

# SESA1015

# Astronautics

# Environment

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Natural and Induced Environments

Our Solar System

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Environmental Effects on Design/Man

Launch Environment

Summary

# Natural and Induced Environments

Possible Environments Include:

Fabrication, Assembly, Test,  
Terrestrial transportation, Launch.  
Space Environment

} *Induced*  
} *Natural*

Space Environment:

... hostile or friendly?



# Our Solar System

## What is the Space Environment?



# Our Solar System

## What is the Space Environment?



Temperature: 2.7° K

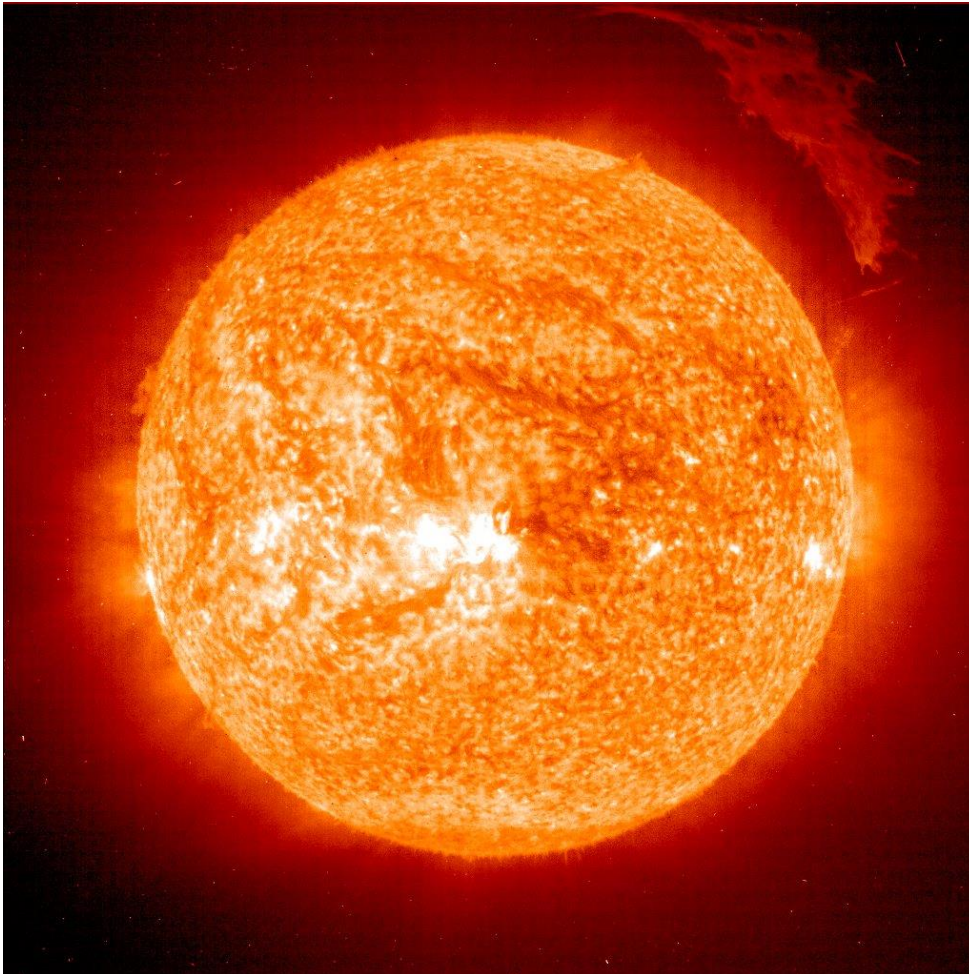
Vacuum

Microgravity

In our solar system, the main input to the space environment is... the sun.

# Our Solar System

## The Sun



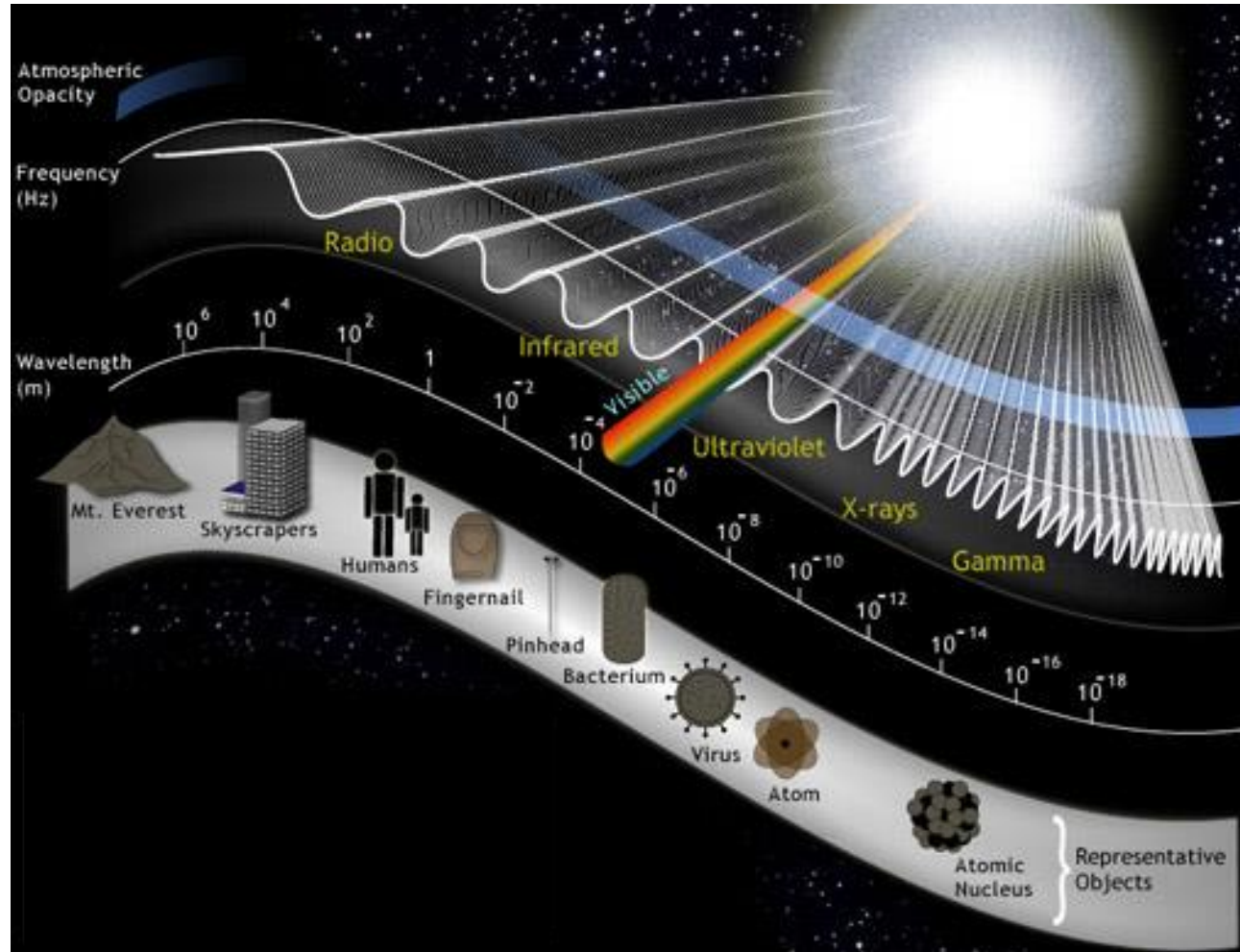
Diameter: 1,400,000 km  
Source of energy:  
Nuclear fusion

