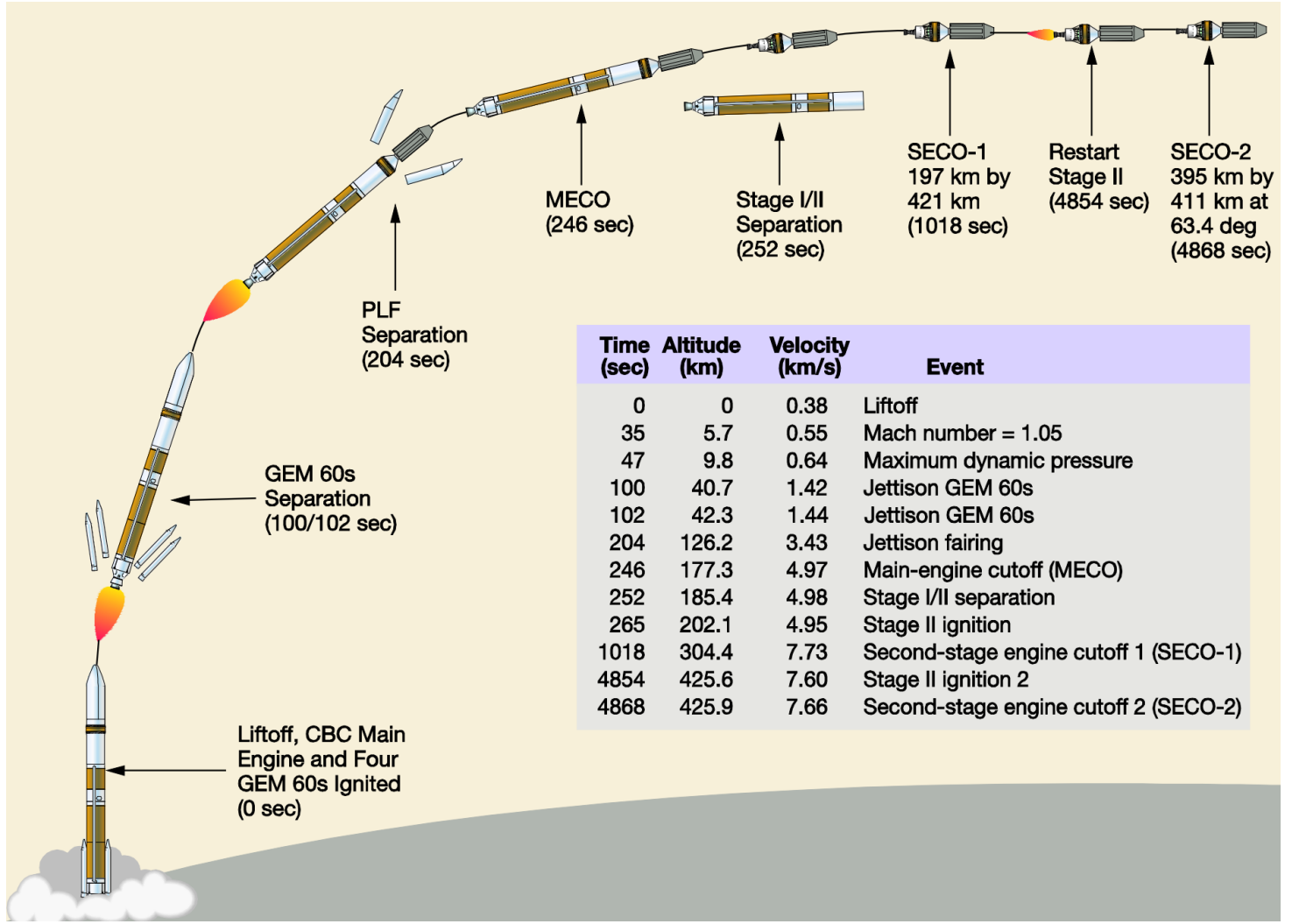
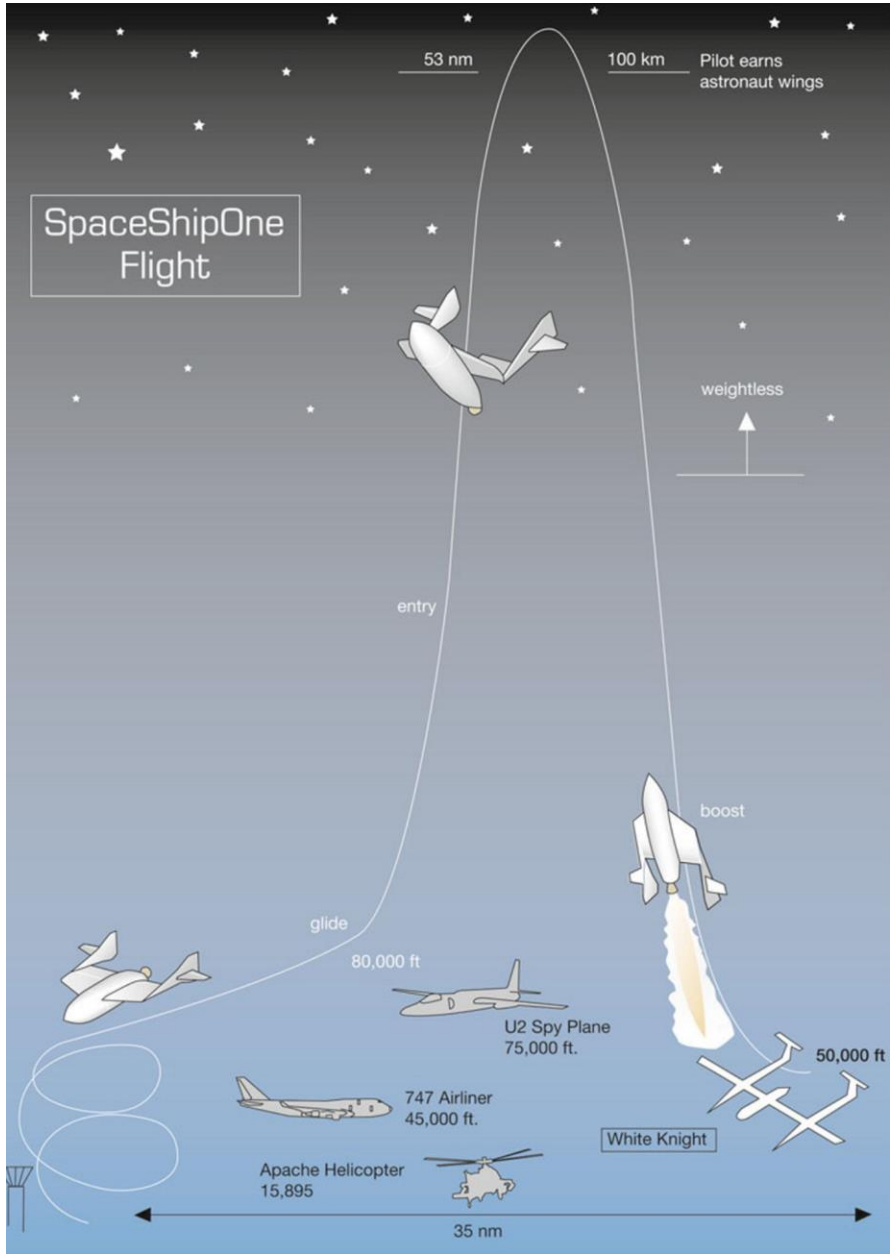


Figure T-2. Transfer Performance DeltaV





A GIANT

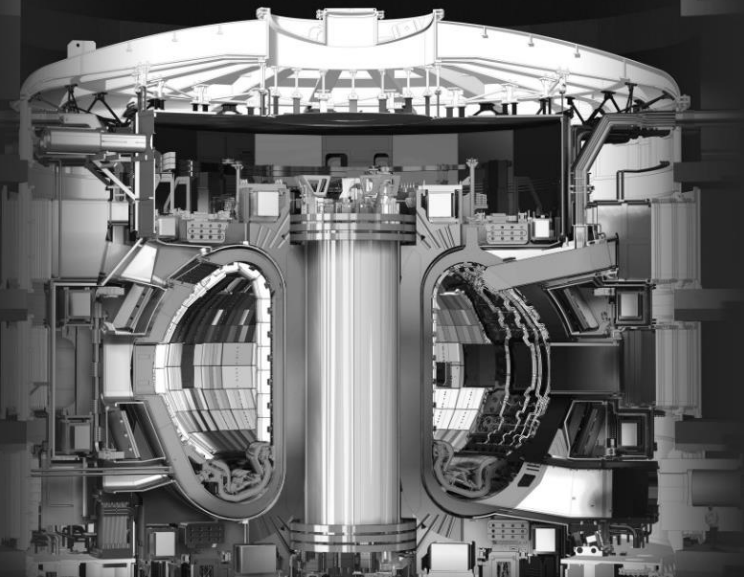
23000_t
Machine weight

10X THE CORE OF THE SUN

150_{million°C}
Plasma temperature

FUSION ENERGY

500_{MW}
Output power

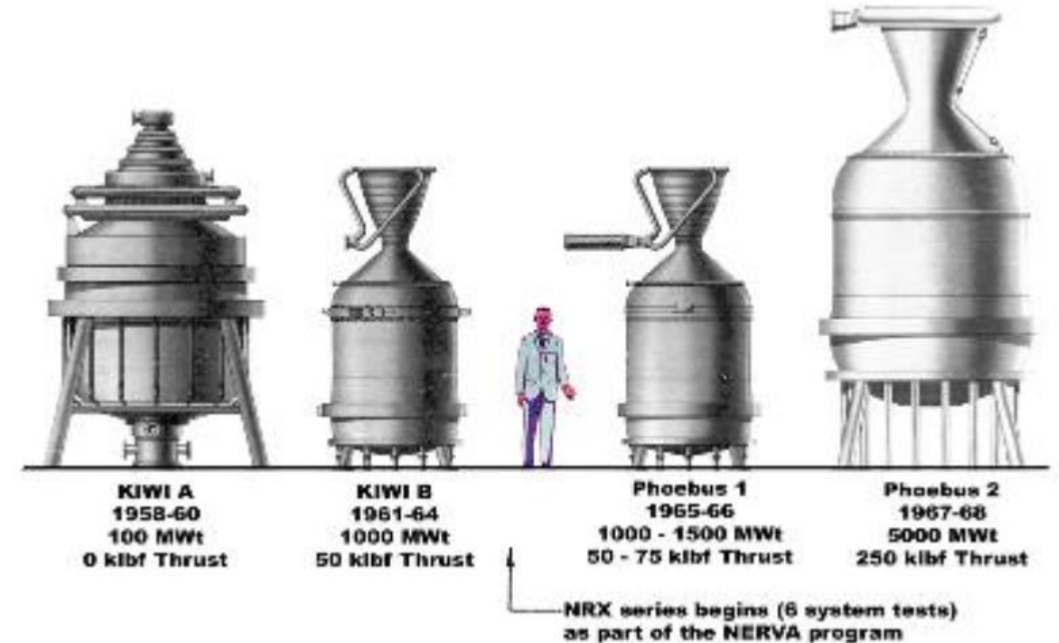
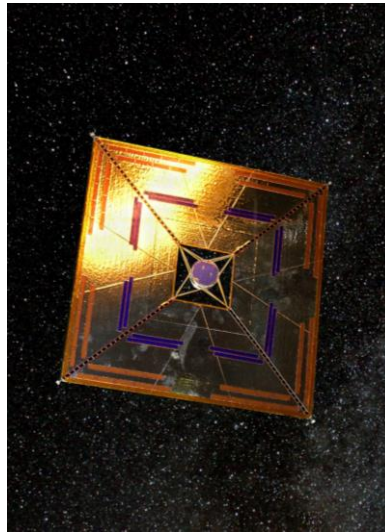
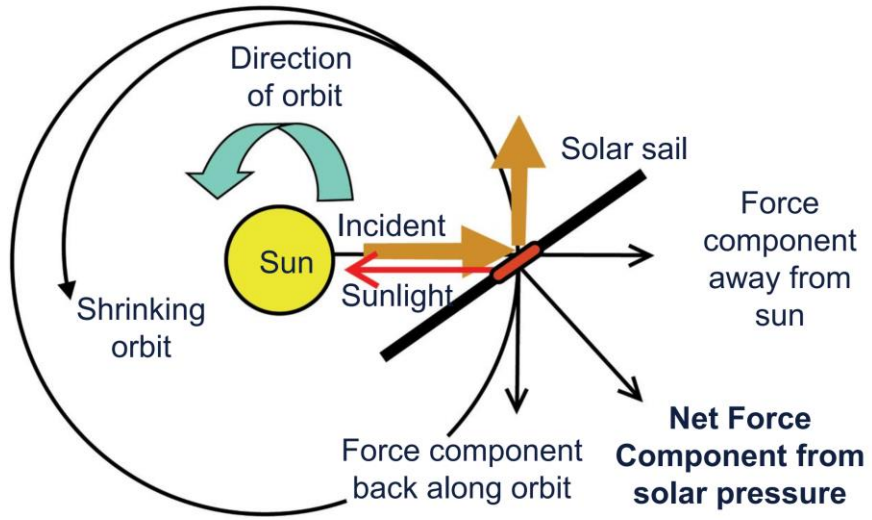
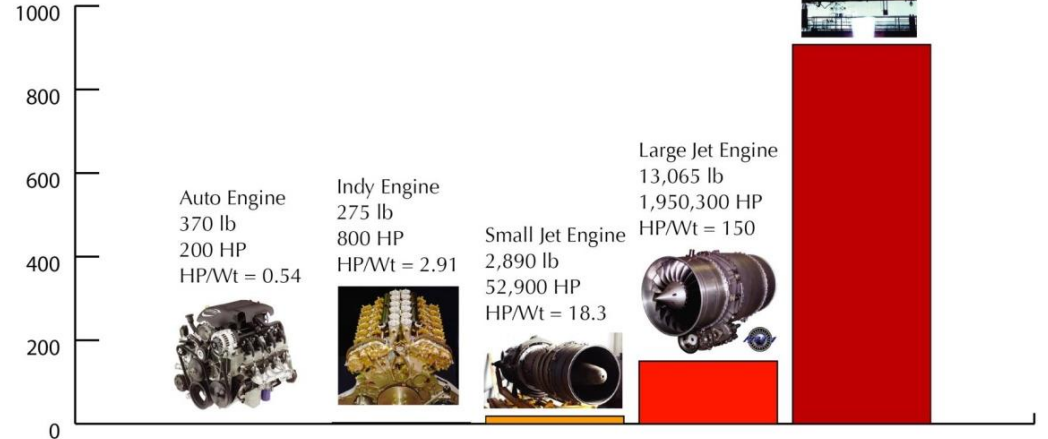


