

# Chapter 5: Mission Analysis

## Lecture 2 – Kepler's Laws

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# Overview of lecture 2

- This is a short lecture focused on:
  1. Kepler's Laws of Planetary Motion:
    - These are three laws identified by Kepler using observations of
      - The shape of planetary orbits
      - The nature of the motion of planets around the Sun
      - The time taken for planets to move around the Sun
  2. How Kepler's Laws became the basis for the modern, mathematical understanding of orbital motion

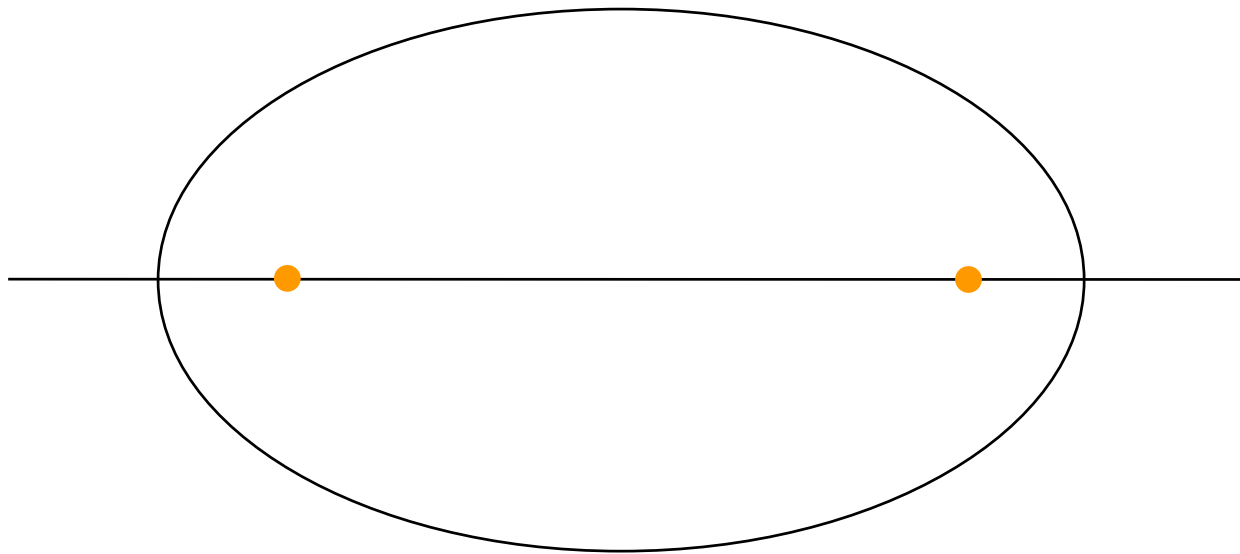
# Kepler's laws of planetary motion

- **Johannes Kepler**
  - December 1571 – November 1630
  - A key figure in the 17th-century scientific revolution, best known for his laws of planetary motion, and his books *Astronomia nova*, *Harmonices Mundi*, and *Epitome Astronomiae Copernicanae*.
  - *"I demonstrate by means of philosophy that the earth is round, and is inhabited on all sides; that it is insignificantly small, and is borne through the stars"*



