



This work has been supported by MEXT and JICFuS as a priority issue (Elucidation of the fundamental laws and evolution of the universe) to be tackled by using Post 'K' Computer

After mastering the shell model,
three (possible) pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
dimensionality

massive
parallel
computers

Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
8+8 on ^{80}Zr core (Sn)
8+10 on ^{132}Sn core
(Sm)

...
island of stability

+

χ EFT based
(multi-)shell int.

many-body dynamics

Shell evolution
(Type I & II)

Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Hamiltonian in shell model calculations

ε_i : Single Particle Energy (SPE)

$\langle j_1, j_2, J, T | V | j_3, j_4, J, T \rangle$

: Two-Body Matrix Element (TBME)

$$H = \sum_i \boxed{\varepsilon_i} n_i + \sum_{i,j,k,l} v_{ij,kl} \boxed{a_i^\dagger a_j^\dagger a_l a_k}$$

Step 1: Calculate matrix elements

$$\langle \phi_1 | H | \phi_1 \rangle, \quad \langle \phi_1 | H | \phi_2 \rangle, \quad \langle \phi_1 | H | \phi_3 \rangle, \dots$$

where ϕ_1, ϕ_2, ϕ_3 are Slater determinants

In the second quantization,

$$\phi_1 = a_\alpha^+ a_\beta^+ a_\gamma^+ \dots \dots |0\rangle$$

closed shell

$$\phi_2 = a_{\alpha'}^+ a_{\beta'}^+ a_{\gamma'}^+ \dots \dots |0\rangle$$

$$\phi_3 = \dots$$

$$H = \sum_i \epsilon_i n_i + \sum_{i,j,k,l} v_{ij,kl} a_i^\dagger a_j^\dagger a_l a_k$$

Step 2: we solve the eigenvalue problem : $\mathbf{H} \Psi = \mathbf{E} \Psi$

$$\begin{bmatrix} <\phi_1 | \mathbf{H} | \phi_1> & <\phi_1 | \mathbf{H} | \phi_2> & . & . & . \\ <\phi_2 | \mathbf{H} | \phi_1> & <\phi_2 | \mathbf{H} | \phi_2> & . & . & . \\ <\phi_3 | \mathbf{H} | \phi_1> & . & . & . & . \\ <\phi_4 | \mathbf{H} | \phi_1> & . & . & . & . \\ . & . & . & . & . \end{bmatrix} \begin{bmatrix} c_1 \\ c_2 \\ c_3 \\ c_4 \\ . \\ . \end{bmatrix} = \mathbf{E} \begin{bmatrix} c_1 \\ c_2 \\ c_3 \\ c_4 \\ . \\ . \end{bmatrix}$$

$$\Psi = c_1 \phi_1 + c_2 \phi_2 + c_3 \phi_3 + \dots$$

c_i probability amplitudes

With Slater determinants $\phi_1, \phi_2, \phi_3, \dots$,
the eigen wave function is expanded as

$$\Psi = c_1 \phi_1 + c_2 \phi_2 + c_3 \phi_3 + \dots$$

c_i probability amplitudes

With this, we can calculate various physical quantities.

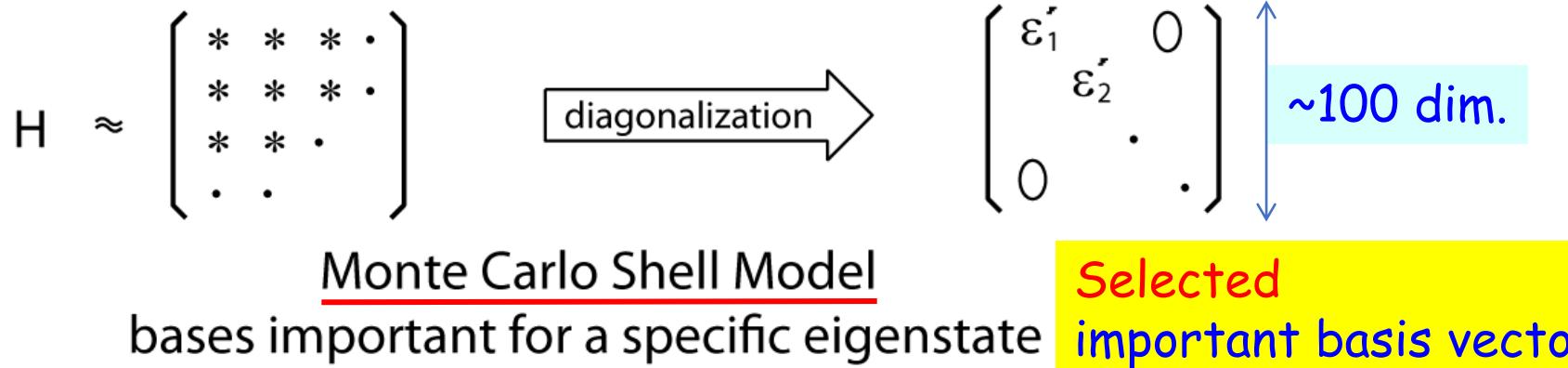
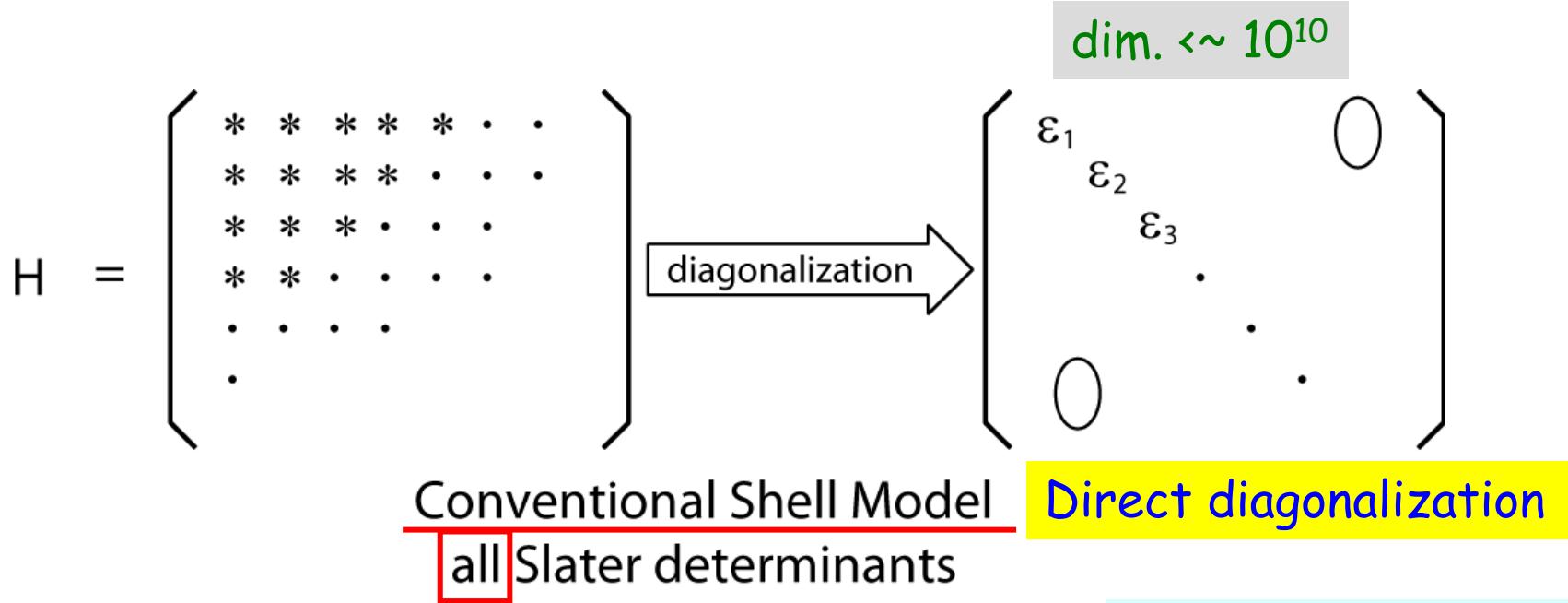
for instance, E2, M1, ... matrix elements

(transitions and moments)

spectroscopic factors for transfer
reactions

$$\langle \Psi' | T | \Psi \rangle$$

Two types of shell-model calculations



Monte Carlo Shell Model

Auxiliary-Field Monte Carlo (AFMC) method

general method for quantum many-body problems

For nuclear physics, **Shell Model Monte Carlo (SMMC)** calculation has been introduced by **Koonin et al.** Good for finite temperature.

- minus-sign problem
- only ground state, not for excited states in principle.

Quantum Monte Carlo Diagonalization (QMCD) method

No sign problem. Symmetries can be restored.

Excited states can be obtained.

→ **Monte Carlo Shell Model (MCSM)**

Background of Monte Carlo Shell Model (I)

Two-body interaction can be rewritten as $V = (1/2) \sum_{\alpha} v_{\alpha} O_{\alpha}^2$

α : index

O_{α} : one-body operators (rearranged by diagonalization)

Hubbard-Stratonovich transformation

True eigenstate : $\psi = \sum_{\sigma} e^{-\beta h(\sigma)} e^{-\beta h(\sigma')} \dots \psi_0$

imaginary time (β) evolution by one-body field $h(\sigma)$

One-body operator is introduced as $h(\sigma) = \sum_{\alpha} s_{\alpha} \sigma_{\alpha} v_{\alpha} O_{\alpha}$

σ_{α} : random number (Gaussian) σ : set of σ_{α} 's

s_{α} : phase (1 for $v_{\alpha} < 0$, i for $v_{\alpha} > 0$)

Background of Monte Carlo Shell Model (II)

True eigenstate : $\psi = \sum_{\sigma, \sigma'} \dots e^{-h(\sigma)} e^{-h(\sigma')} \dots \psi_0$

Use $\phi(\sigma, \sigma', \dots) = e^{-\beta h(\sigma)} e^{-\beta h(\sigma')} \dots \psi_0$
as a basis for shell-model diagonalization

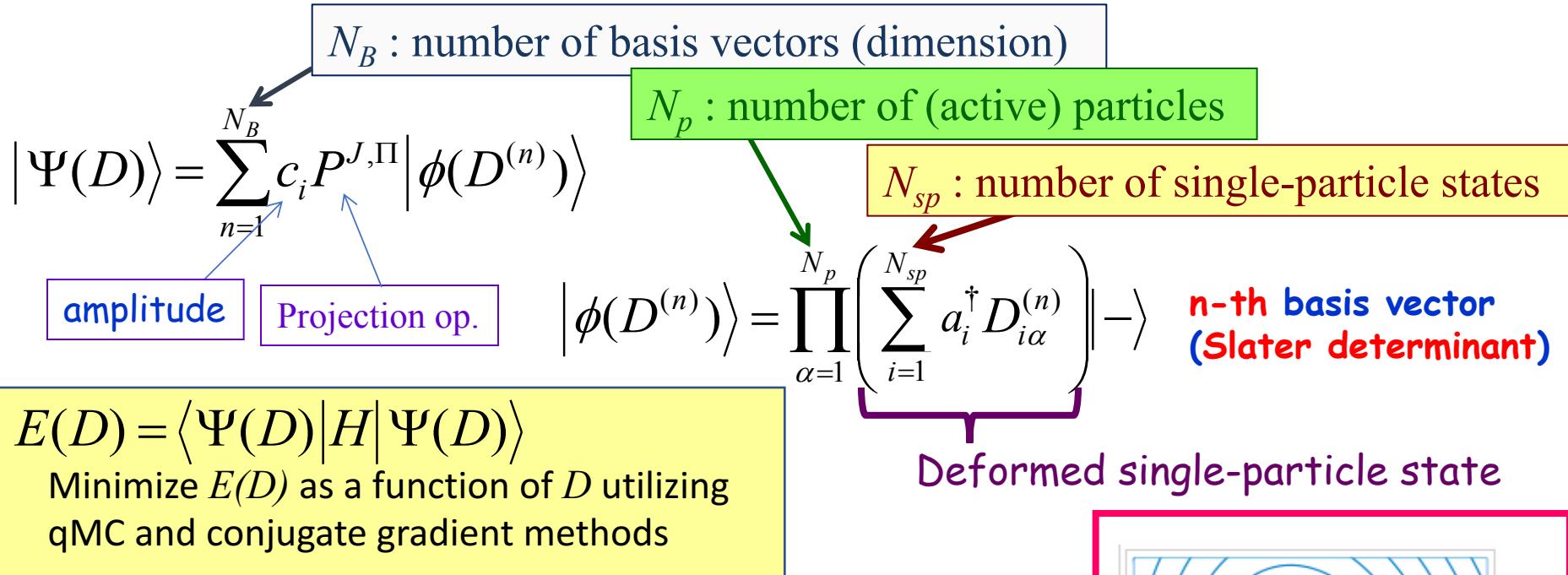
$\phi(\sigma, \sigma', \dots)$ are selected and refined :

- (i) Random sampling -> only those lowering energy are kept
- (ii) Polished by varying σ, σ', \dots gradually (random noise reduced)
- (iii) Symmetry restoration (Angular momentum, parity)

Slater determinants (or Cooper-pair type wave functions)
are used

Usually, 20~50 bases are kept (many more thrown away)

Advanced Monte Carlo Shell Model (currently used)



Step 1 : stochastic generation of candidates of the n -th MCSM basis vector

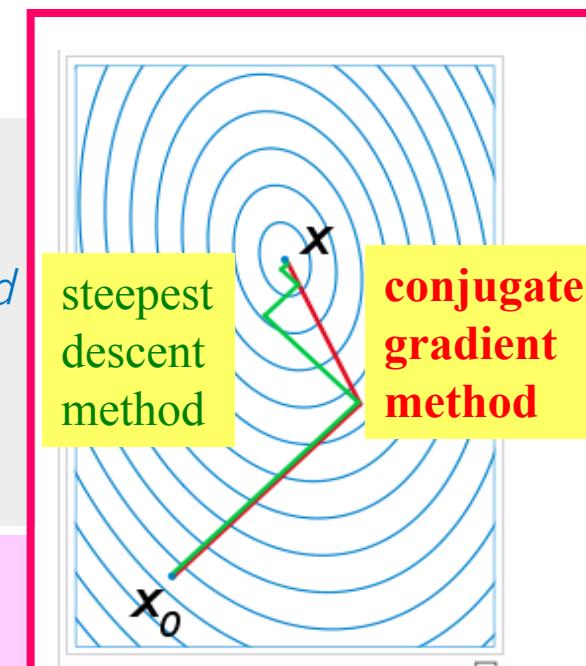
~~$|\phi(\sigma)\rangle = \prod_i e^{\Delta \rho_i h(\sigma)} \cdot |\phi^{(0)}\rangle$~~ only theoretical background

Shift randomly matrix elements of the matrix D .

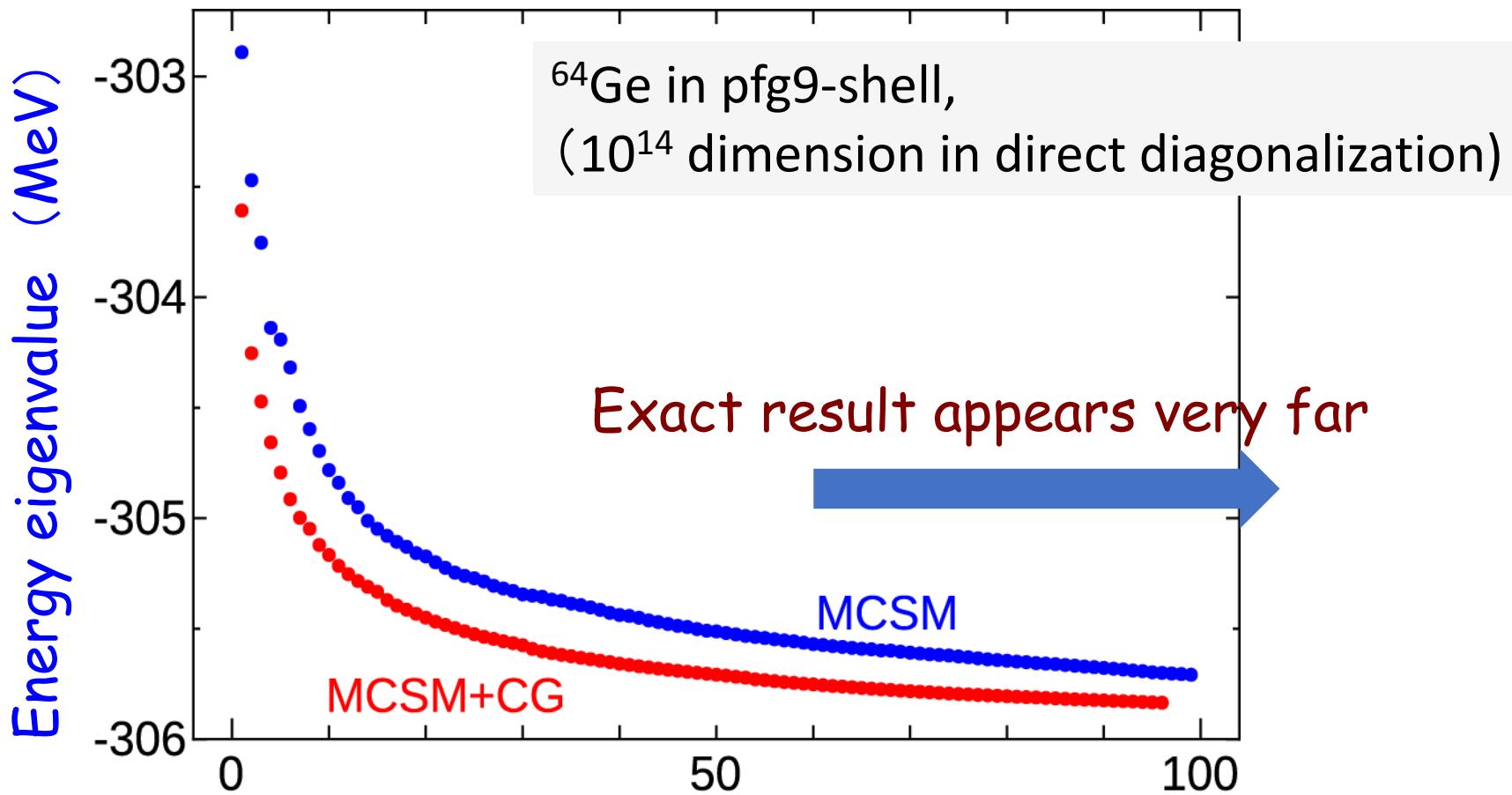
(The very initial one can be a Hartree-Fock state.)

Select the one producing the lowest $E(D)$ (rate < 0.1 %)

Step 2 : polish D by means of the conjugate gradient (CG) method “variationally”.

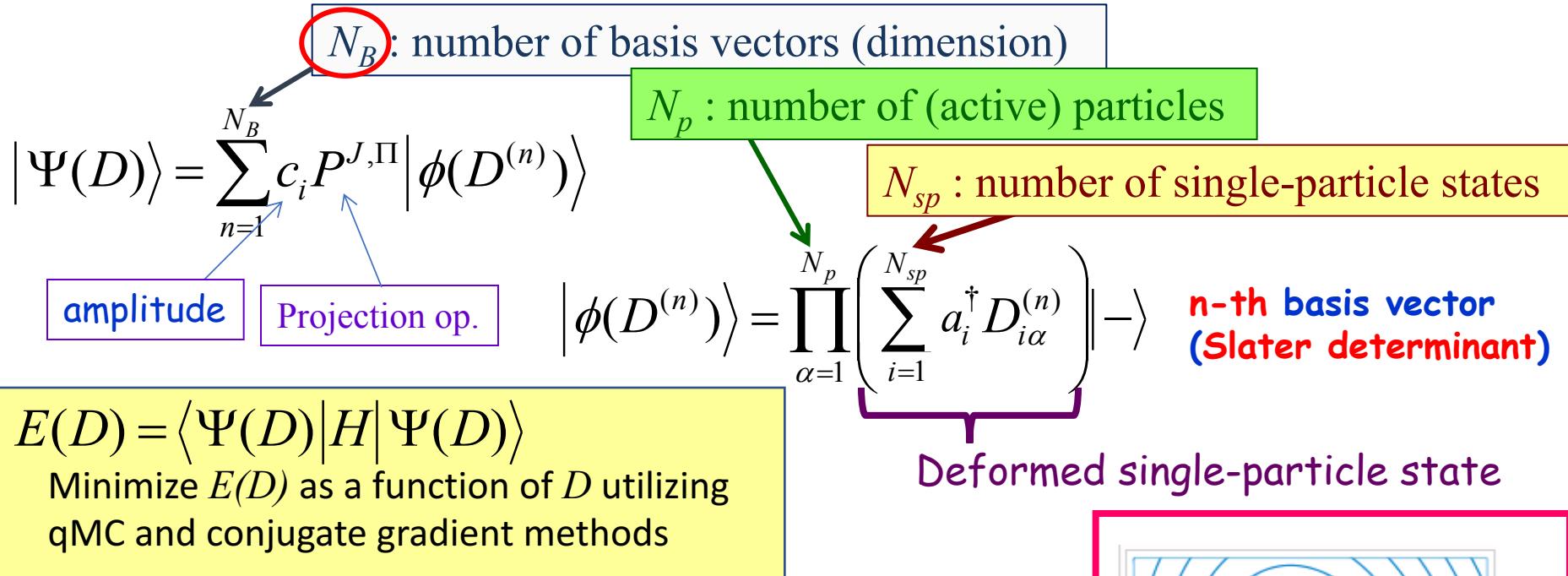


Example of MCSM calculation



Numerous MC trials and CG optimization for each basis vector

Advanced Monte Carlo Shell Model (currently used)