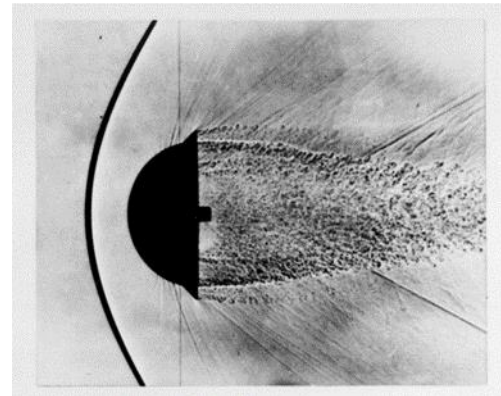
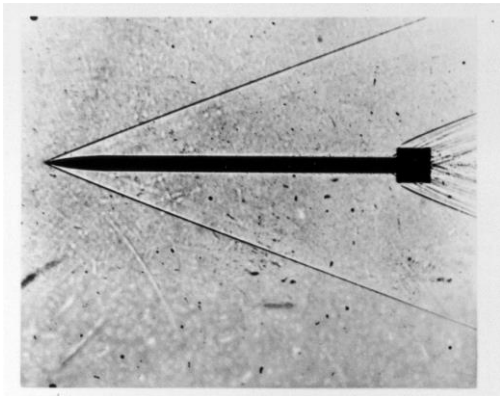
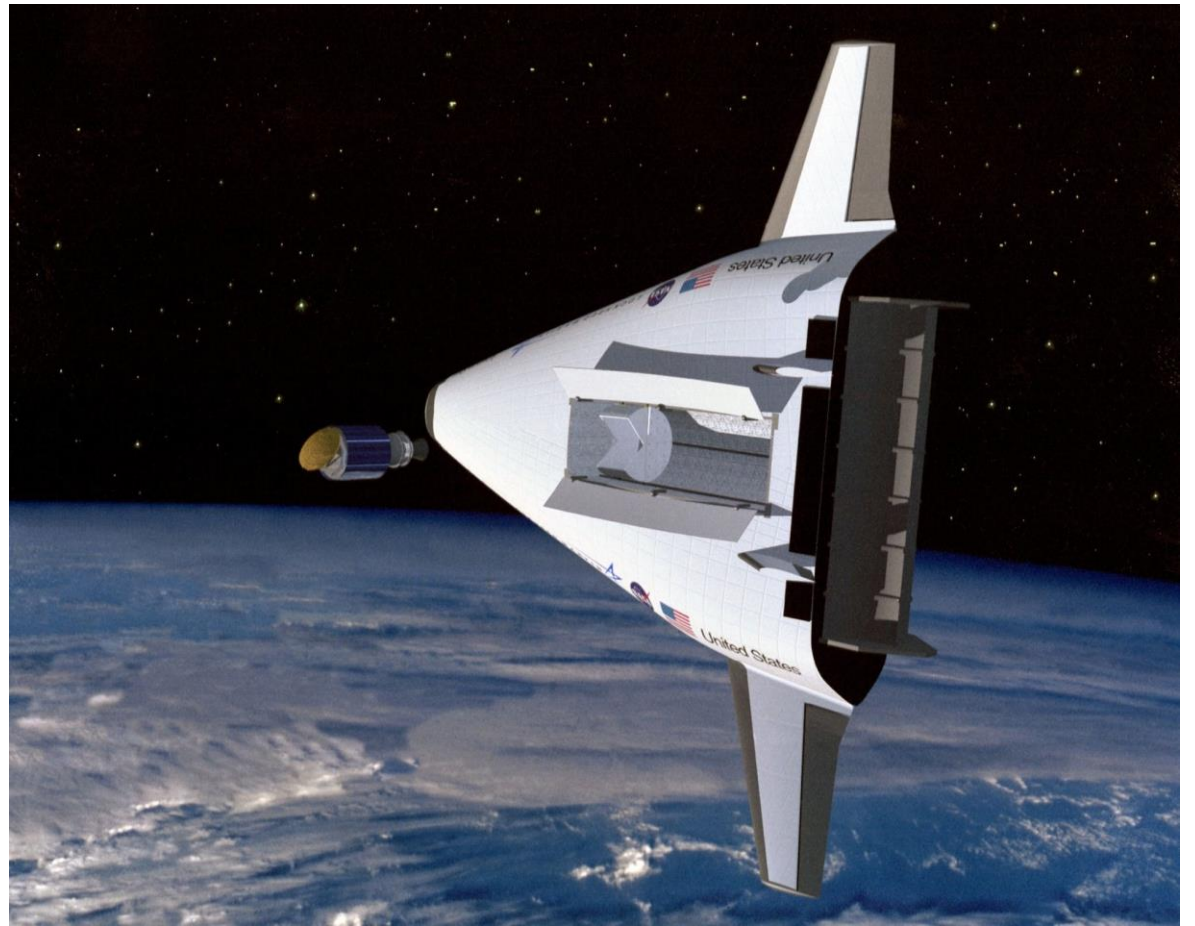
THE ITER TOKAMAK

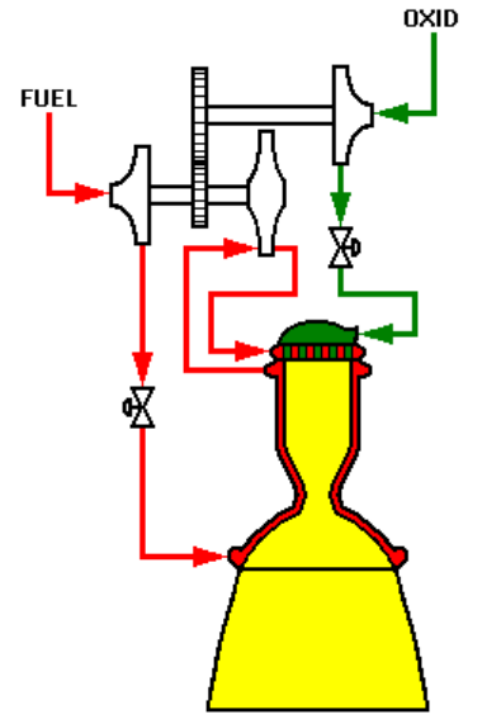
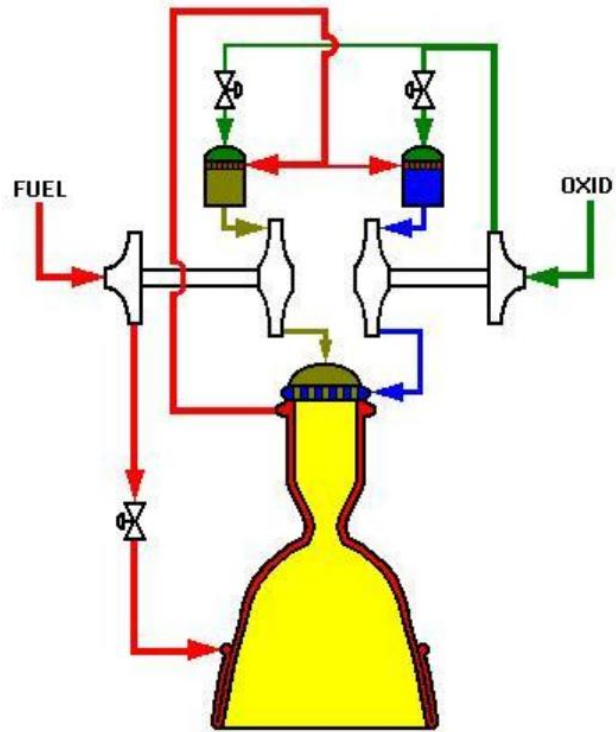
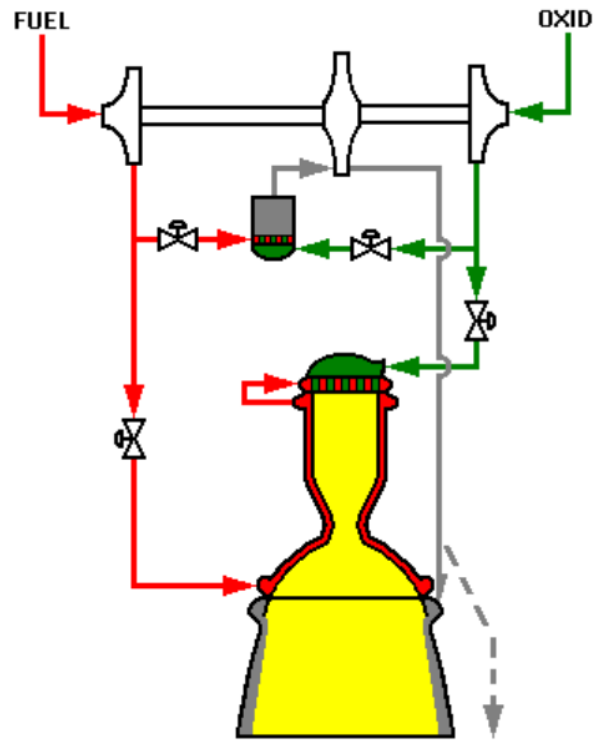
The tokamak is an experimental machine designed to harness the energy of fusion. ITER will be the world's largest tokamak, with a plasma radius (R) of 6.2 m and a plasma volume of 840 m³.