- Electromagnetic Radiation
- Particle Radiation



# Our Solar System

## The Sun – Electromagnetic Radiation – The Electromagnetic Spectrum

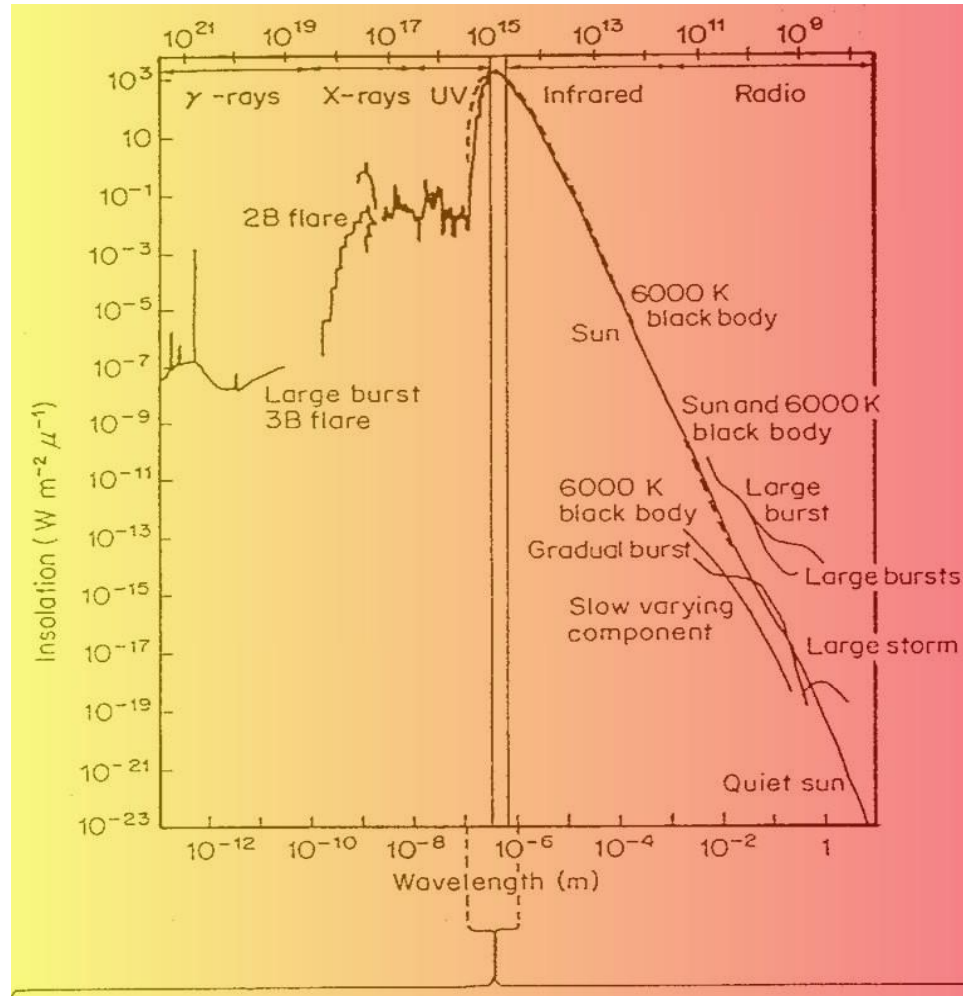


Average Energy  
output at 1 AU :  
 $1370 \text{ W/m}^2$

Solar Cycle – 11 year  
cycle

# Our Solar System

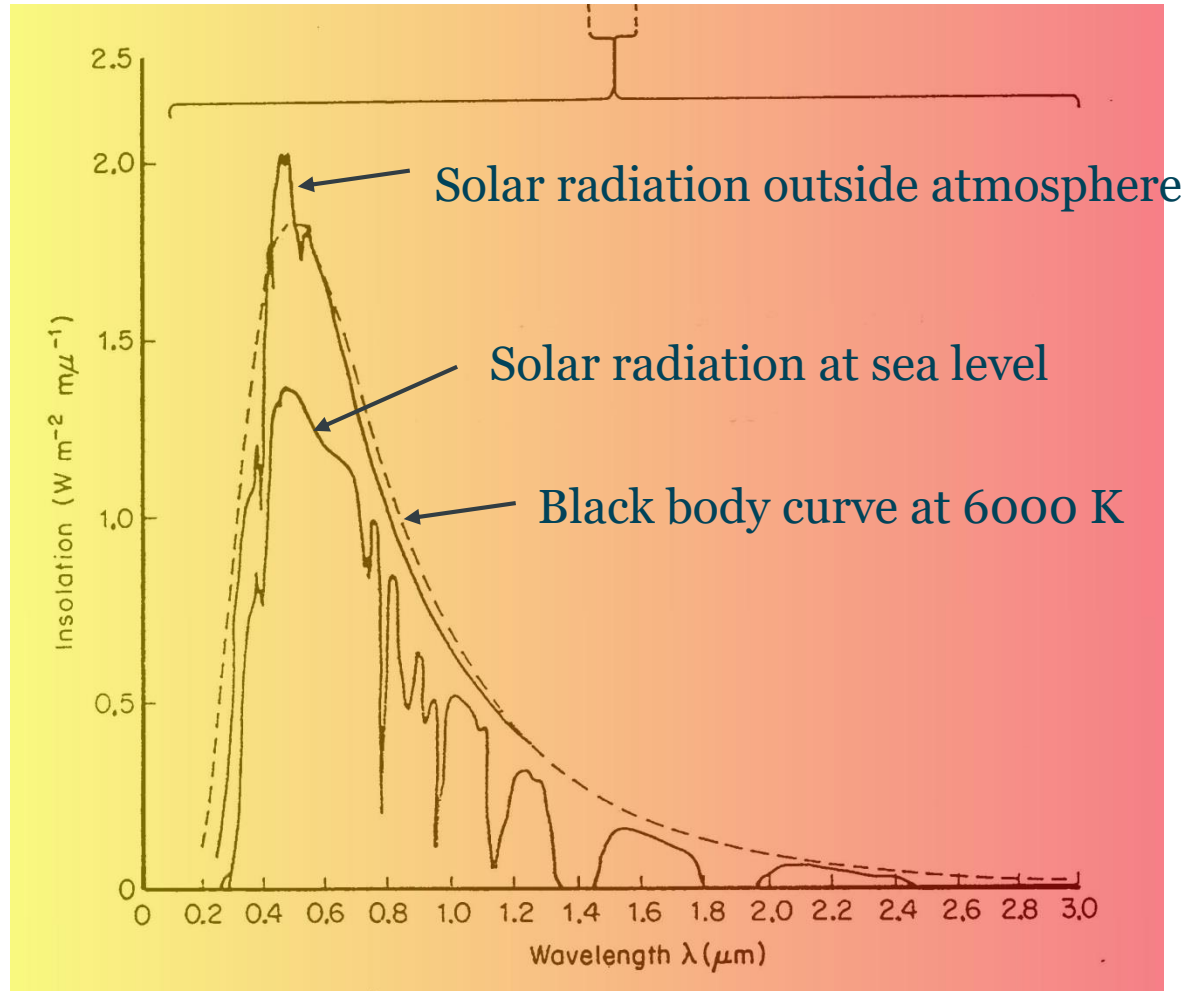
## The Sun – Electromagnetic Radiation





# Our Solar System

## The Sun – Electromagnetic Radiation – UV/visable wavelengths



# Our Solar System

## The Sun – Electromagnetic Radiation

- Solar spectrum approximates to a black body at  $\sim 6000\text{K}$  in UV, visible and IR.
- Peak output in visible at  $\sim 0.5\ \mu\text{m}$  wavelength
- Solar Variability:

Radio  $\times 100$ ;

IR small;

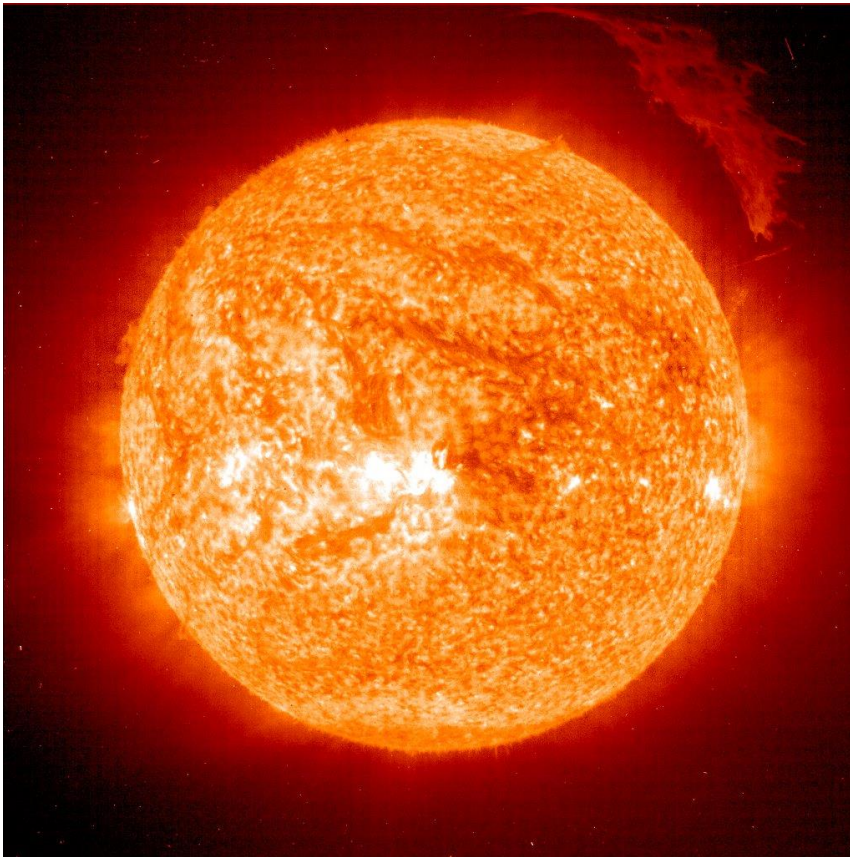
Visible  $< 1\%$ ;

UV  $\times$  few;

X-ray  $\times 100$ .

# Our Solar System

## The Sun – Particle Radiation



‘stream of energetic (high speed) subatomic particles called the *solar wind*’

Source: violent eruptions that take place on the Sun’s surface and in its atmosphere

Material is thrown into space as a result, consisting of *protons*, *electrons* and *ions* (the nuclei of atoms stripped of their electrons)

At Earth (depending on solar activity)

$$n \sim 5 \times 10^6 \text{ protons/m}^3,$$

$$v \sim 200 - 800 \text{ km/s}$$

Effects on spacecraft materials + systems

# Our Solar System

## The Sun – Solar Proton Events (SPE), ‘*Solar Storm*’

At times of a solar maximum the frequency and violence of the outburst's on the sun's surface ('solar flares') increase. These events can discharge a cloud of energetic charged particles towards the Earth.

Increase in high energy protons ( $\sim 1$  MeV to 1 GeV) for a few hours to a few days.

Frequency – a few / year, peaking at solar max.

## Other Particle Radiation – Galactic Cosmic Rays (GCR)

Extremely high energy particles emanating from outside the solar system ( $10^9 - 10^{12}$  eV).

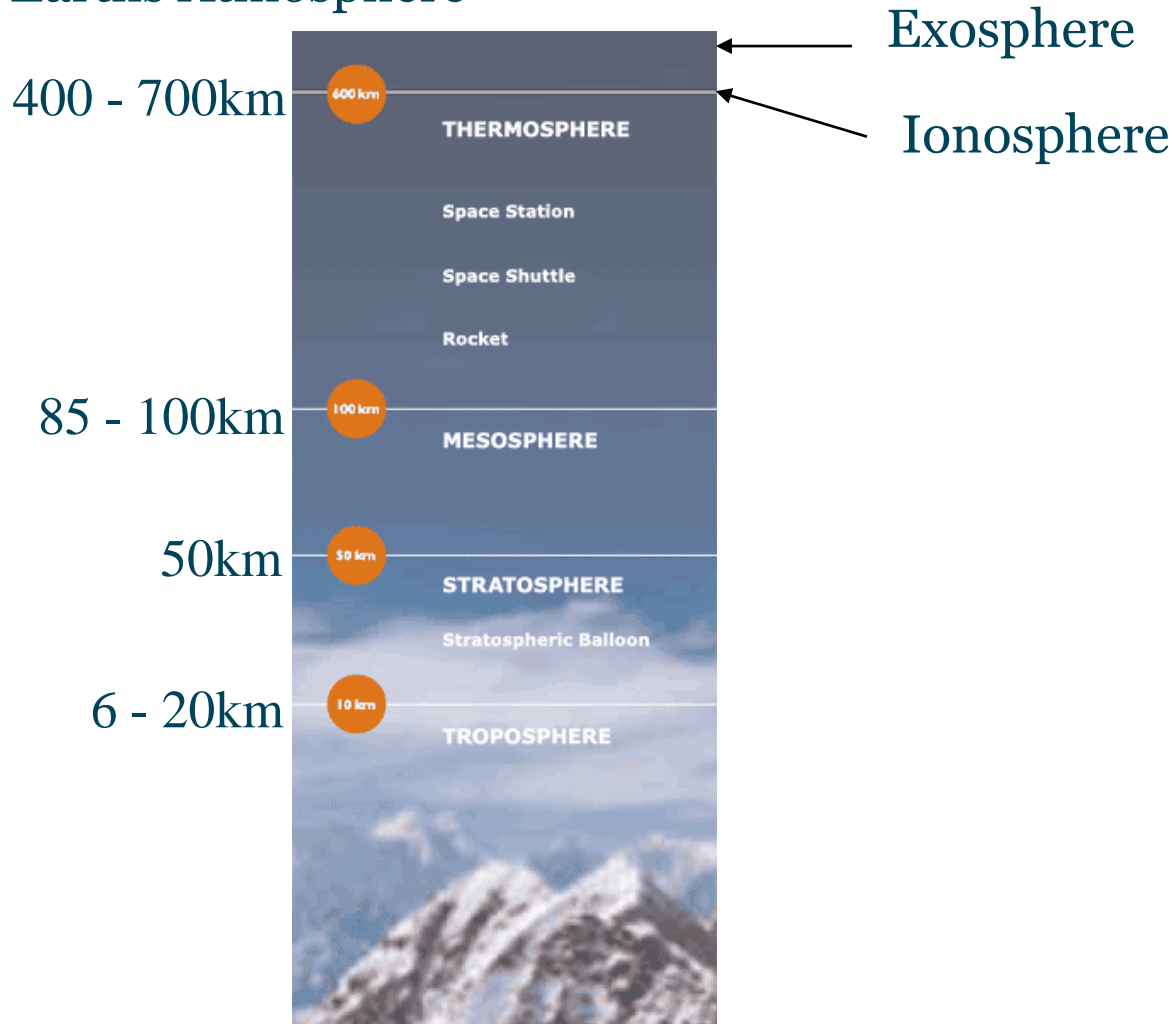
# Earth and the Earth Orbit Environment

The Earth has some degree of natural protection from the Sun's output...