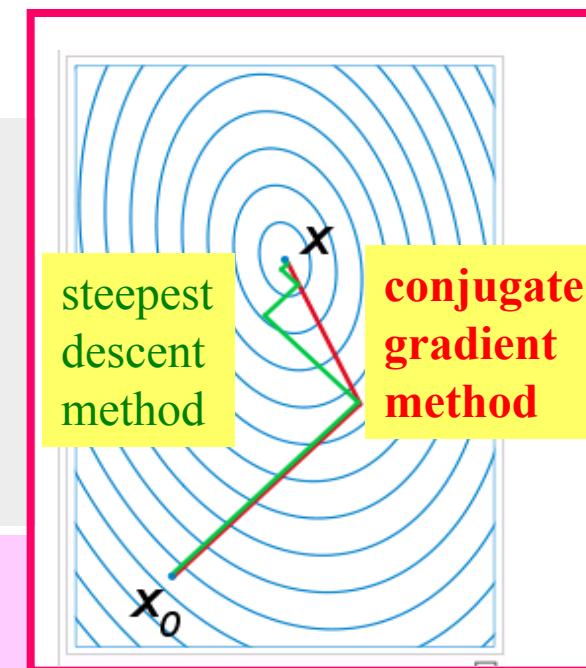
Step 1 : stochastic generation of candidates of the n -th MCSM basis vector

$|\phi(\sigma)\rangle = \prod_i e^{\Delta \rho_i h_i(\sigma)} |\phi^{(0)}\rangle$ only theoretical background

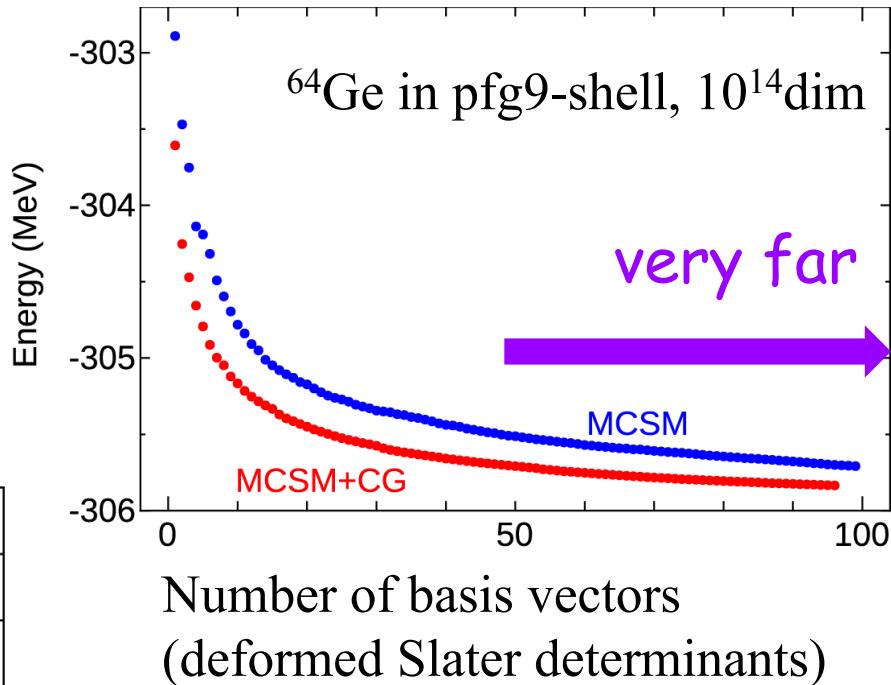
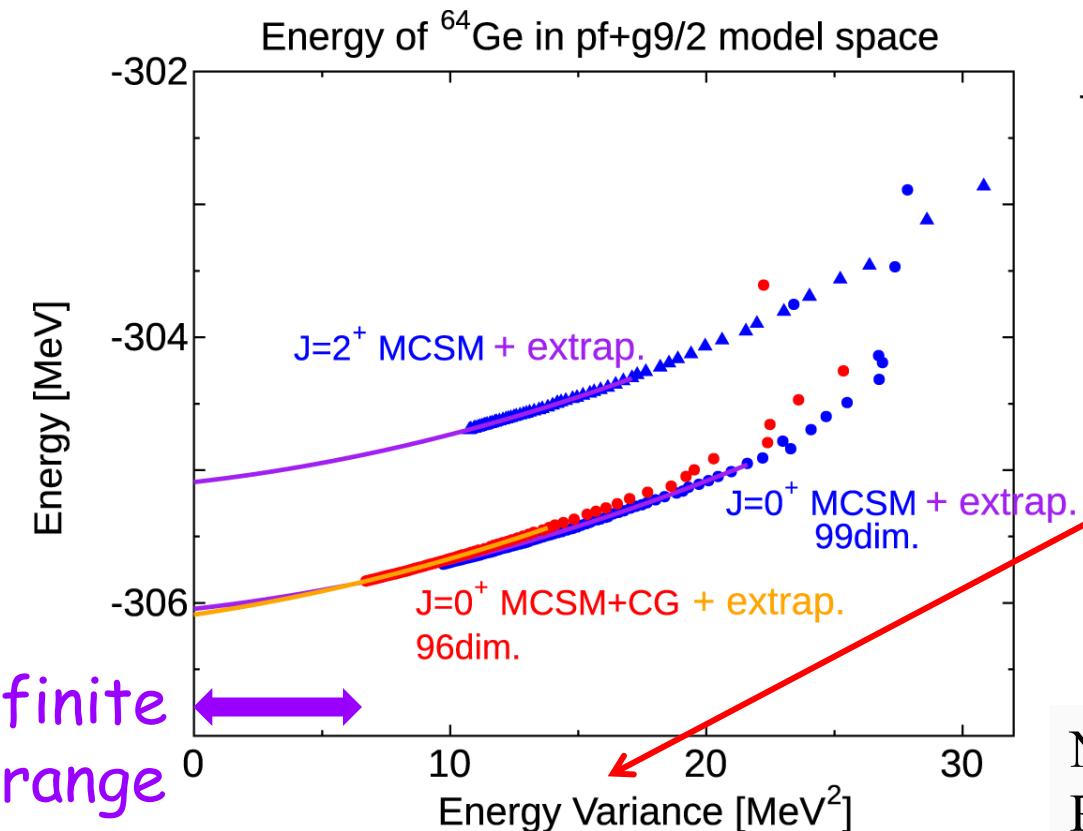
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Extrapolation by Energy Variance



Variance : $\langle \Delta H^2 \rangle = \langle H^2 \rangle - \langle H \rangle^2$

$$\langle H \rangle = E_0 + a\langle \Delta H^2 \rangle + b\langle \Delta H^2 \rangle^2 + \dots$$

N. Shimizu, et al.,
Phys. Rev. C **82**, 061305(R) (2010).

MCSM (Monte Carlo Shell Model -Advanced version-)

1. Selection of important many-body basis vectors

by quantum Monte-Carlo + diagonalization methods

basis vectors : about 100 selected Slater determinants

composed of "deformed" single-particle states

2. Variational refinement of basis vectors

conjugate gradient method

3. Variance extrapolation method → exact eigenvalues

+ innovations in algorithm and code (=> now moving to GPU)



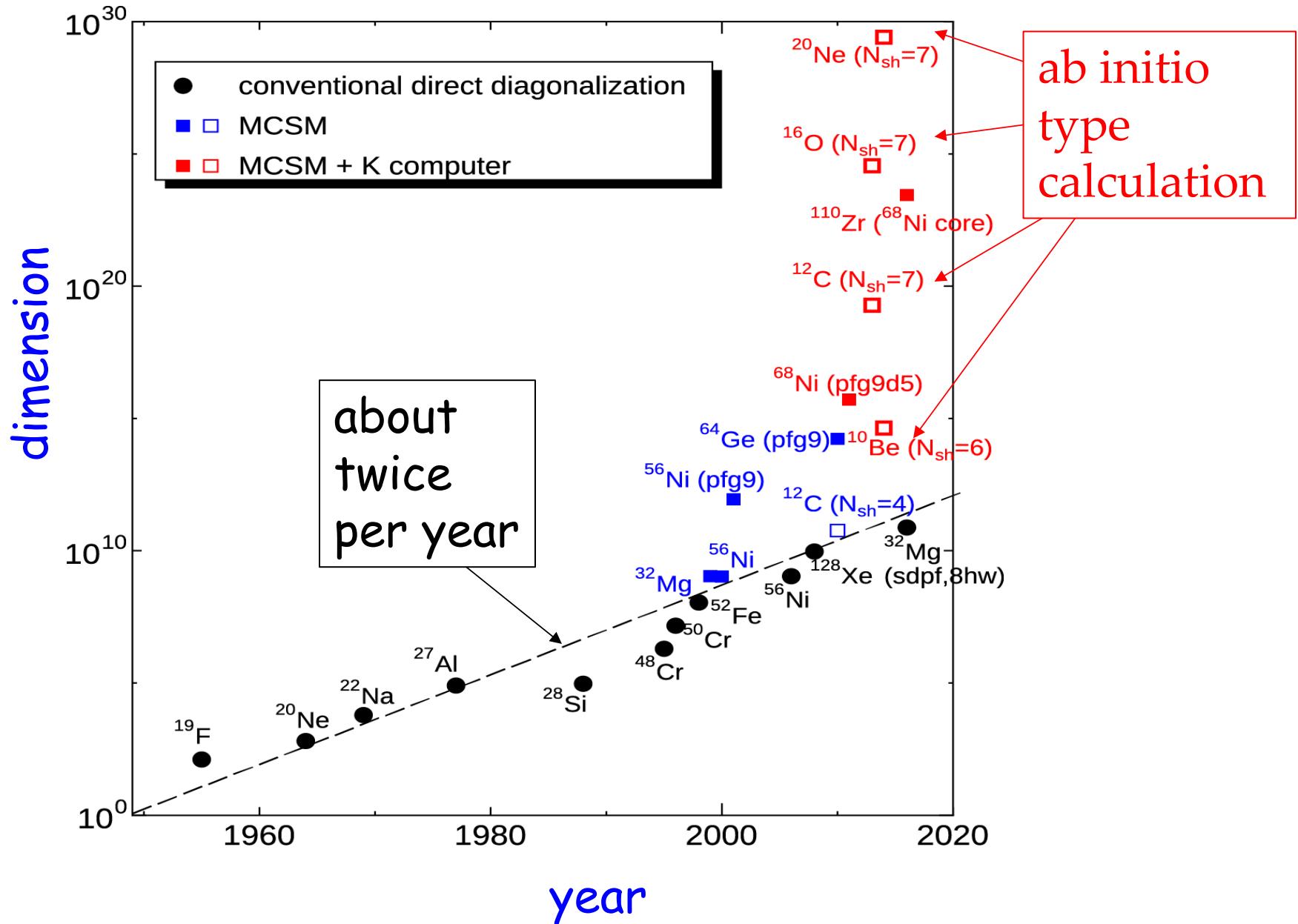
K computer (in Kobe) 10 peta flops machine

⇒ *Projection of basis vectors*

Rotation with three Euler angles
with about 50,000 mesh points

Example : $8^+ {}^{68}\text{Ni}$ 7680 core × 14 h

Dimension of the shell-model many-body Hilbert space



MCSM basis vectors on Potential Energy Surface

eigenstate

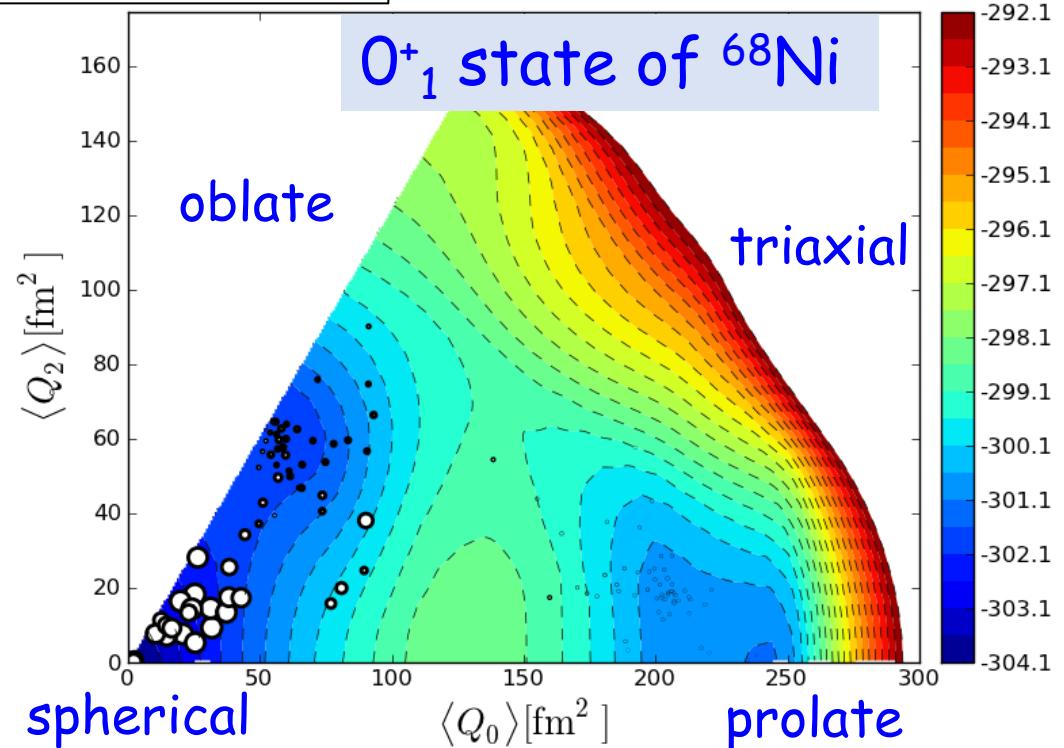
$$\Psi = \sum_i c_i P[J^\pi] \Phi_i$$

Slater determinant of deformed
s. p. states \rightarrow intrinsic shape

amplitude

projection onto J^π

- PES is calculated by CHF for the shell-model Hamiltonian
- Location of circle : quadrupole deformation of unprojected MCSM basis vectors
- Area of circle : overlap probability between each projected basis and eigen wave function



Called **T-plot** in reference to
Y. Tsunoda, *et al.*
PRC 89, 031301 (R) (2014)



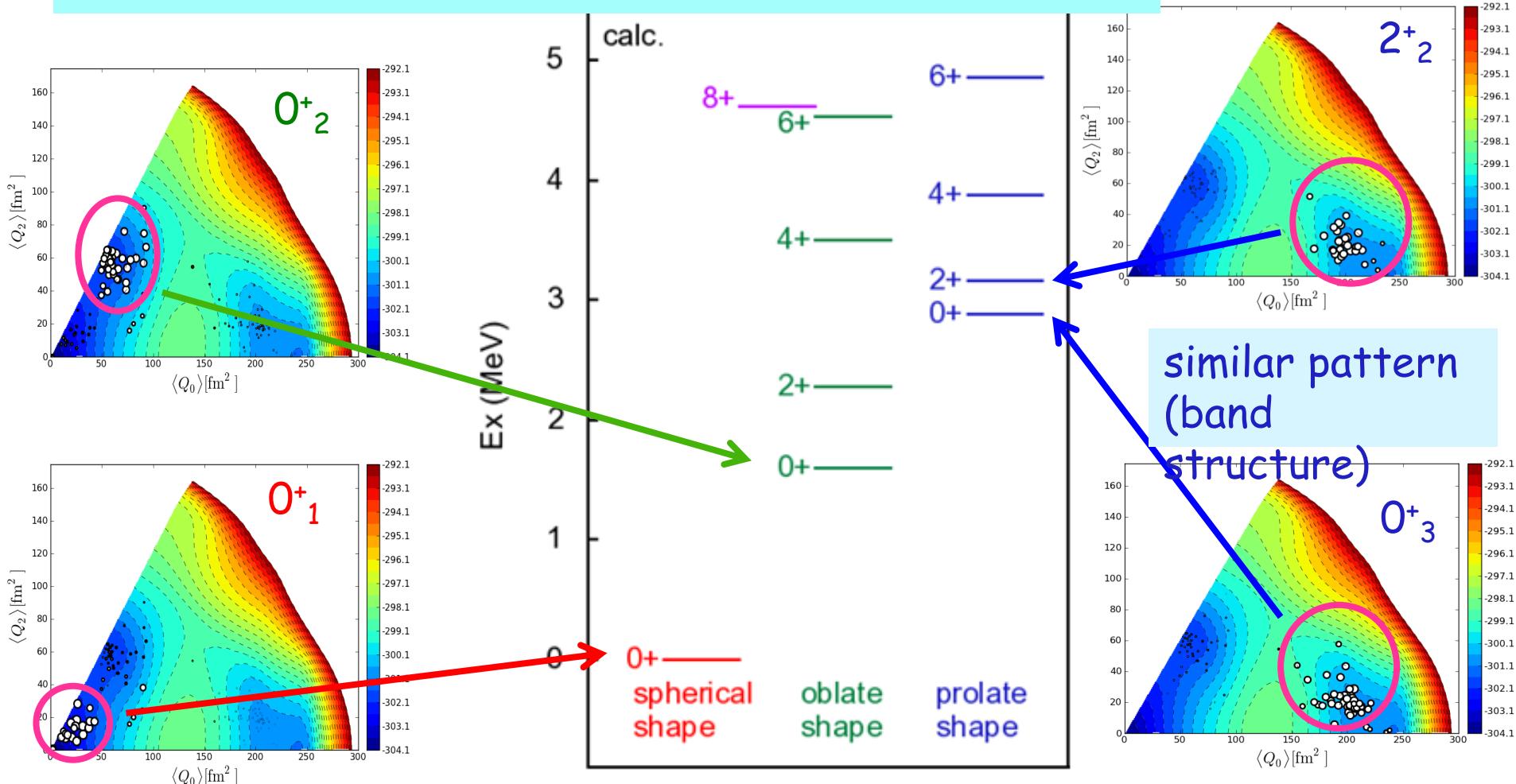
Y. Tsunoda

General properties of T-plot :

Certain number of large circles in a small region of PES
 \Leftrightarrow pairing correlations

Spreading beyond this can be due to shape fluctuation

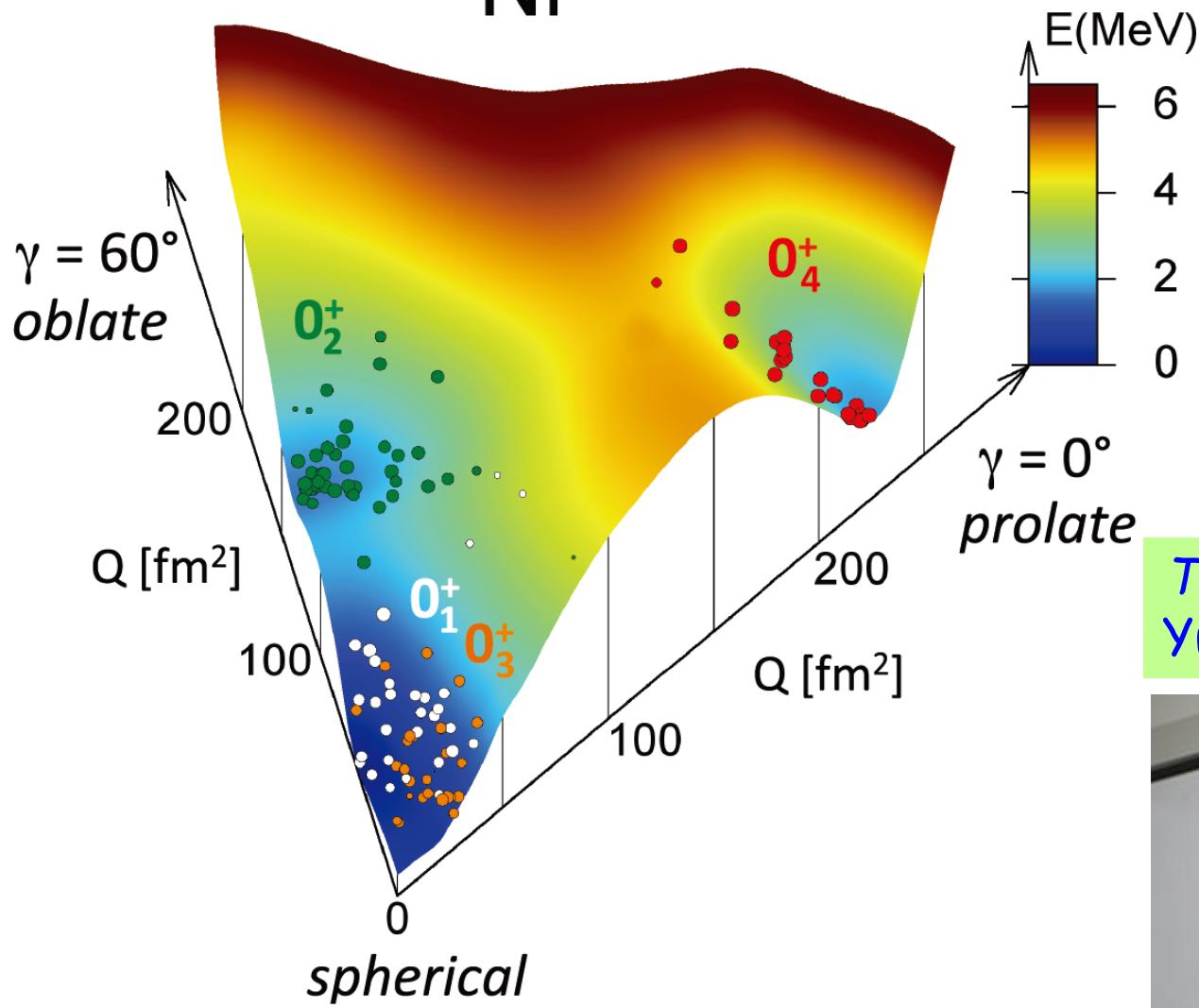
Example : shape assignment to various 0^+ states of ^{68}Ni



3-dimensional T-plot

~ Alpine villages

^{66}Ni



T-plot invented by
Yusuke Tsunoda



Three pillars combined for future

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pf
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(Sm)

...

island of stability

+
 χ EFT based
multi-shell int.