READ MORE

HP/Wt

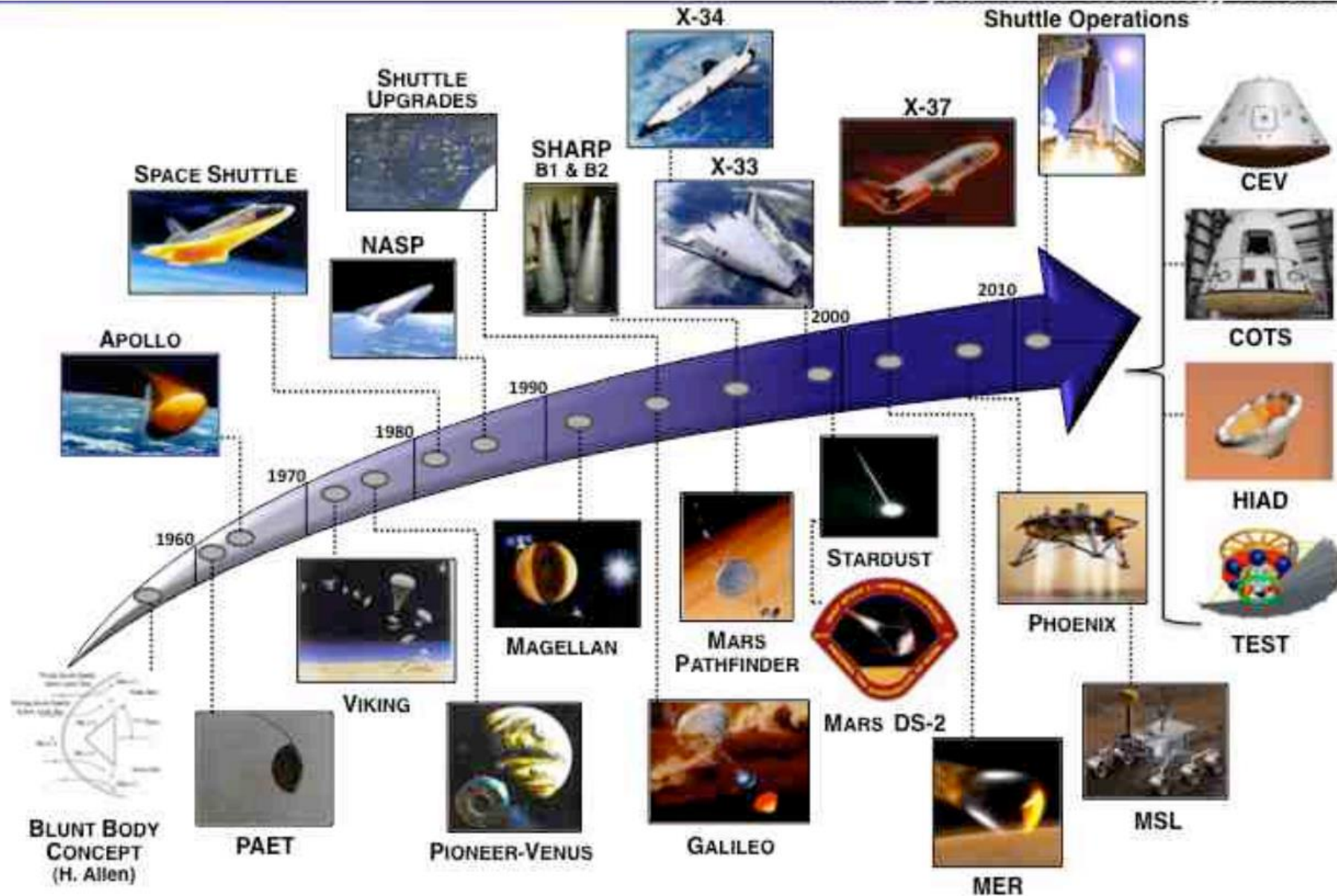




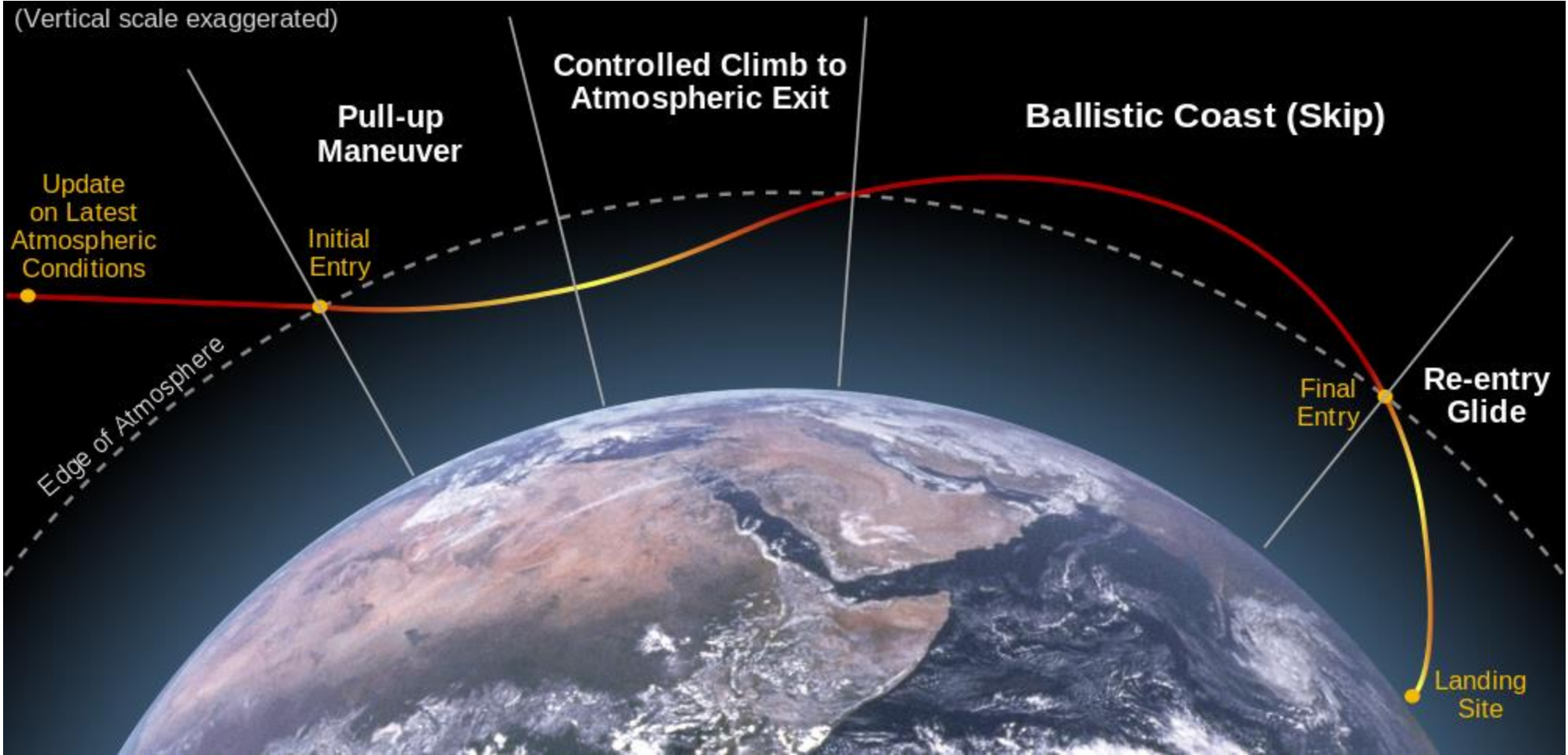




NASA Entry Vehicles / Missions Supported by Ames

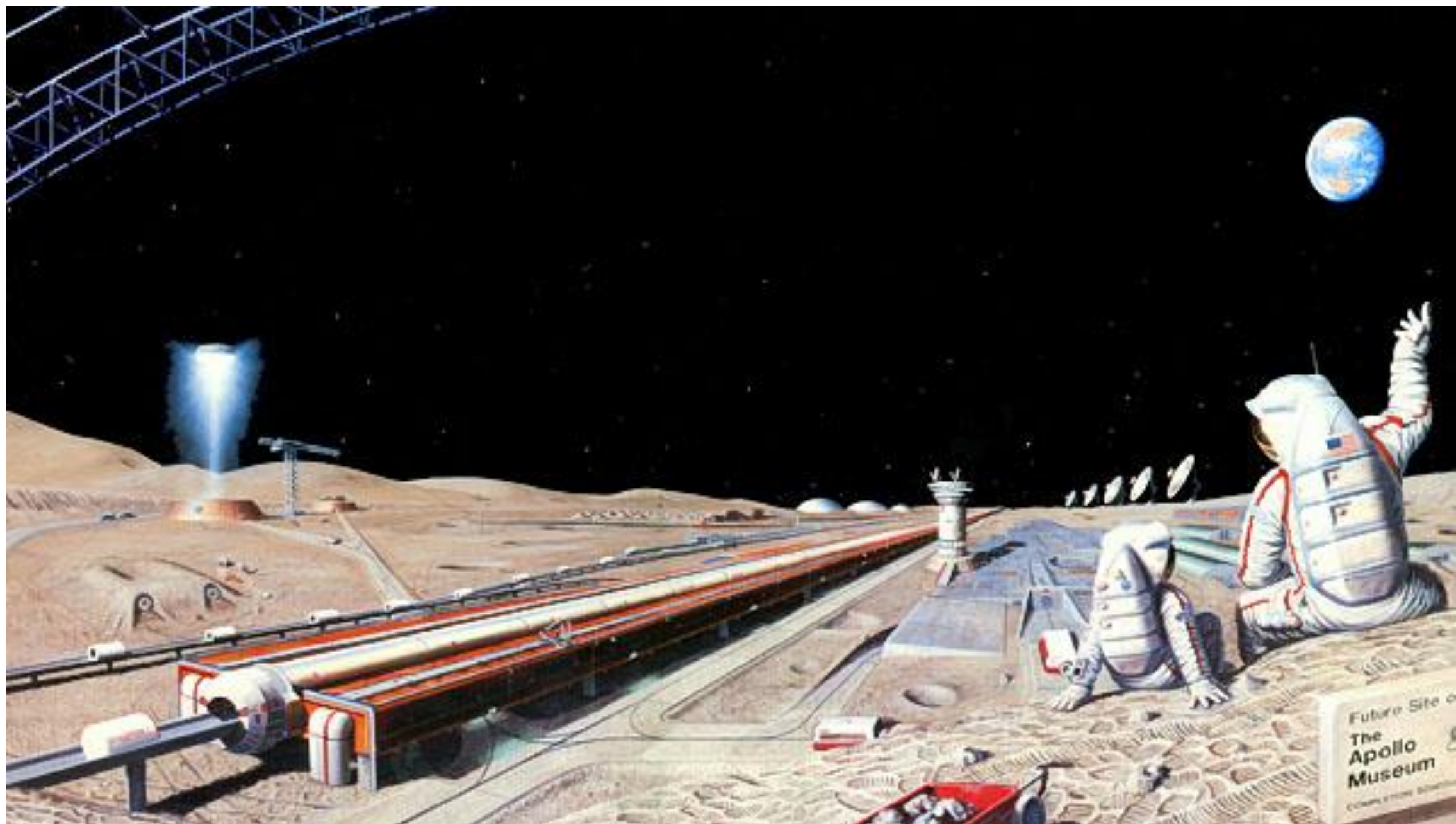


(Vertical scale exaggerated)

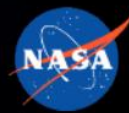


Booster	Payload	Stage 1	Stage 2	Other	Total	Price
Conventional	20 tons	\$90M	\$30M	\$80M	\$200M	\$10,000/kg
Falcon 9 Expendable	20	60	20	40	\$120M	\$6,000/kg
Falcon 9 Reusable Stage 1	20	0	20	40	\$60M	\$3,000/kg
Falcon 9 Reusable	20	0	0	40	\$40M	\$2,000/kg
Starship	160	0	0	40	\$40M	\$250/kg
Falcon 9 Reusable 10X Flights	20	0	0	4	\$4M	\$200/kg
Starship 10X Flights	160	0	0	0	\$4M	\$25/kg





SLS Offers Unrivaled Payload Volume



- ◆ SLS is investigating utilizing existing fairings for early cargo flights, offering payload envelope compatibility with design for current EELVs
- ◆ Phase A studies in work for 8.4m and 10 m fairing options



4m x 12m
(100 m³)



5m x 14m
(200 m³)



5m x 19m
(300 m³)



8.4m x 31m
(1200 m³)



10m x 31m
(1800 m³)