- The atmosphere
- The Earth's Magnetic field

# Earth and the Earth Orbit Environment

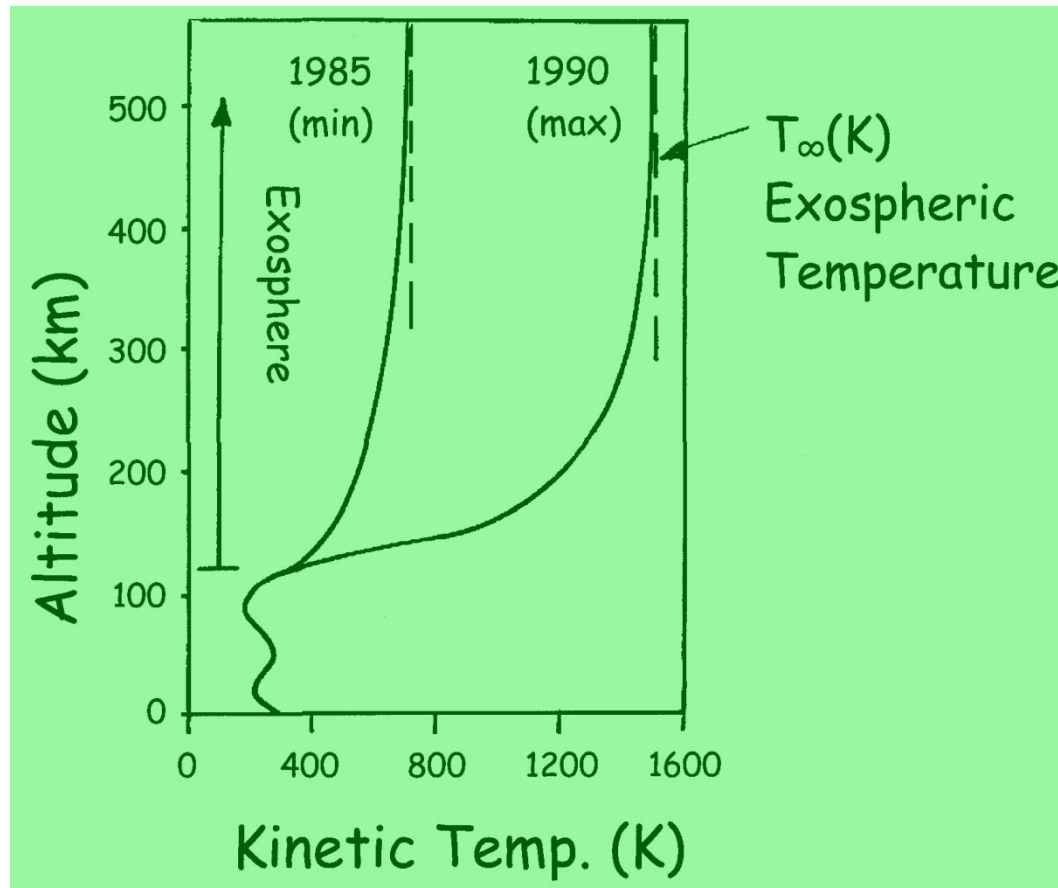
## Earth's Atmosphere





# Earth and the Earth Orbit Environment

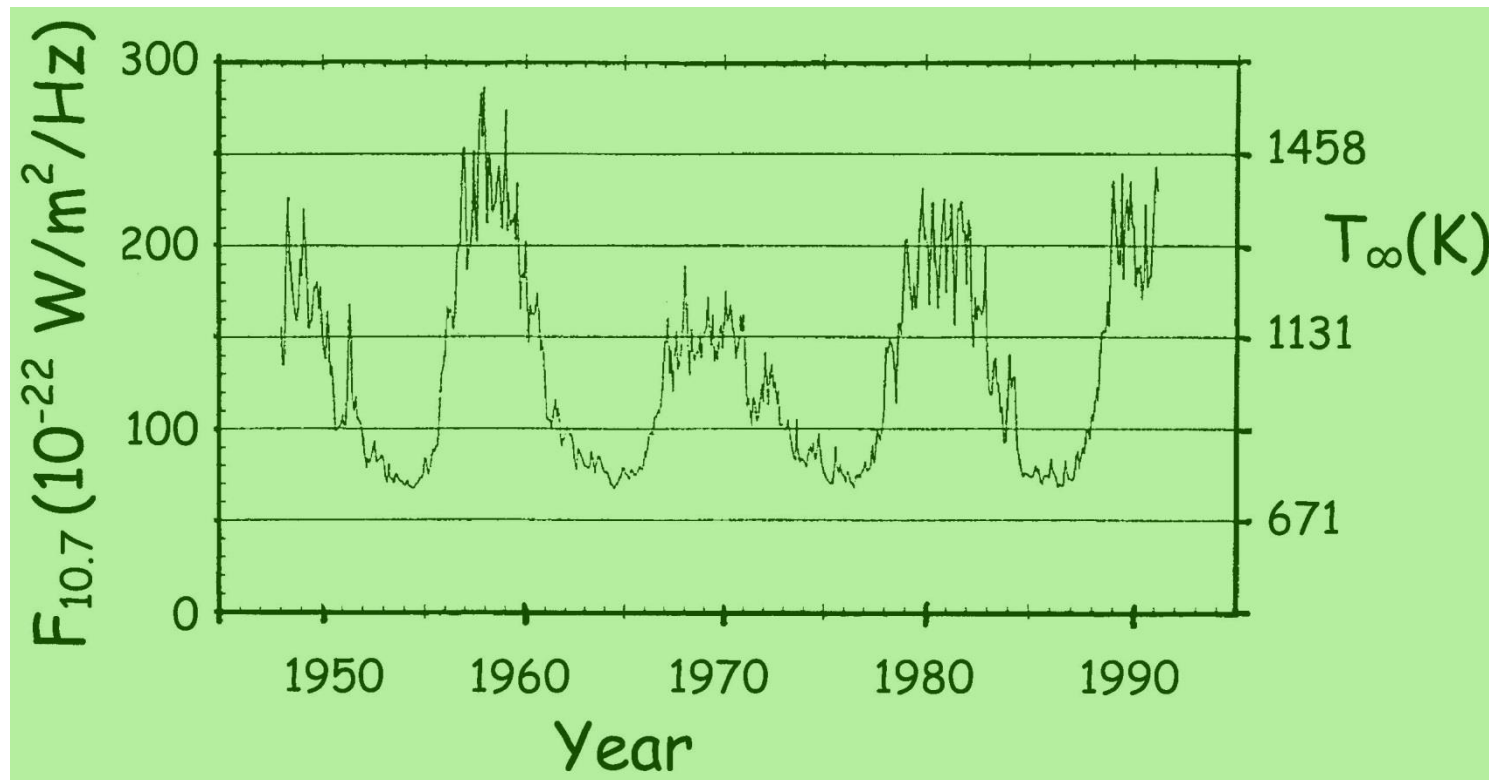
Earth's Atmosphere – the atmosphere and the electromagnetic radiation



# Earth and the Earth Orbit Environment

Earth's Atmosphere – the atmosphere and the electromagnetic radiation

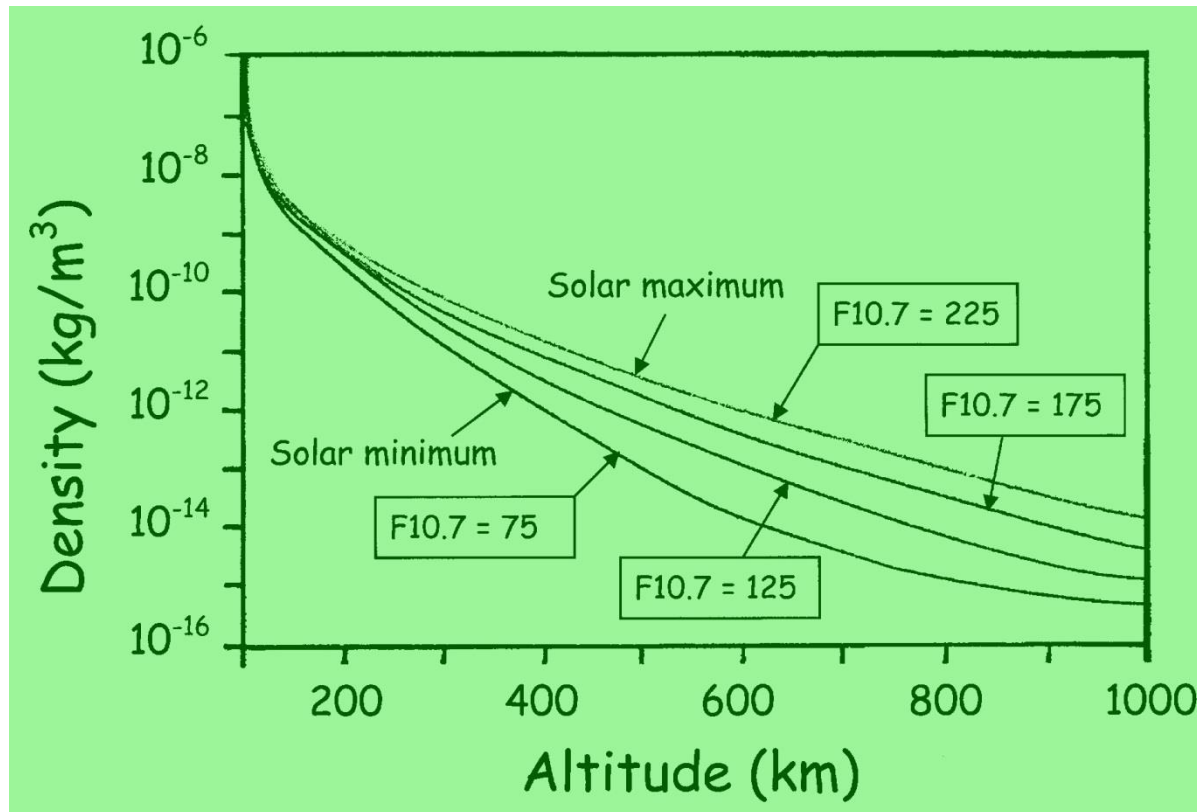
Exospheric Temperature with Solar Cycle



# Earth and the Earth Orbit Environment

Earth's Atmosphere – the atmosphere and the electromagnetic radiation

## Atmospheric Density



# Earth and the Earth Orbit Environment

## Atmosphere – Density

Density above some reference altitude  $h_o$  is give by:

$$\rho_i = \rho_{i0} \exp \left\{ \frac{-gM_i}{RT} (h - h_o) \right\}, \quad (2.1)$$