many-body dynamics

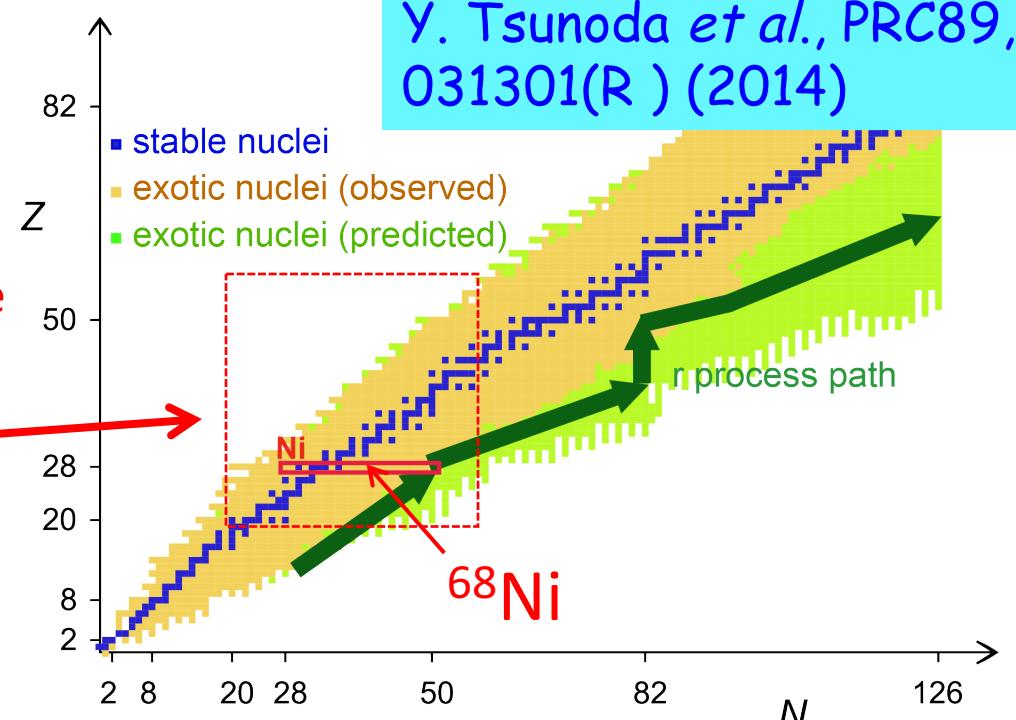
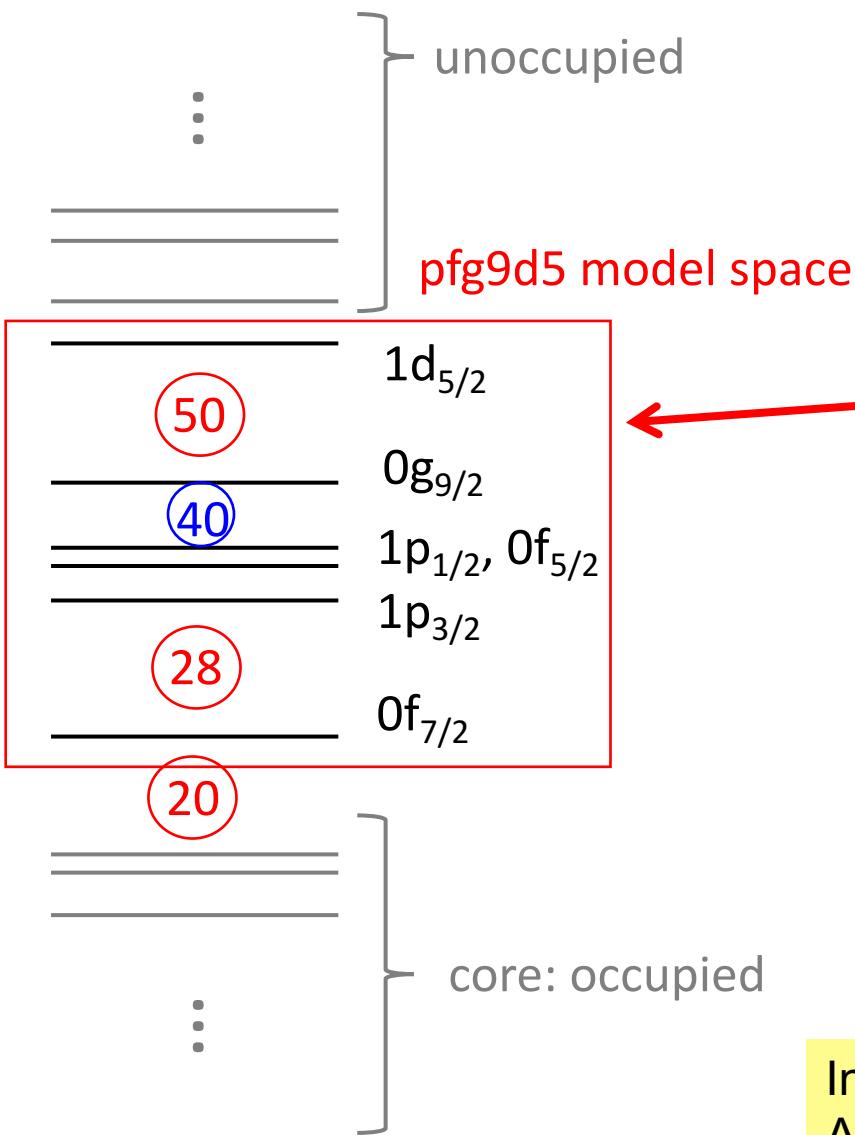
Shell evolution
(Type I & II)

Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Monte Carlo Shell Model (MCSM) calculation on Ni isotopes

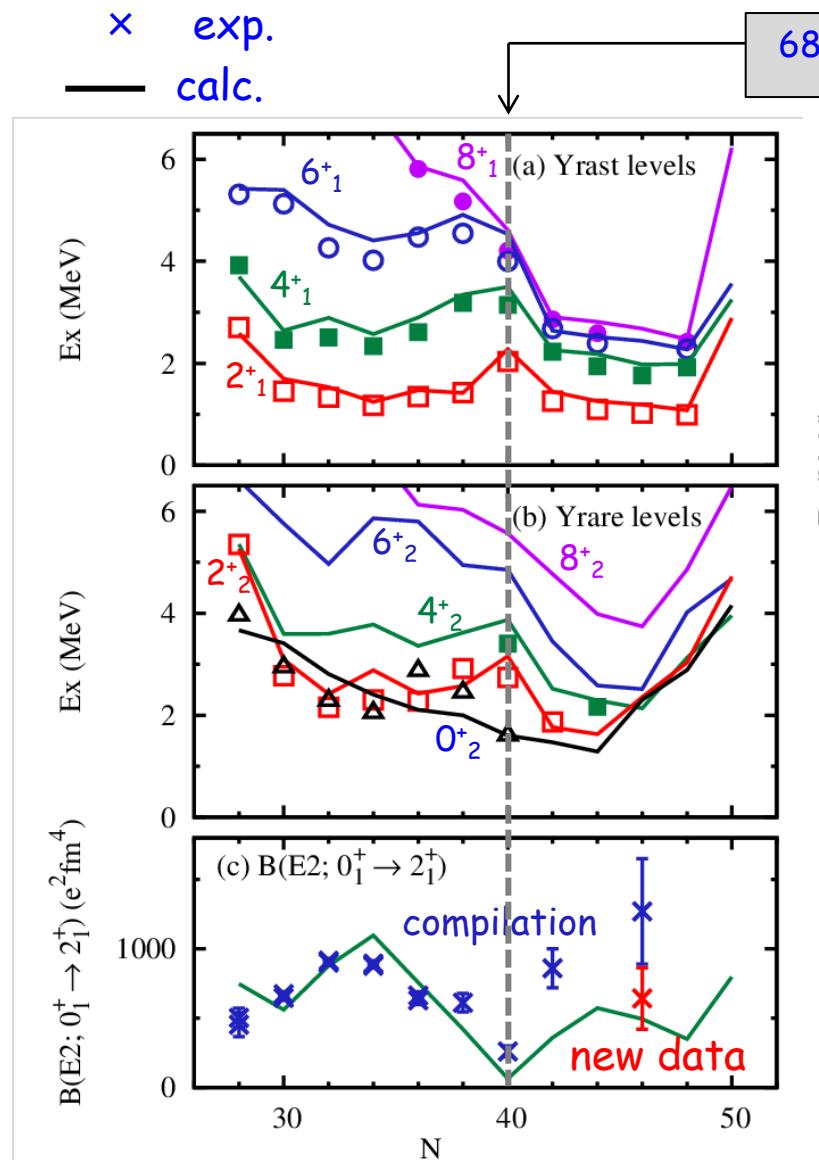


This model space is wide enough to discuss how **magic numbers 28, 50** and **semi-magic number 40** are visible or smeared out.

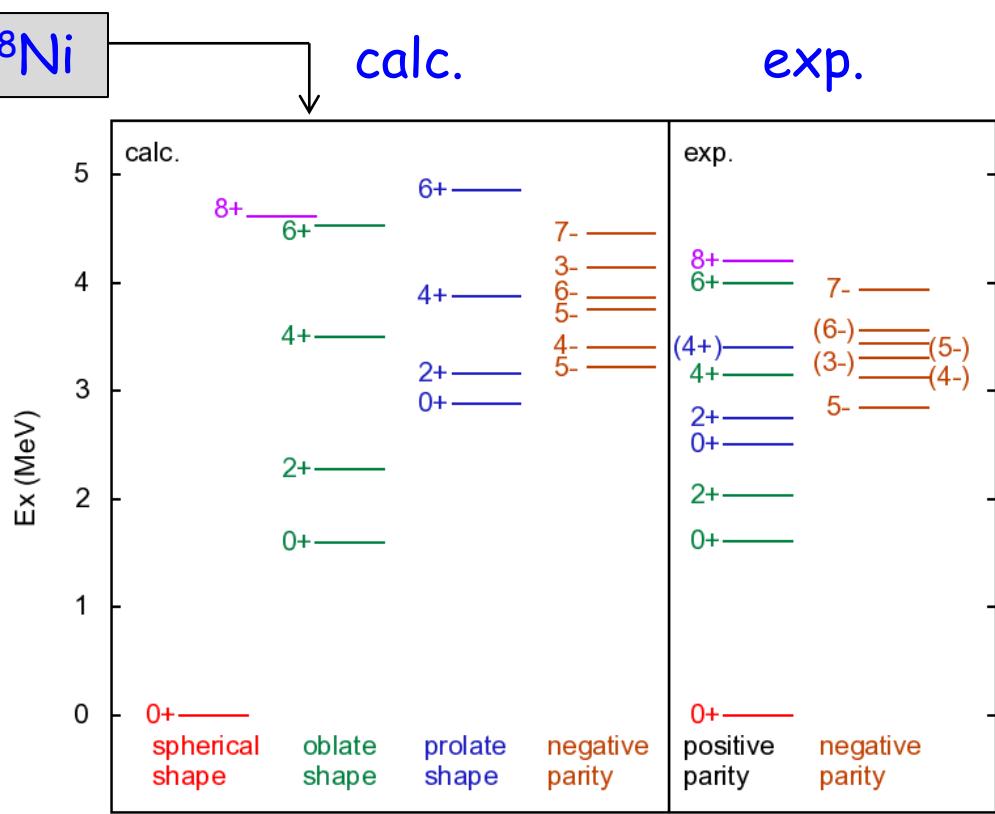
Interaction:
A3DA interaction is used with minor corrections

Energy levels and B(E2) values of Ni isotopes

Description by the same Hamiltonian



Shape coexistence in ^{68}Ni

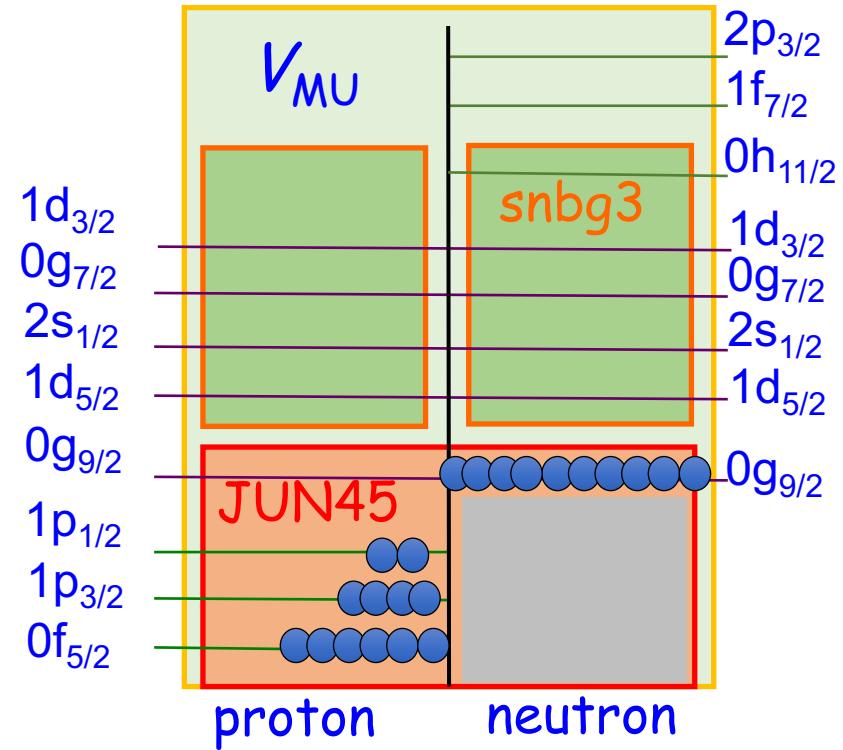


Y. Tsunoda, TO, Shimizu, Honma and Utsuno, PRC 89, 031301 (R) (2014)

An example : shapes of Zr isotopes by Monte Carlo Shell Model

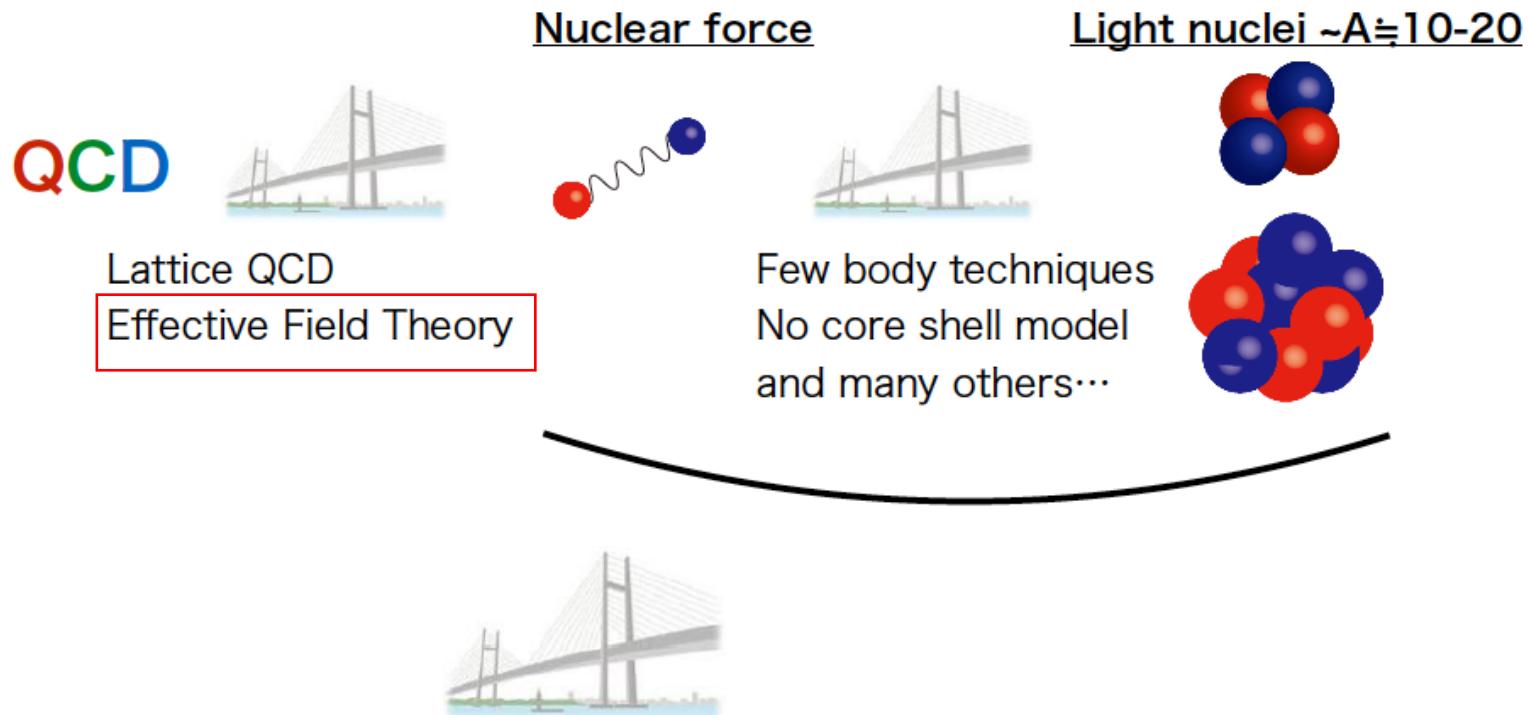
- Effective interaction:
 $JUN45 + snbg3 + V_{MU}$
known effective interactions
+ minor fit for a part of
T=1 TBME's
- Nucleons are excited fully
within this model space
(no truncation)

We performed Monte Carlo Shell Model (MCSM) calculations, where the largest case corresponds to the diagonalization of 3.7×10^{23} dimension matrix.

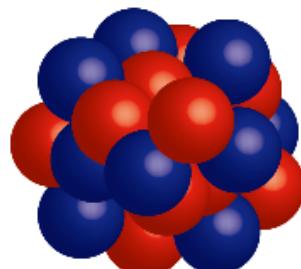


Togashi, Tsunoda, TO *et al.* PRL
117, 172502 (2016)

Part I : shell model powered by modern nuclear forces



Medium mass nuclei~A≈20-100



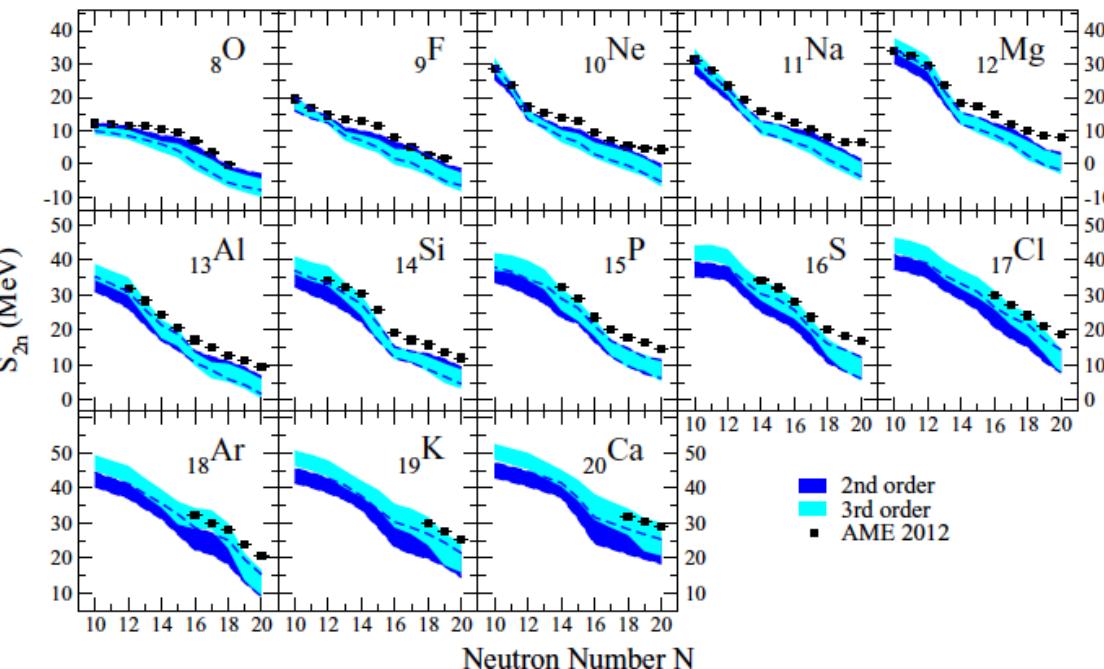
shell model with core
via the **effective interaction**
derived from nuclear force

Courtesy from N. Tsunoda

Examples of structure calculations starting from chiral EFT forces

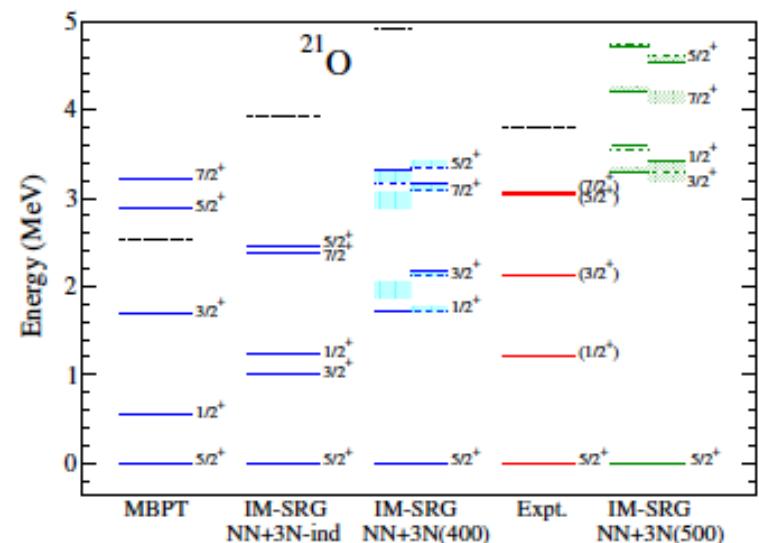
NN force: N3LO + 3N force: N2LO
 → valence shell interaction

MBPT
 (Many-body Perturbation Theory)
 applicable to one major shell



Simonis et al. PRC 93, 011302(R) (2016)

IM-SRG
 (In-Medium Similarity Renormalization Group)



Bogner et al. PRL 113, 142501 (2014)

+ Coupled Cluster calculation + N2LOsat potential + ...