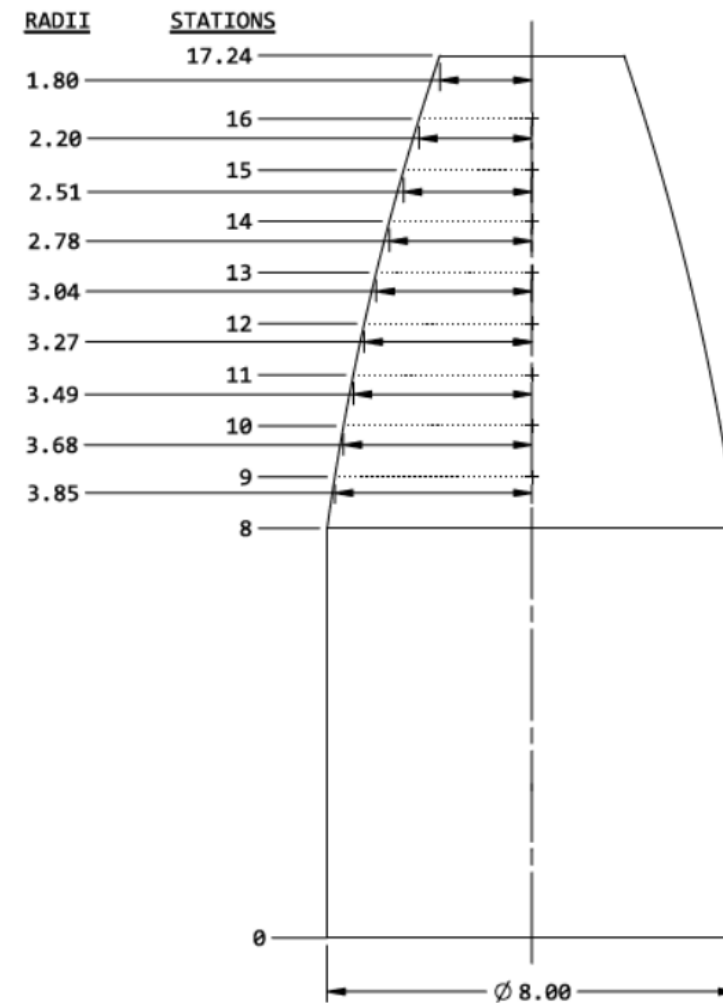
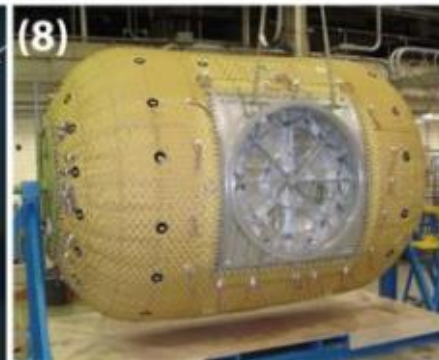
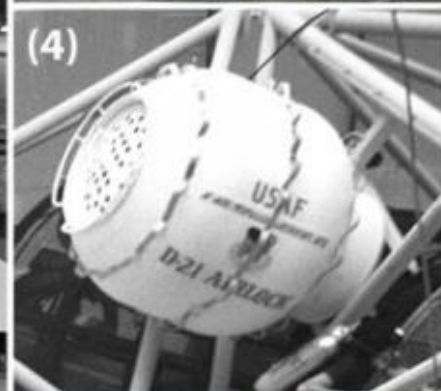
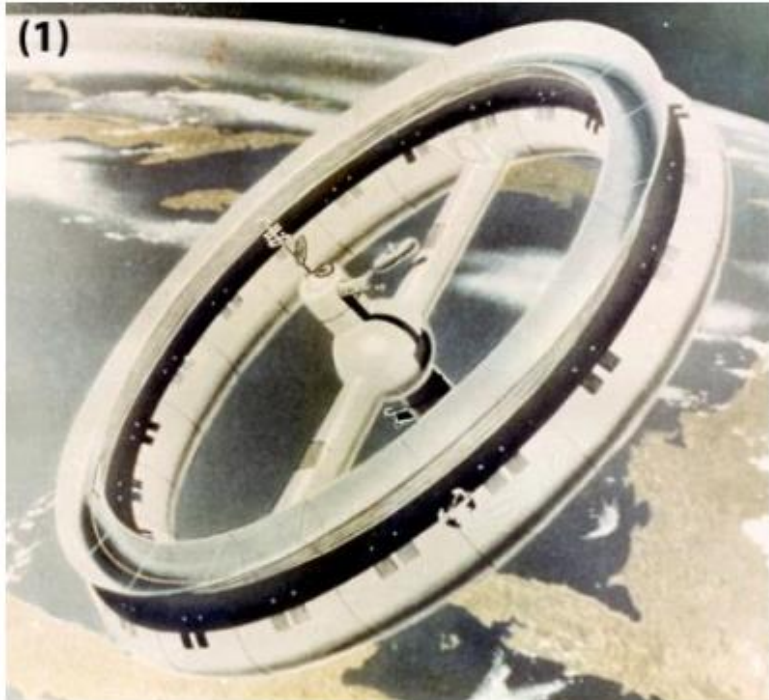
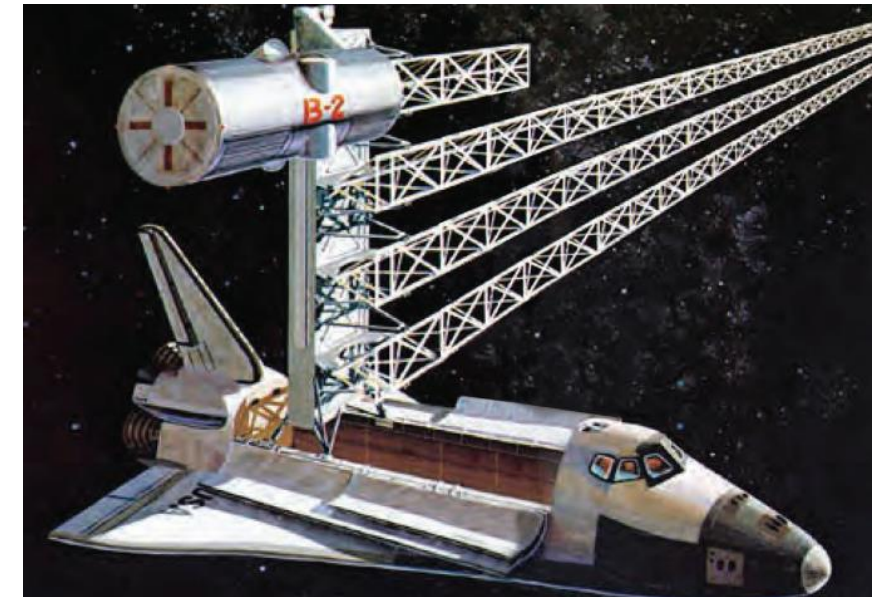
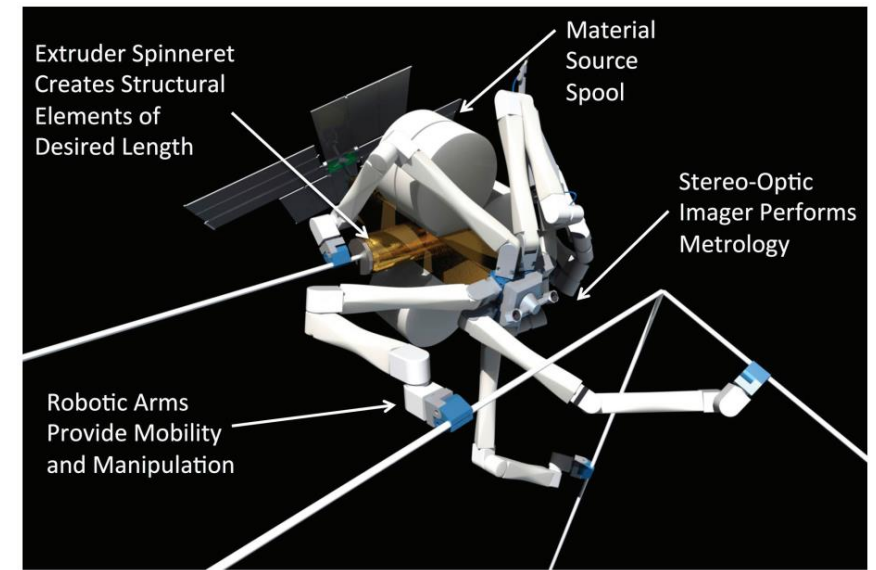
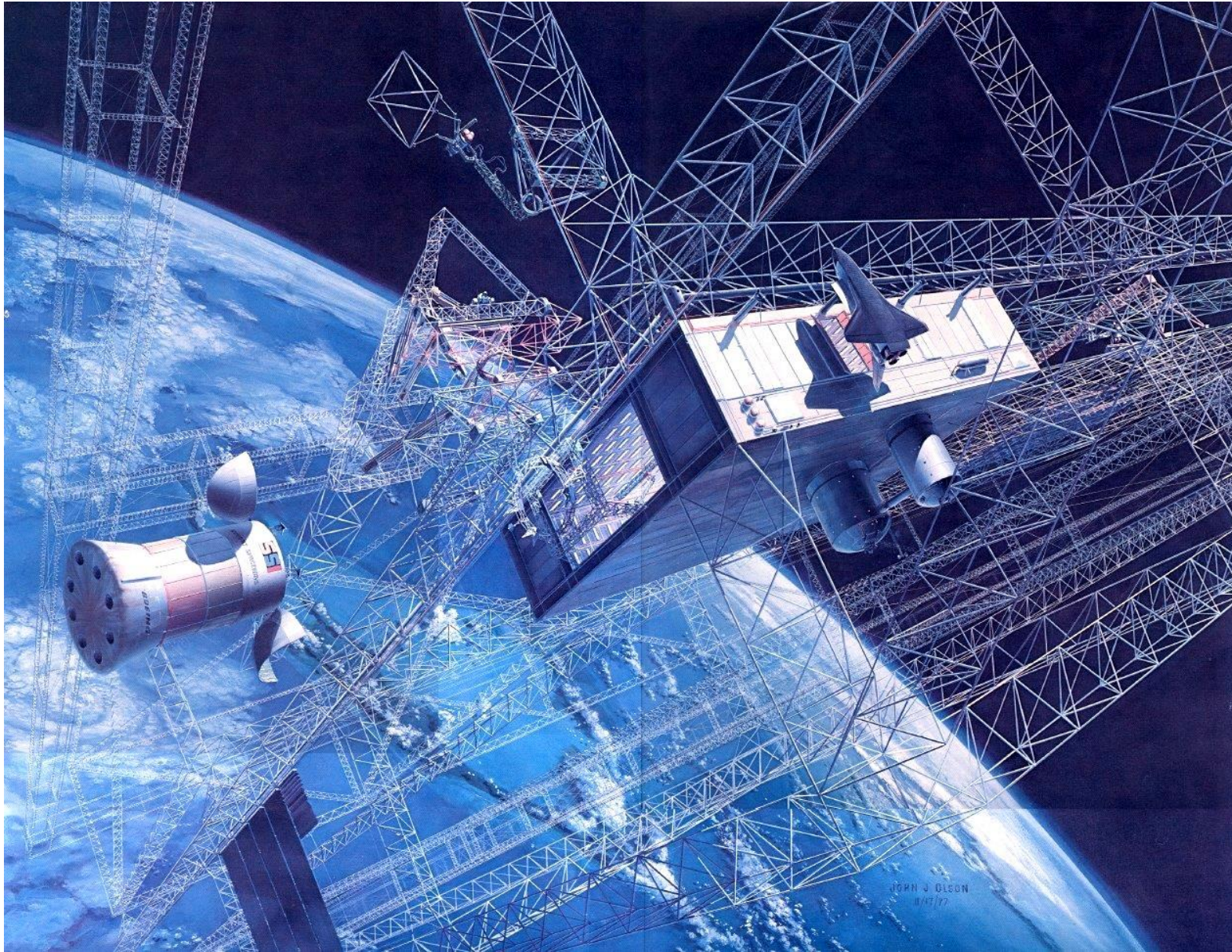
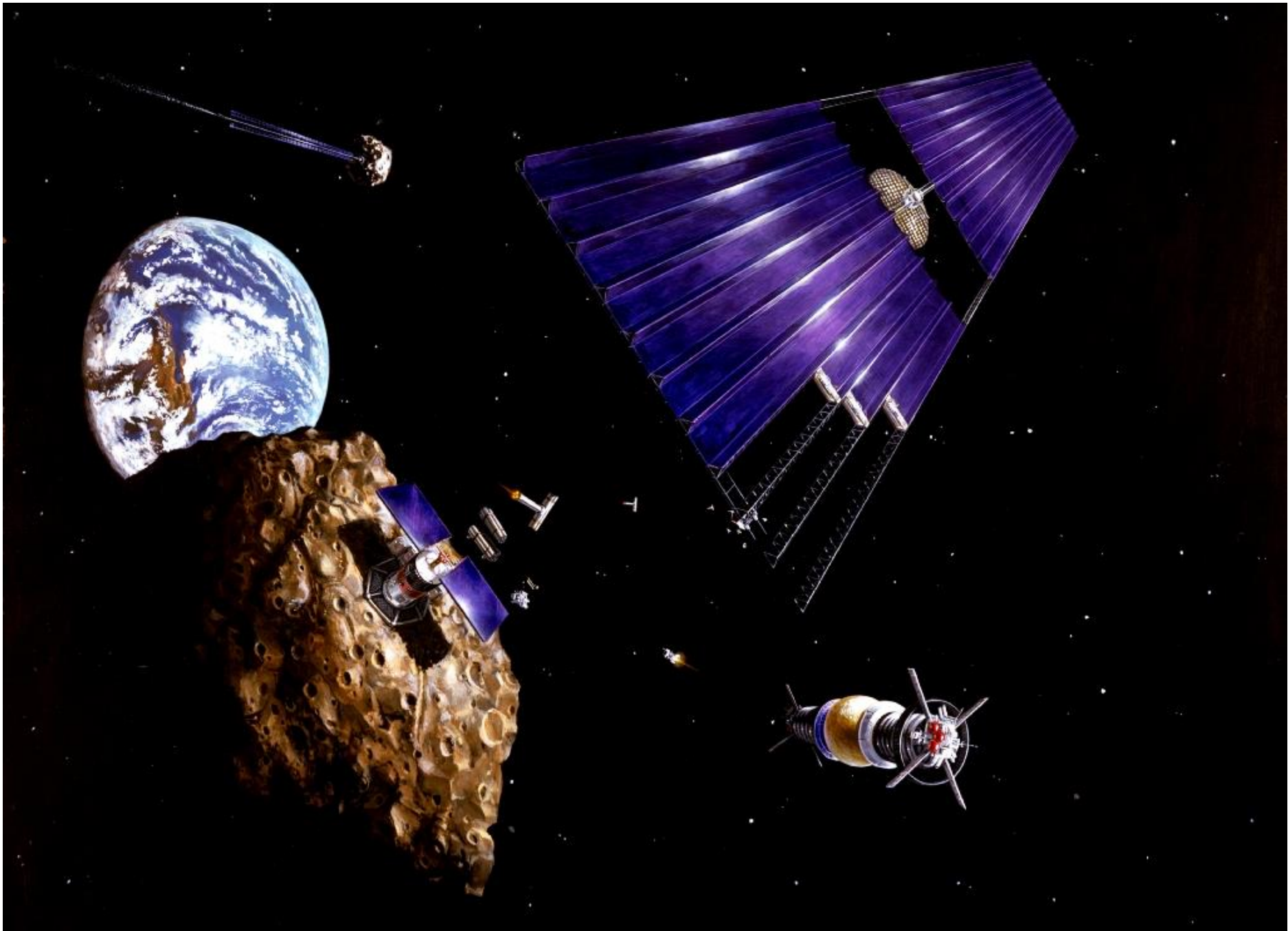