where

$\rho_i$  = density of species  $i$

$\rho_{i0}$  = density of species  $i$  at height  $h_o$

$g$  = acceleration due to gravity

$M_i$  = molecular weight of species  $i$

$R$  = Universal Gas Constant

$T$  = atmospheric temperature

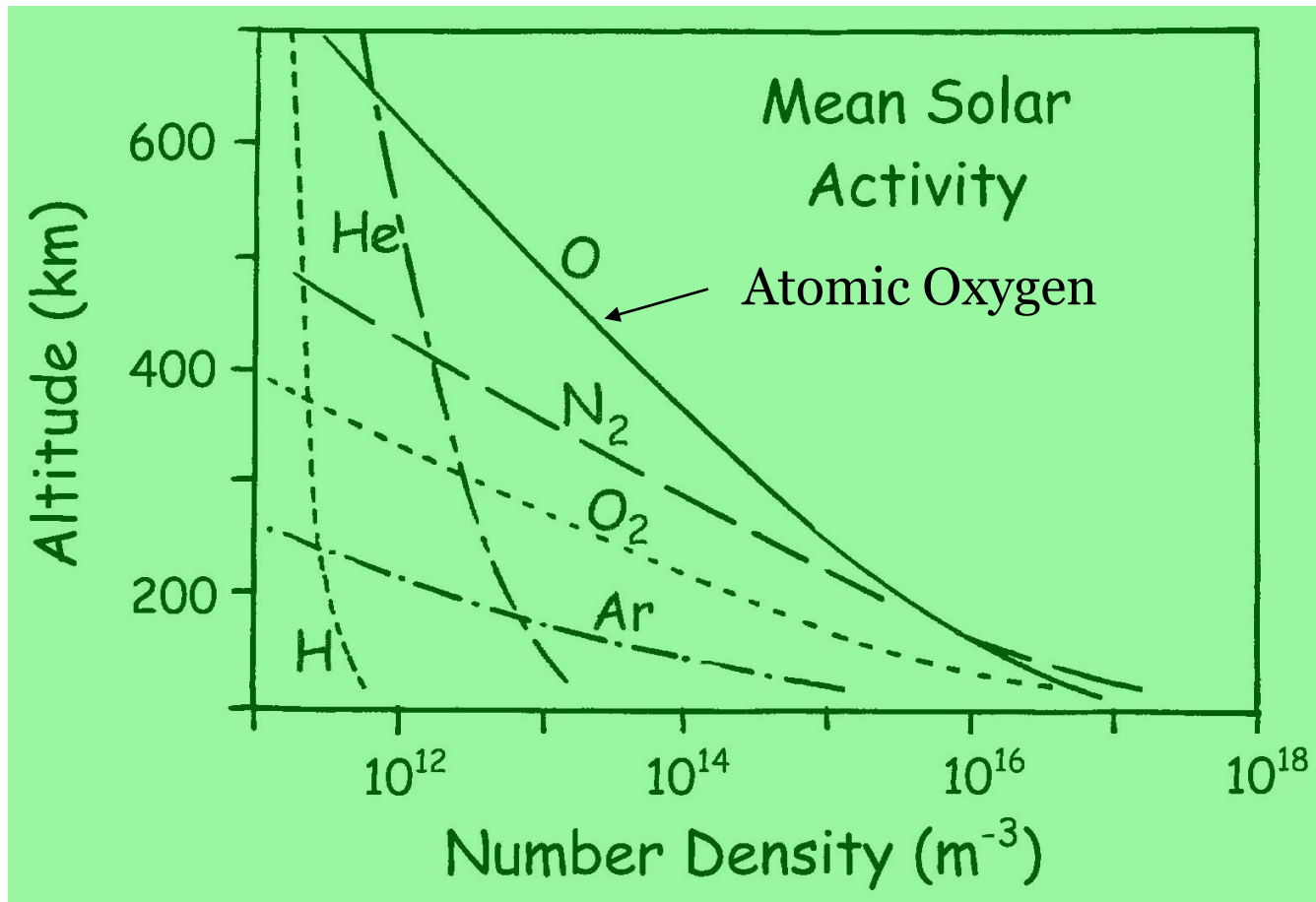
Total density is

$$\rho = \rho_0 \exp \left\{ - \frac{(h - h_0)}{H} \right\}, \quad (2.2)$$

where  $H$  = scale height ( $\propto T/M_i$ )

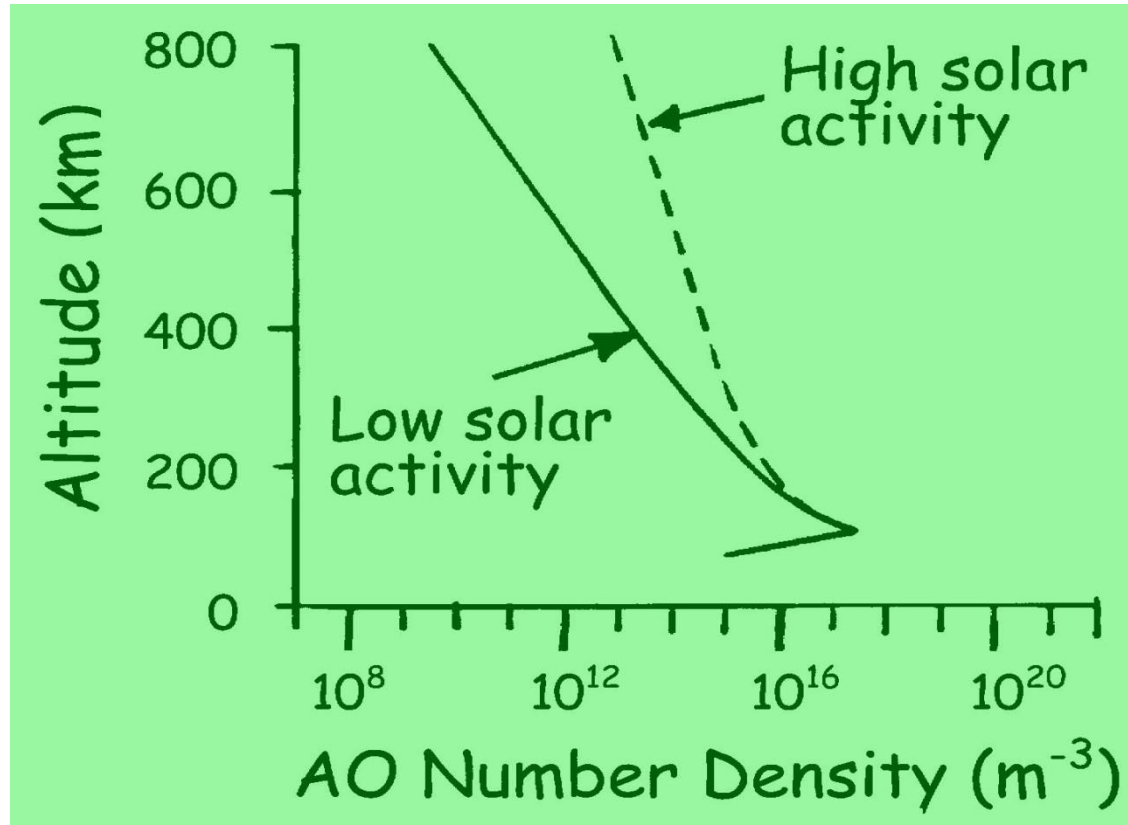
# Earth and the Earth Orbit Environment

## Earth's Atmosphere – Constituents



# Earth and the Earth Orbit Environment

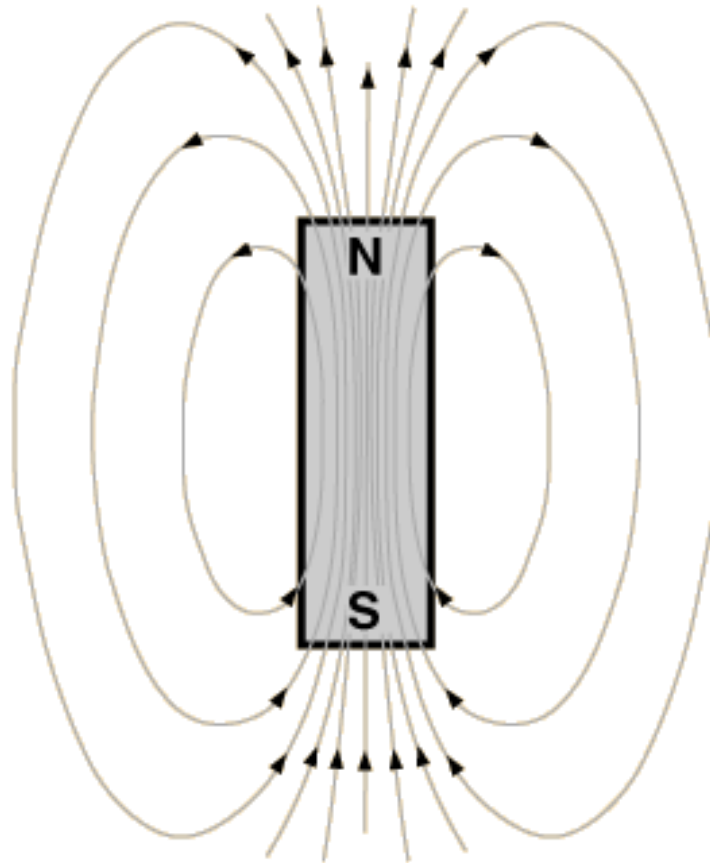
## Earth's Atmosphere – Atomic Oxygen (AO)