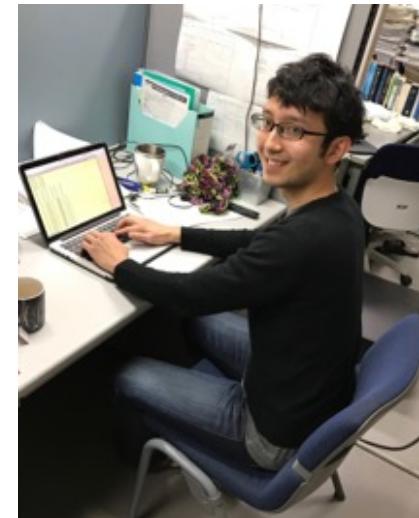
A recent development starting from *chiral EFT + 3NF*

EKK method to handle consistently
two (or more) major shells

-> Effective shell-model interaction

- (i) without fit of two-body m. e.,
- (ii) applicable to broken magicity,
or fusion of two shells,

both are crucial for exotic nuclei.



PHYSICAL REVIEW C 95, 021304(R) (2017)

Exotic neutron-rich medium-mass nuclei with realistic nuclear forces

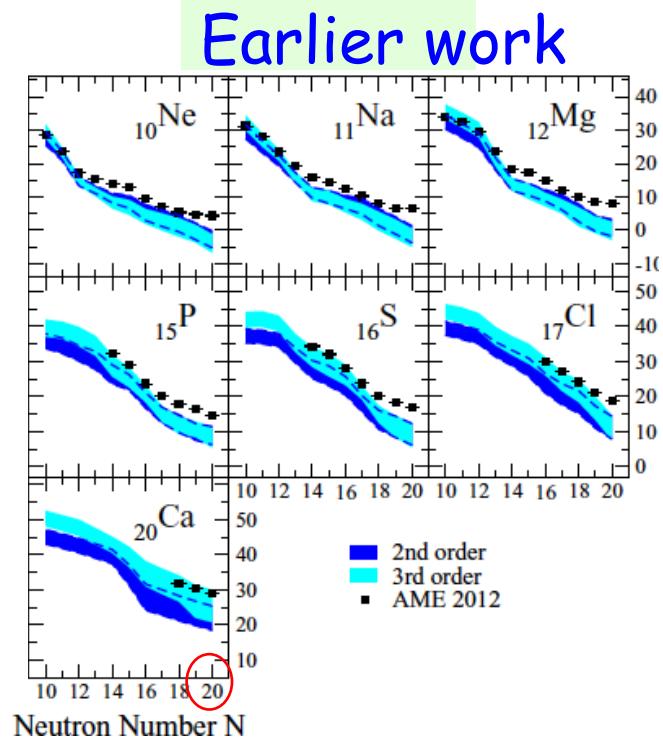
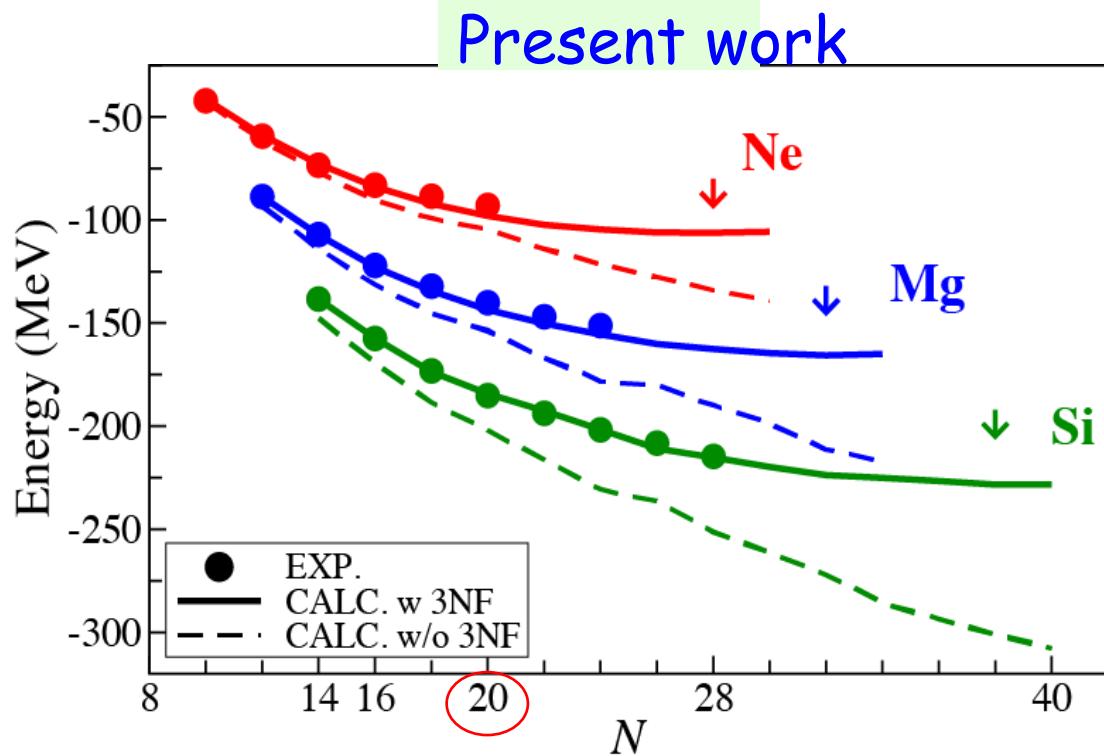
Naofumi Tsunoda,¹ Takaharu Otsuka,^{1,2,3,4} Noritaka Shimizu,¹ Morten Hjorth-Jensen,^{5,6}
Kazuo Takayanagi,⁷ and Toshio Suzuki⁸

* E. M. Krenciglowa and T. T. S. Kuo, Nucl. Phys. A 235, 171 (1974).

Re-visit to the "Island of Inversion" with ab initio TBMEs

Calculations with full $sd + pf$
shell

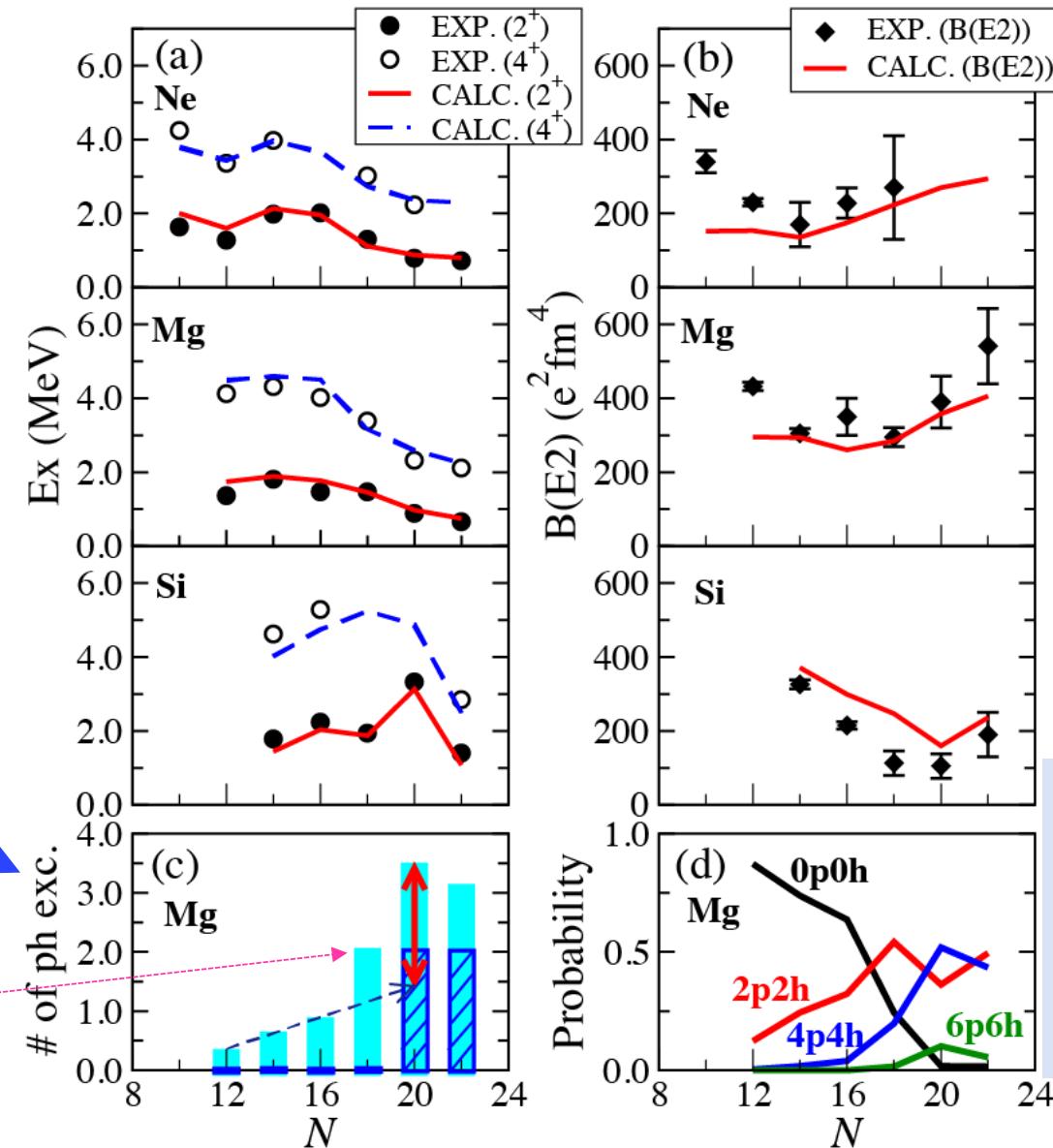
ground-state energies



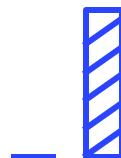
Ne-Mg-Si

2⁺ & 4⁺ levels
and B(E2)
values

of particle-hole
excitations across
N=20 gap :
(modest) steady
increase
+
abrupt increase
after N=18

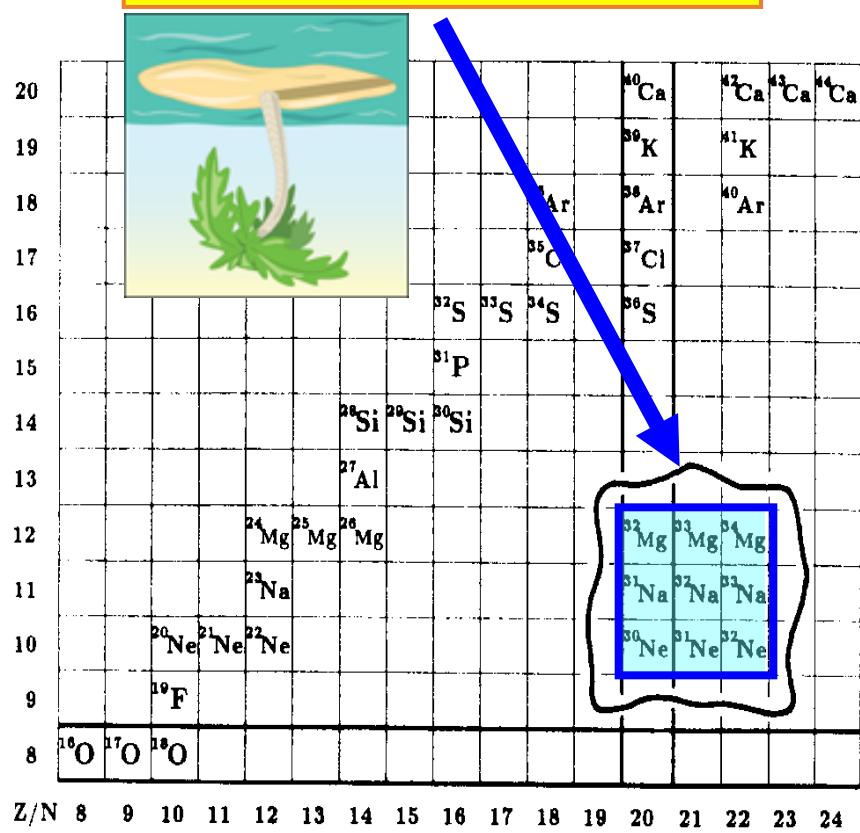


ground
states
of Mg
isotopes



Early idea of the Island of Inversion (WBB)
Op-Oh or 2p-2h (discrete)

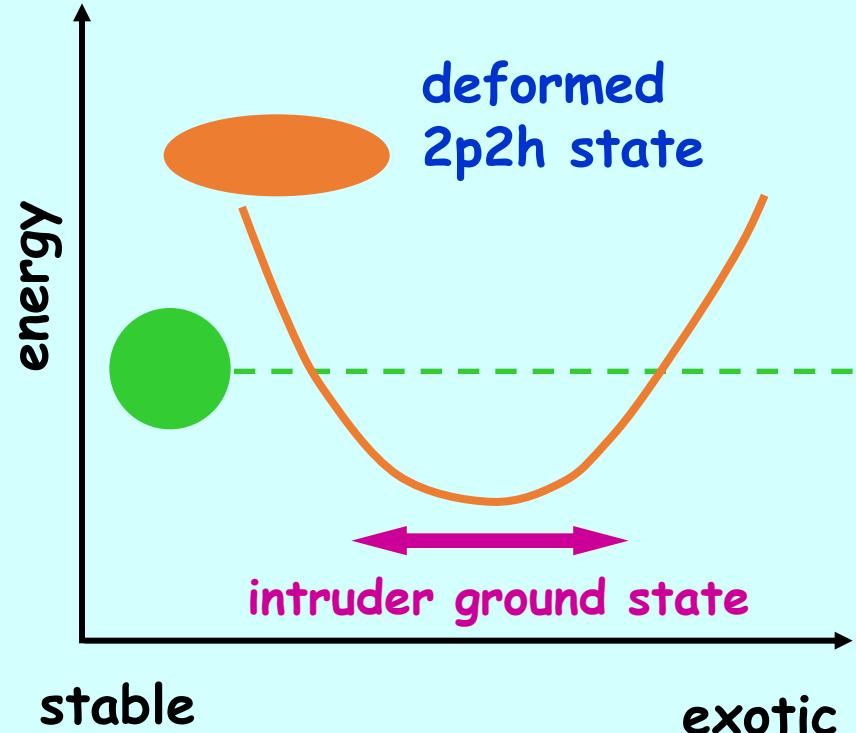
Island of Inversion



9 nuclei:
Ne, Na, Mg with N=20-22

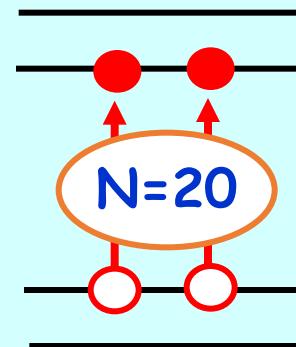
Phys. Rev. C 41, 1147 (1990),
Warburton, Becker and
Brown

Basic picture was



pf shell

sd shell



$^{31}_{12}\text{Mg}_{19}$: very difficult to fit by the shell model

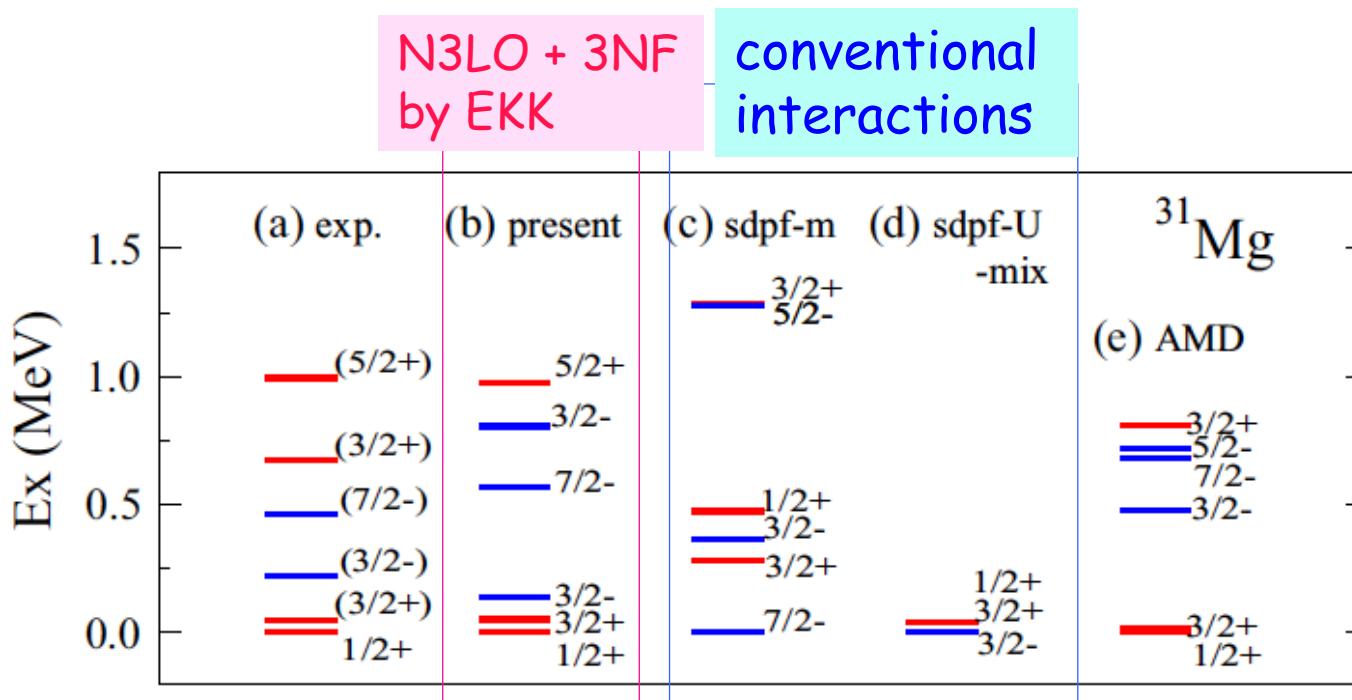


FIG. 4. Energy levels of ^{31}Mg . (a) experimental values, (b) present work, (c) sdpf-m [16], (d) sdpf-U-mix [17], and (e) AMD+GCM calculation [52].

exp. : Phys. Rev. Lett. 94, 022501 (2005), G. Neyens, et al.

Mixing between sd and pf shells is crucial

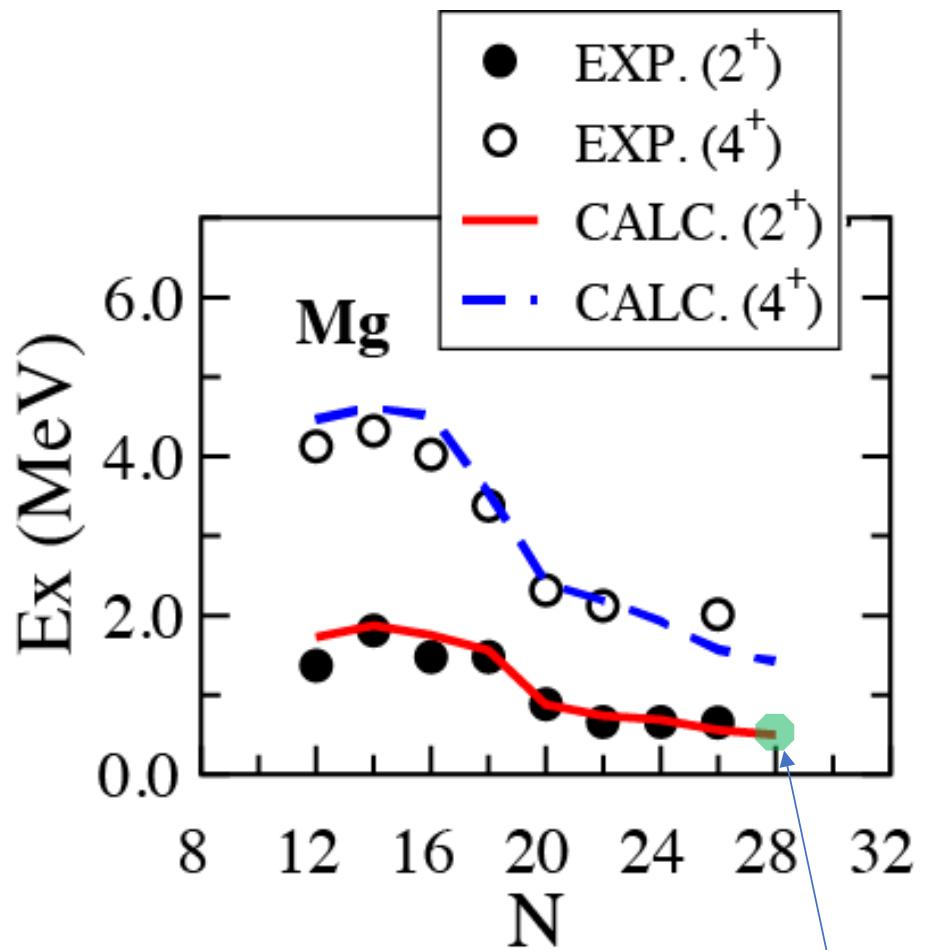
Fit of relevant TBME's is infeasible : too many TBME's but too few data

[16] Y. Utsuno, T. Otsuka, T. Mizusaki, and M. Honma, Phys. Rev. C **60**, 054315 (1999).

[17] E. Caurier, F. Nowacki, and A. Poves, Phys. Rev. C **90**, 014302 (2014).