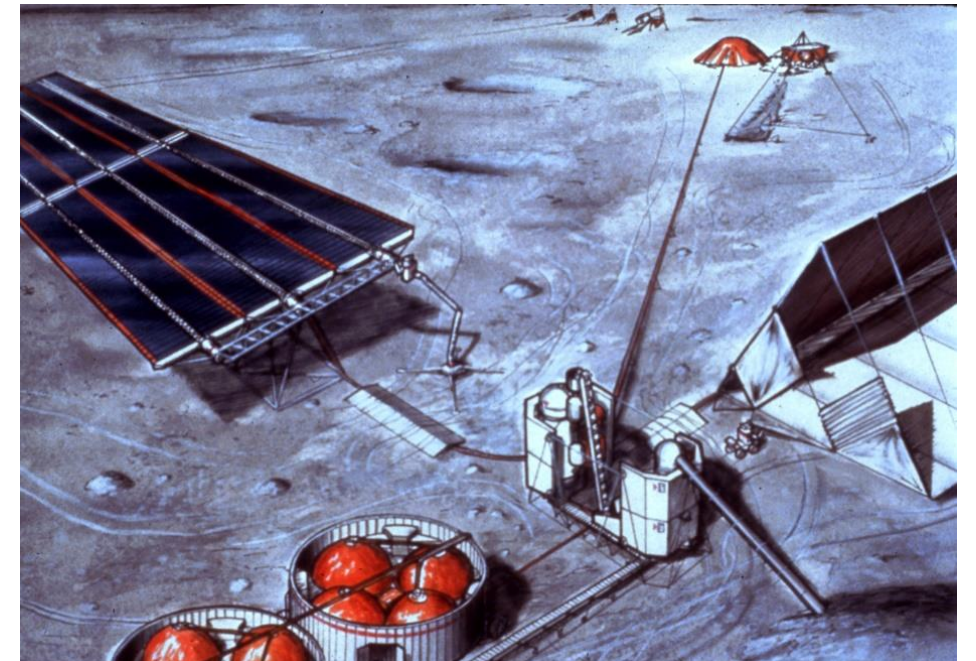
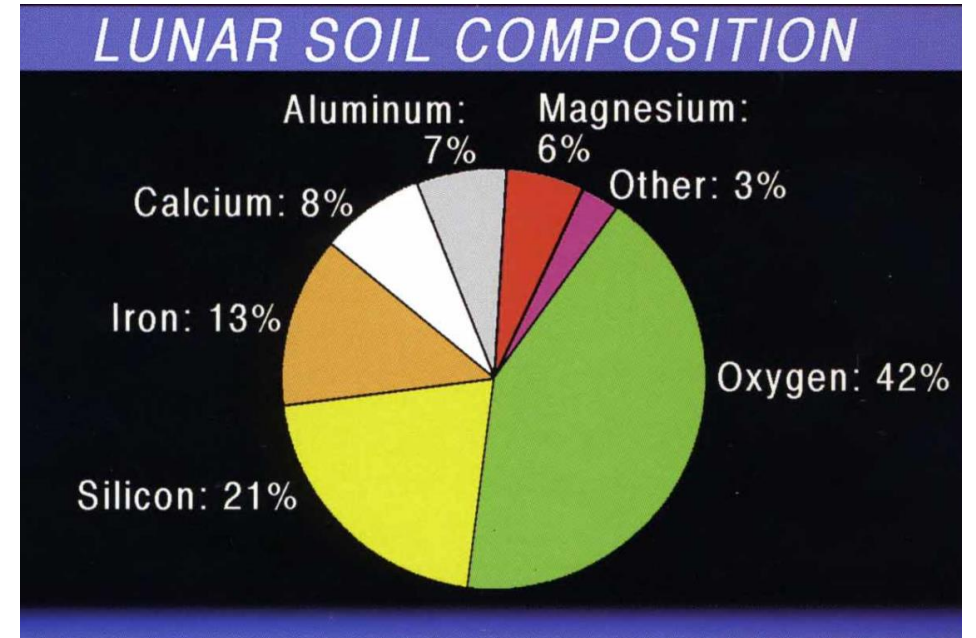
Figure 4: Starship payload volume (dimensions in m)

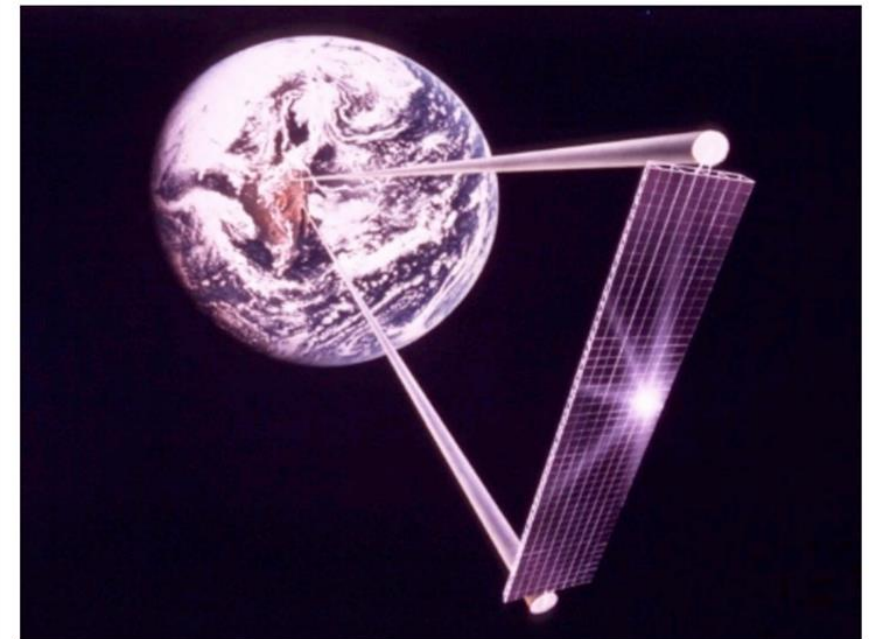
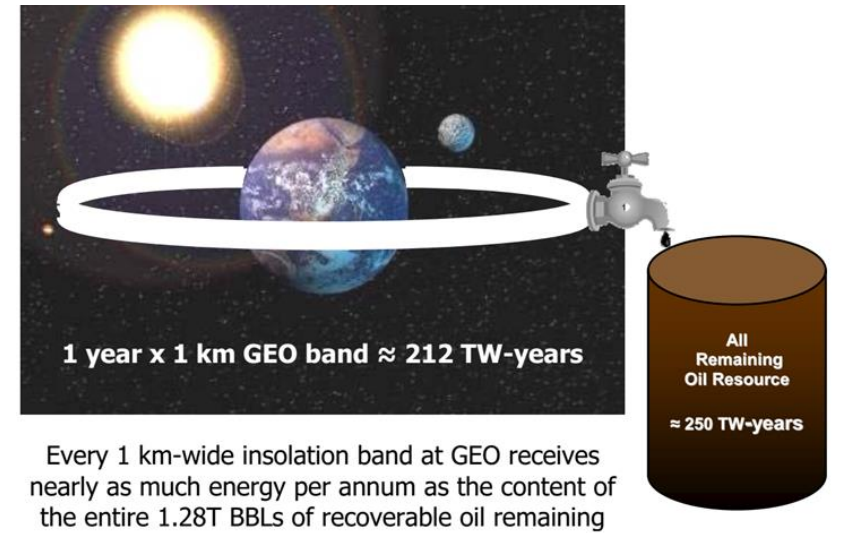
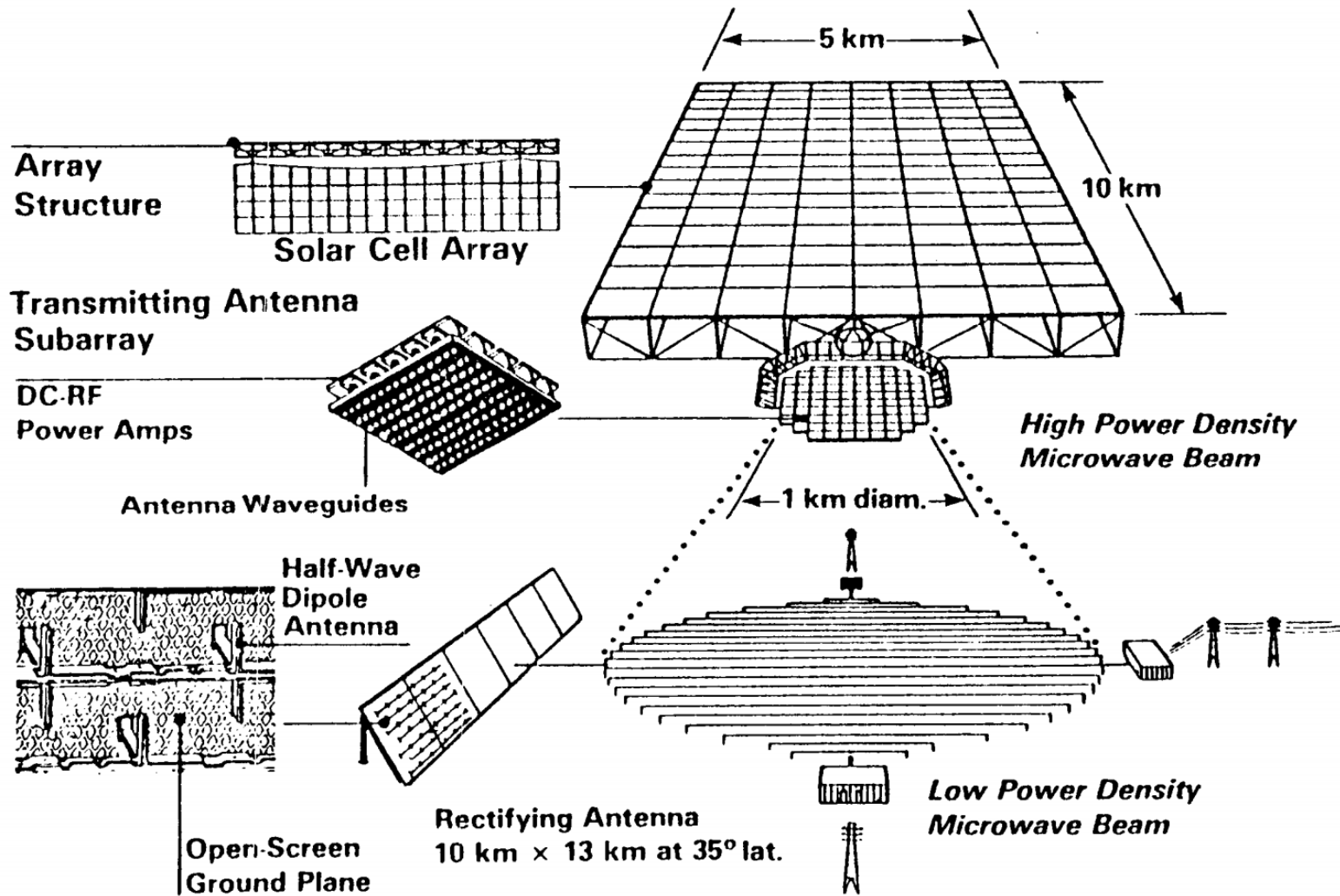