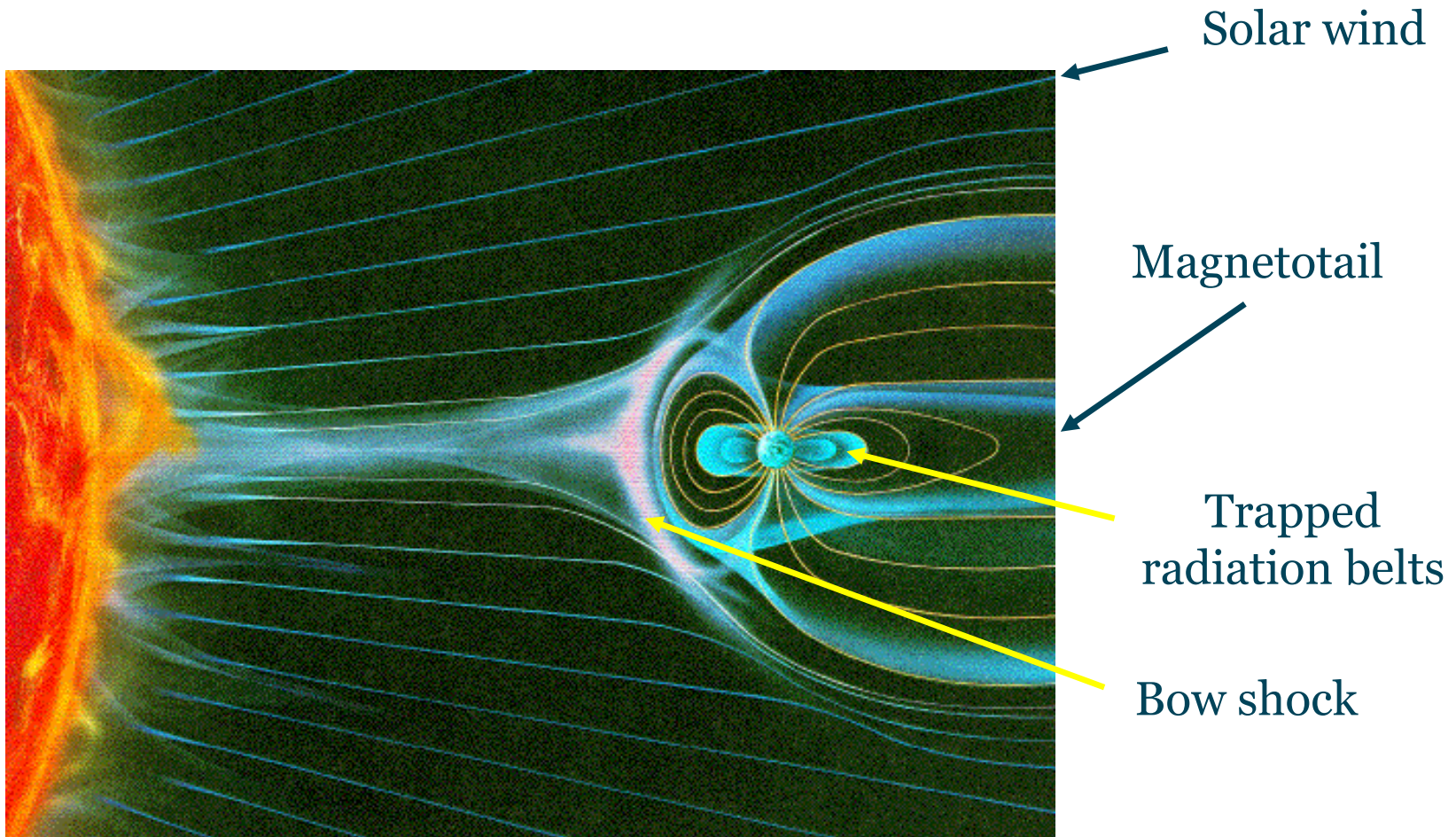
# Earth and the Earth Orbit Environment

## Earths Magnetic Field



# Earth and the Earth Orbit Environment

## Earth's Magnetic Field – The Magnetosphere



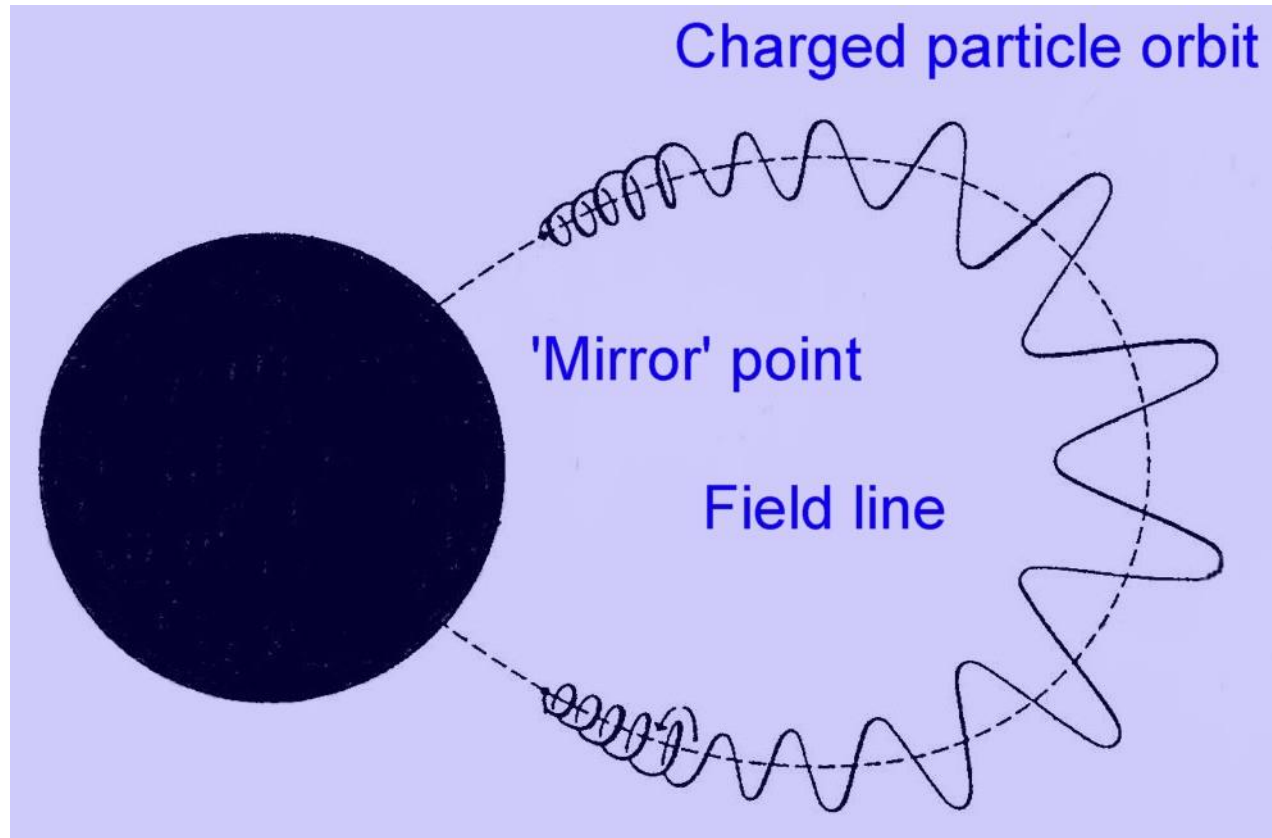
# Earth and the Earth Orbit Environment

## Earth's Magnetosphere and Ionizing Particle Radiation

- Structure of magnetosphere formed by interaction of solar wind and Earth's (approximately) dipole field
- Particles penetrate magnetosphere at North and South magnetic poles, and become trapped (Van Allen trapped radiation belts)
- Magnetosphere boundary is dynamic and variable, depending on intensity of the solar wind ('magnetic storms')

# Earth and the Earth Orbit Environment

## Ionizing Particle Radiation – Trapped Radiation

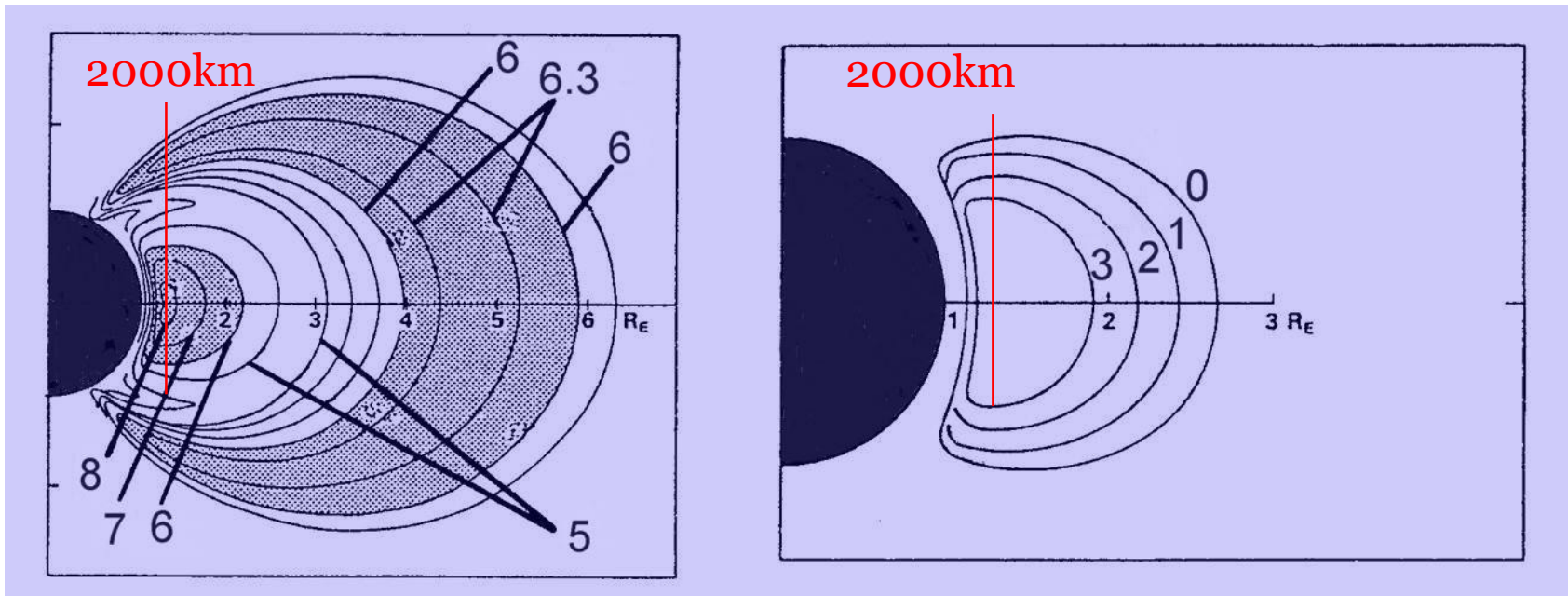


# Earth and the Earth Orbit Environment

## Ionizing Particle Radiation – Trapped Radiation

Structure of the Van Allen radiation belts.

Numerical values are x, where flux is  $10^x \text{ cm}^{-2} \text{ s}^{-1}$ .



Electrons,  $E > 0.5 \text{ MeV}$

High energy protons,  $E \geq 100 \text{ MeV}$



# Earth and the Earth Orbit Environment

## Debris

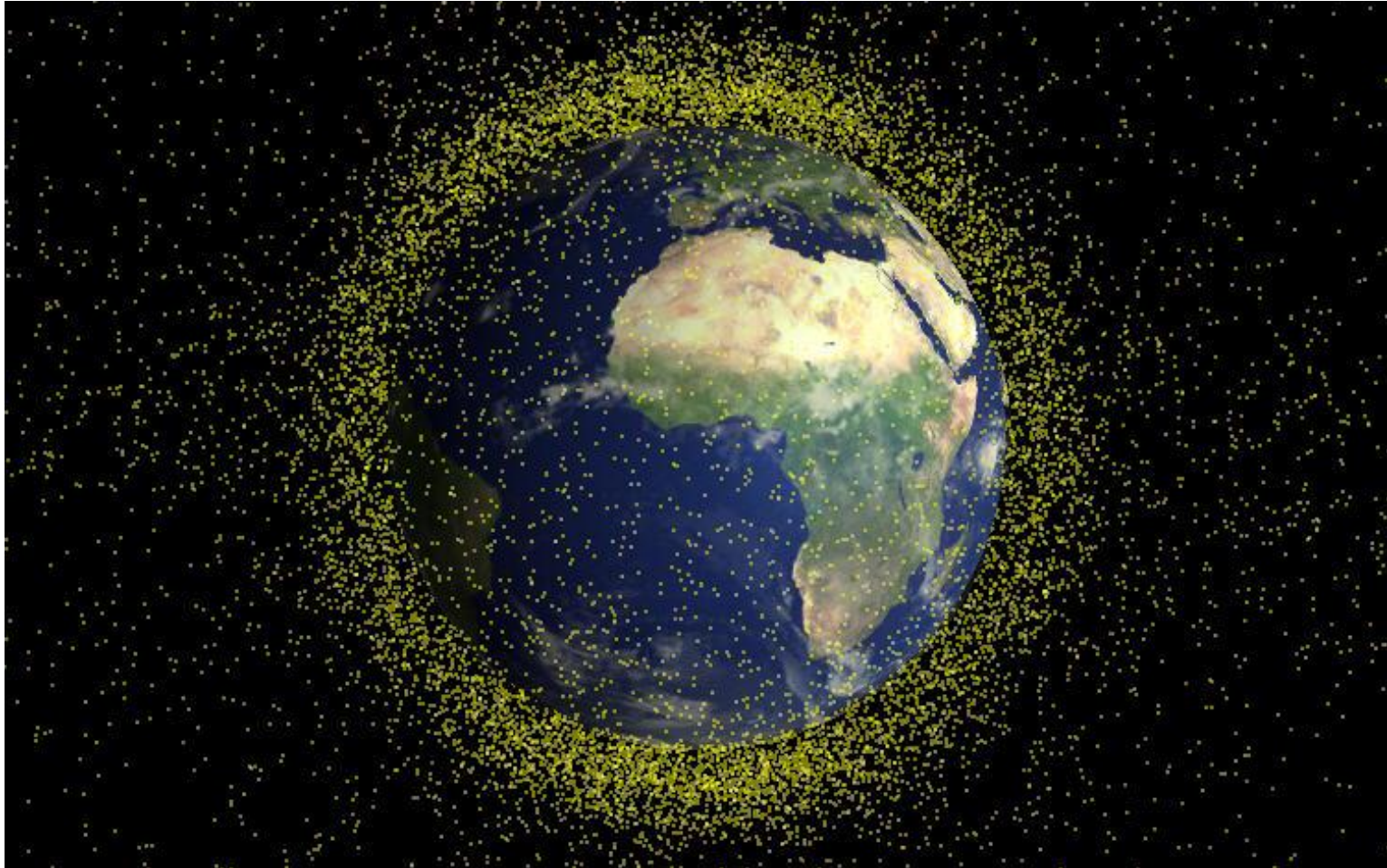
Divided into:

- *Natural Space Debris* – meteoroids that the earth encounters through its orbit (estimated that ~100 metric tonnes of this material rains down into the Earth atmosphere each day)
- *Artificial Debris* – man made objects  
(in 2009 Earths orbital environment contained 3,371 operational satellites and 11,492 debris objects  $\geq 10\text{cm}$ , NASA)



# Earth and the Earth Orbit Environment

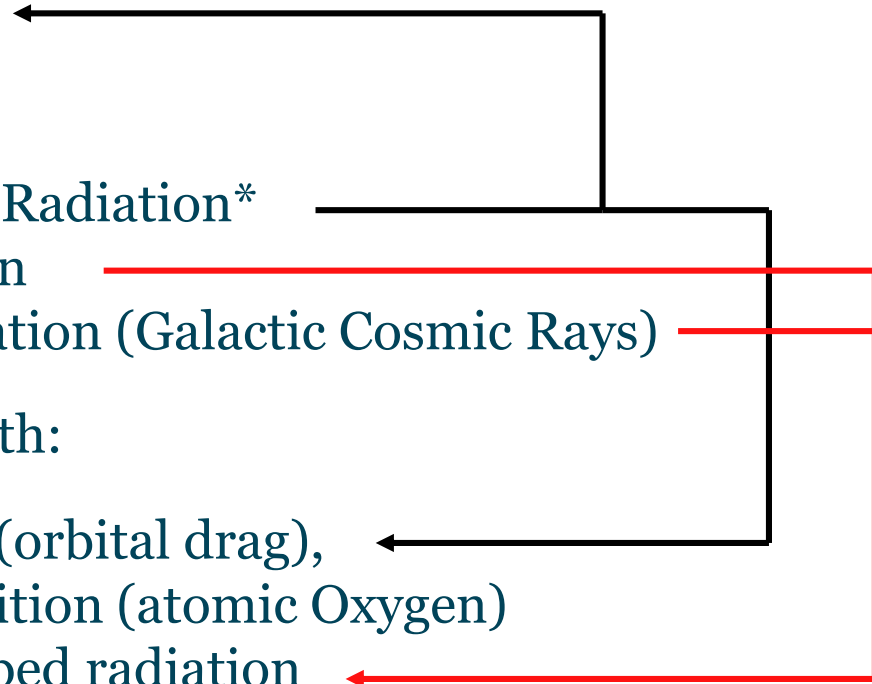
## Debris



# Summary of Space Environment

## In the Solar System:

- Low Temperatures\*
- Vacuum
- Microgravity
- Sun's Electromagnetic Radiation\*
- Sun's Particle Radiation
- External Particle Radiation (Galactic Cosmic Rays)



## Environment Around the Earth:

- Atmosphere – density (orbital drag),  
composition (atomic Oxygen)
- Magnetic Field – Trapped radiation
- Debris

# Environmental Effects on Design/Humans

Is the environment hostile or friendly?

- Hostile
  - High vacuum
  - Particle radiation
  - High energy Electromagnetic radiation
  - Atomic oxygen (Earths Atmosphere)
  - Debris
  
- Friendly
  - No water vapour/rain
  - No wind
  - Clean environment
  - Zero effective gravity

# Environmental Effects on Design/Humans

## Vacuum

Some materials evaporate in hard vacuum (known as ‘outgassing’)

- resins
- adhesives
- lubricants
- some metals

### Temperature for given outgassing rate

Material	0.1 $\mu\text{m}/\text{yr}$	1 mm/yr
Cadmium	38°C	122°C
Magnesium	110°C	233°C
Gold	660°C	950°C
Titanium	920°C	1250°C

Generally not an issue of structural integrity, but one of contamination.

# Environmental Effects on Design/Humans

## Micro gravity

- Structural design
- Difficulty in mechanism deployment testing on the ground
- No convective heat transport
- Propellant management problems
- (Physiological effects on humans)

# Environmental Effects on Design/Humans

## Solar Radiation – Electromagnetic Output

Effects of solar e.m. radiation (UV):

- Changes in optical and electrical properties
- Embrittlement in polymers
- Darkening of solar array cover-glass and adhesives



# Environmental Effects on Design/Humans

## Ionizing Particle Radiation – Effects on materials and systems

- Degradation of semi-conductor devices (solar arrays)
- On-board software errors by single particle event (Single Event Upsets, SEU)
- Spacecraft charging
- (Physiological effects on humans)

# Environmental Effects on Design/Humans

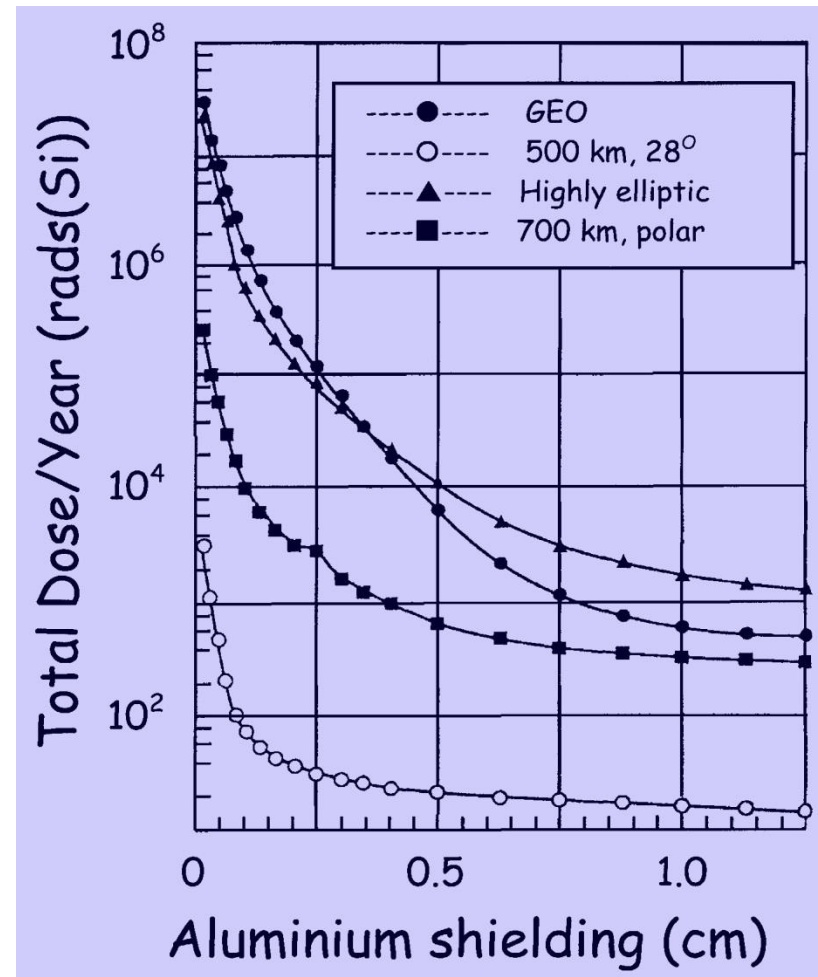
## Ionizing Particle Radiation – Total Dose as a Function of Orbit

Standard Unit:

A rad (Si) (Radiation Absorbed Dose) is that amount of radiation which deposits 100 ergs of energy per gram of silicon semi-conducting material.

(1 erg =  $10^{-7}$  J,

→ 100 rads = 1 J/kg)



# Environmental Effects on Design/Humans

## Earth Atmosphere - Effects

- aerodynamic drag, lift and heating (re-entry)
- erosion (AO)

## Atomic Oxygen (AO) Erosion Example

Estimate the depth of erosion of kapton for a typical space shuttle mission of 10 days duration during solar maximum.

### Kapton:

- material commonly used in spacecraft insulation blankets.
- erosion rate  $\sim 2.8 \mu\text{m}$  for every  $10^{24}$  atoms/ $\text{m}^2$  of AO fluence

# Environmental Effects on Design/Humans

## Earth Atmosphere – Atomic Oxygen (AO) Erosion Example

AO fluence over time  $t$  is

$$\boxed{F_0 = n_{AO} V t \quad (\text{m}^{-2})} \quad (2.3)$$

where  $n_{AO}$  = number density of AO ( $\text{m}^{-3}$ )  
 $V$  = spacecraft velocity ( $\text{m s}^{-1}$ )

Typical shuttle altitude  $\sim 300$  km, so

$$F_0 = (10^{15} / \text{m}^3) (7.7 \times 10^3 \text{ m/s}) (8.64 \times 10^5 \text{ s})$$

$$\sim 6.7 \times 10^{24} \text{ atoms} / \text{m}^2$$

$$\therefore \text{Depth of erosion} \quad \sim 2.8 \mu\text{m} (6.7 \times 10^{24}) / 10^{24}$$

$$\sim 19 \mu\text{m}$$

# Environmental Effects on Design/Man

## Earth Atmosphere – Atomic Oxygen (AO) Erosion rates

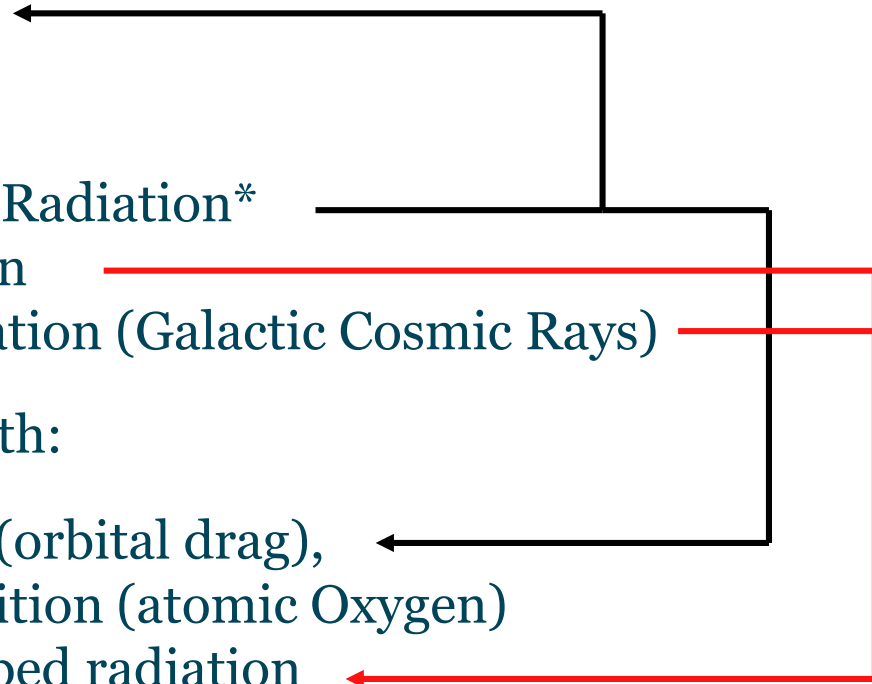
Material	Surface Recession* ( $\mu\text{m}$ )
Kapton	2.8
Teflon	0.029
Aluminium	0.00035
Silver	9.8

\* Surface recession for every  $10^{24}$  atoms /  $\text{m}^2$  of AO fluence.