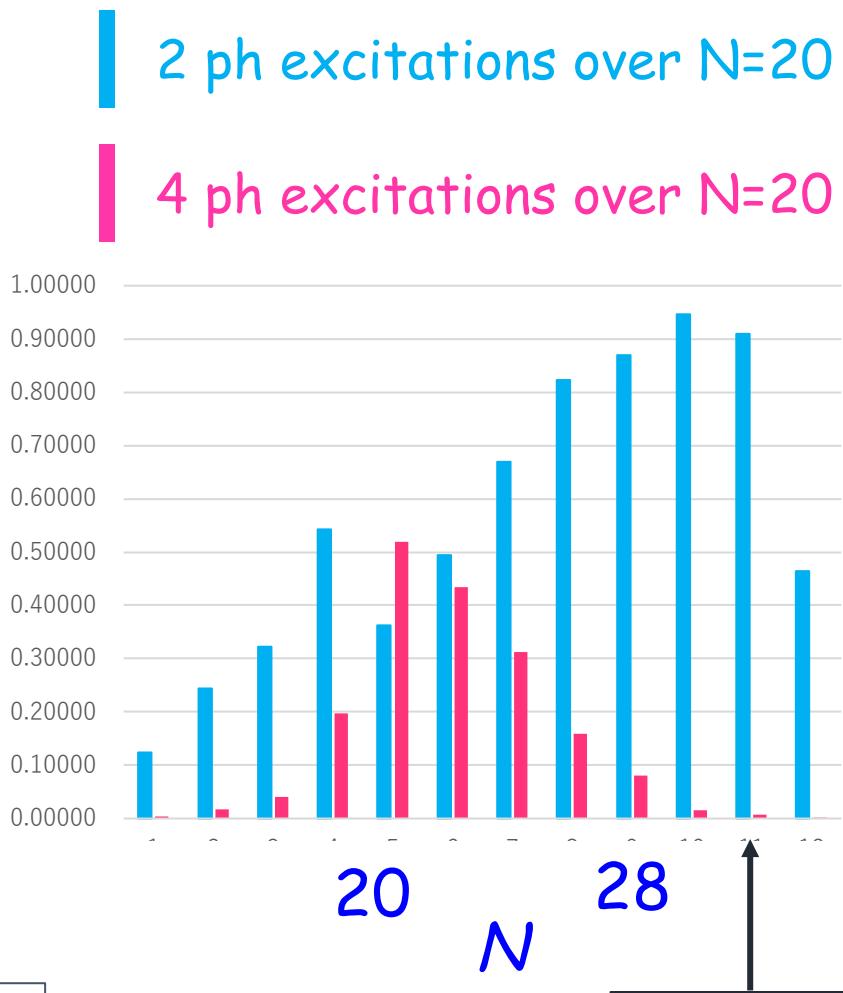
[52] M. Kimura, Phys. Rev. C **75**, 041302 (2007).

Mg up to N=28 2⁺ & 4⁺ levels



Recent data
Crawford *et al.*
(ARIS 2017 + priv. com.)

Probability of ph configurations over Z=N=20 (ground state)



our
drip line

We also found

χ EFT based effective interaction seems to produce more particle-hole excitations than conventional phenomenological effective interactions.

Three pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
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parallel
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Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
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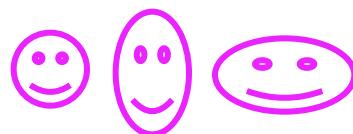
Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Difference between stable and exotic nuclei

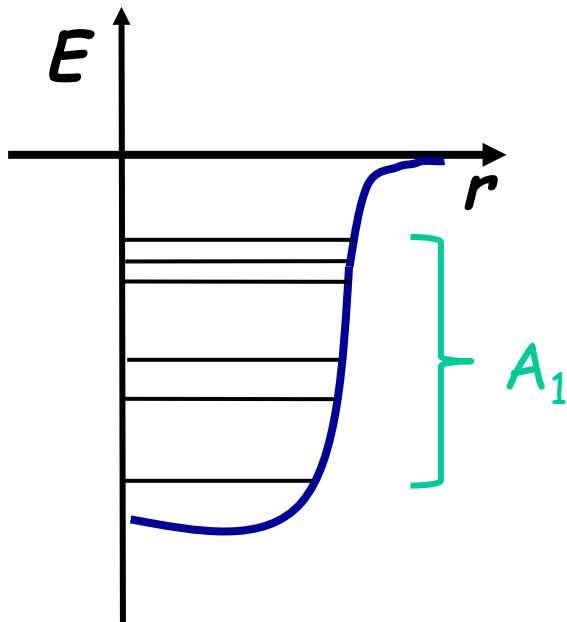
	stable nuclei	exotic nuclei
life time	infinite or long	short
number	~300	7000 ~ 10000
properties		
density	constant inside (density saturation)	low-density surface (halo, skin)
shell	same magic numbers (2,8,20,28, ... (1949))	?
shape	shape transition	



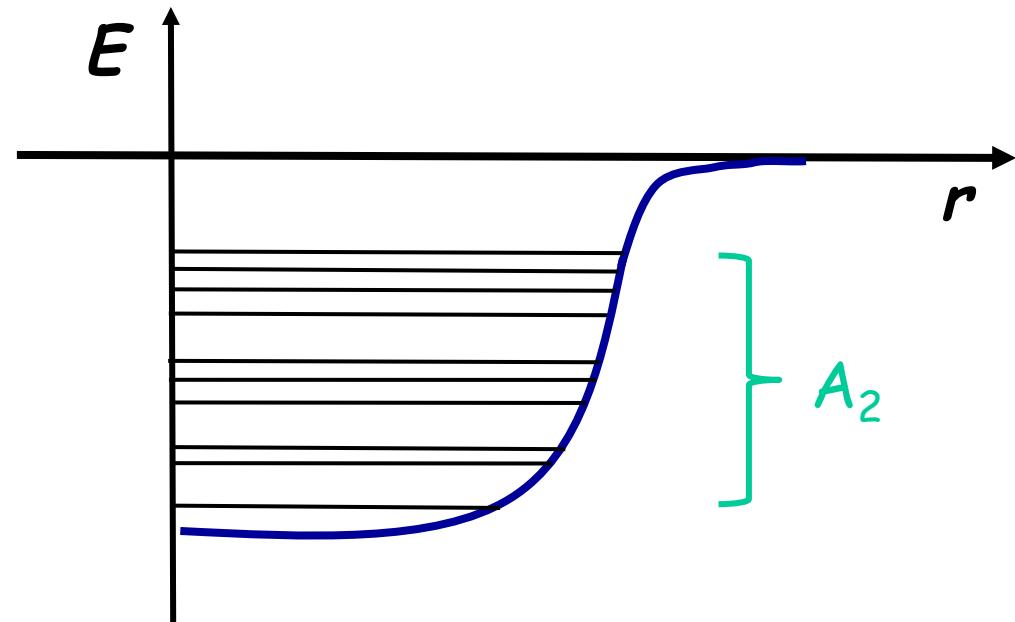
Single-particle states - starting point -

Mean potential becomes wider so as to cast A nucleons with the same separation energy.

light nuclei



heavy nuclei



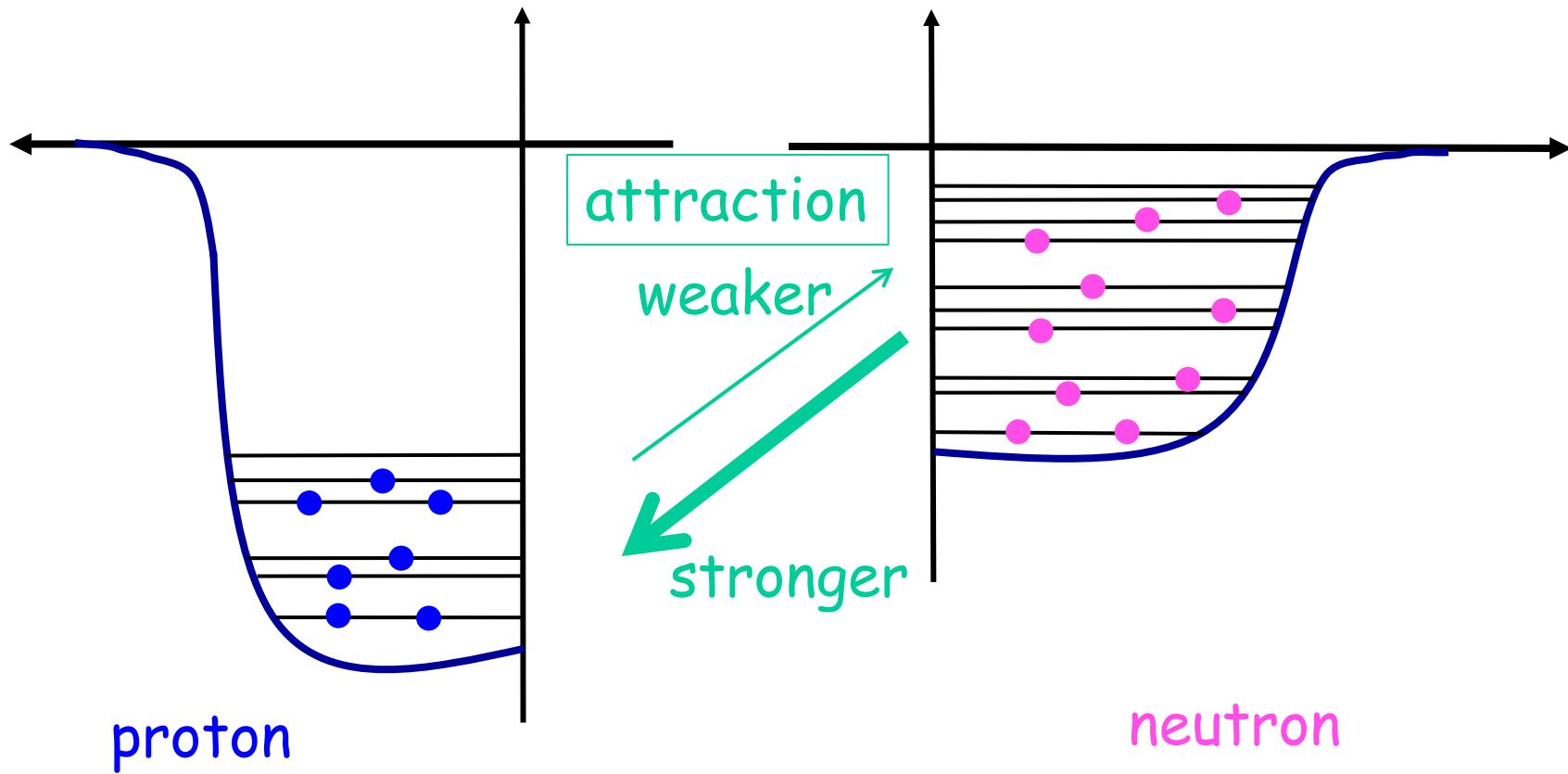
But, this is a story for **stable** nuclei.

proton-neutron interaction

>> proton-proton or neutron-neutron

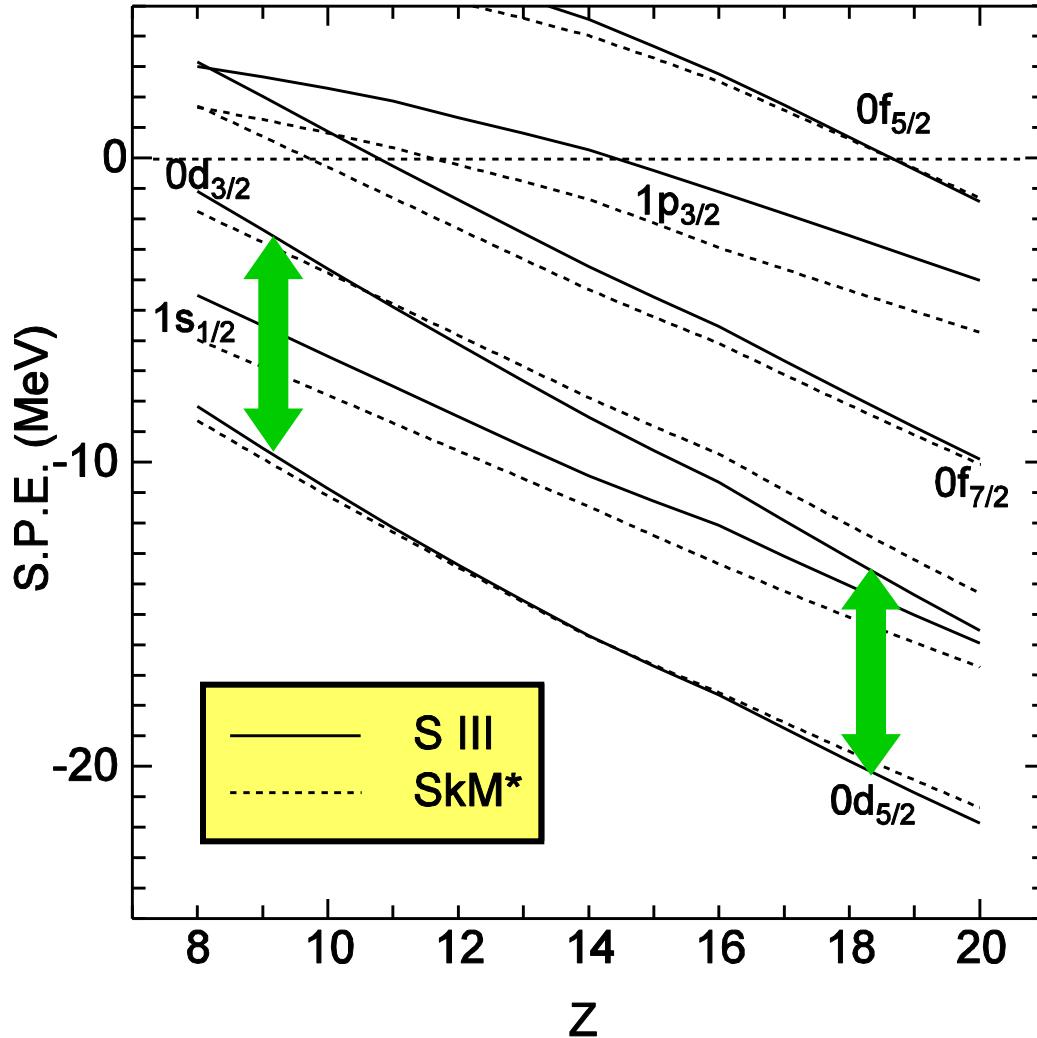
If $Z \ll N$, protons are more bound.

*Relative relations are preserved,
because only the depth changes.*



Realization in Hartree-Fock energies by Skyrme model

Neutron Single-Particle Energies at $N=20$



The shell structure remain rather unchanged

-- orbitals shifting together

-- change of potential depth

~ Woods-Saxon.

From simple but general properties as,

density saturation
+ short-range NN interaction
+ spin-orbit splitting

→ Mayer-Jensen's magic number
with rather constant gaps
(except for gradual A dependence)

robust feature -> no way out ???

This has been one of the major objectives of nuclear structure physics in recent years.

Let's see what occurs
in the shell structure
of exotic nuclei.

Monopole matrix element between orbits j and j'

$$V_{nn}^m(j, j') = \frac{\sum_{(m,m')} \langle j, m ; j', m' | \hat{v}_{nn} | j, m ; j', m' \rangle}{\sum_{(m,m')} 1},$$

v_{nn} is interaction; m, m' are magnetic substates

Monopole matrix element between orbits j and j'

$$\begin{aligned} & \langle \text{red circle up, blue circle up} | v | \text{red circle up, blue circle up} \rangle + \langle \text{red circle up, blue circle cross} | v | \text{red circle up, blue circle cross} \rangle + \langle \text{red circle up, blue circle empty} | v | \text{red circle up, blue circle empty} \rangle + \dots \\ & + \langle \text{red circle down, blue circle up} | v | \text{red circle down, blue circle up} \rangle + \dots \dots \dots \dots + \langle \text{red circle down, blue circle empty} | v | \text{red circle down, blue circle empty} \rangle \\ = & \frac{\text{number of matrix elements in the summation}}{\text{number of matrix elements in the summation}} \end{aligned}$$

 : magnetic substates of the orbit j

 : magnetic substates of the orbit j'

Monopole matrix elements can be written equivalently by usual TBMEs as

$$V_T^m(j, j') = \frac{\sum_J (2J+1) \langle j, j'; J, T | \hat{v} | j, j'; J, T \rangle}{\sum_J (2J+1)}.$$

for $T = 0$ and 1 ,

Monopole interaction for n-n (or p-p) interaction is defined as

$$\hat{v}_{nn}^m(j, j') = \begin{cases} V_{nn}^m(j, j) \frac{1}{2} \hat{n}_j (\hat{n}_j - 1) & \text{for } j = j' \\ V_{nn}^m(j, j') \hat{n}_j \hat{n}_{j'} & \text{for } j \neq j' \end{cases}$$

$T=1$ part of the p-n monopole interaction is given as

$$\begin{aligned}\hat{v}_{pn,mono,T=1} = & \sum_{j,j'} V_{T=1}^m(j, j') \frac{1}{2} \hat{n}_j^p \hat{n}_{j'}^n \\ & + \sum_{j < j'} V_{T=1}^m(j, j') \frac{1}{2} \left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\} \\ & + \sum_j V_{T=1}^m(j, j) \frac{1}{2} : \hat{\tau}_j^+ \hat{\tau}_j^- : .\end{aligned}$$

Finally, we get the full expression for the p-n interaction

$$\begin{aligned}\hat{v}_{pn,mono} = & \sum_{j,j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') + V_{T=1}^m(j, j') \right\} \boxed{\hat{n}_j^p \hat{n}_{j'}^n} \\ & - \sum_{j < j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') - V_{T=1}^m(j, j') \right\} \\ & \quad \left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\} \\ & - \sum_j \frac{1}{2} \left\{ V_{T=0}^m(j, j) - V_{T=1}^m(j, j) \right\} : \hat{\tau}_j^+ \hat{\tau}_j^- : .\end{aligned}$$

This is consistent with the final form of the monopole interaction of Poves and Zuker (Phys. Rep. 70, 235 (1981)) besides different formulation.

$$\hat{v}_{pn,mono} = \sum_{j,j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') + V_{T=1}^m(j, j') \right\} \hat{n}_j^p \hat{n}_{j'}^n$$

$$- \sum_{j < j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') - V_{T=1}^m(j, j') \right\}$$

$$\left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\}$$

$$- \sum_j \frac{1}{2} \left\{ V_{T=0}^m(j, j) - V_{T=1}^m(j, j) \right\} : \hat{\tau}_j^+ \hat{\tau}_j^- :$$

The second and third terms can be visualized as shown in the figure.

They are still monopole, but proton and neutron must exchange their states !

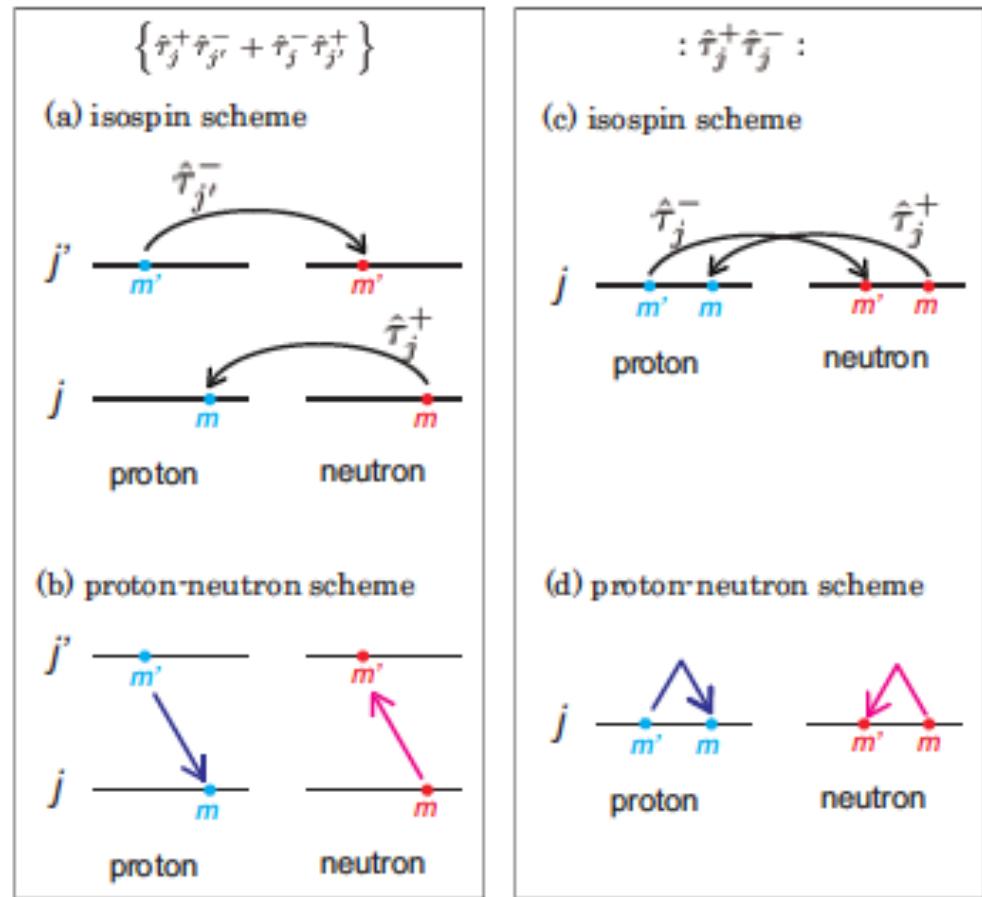


Figure 16 Implication of $\hat{\tau}_j^+ \hat{\tau}_{j'}^-$ terms. Panels (a) and (c) are for the $\{\hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+\}$ and $:\hat{\tau}_j^+ \hat{\tau}_j^-:$ cases in the isospin scheme, respectively. Panels (b) and (d) are similar to (a) and (c), respectively, in the proton-neutron scheme. The magnetic substates are indicated by m and m' .

Effective single-particle energy (ESPE)

$$\epsilon_j = \epsilon_j^0 + \hat{\epsilon}_j .$$

Contribution from valence nucleons through the monopole interaction

contribution from the inert core (closed shell) ; constant within a given nucleus

ESPEs are given for protons and neutrons as

$$\hat{\epsilon}_j^p = \sum_{j'} V_{T=1}^m(j, j') \hat{n}_{j'}^p + \sum_{j'} [V_{pn}^m(j, j')] \hat{n}_{j'}^n$$

$$\hat{\epsilon}_j^n = \sum_{j'} V_{T=1}^m(j, j') \hat{n}_{j'}^n + \sum_{j'} [V_{pn}^m(j', j)] \hat{n}_{j'}^p .$$

where $[V_{pn}^m(j, j')] = \frac{1}{2} \{ \tilde{V}_{T=0}^m(j, j') + \tilde{V}_{T=1}^m(j, j') \}$ and

$$[\tilde{V}_{T=0,1}^m(j, j')] = V_{T=0,1}^m(j, j') \quad \text{for } j \neq j' ,$$

$$[\tilde{V}_{T=0}^m(j, j)] = V_{T=0}^m(j, j) \frac{2j+2}{2j+1} ,$$

and

$$[\tilde{V}_{T=1}^m(j, j)] = V_{T=1}^m(j, j) \frac{2j}{2j+1} .$$

with approximation $- : \hat{\tau}_j^+ \hat{\tau}_j^- : \sim \frac{\hat{n}_j^p \hat{n}_j^n}{2j+1}$

Usually ESPEs are discussed for their differences.