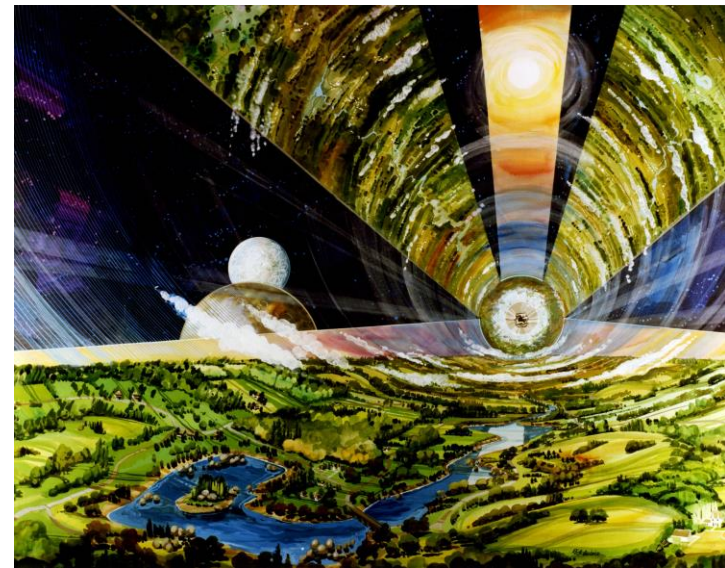
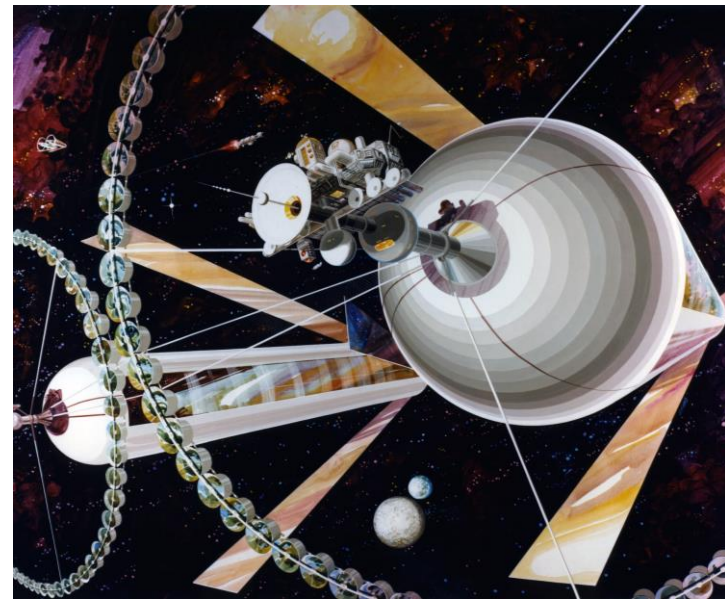
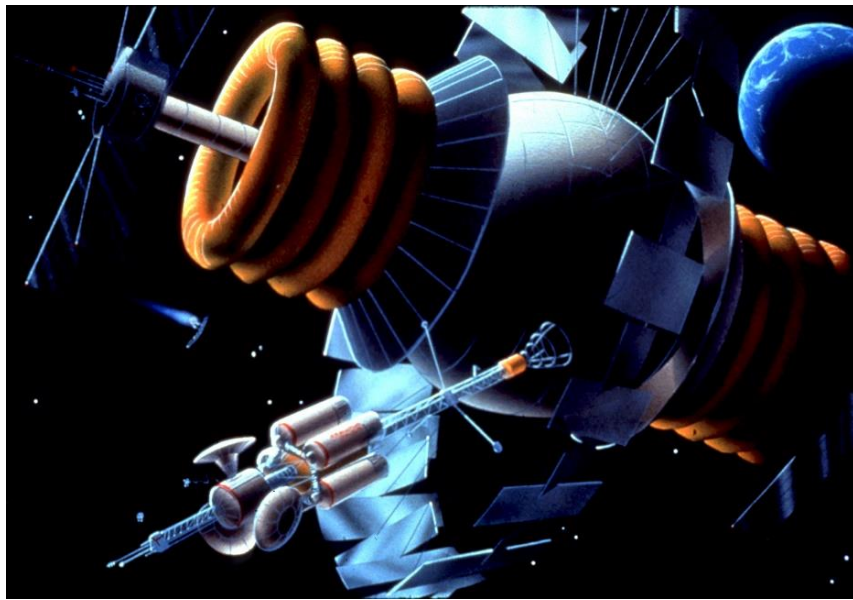
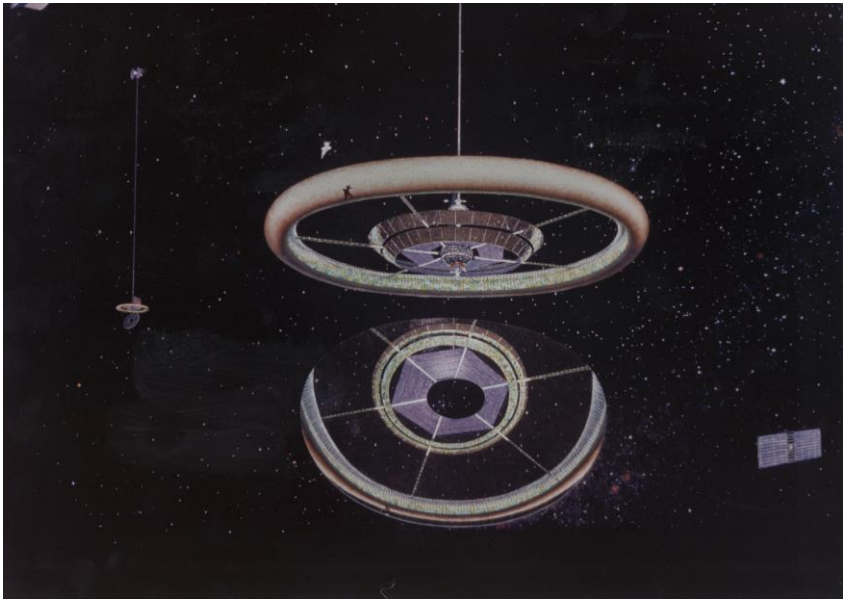




Volatile	Concentration ppm ($\mu\text{g/g}$)	Average mass per m^3 of regolith (g)
H	46 ± 16	76
^3He	0.0042 ± 0.0034	0.007
^4He	14.0 ± 11.3	23
C	124 ± 45	206
N	81 ± 37	135
F	70 ± 47	116
Cl	30 ± 20	50









WE NEED YOU

$\tau = 0.94$

2.9

4.1

3.8

4.7

1205
11:14

1220
11:04

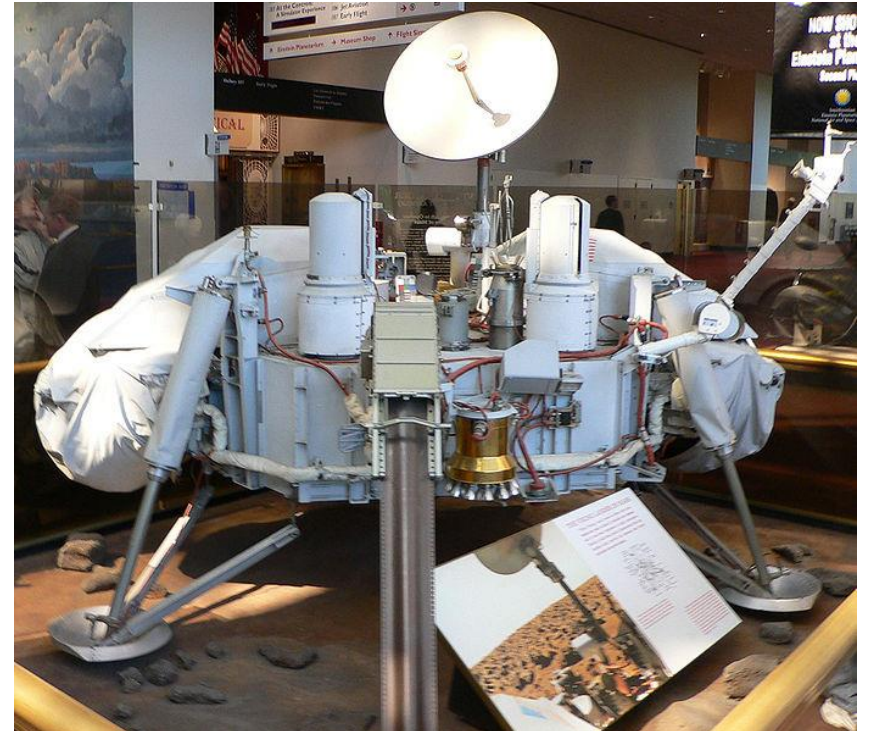
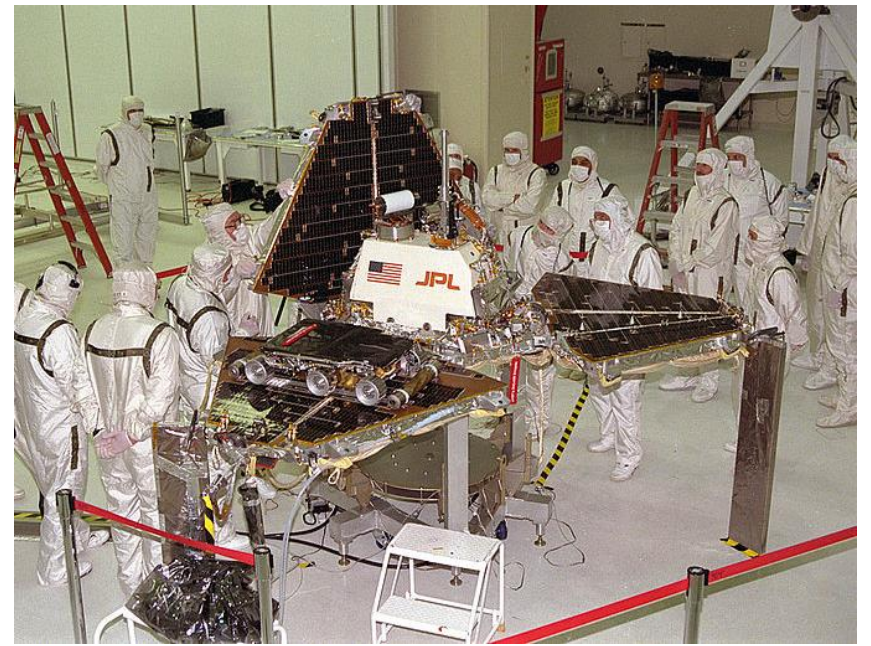
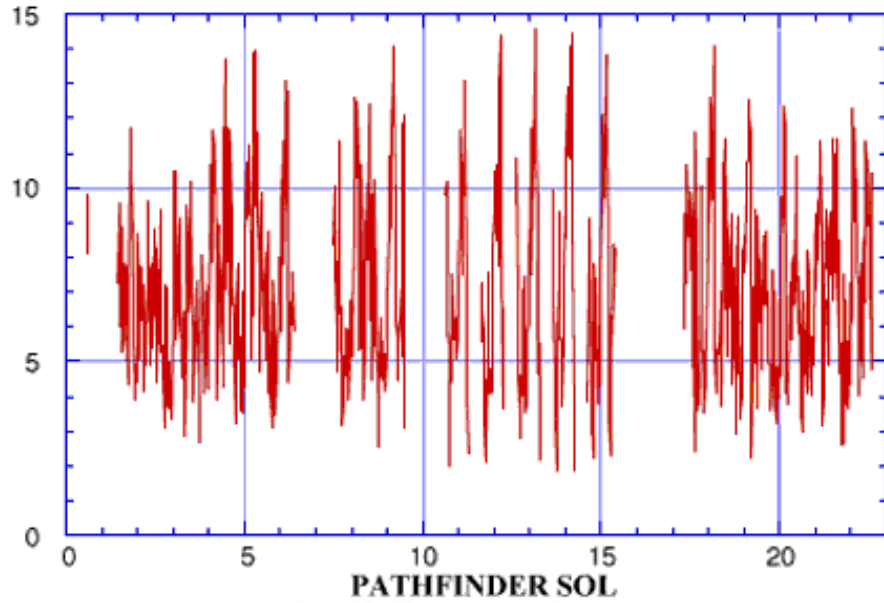
1225
11:30

1233
10:55

1235
10:53

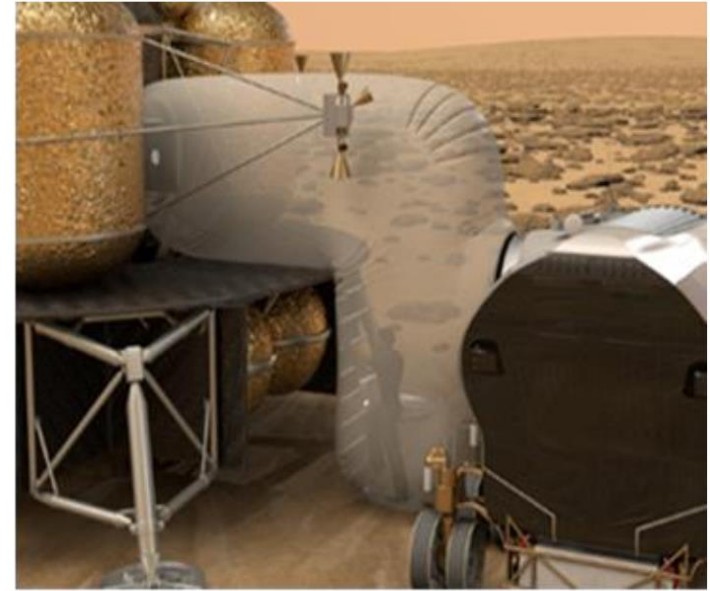
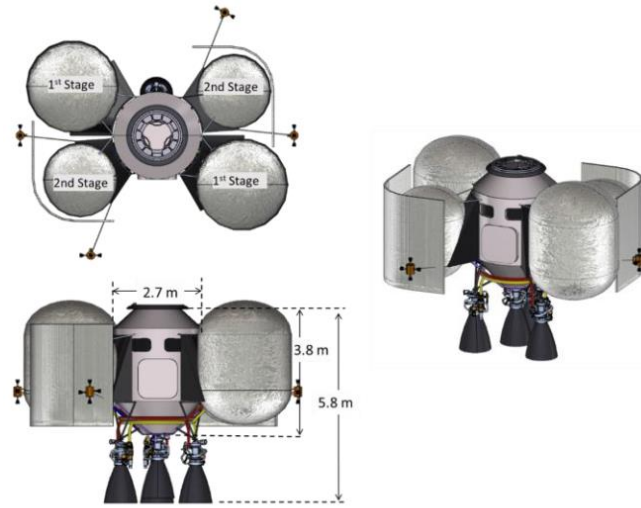
Opportunity Sol Number and Local True Solar Time