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- Sun's Particle Radiation
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# Launch Environment

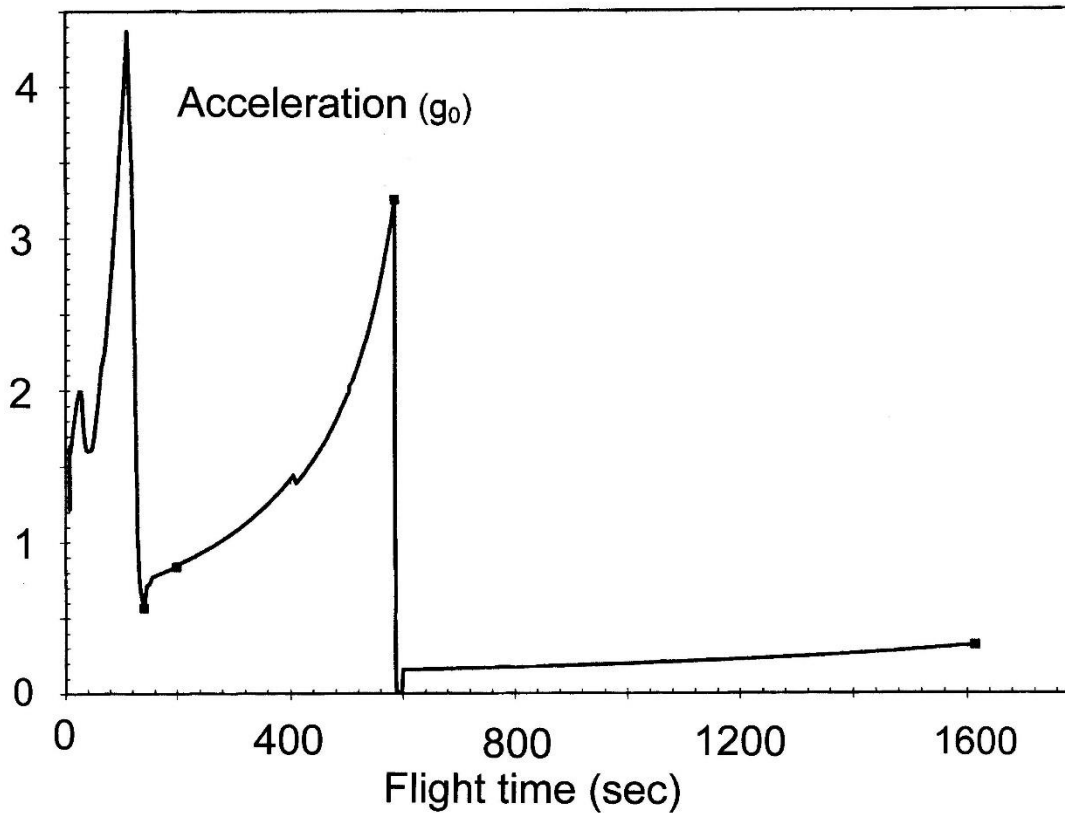
- Most Severe Environment (usually):
  - main driver for structural design of spacecraft
- ‘Static’ Acceleration:
  - launch vehicle dependent
    - Ariane  $5 \leq 4g_0$
    - Shuttle  $\leq 3g_0$
    - Saturn  $\leq 8g_0$
  - peak occurs at stage burnout



# Launch Environment

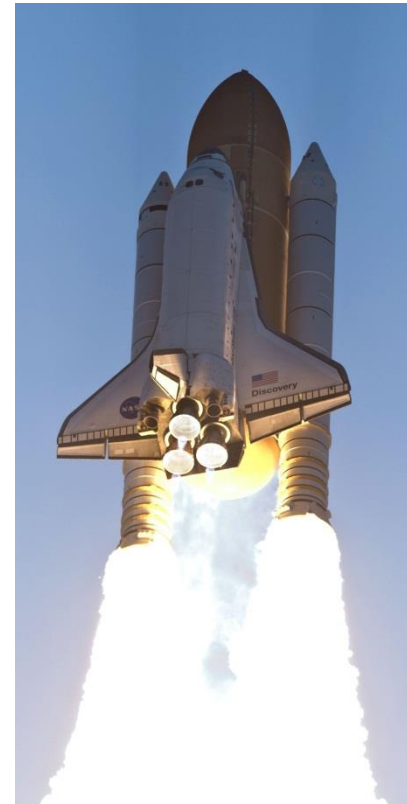
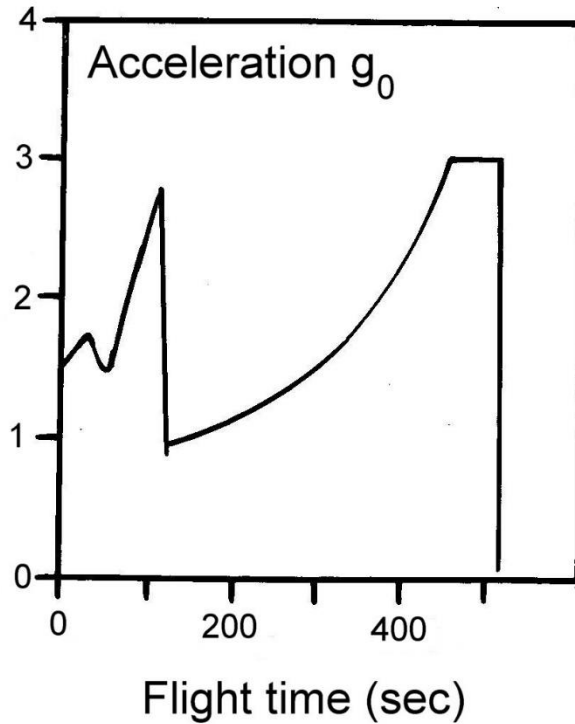
## Static Acceleration

### Ariane 5 static acceleration on launch



# Launch Environment

## Static Acceleration



## Space shuttle static acceleration

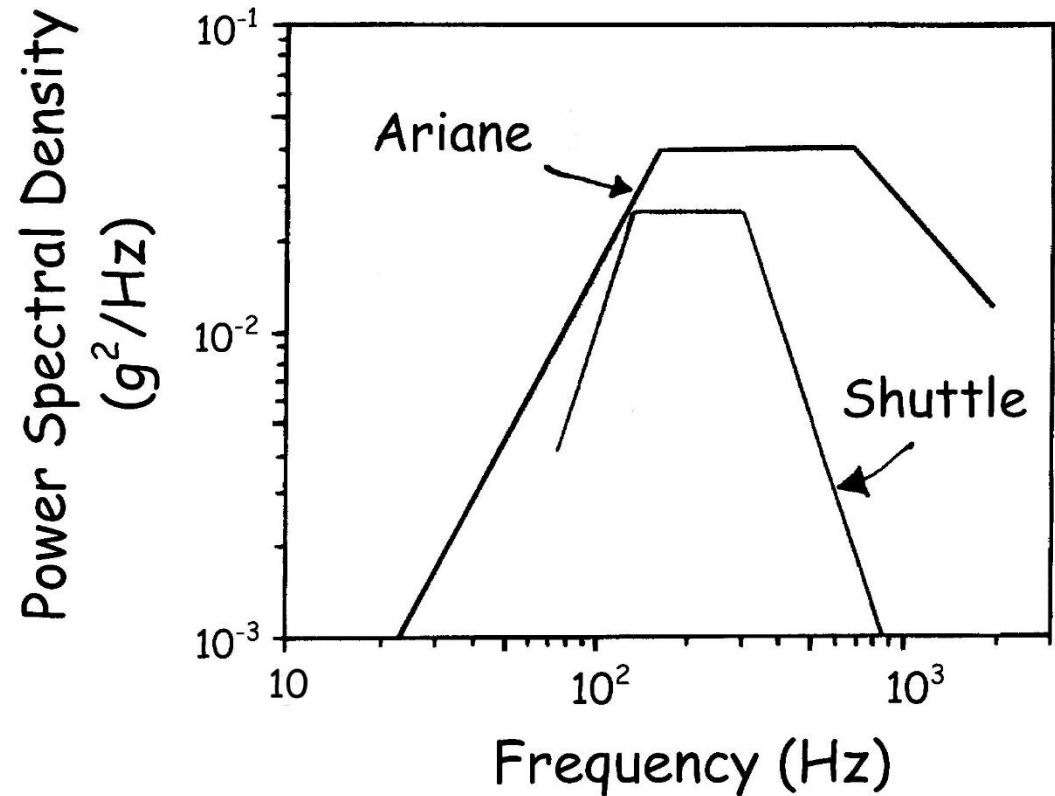
# Launch Environment

## Dynamic Acceleration

(or “oscillatory acceln.” or “random vibration”)

Generated by:

- Functioning machinery
- Combustion phenomena
- Structural excitation in acoustic field



PSD Envelopes

# Launch Environment

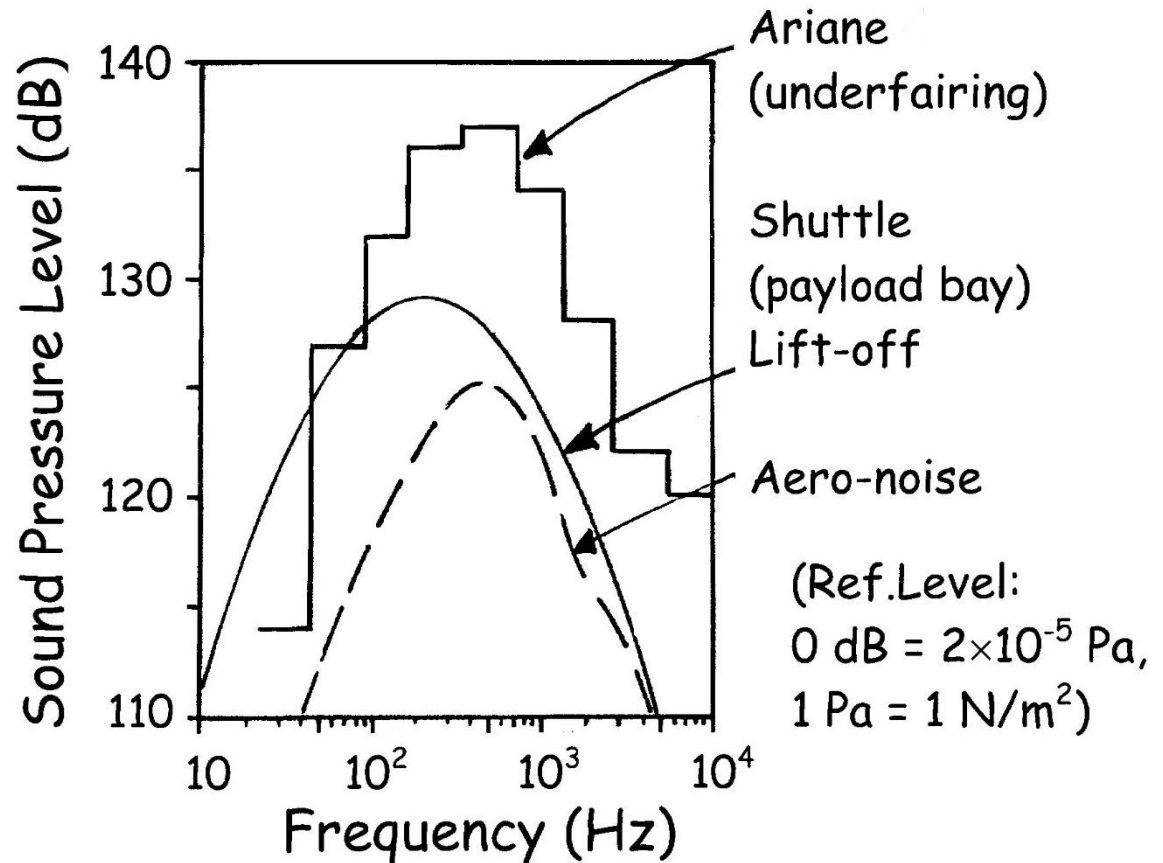
## Acoustic Environment

Intense acoustic field generated by:

- engine noise
- aerodynamic noise

Greatest intensity at:

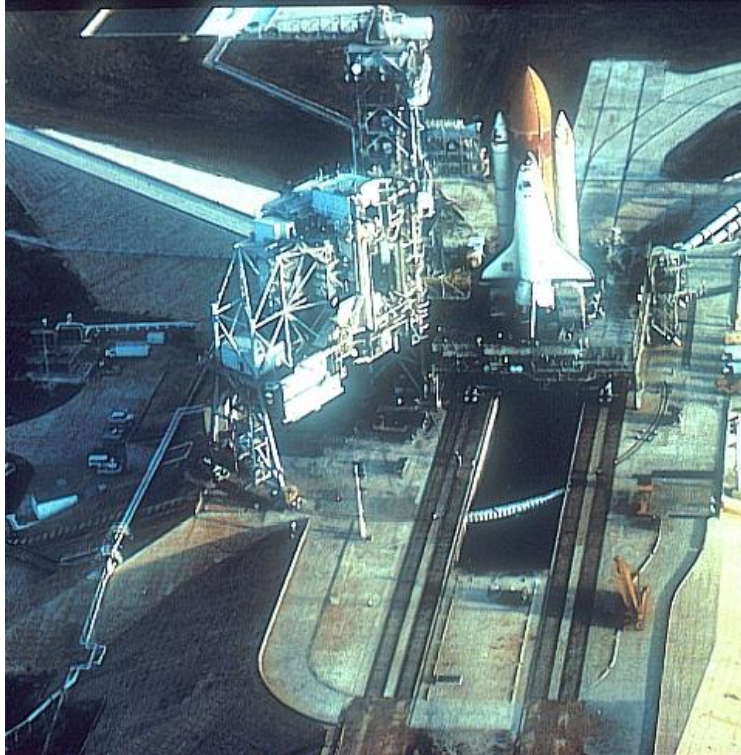
- launcher release
- transonic speed





# Launch Environment

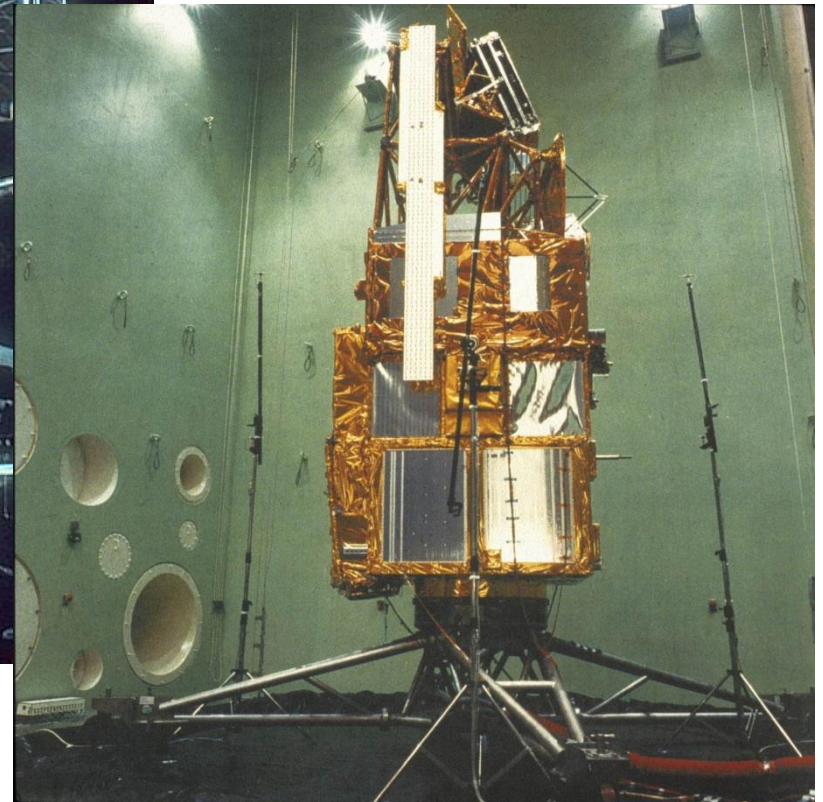
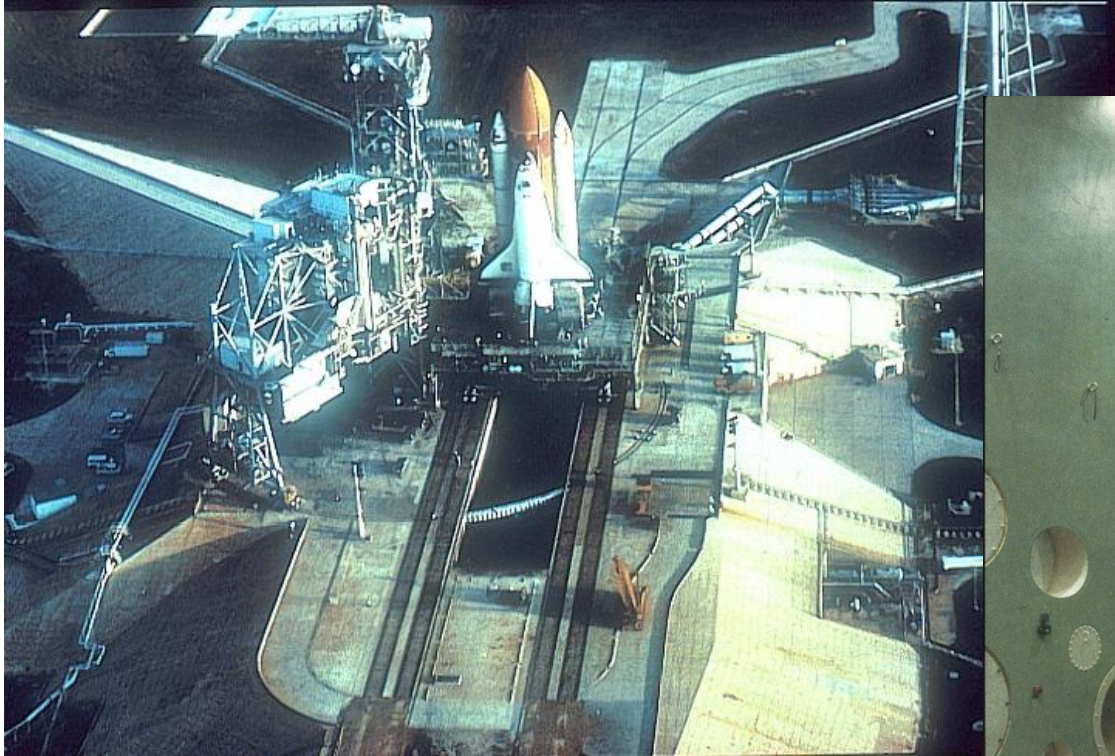
## Acoustic Environment





# Launch Environment

## Acoustic Environment



# Launch Environment

## Effects on Spacecraft

- Main input to structural design
- Ground testing to ensure spacecraft survival on launch



# Environment Summary

## Key points:

Natural and Induced  
Environments



- Categories and types of environments that the satellite will have to survive

Our Solar System



- The important environmental inputs to our Solar system and their basic properties

Earth and the Earth Orbit  
Environment



- The important properties of the space environment around the Earth
- How these properties interact with the inputs from the Solar system and
- How these properties generally vary proportional to orbital altitude/location

Environmental Effects on  
Design/Man



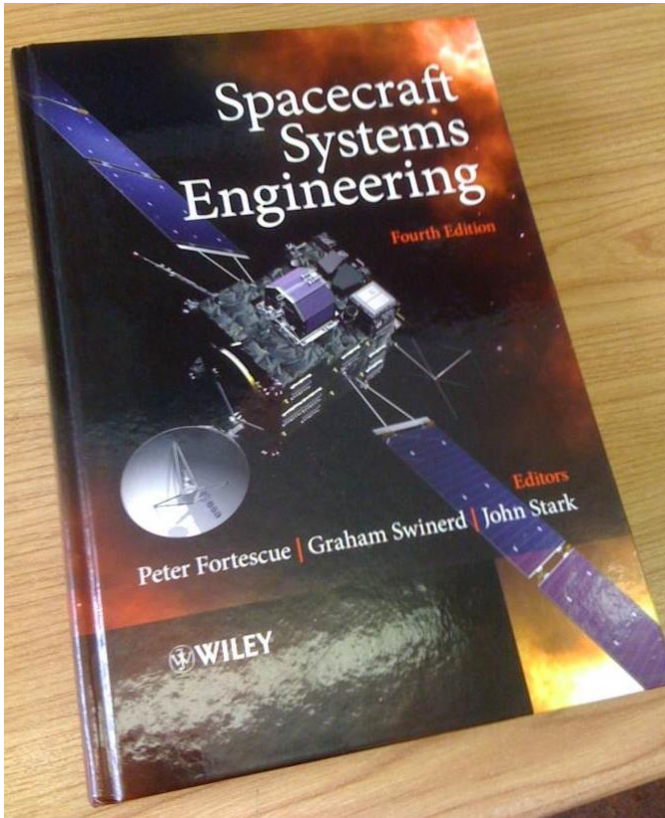
- How the space environmental properties, both in the Solar system and around the Earth, most significantly affect the design of spacecraft

Launch Environment

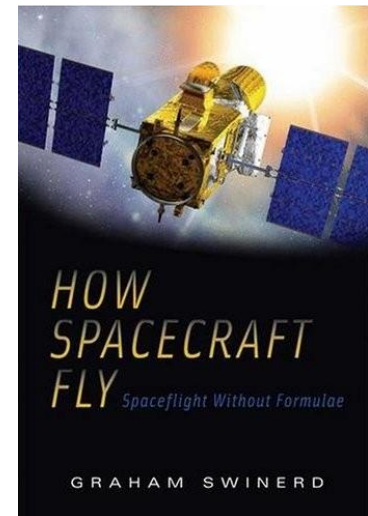


- The important environmental inputs from launch
- How these environments are quantified
- How they affect the design of spacecraft

# Environment Summary



Read Chapter 2 of  
Fortescue, Stark &  
Swinerd



Read Chapter 6 of  
'How S/C Fly'