The differences of ESPEs are given
for protons

$$\Delta\hat{\epsilon}_j^p = \sum_{j'} V_{T=1}^m(j, j') \Delta\hat{n}_{j'}^p + \sum_{j'} V_{pn}^m(j, j') \Delta\hat{n}_{j'}^n ,$$

for neutrons

$$\Delta\hat{\epsilon}_j^n = \sum_{j'} V_{T=1}^m(j, j') \Delta\hat{n}_{j'}^n + \sum_{j'} V_{pn}^m(j', j) \Delta\hat{n}_{j'}^p$$

For a chain of isotopes, $\Delta\hat{n}_{j'}^n$ is the change of the neutron number.

Shell evolution

Monopole matrix element of the central force with a gaussian dependence on the distance.

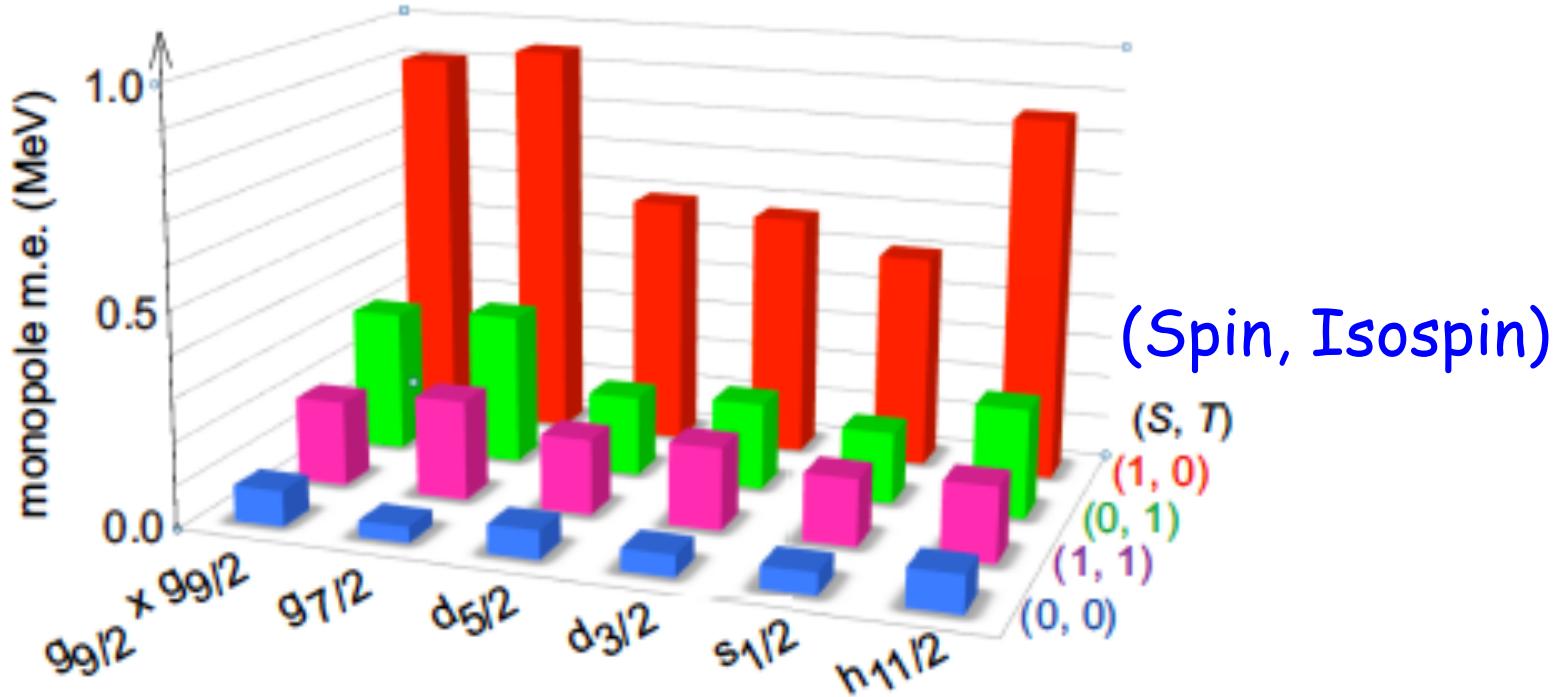


Figure 24 Monopole matrix elements of central gaussian interactions for (S, T) channels with an equal strength parameter (see the text). One of the orbit is $1g_{9/2}$, and the other is shown.

Monopole interaction from a central force : $A=100$

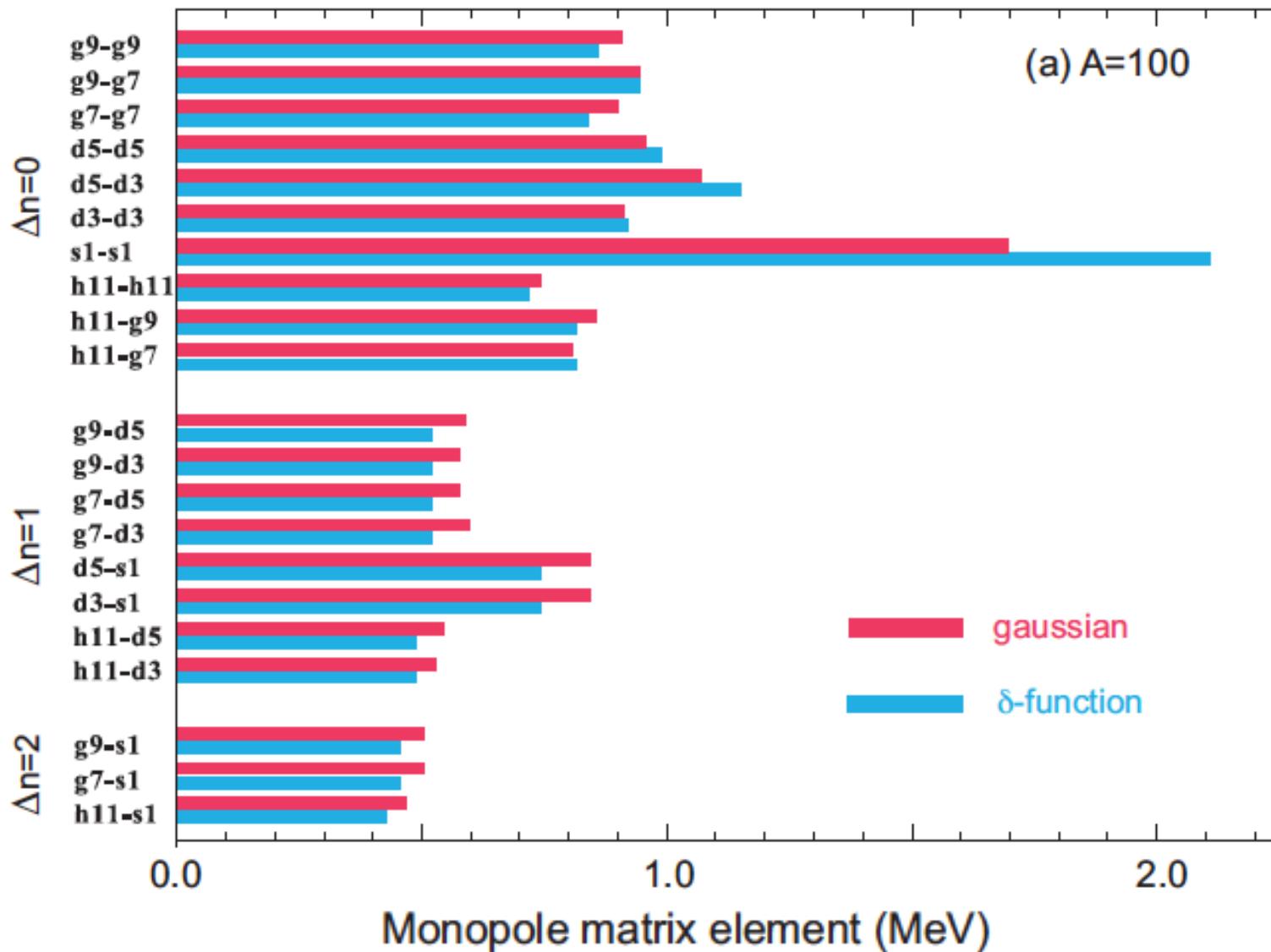


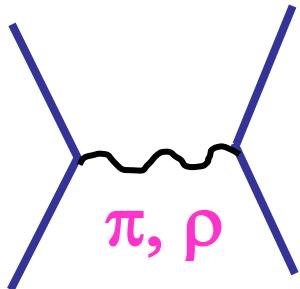
Figure 26 Monopole matrix elements of central gaussian and delta interactions for ($S = 1, T = 0$) channel. The orbit labeling is abbreviated like g9 for $1g_{9/2}$, etc. The orbits are from valence shell for (a) $A = 100$ and (b) $A = 70$.

Besides central corce,
another important contribution
comes from the tensor force

Tensor Force

π meson : primary source

ρ meson ($\sim \pi + \pi$) : minor ($\sim 1/4$) cancellation



Ref: Osterfeld, Rev. Mod. Phys. 64, 491 (92)

Multiple pion exchanges

→ strong effective central forces in NN interaction

(as represented by σ meson, etc.)

→ nuclear binding

Presently : First-order tensor-force effect
(at medium and long ranges)

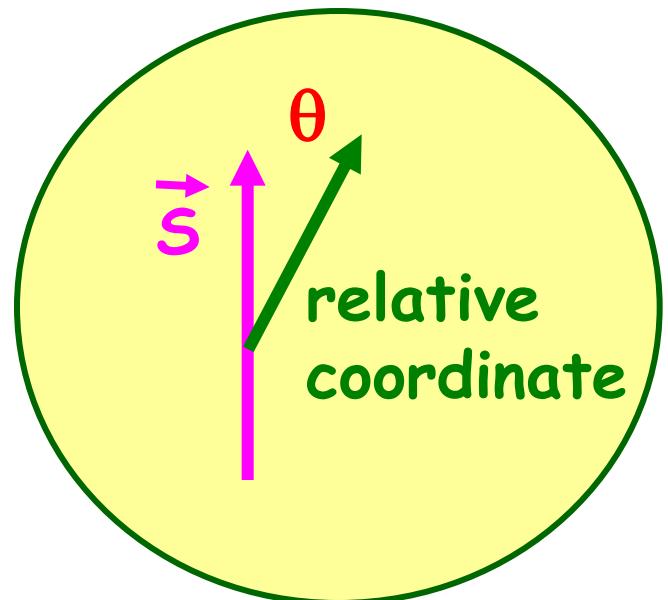
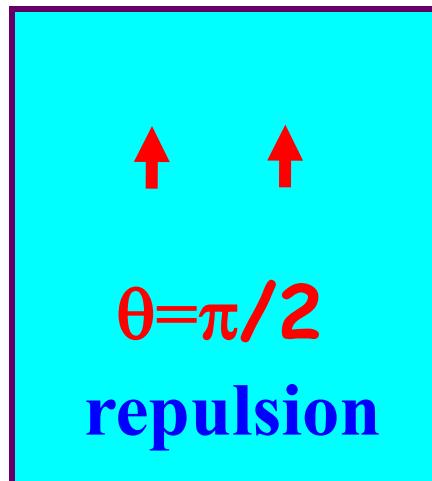
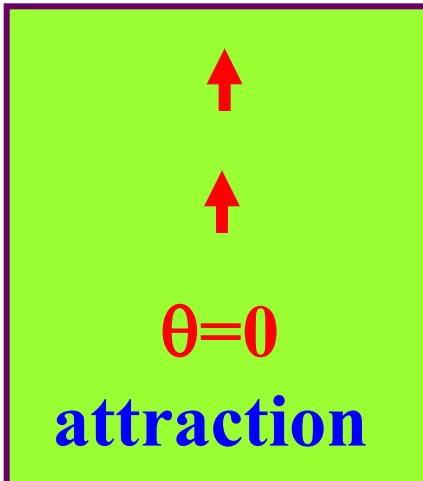
One pion exchange → Tensor force

How does the tensor force work ?

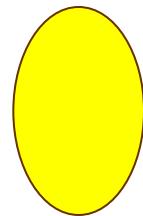
Spin of each nucleon \uparrow is parallel, because the total spin must be $S=1$

The potential has the following dependence \rightarrow on the angle θ with respect to the total spin S .

$$V \sim Y_{2,0} \sim 1 - 3 \cos^2 \theta$$



Monopole effects due to the tensor force



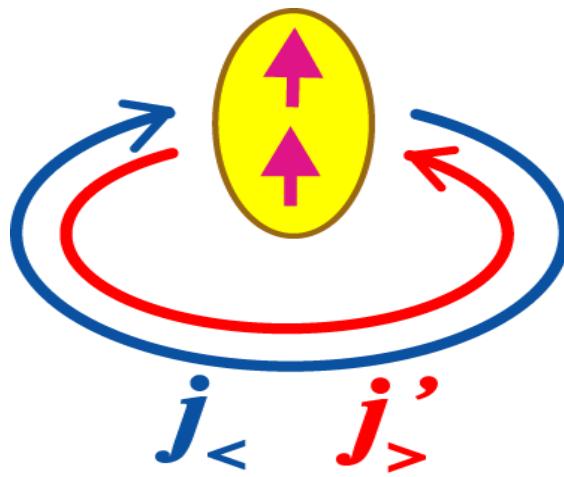
wave function of relative motion



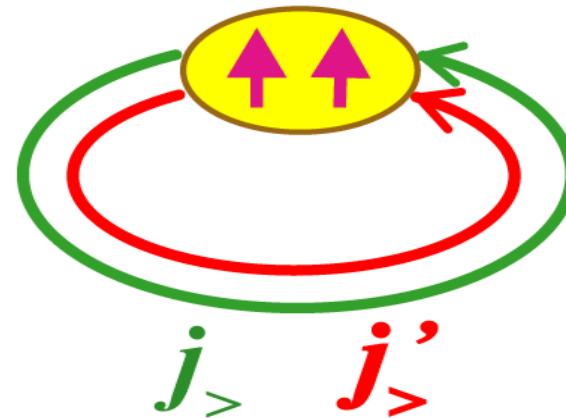
spin of nucleon

large relative momentum

small relative momentum



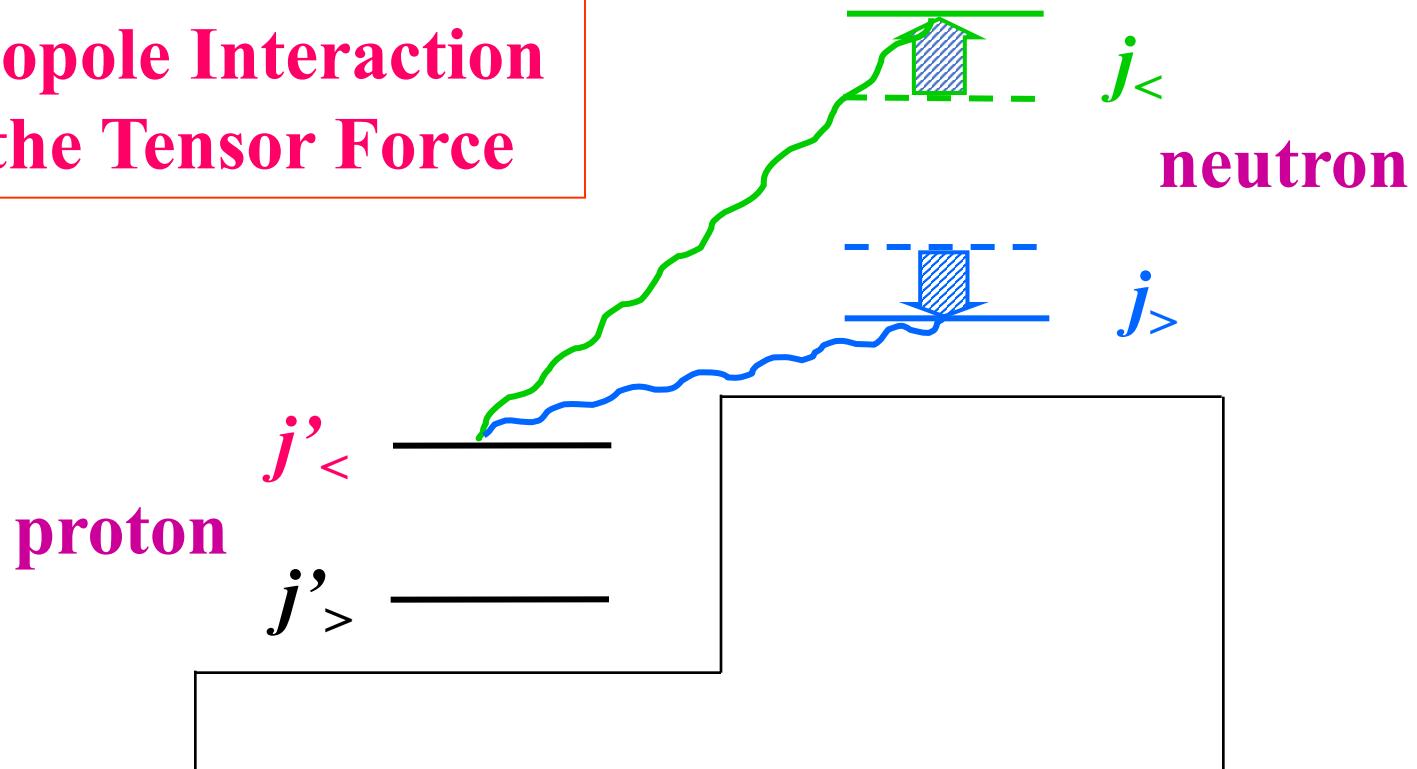
attractive



repulsive

$$j_> = l + \frac{1}{2}, \quad j_< = l - \frac{1}{2}$$

Monopole Interaction of the Tensor Force



Identity for tensor monopole interaction

$$(2j_>+1) \ v_{m,T}^{(j'j_>)} + (2j_<+1) \ v_{m,T}^{(j'j_<)} = 0$$

$v_{m,T}$: monopole strength for isospin T

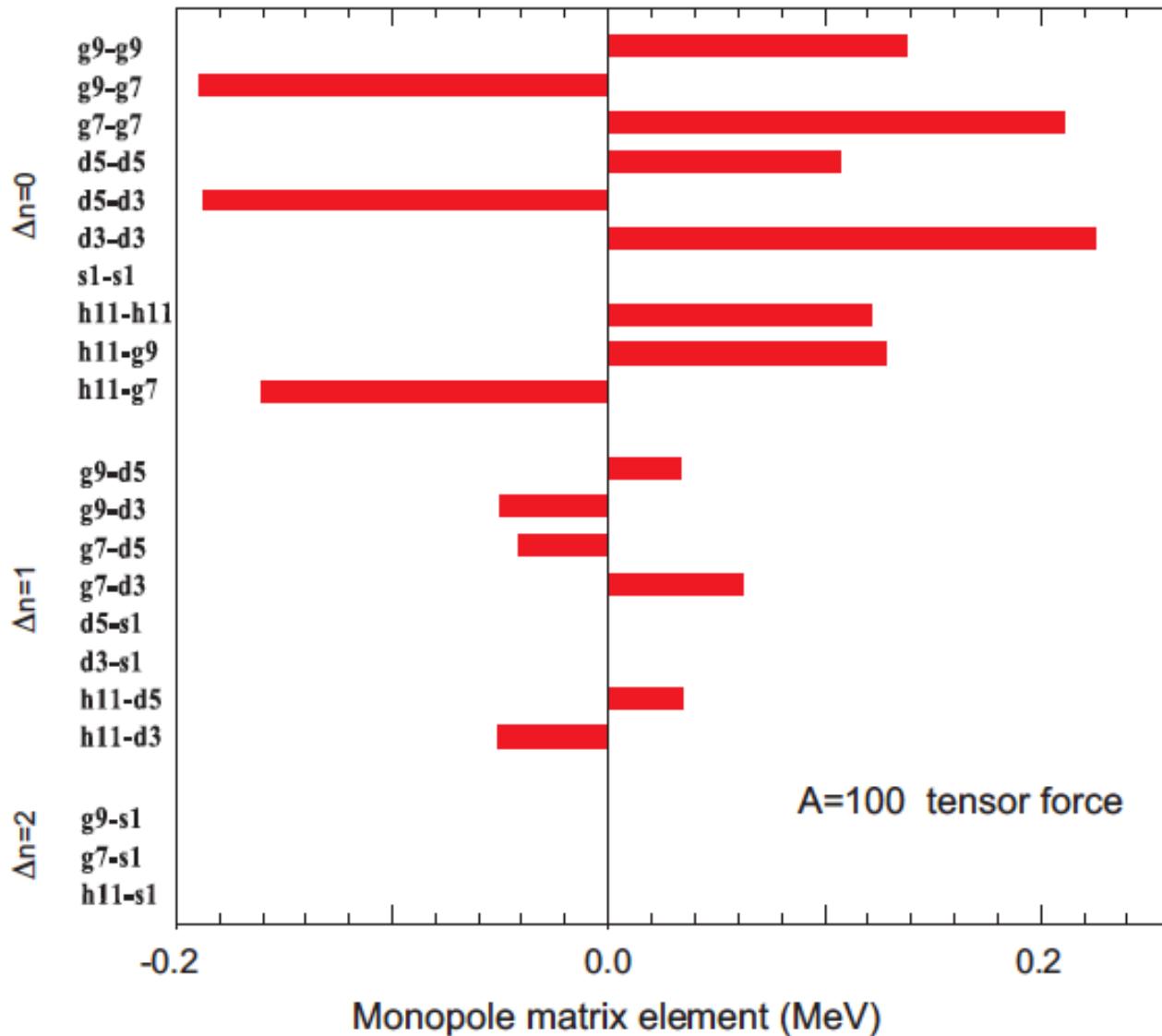
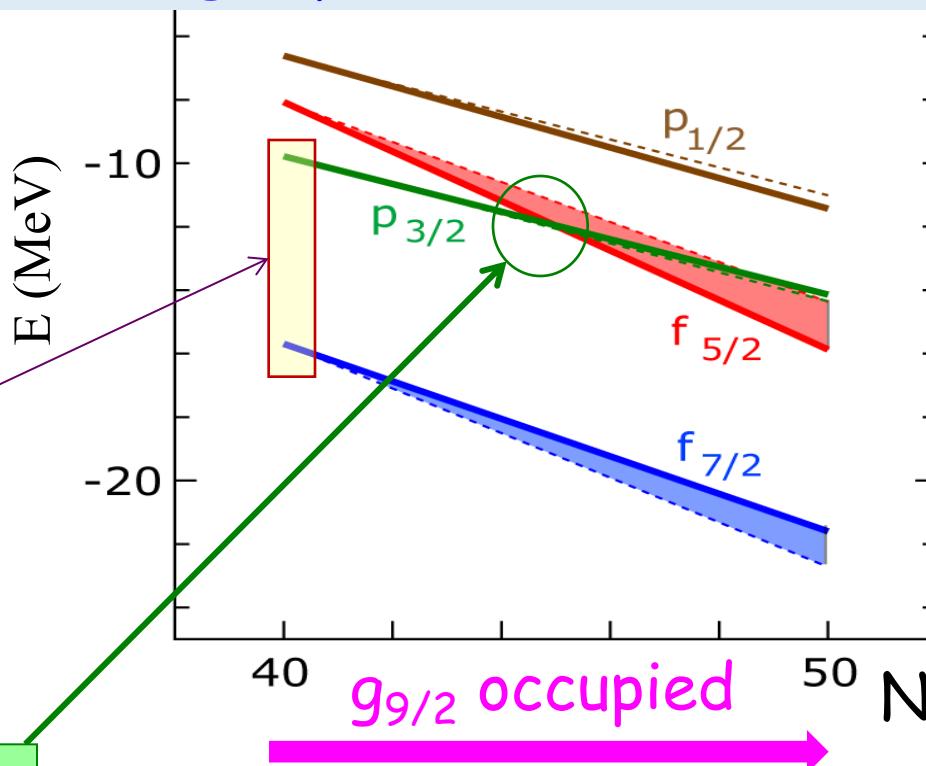


Figure 33 Monopole matrix elements of the tensor force in the $T=0$ channel. The orbit labeling is abbreviated like g₉ for $1g_{9/2}$, etc. The orbits are from valence shell for $A = 100$.