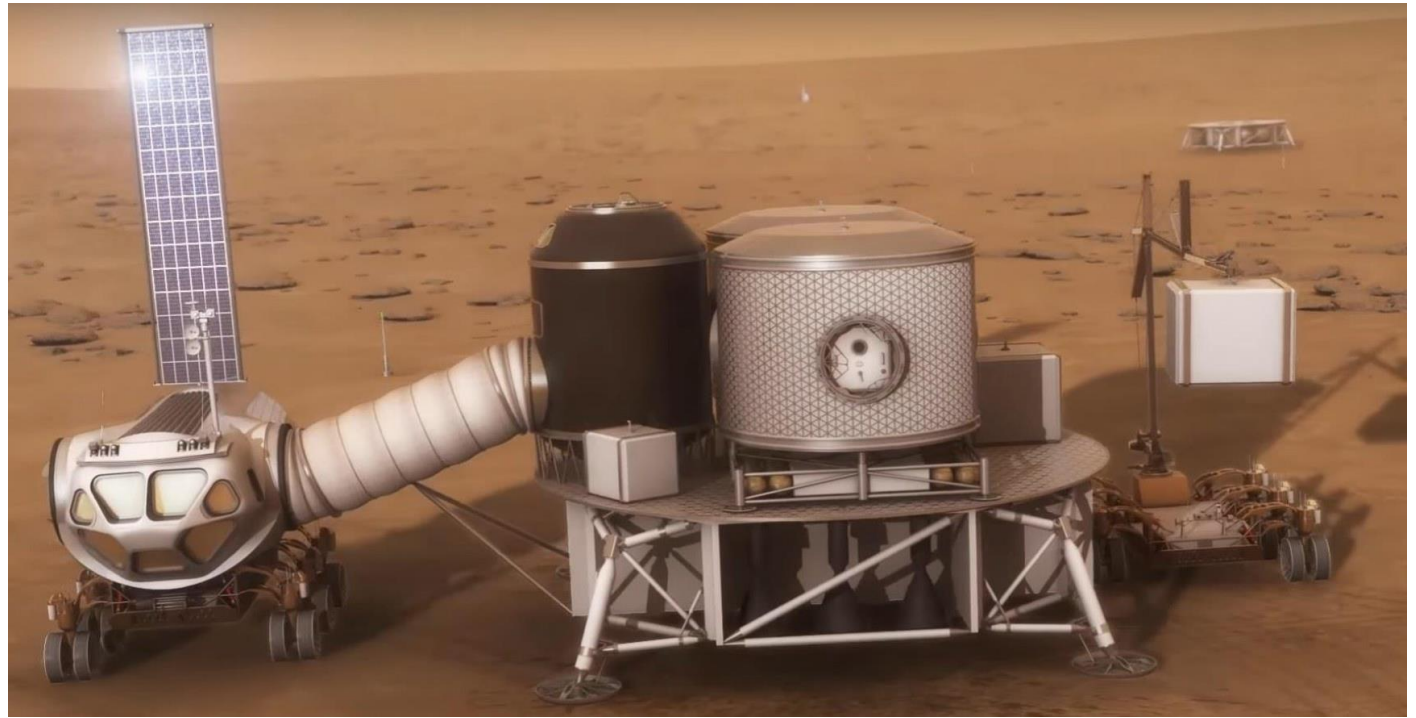
MARS PATHFINDER QUALITATIVE WIND SPEED

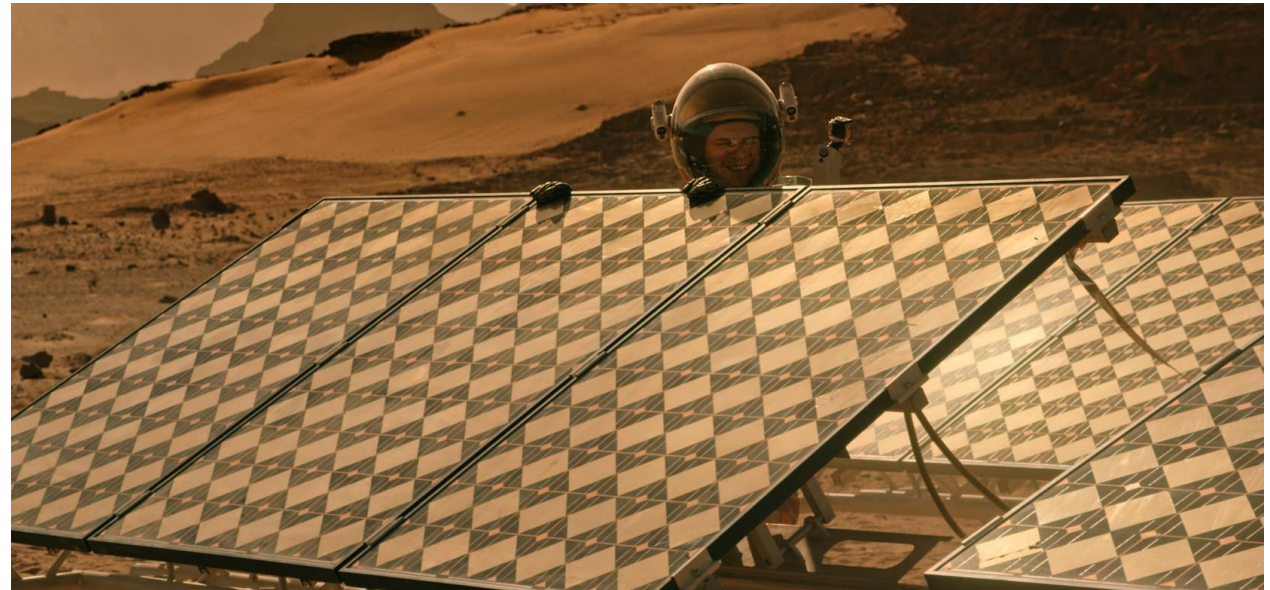


Wind speed on Mars		R. Pappa	"Feels like" wind speed on Earth (at STP)	
mph	m/s	dynamic pressure (Pa)	m/s	mph
10	4.5	0.2367	0.6	1.4
50	22.4	5.9169	3.0	6.8
67	30.0*+	10.6587	4.1	9.1
100	44.7	23.6677	6.1	13.5
150	67.1	53.2523	9.1	20.3
200	89.4	94.6708	12.1	27.1
224	100.0	118.4304	13.5	30.3
447	200.0	473.7216	27.1	60.6
500	223.5	591.6924	30.3	67.7

* Highest measured wind speed from Viking Lander (1.6 m sensor elevation)







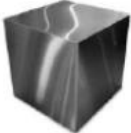








Primary Resources of Interest for Human Exploration

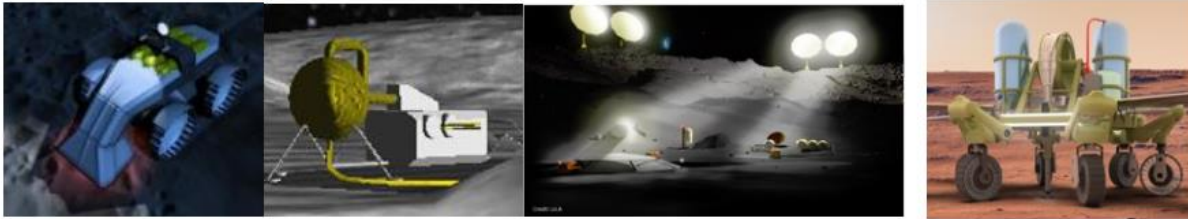
	 Moon	 Mars	 Asteroids	Uses
Water (Hydrogen) 	Icy Regolith in Permanently Shadowed Regions (PSR) Solar wind hydrogen with Oxygen	Hydrated Soils/Minerals: Gypsum, Jarosite, Phyllosilicates, Polyhydrated Sulfates Subsurface Icy Soils in Mid-latitudes to Poles	Subsurface Regolith on C-type Carbonaceous Chondrites	<ul style="list-style-type: none"> Drinking, radiation shielding, plant growth, cleaning & washing Making Oxygen and Hydrogen Breathing Oxidizer for Propulsion and Power Fuel Production for Propulsion and Power Plastic and Petrochemical Production <i>In situ</i> fabrication of parts Electical power generation and transmission
Oxygen 	Minerals in Lunar Regolith: Ilmenite, Pyroxene, Olivine, Anorthite	Carbon Dioxide in the atmosphere (~96%)	Minerals in Regolith on S-type Ordinary and Enstatite Chondrites	
Carbon (Gases) 	<ul style="list-style-type: none"> CO, CO₂, and HC's in PSR Solar Wind from Sun (~50 ppm) 	Carbon Dioxide in the atmosphere (~96%)	Hydrocarbons and Tars (PAHs) in Regolith on C-type Carbonaceous Chondrites	
Metals 	Minerals in Lunar Regolith <ul style="list-style-type: none"> Iron/Ti: Ilmenite Silicon: Pyroxene, Olivine, Anorthite Magnesium: Mg-rich Silicates Al: Anorthitic Plagioclase 	Minerals in Mars Soils/Rocks <ul style="list-style-type: none"> Iron: Ilmenite, Hematite, Magnetite, Jarosite, Smectite Silicon: Silica, Phyllosilicates Aluminum: Laterites, Aluminosilicates, Plagioclase Magnesium: Mg-sulfates, Carbonates, & Smectites, Mg-rich Olivine 	Minerals in Regolith/Rocks on S-type Stony Iron and M-type Metal Asteroids	

Note: Rare Earth Elements (REE) and Platinum Group Metals (PGM) are not driving Resources of interest for Human Exploration

Resource Prospecting – Looking for Water Hydrated minerals & subsurface ice on Mars



Mining Polar Water & Volatiles Mining near surface ice on Mars



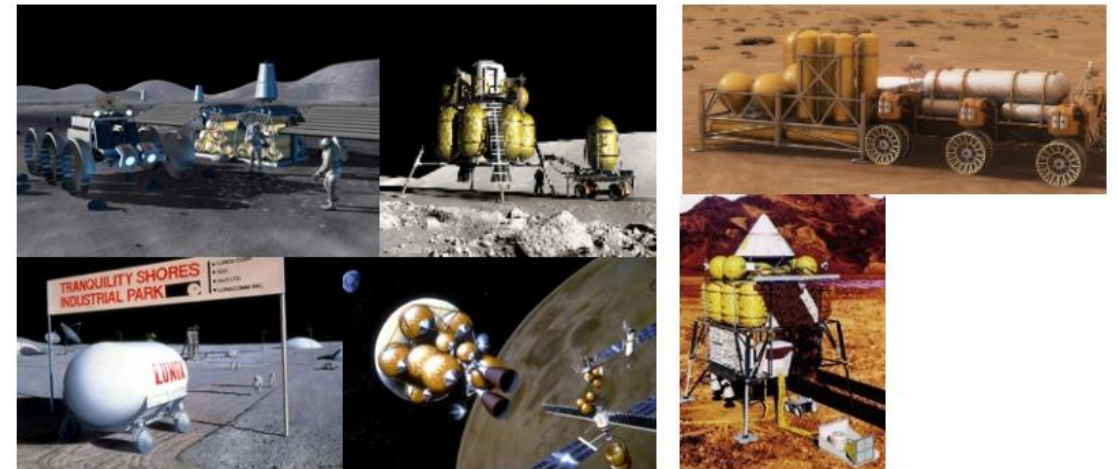
Landing Pads, Berms, Roads, and Structure Construction