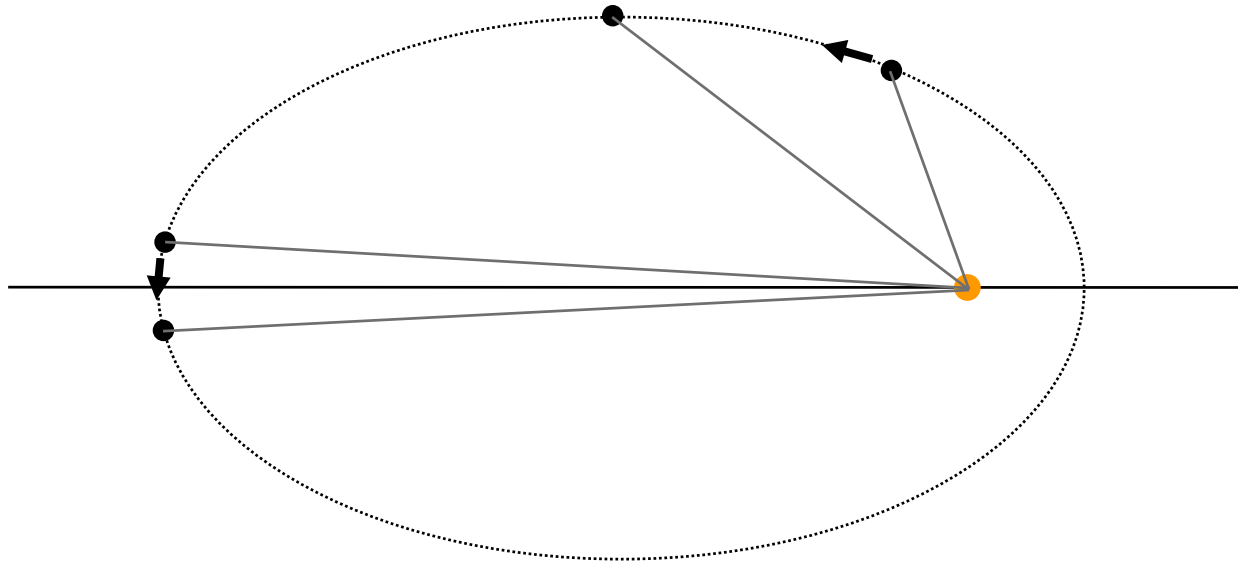
# Kepler's laws of planetary motion

1. The orbit of each planet is an ellipse with the Sun at one focus (1609)



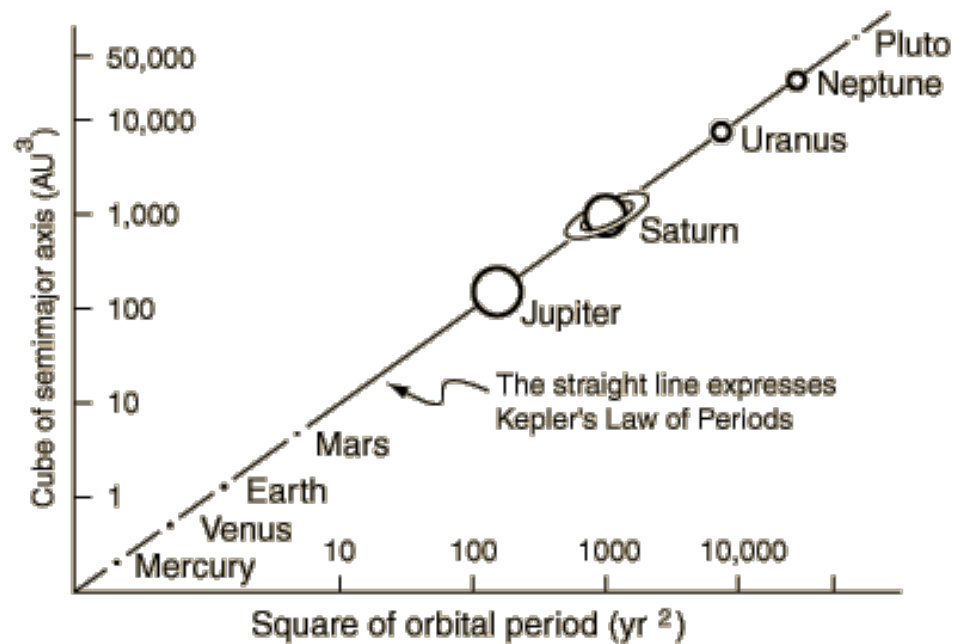
# Kepler's laws of planetary motion

2. The line joining the planet to the Sun sweeps out equal areas in equal times (1609)



# Kepler's laws of planetary motion

3. The square of the period of the planet is proportional to the cube of its mean distance from the Sun (1619)



# Kepler's laws of planetary motion

- Kepler's laws originated from observations of the solar system
- In his book 'Philosophiae Naturalis Principia Mathematica' (1687) Isaac Newton established that Kepler's laws follow mathematically from his Law of Universal Gravitation and his Laws of Motion
  - He proved using calculus that orbits are elliptical if the gravitational force is inverse square

We will look at Kepler's 1<sup>st</sup> Law initially: we will first understand how to describe an ellipse using mathematics (the ellipse equation)