Proton single-particles levels of Ni isotopes

From
Grave,
EPJA25,
357



Central Gaussian
+ Tensor

solid line:
full effect

dotted line:
central only

shaded area :
effect of
tensor force

PRL 103, 142501 (2009)

PHYSICAL REVIEW LETTERS

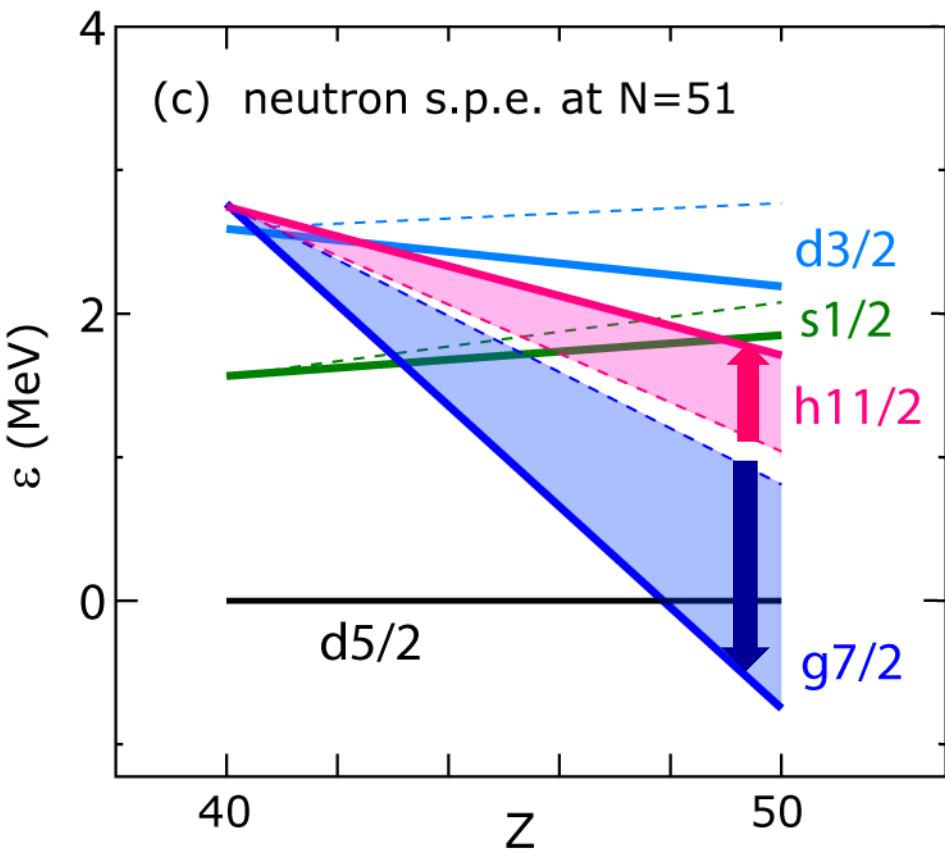
week ending
2 OCTOBER 2009

Nuclear Spins and Magnetic Moments of $^{71,73,75}\text{Cu}$: Inversion of $\pi 2p_{3/2}$ and $\pi 1f_{5/2}$ Levels in ^{75}Cu

K. T. Flanagan,^{1,2} P. Vingerhoets,¹ M. Avgoulea,¹ J. Billowes,³ M. L. Bissell,¹ K. Blaum,⁴ B. Cheal,³ M. De Rydt,¹ V. N. Fedosseev,⁵ D. H. Forest,⁶ Ch. Geppert,^{7,8} U. Köster,¹⁰ M. Kowalska,¹¹ J. Krämer,⁹ K. L. Kratz,⁹ A. Krieger,⁹ E. Mané,³ B. A. Marsh,⁵ T. Matema,¹⁰ L. Mathieu,¹² P. L. Molkanov,¹³ R. Neugart,⁹ G. Neyens,¹ W. Nörtershäuser,^{9,7} M. D. Seliverstov,^{13,16} O. Serot,¹² M. Schug,⁴ M. A. Sjoedin,¹⁷ J. R. Stone,^{14,15} N. J. Stone,^{14,15} H. H. Stroke,¹⁸ G. Tungate,⁶ D. T. Yordanov,⁴ and Yu. M. Volkov¹³

¹Instituut voor Kern- en Stralingsfysica, Katholieke Universiteit Leuven, B-3001 Leuven, Belgium

Shell structure of a key nucleus ^{100}Sn



solid line : full
(central + tensor)

dashed line : central only
Fedderman-Pittel (1977)

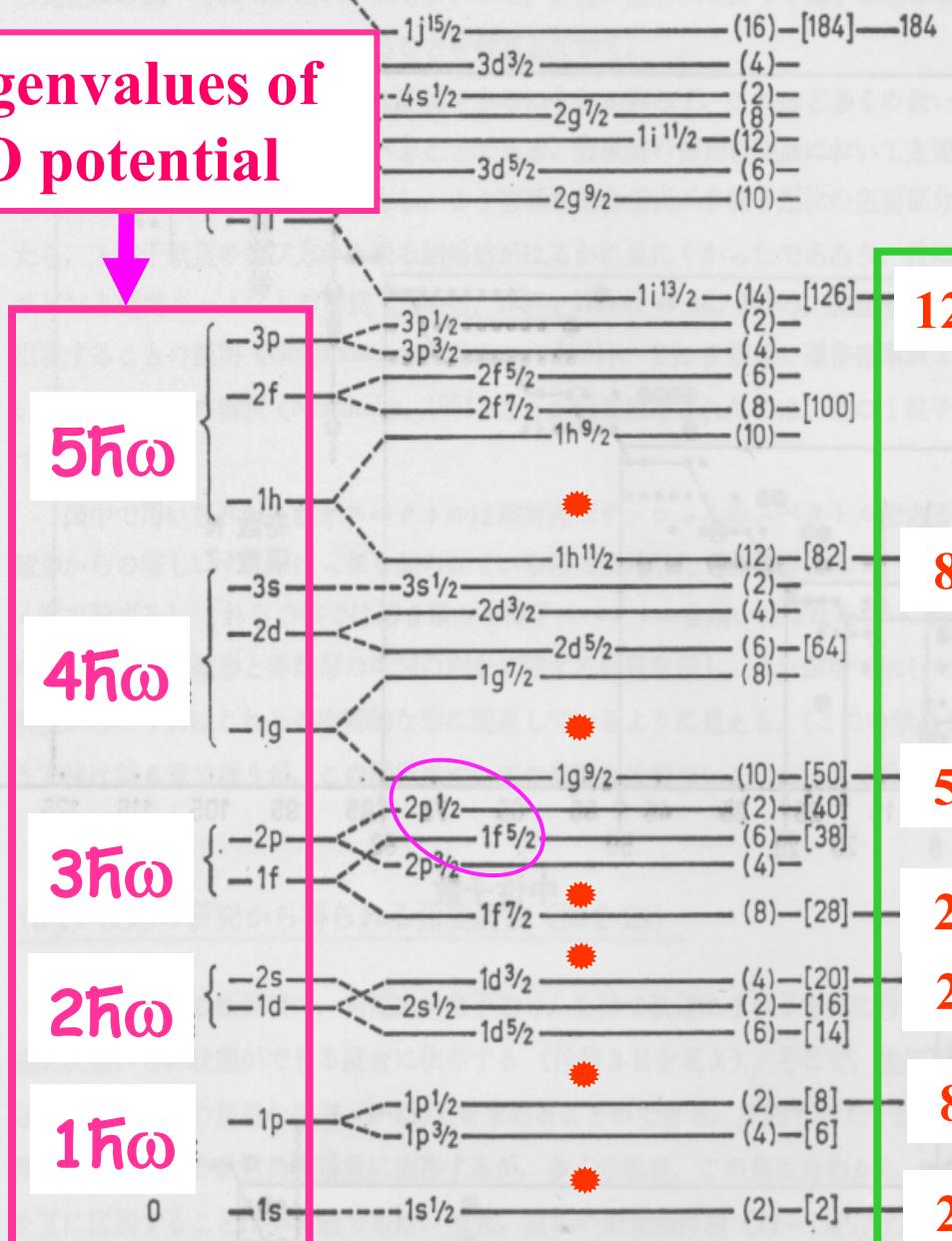
shaded area :
effect of tensor force

Exp. $d5/2$ and $g7/2$ should be close
Seweryniak et al.

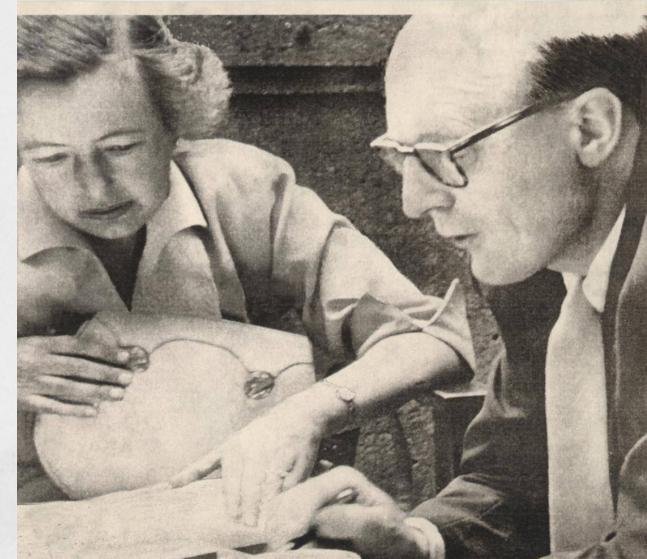
Phys. Rev. Lett. 99, 022504 (2007)
Gryzwacz et al.

New magic number ?

Eigenvalues of HO potential



Magic numbers
Mayer and
Jensen (1949)



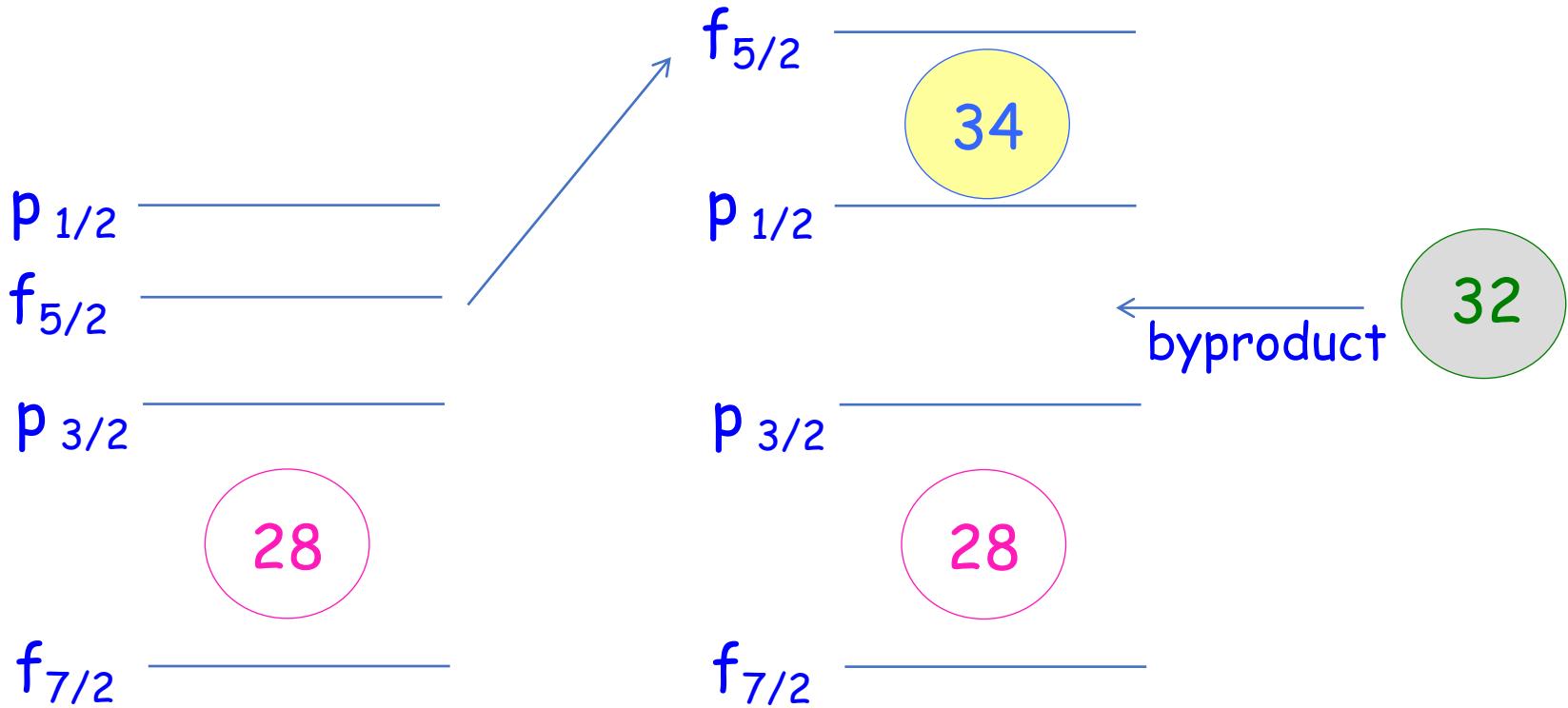
R SHELL MODEL
AND THE NUCLEUS

図 2-23 1粒子軌道の順序。図は M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure*, p. 58, Wiley, New York, 1955 からとった。

Basic picture

shell structure
for neutrons
in Ni isotopes
($f_{7/2}$ fully occupied)

N=34 magic number may appear
if proton $f_{7/2}$ becomes vacant (Ca)
($f_{5/2}$ becomes less bound)



Appearance of new magic number $N=34$

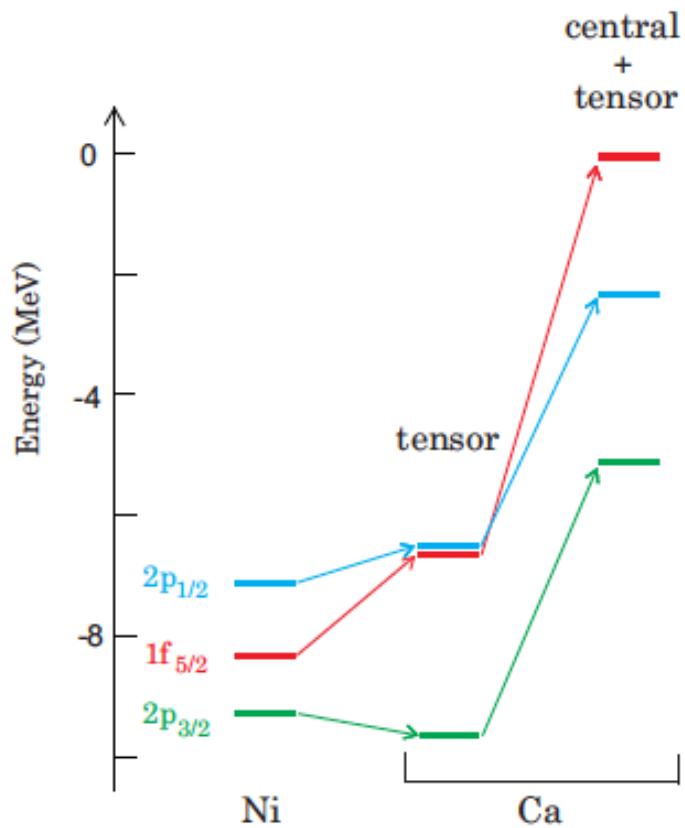


Figure 41 Change of single-particle energies from ^{56}Ni to ^{48}Ca .

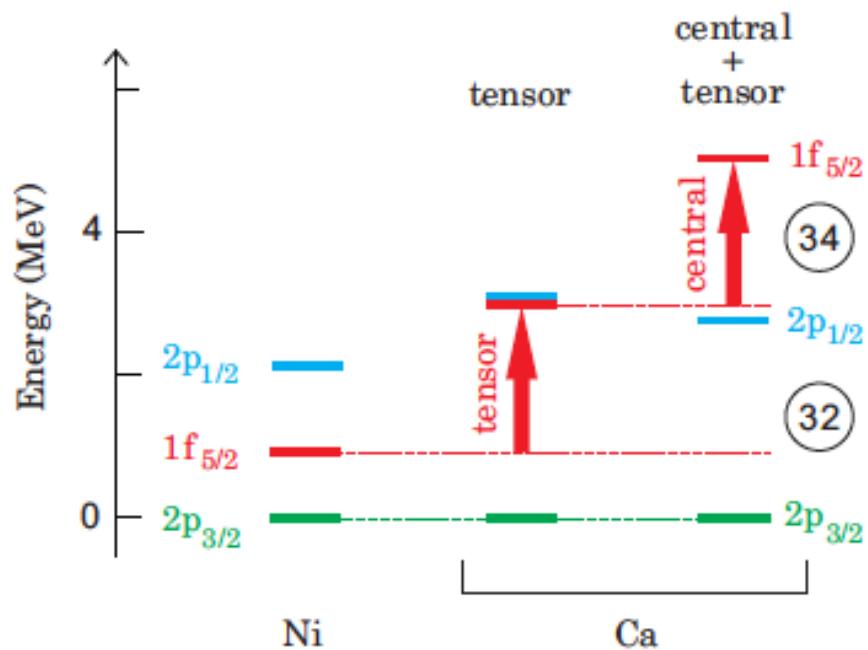
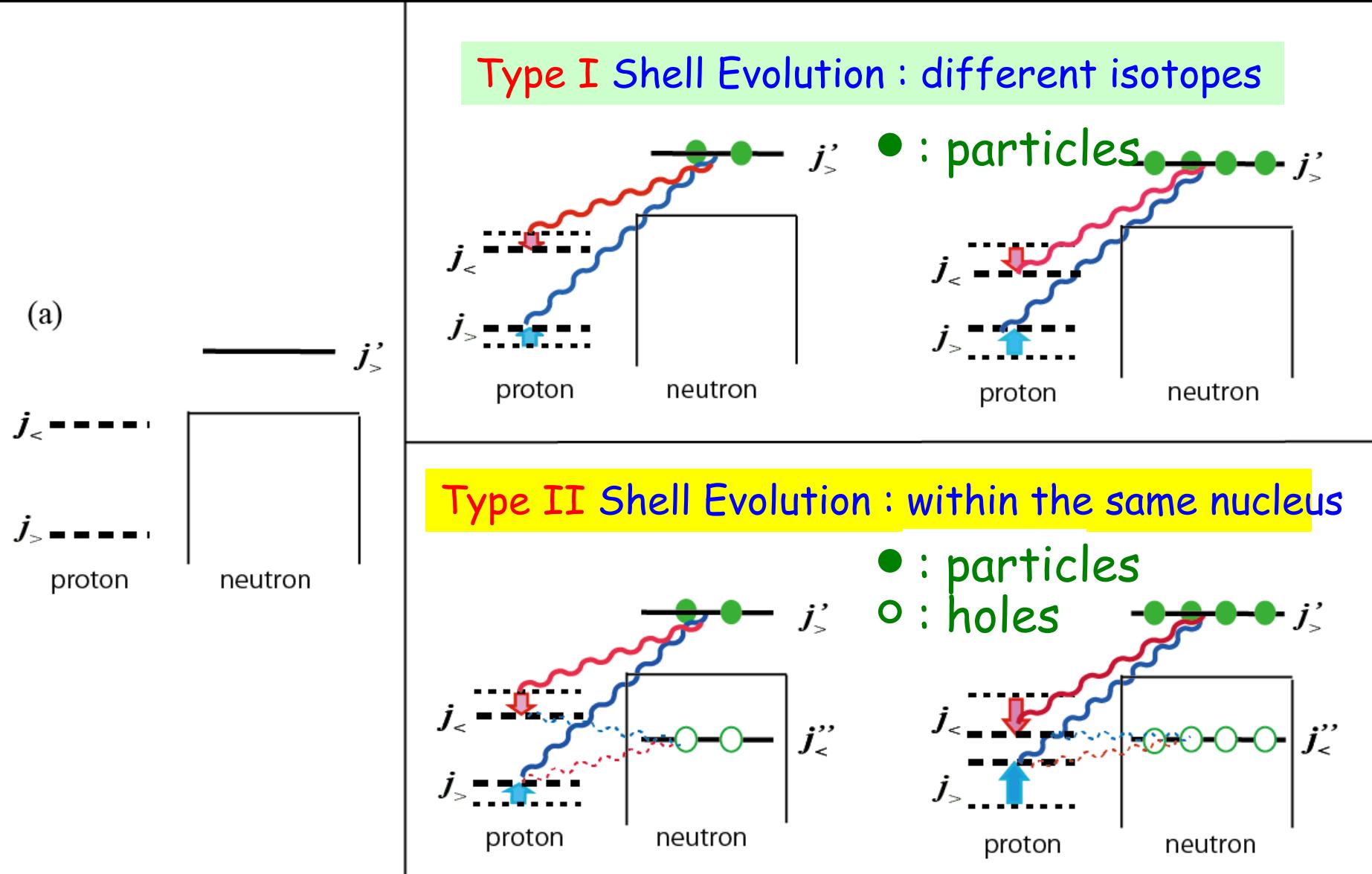


Figure 42 Change of single-particle energies from ^{56}Ni to ^{48}Ca relative to the $2p_{3/2}$ orbit. The red arrows indicate the change of the $1f_{5/2}$ ESPE. The arising magic numbers, $N=32$ and 34 , are shown in black circles.