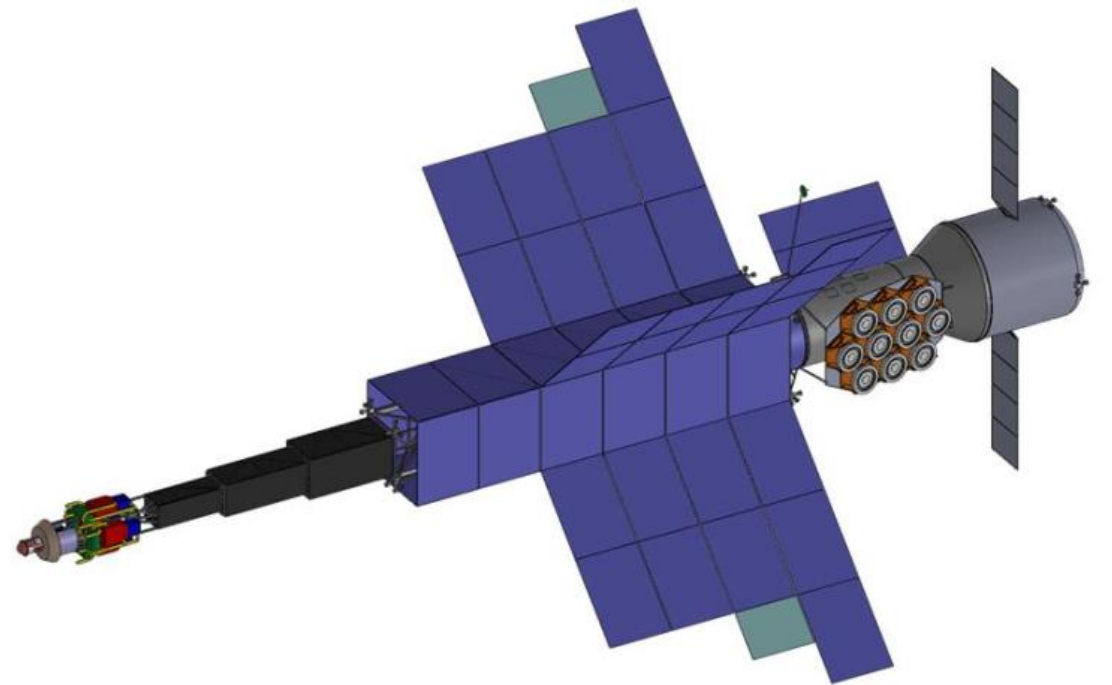
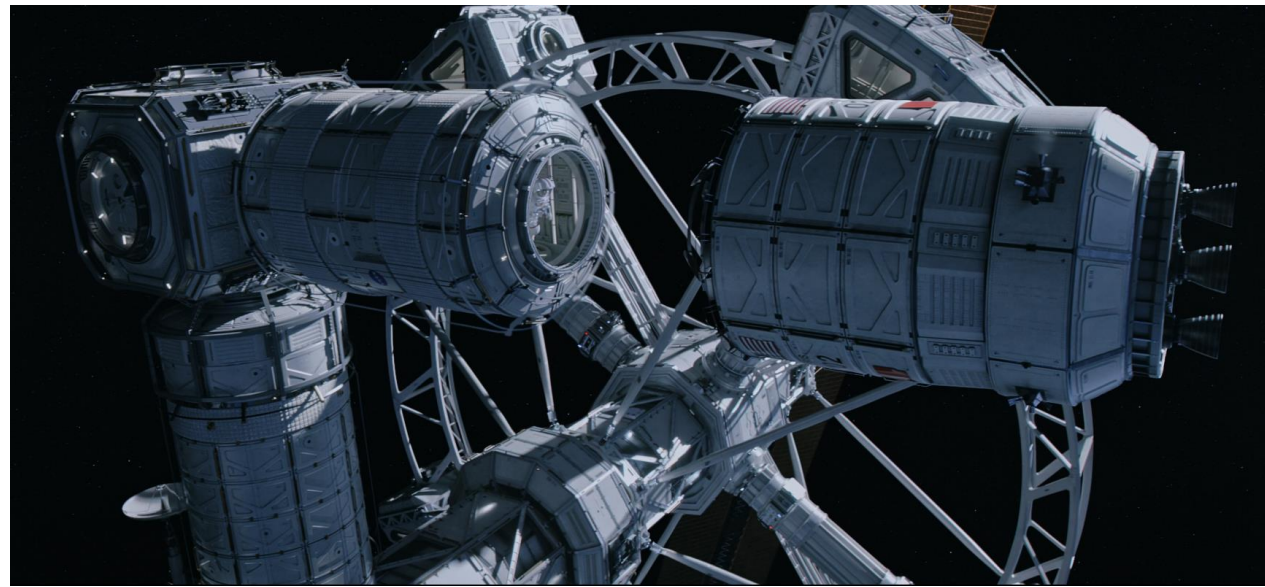


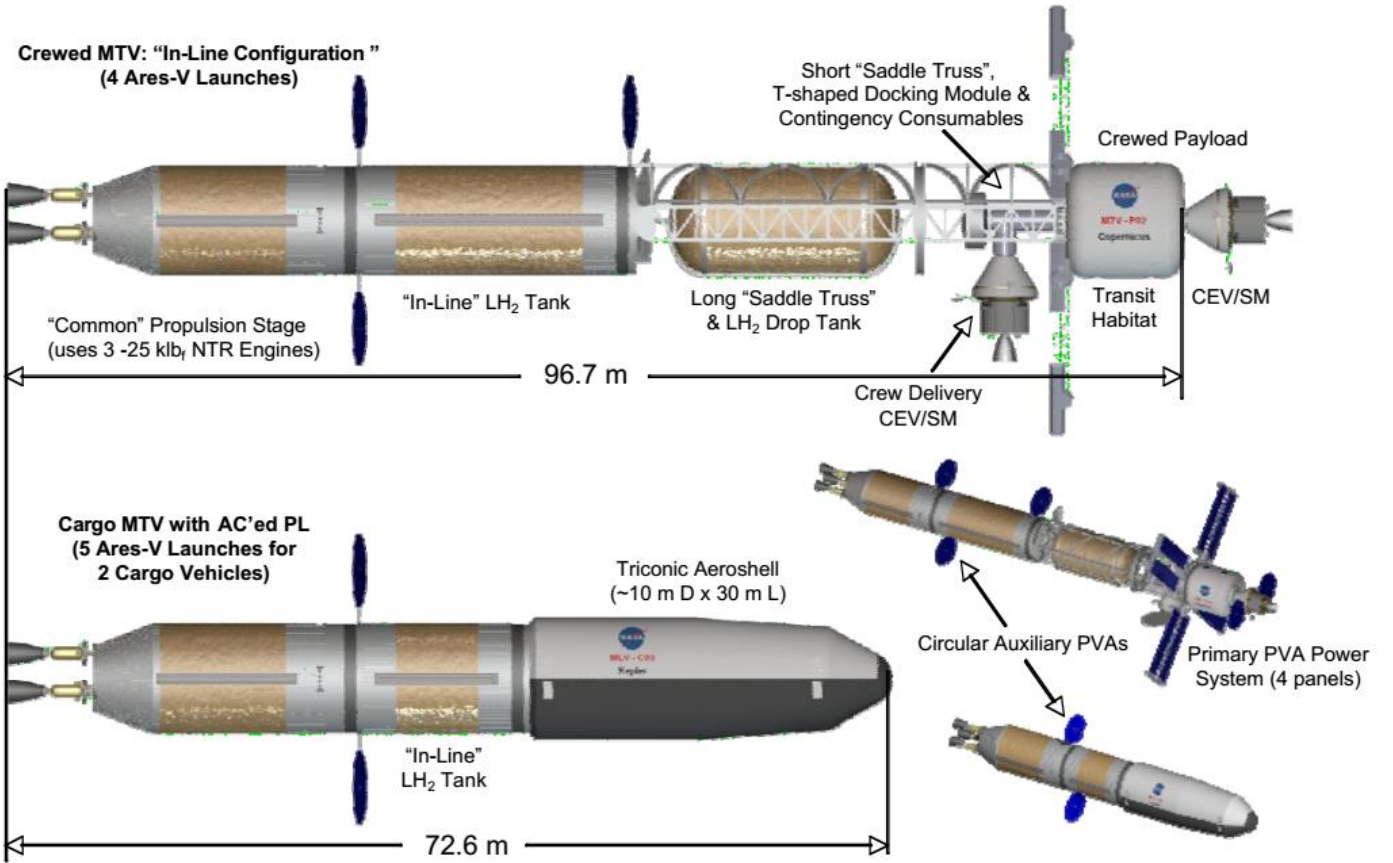
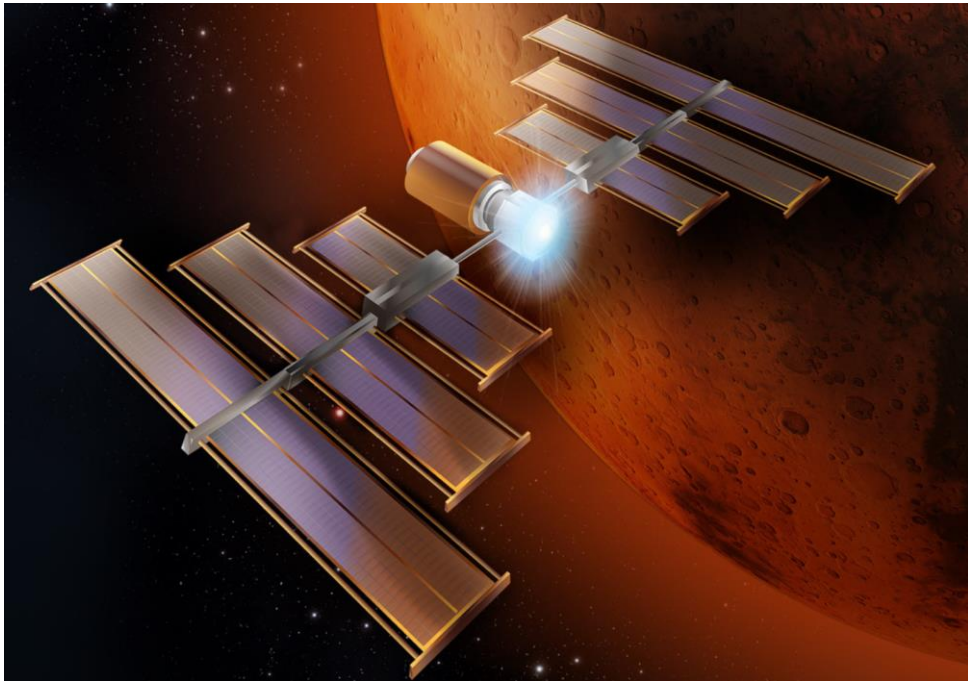
Excavation & Regolith Processing for O₂ Production Excavation & Processing for H₂O Extraction



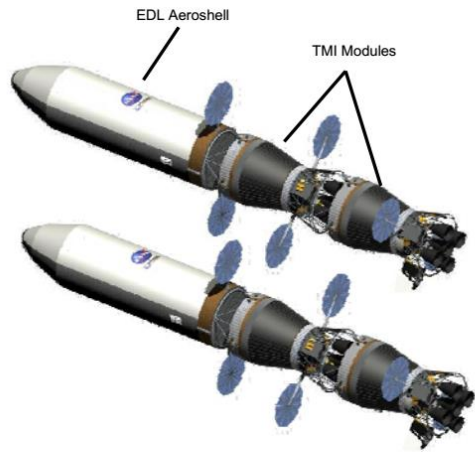
Refueling and Reusing Landers & Rovers







2 Cargo Vehicles:
7 launches
170-day assembly time in LEO



Crew Vehicle:
5 launches
120-day assembly time in LEO

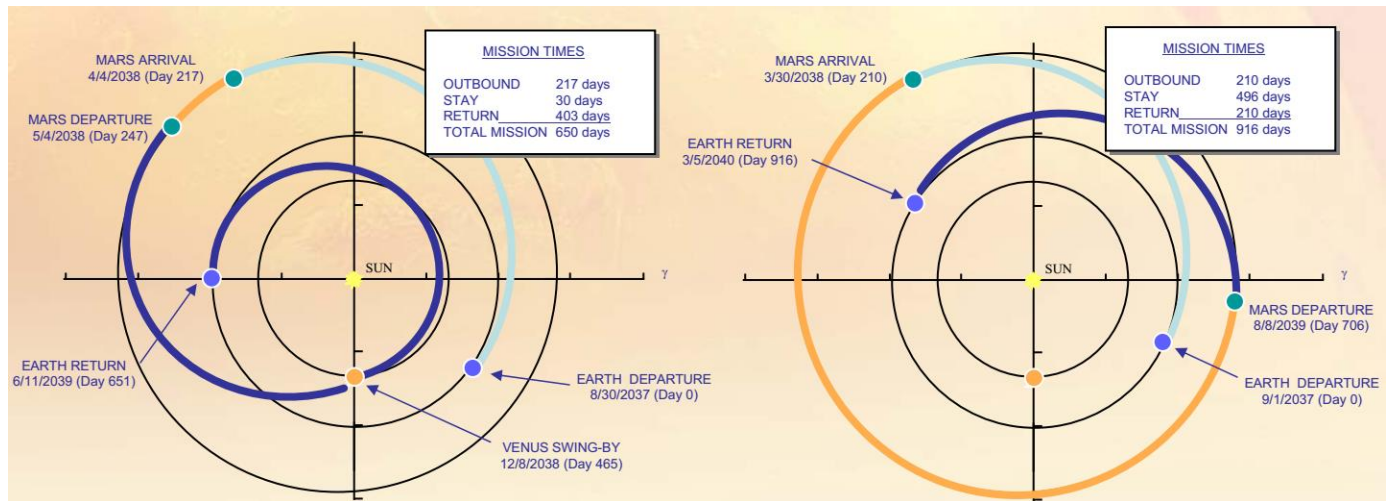
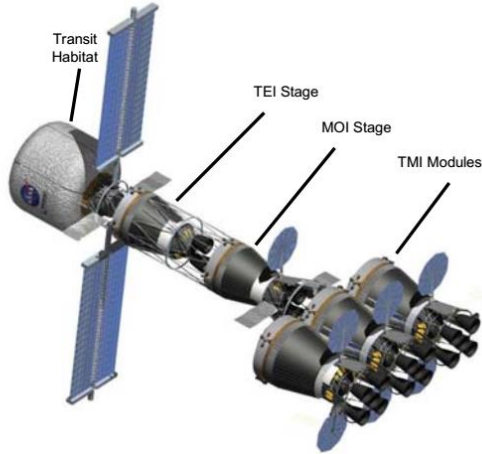


Figure 4-6. Chemical/aerobrake cargo and crewed MTV concepts.