Theoretical Physics I Stanis les Hubin IPSP 3720433 11 Problem 2.1 Easeth orbit around the Sun. a) $c \approx 3.10^8 \text{m/s}^{-1}$, $t = 8 \text{mim} + 198 \approx 500 \text{s} = 600 \text{use}$ it is easier to coloulste $= 5.10^2 \text{s}$ with powers of 10 and simple numbers) 14 = 1.5.10 11 m. $G = 6.7 \cdot 10^{-11} \, \text{m}^3 \, \text{kg}^{-1} \, \text{s}^{-2} \qquad [G] = \, \text{m}^3 \, \text{kg}^{-1} \, \text{s}^{-2}$ $M = 2 \cdot 10^{30} \, \text{kg} \qquad [M] = \text{kg}$ $D = 1.5 \cdot 10^{11} \, \text{m} \qquad [D] = \text{m}$ $[T] = s = m^{0}kg^{0}s^{1} = [G^{Ro}], [M^{Rm}], [\mathcal{D}^{Po}] =$ = (m3kg-15-2)P6, kg PM, mPD $\begin{cases}
0 = 3p_6 + p_0 \\
0 = -p_6 + p_m \leftrightarrow \int p_m = -\frac{3}{2} \\
1 = -2p_6
\end{cases}$ $\begin{cases}
p_6 = -\frac{1}{2} \\
p_6 = -\frac{1}{2}
\end{cases}$ Therefore $T \approx \sqrt{\frac{B^3}{6M}} = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{-11} \cdot 2 \cdot 10^{30}}} \leq = \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{33}}} \leq \sqrt{\frac{1.5^3 \cdot 10^{33}}{67 \cdot 10^{33}}} = \sqrt{\frac{1.5^3 \cdot 10^{3$ $= \sqrt{\frac{15 \cdot 10^{14} (1.5)^{2}}{6.7 \cdot 2}} S = \frac{1}{2} \cdot 10^{7} S = 5 \cdot 10^{6} S,$ 0) 5.10^6 s $\approx 5.10^6$ s, $\frac{year}{365d}$, $\frac{d}{24h}$, $\frac{h}{3600s} = \frac{5.10^6}{365.24.3600}$ year $\approx \frac{50}{365}$ year $\approx \frac{5}{6}$ year.

It is ≈ 2.5 t smaller than actual period. 17 d) $\mathcal{E}(t) = \frac{\chi(t)}{L} = \left(\frac{\chi(t)}{L}, \frac{\chi_1(t)}{L}, \frac{\chi_2(t)}{L}\right)$ Most likely L is the surage distance of planet to Sun.
But if we will consider dynamics with many planets at once,
L con mean 1A.U. (Earth's distance) or surage " cadius" of solar system. Problem 2.2 Oscillating period of particle attached to a spring.

gmoon = 1.6 m·s⁻², M= 10⁻¹kg, k= 1.6 kg·s⁻².

a) Construct length/time scale for non-dimensionalitation of dynam Using B-It method: $m - \left[g_{M \alpha n}\right] \cdot \left[k^{-1}\right] \cdot \left[M'\right] = \left[g_{K}^{M}\right]$ $kg = \left[M^{\frac{1}{2}}\right] \cdot \left[k^{-\frac{1}{2}}\right] = \left[N_{K}^{M}\right]$ $= \left[N_{K}^{M}\right] \cdot \left[k^{-\frac{1}{2}}\right] = \left[N_{K}^{M}\right]$ So lengths will be sided by $g_{\overline{k}}^{M}$, time by $\sqrt{\underline{k}}$.

Therefore $\xi(t) = \frac{x(t)}{(gM/k)}$, $\xi(t) = \frac{\dot{x}(t)}{(gM/k)} \cdot \sqrt{f_{M}} = \frac{\dot{x}(t)}{g\sqrt{\frac{M}{k}}}$. 8) w = = = = = = 45-1. c) $\mathcal{E} = \mathcal{E}^2(t/T) + \mathcal{E}^2(t/T)$ is const. That means phase space visualization is a circle in $\mathcal{E}(t/T)O\mathcal{E}(t)$ Radius is $\sqrt{\mathcal{E}}$, plane. It is troversed clockwise (relative is positive and goes to o when particle peoches positive coordinate amplitude, etc.).

VE 15(+17) d) (for seminors) Problem 2,3 Conversion Joule-Colorie, 21 a) · P = 2000 W, T x 1 min 405 = 1025, So Q = 2.103.102 J = 2.105 J, on the one hand. • On the other, $\Delta T \approx 80 \, \text{K}$, $V = 10^{-3} \text{m}^3$, $m = 10^{-3} \frac{\text{kg}}{3}$. $10^{-3} \text{m}^3 = 1 \text{kg}$. So $\Omega = c \text{mat} = 1 \text{kg} \cdot 80 \text{K}$, $\frac{1 \text{cal}}{\text{kg} \cdot \text{K}} = 10^{3} \, \text{g} \cdot 80 \, \text{K}$. $\frac{1 \text{cal}}{g \cdot \text{K}} = 8 \cdot 10^{4} \, \text{cal}$. • So $8 \cdot 10^{4} \, \text{cal} = 20 \cdot 10^{4} \, \text{J} \longrightarrow 1 \, \text{cal} = 2.5 \, \text{J}$.

6) From literature col = 47. Possibilities for discrepancy:

- actual Part > P (it was weither 2000-2400W," I took 2000 for singularity) - will gree higher Q1 - school stact < st - well give less & 2 (I took 20°C as to, while in my house it is more) - schuel p of water can be smeller - will give less d2 - (volume is not the resson, it is marked enough accurately for such error) Combination of factors can work even it some go in other direction. Problem 14 Water waves 0=f(1,g) a) [1]=m [g]=m.s-2 [10] = m·s-1, so [10] = [12], [g2], 10=1/19 b) 10 ~ \soft-10 m-s^2 = \square 300.3.10 ms^2 ~ 10 ms-1 c) 1 ft = 304,8·10³ m ₹ 20.3043 feasibility depends only on good. If we have very pacturate ft value and need to maintain accuracy — it is not enough. If we do not need, it is all right.

Also, if we do simple calculations, it is easier to use 0.3 m, then 1/3 m,

Index der Kommentare

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1.1
         2.1: 18/19
         2.1 was done very clear and concise :)
1.2
1.3
         please go more into an argument concerning the margin of error when reasoning for approximations
1.4
1.5
         7/7
1.6
         3/3
1.7
         keen eye :)
1.8
         great explanation:)
         2/2
1.9
2.1
         2.3: 3/4
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For one minute you would need a more powerful kettle