Shell Evolution - from Type I to Type II -

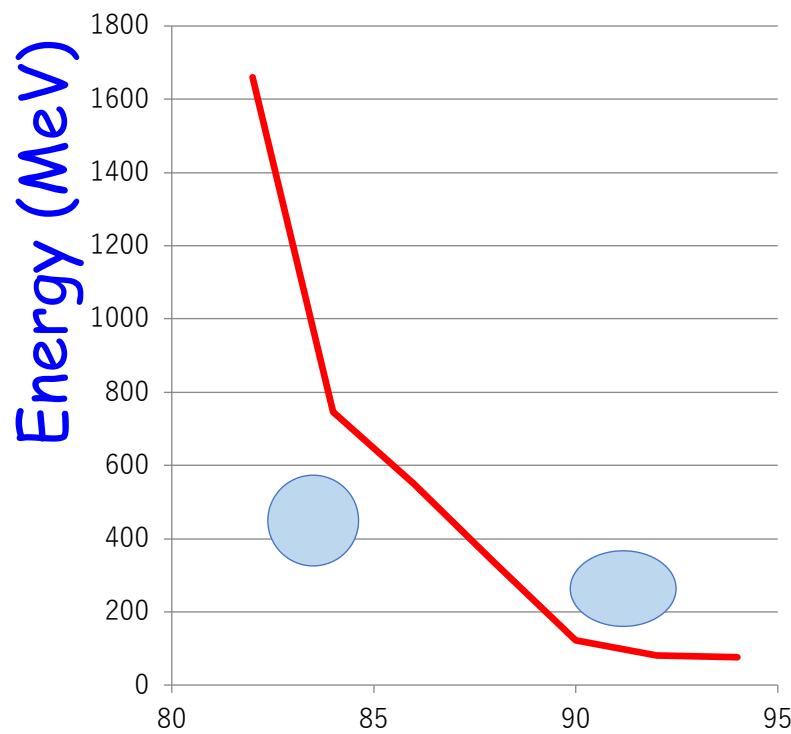


Quantum Phase Transition

Shape change as a function of N (or Z)

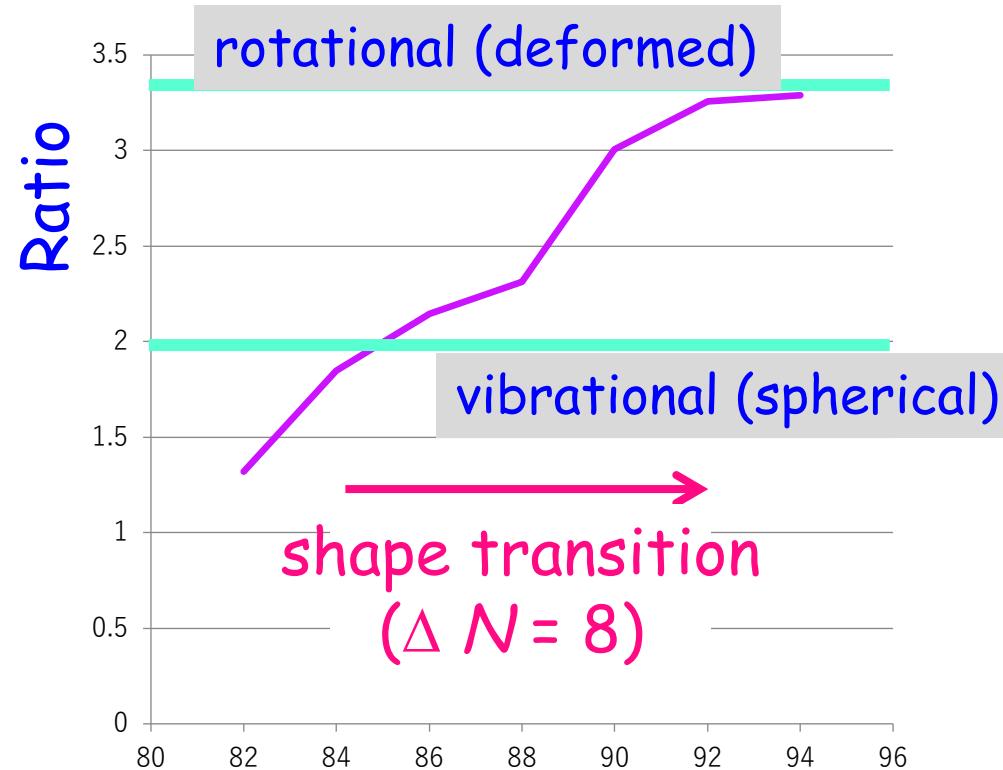
2^+ and 4^+ level properties of Sm isotopes

$\text{Ex } (2^+)$:
excitation energy of first 2^+ state



Neutron number, N

$$R_{4/2} = \text{Ex } (4^+) / \text{Ex}(2^+)$$



Neutron number, N

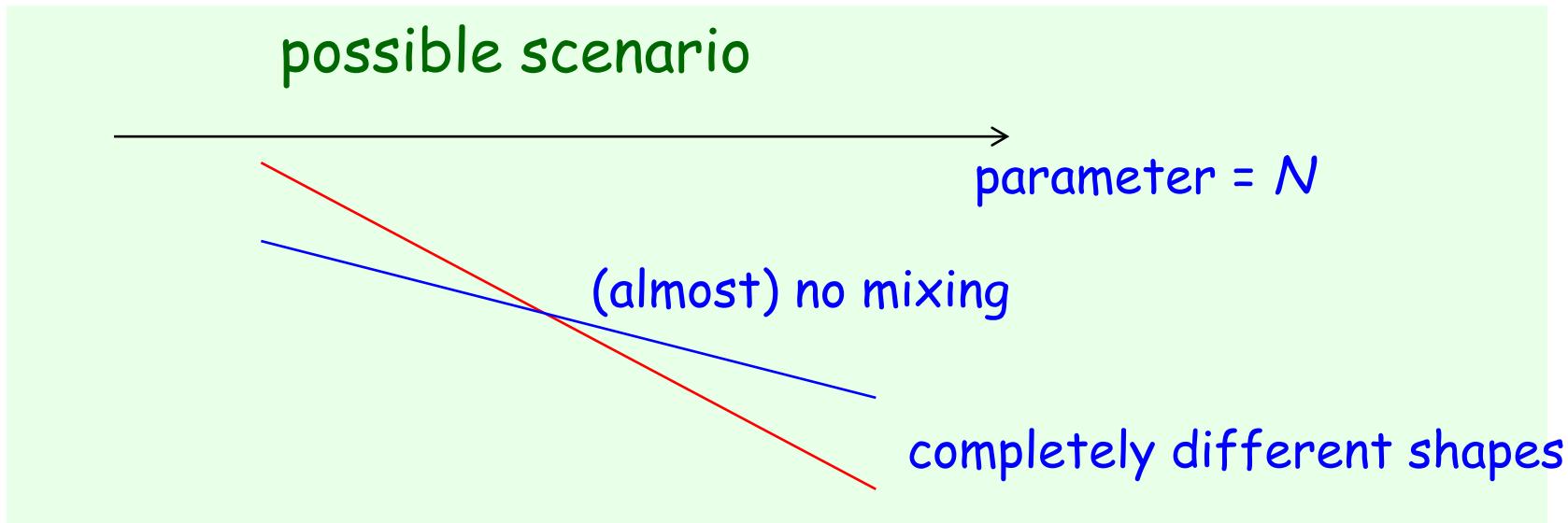
Can this be a kind of Phase Transition?

Can the shape transition be a "Quantum Phase Transition" ?

The shape transition occurs rather gradually.

The definition of Quantum Phase Transition :

an abrupt change in the ground state of a many-body system
by varying a physical parameter at zero temperature. (cf., Wikipedia)

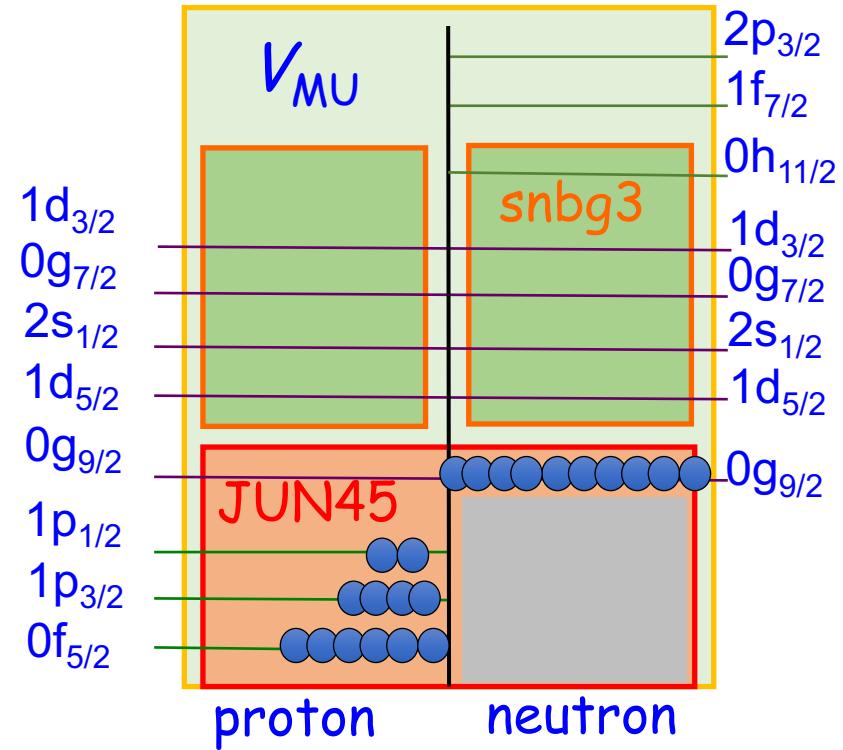


Sizable mixing occurs usually in finite quantum systems.

An example : shapes of Zr isotopes by Monte Carlo Shell Model

- Effective interaction:
 $JUN45 + snbg3 + V_{MU}$
known effective interactions
+ minor fit for a part of
T=1 TBME's
- Nucleons are excited fully
within this model space
(no truncation)

We performed Monte Carlo Shell Model (MCSM) calculations, where the largest case corresponds to the diagonalization of 3.7×10^{23} dimension matrix.



^{56}Ni

Togashi, Tsunoda, TO *et al.* PRL
117, 172502 (2016)

From earlier shell-model works ...

PHYSICAL REVIEW C

VOLUME 20, NUMBER 2

AUGUST 1979

Unified shell-model description of nuclear deformation

P. Federman

Instituto de Física, Universidad Nacional Autónoma de México, Apartado Postal 20-364, México 20, D. F.

S. Pittel

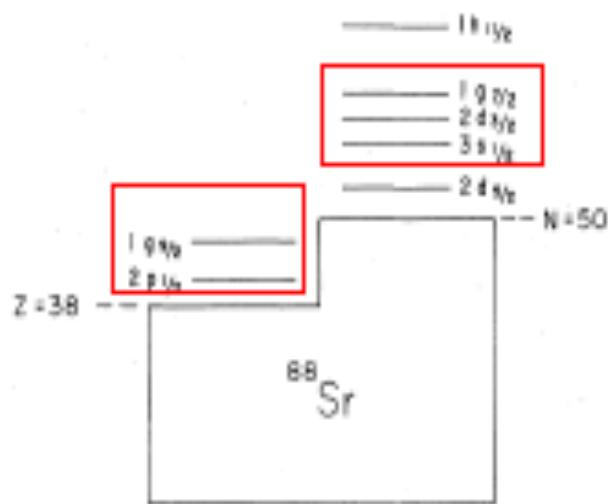


FIG. 3. Single-particle levels appropriate to a description of nuclei in the Zr-Mo region. An ^{88}Sr core is assumed.

PHYSICAL REVIEW C 79, 064310 (2009)

Shell model description of zirconium isotopes

K. Sieja,^{1,2} F. Nowacki,³ K. Langanke,^{2,4} and G. Martínez-Pinedo¹

In this paper, we perform for the first time a SM study of Zr isotopes in an extended model space ($1f_{5/2}$, $2p_{1/2}$, $2p_{3/2}$, $1g_{9/2}$) for protons and ($2d_{5/2}$, $3s_{1/2}$, $2d_{3/2}$, $1g_{7/2}$, $1h_{11/2}$) for neutrons, dubbed hereafter $\pi(r3-g)$, $\nu(r4-h)$.

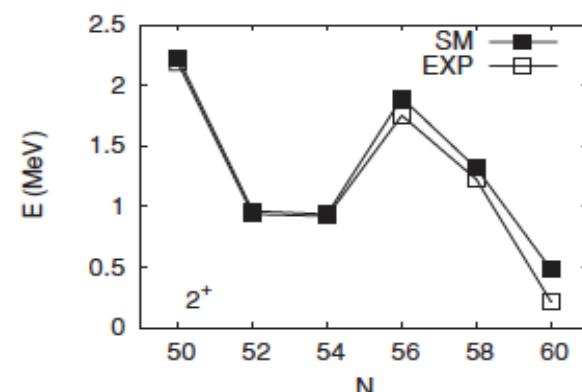
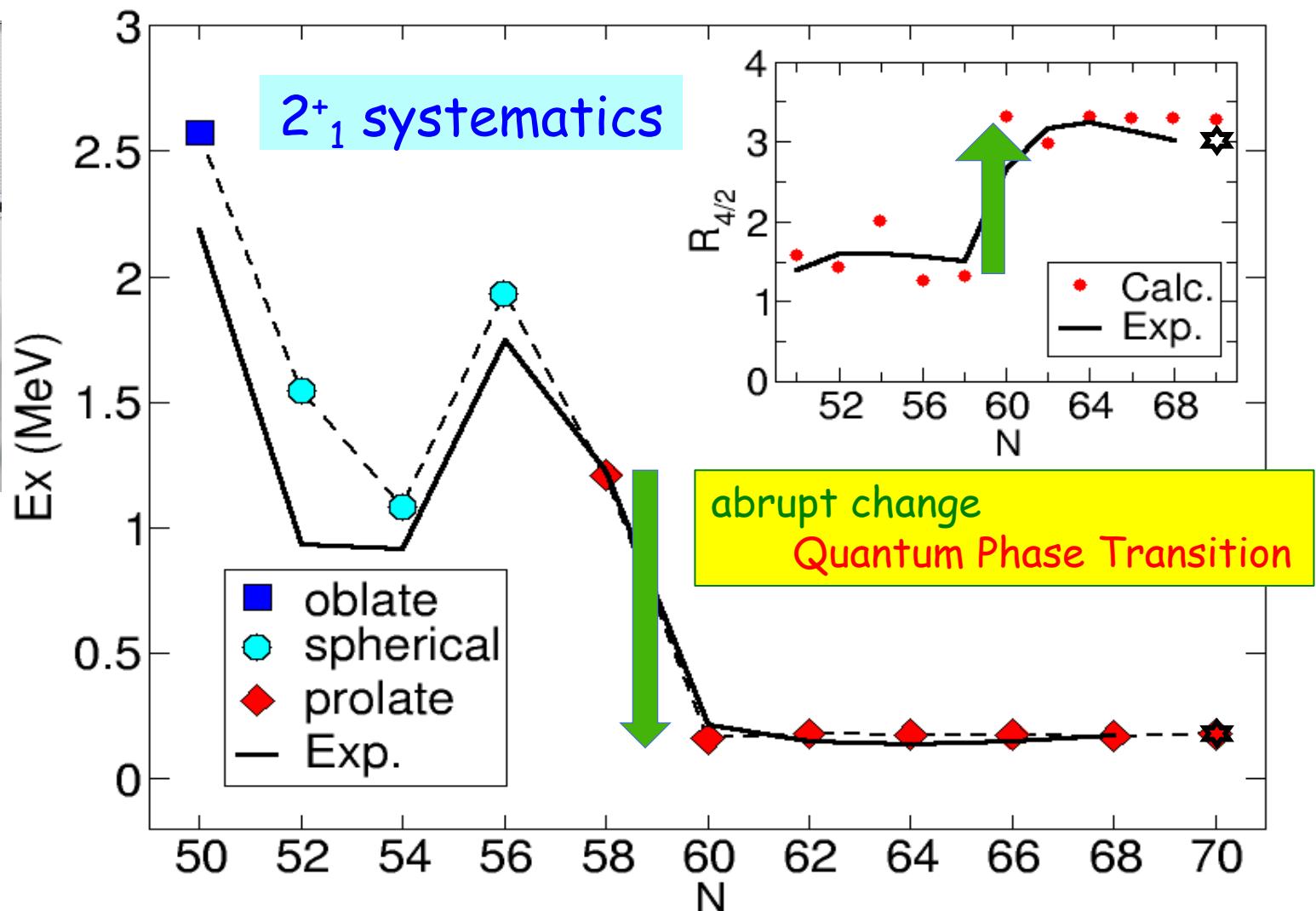
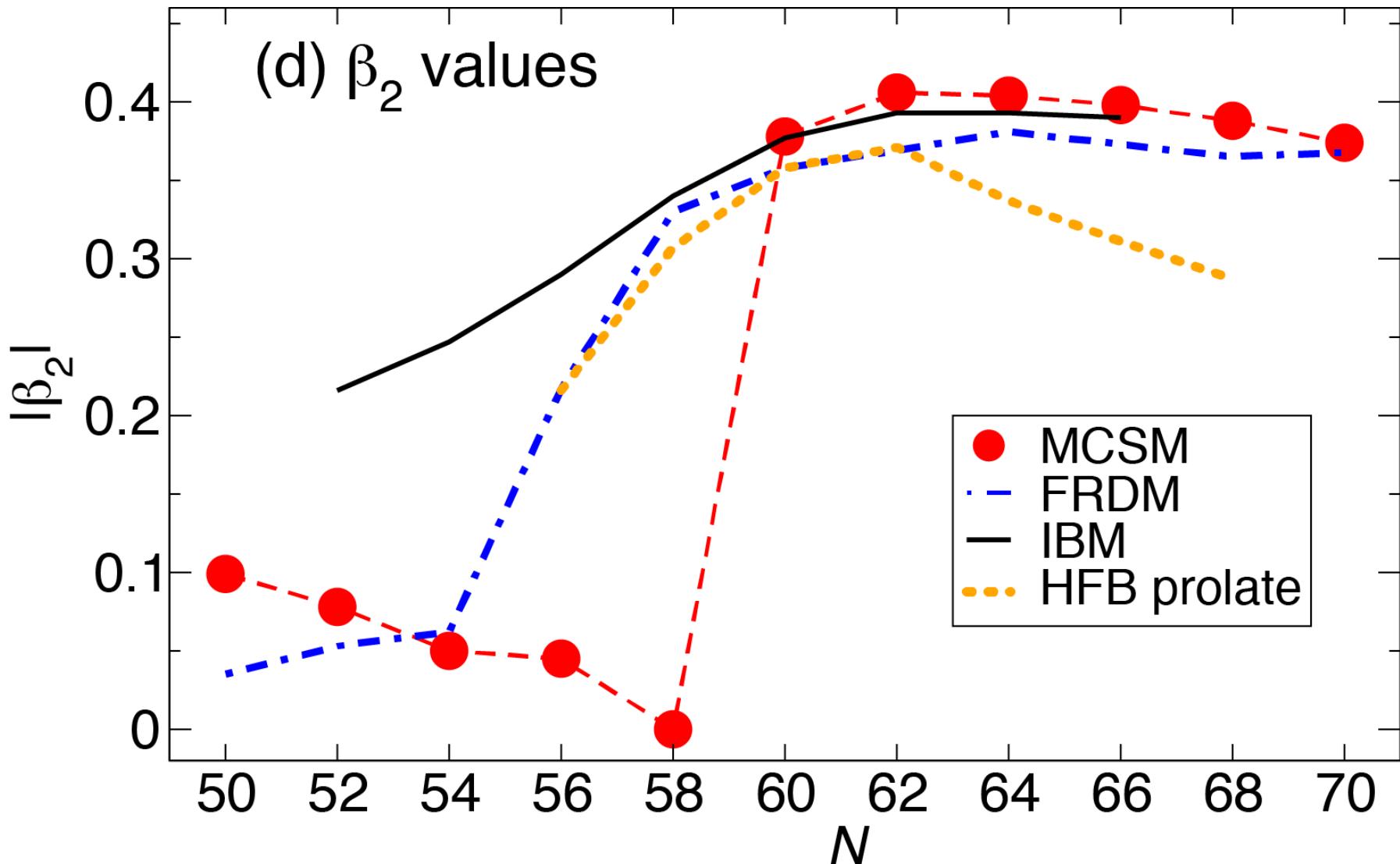


FIG. 12. Systematics of the experimental and theoretical first excited 2^+ states along the zirconium chain.



Quantum Phase Transition in the Shape of Zr isotopes

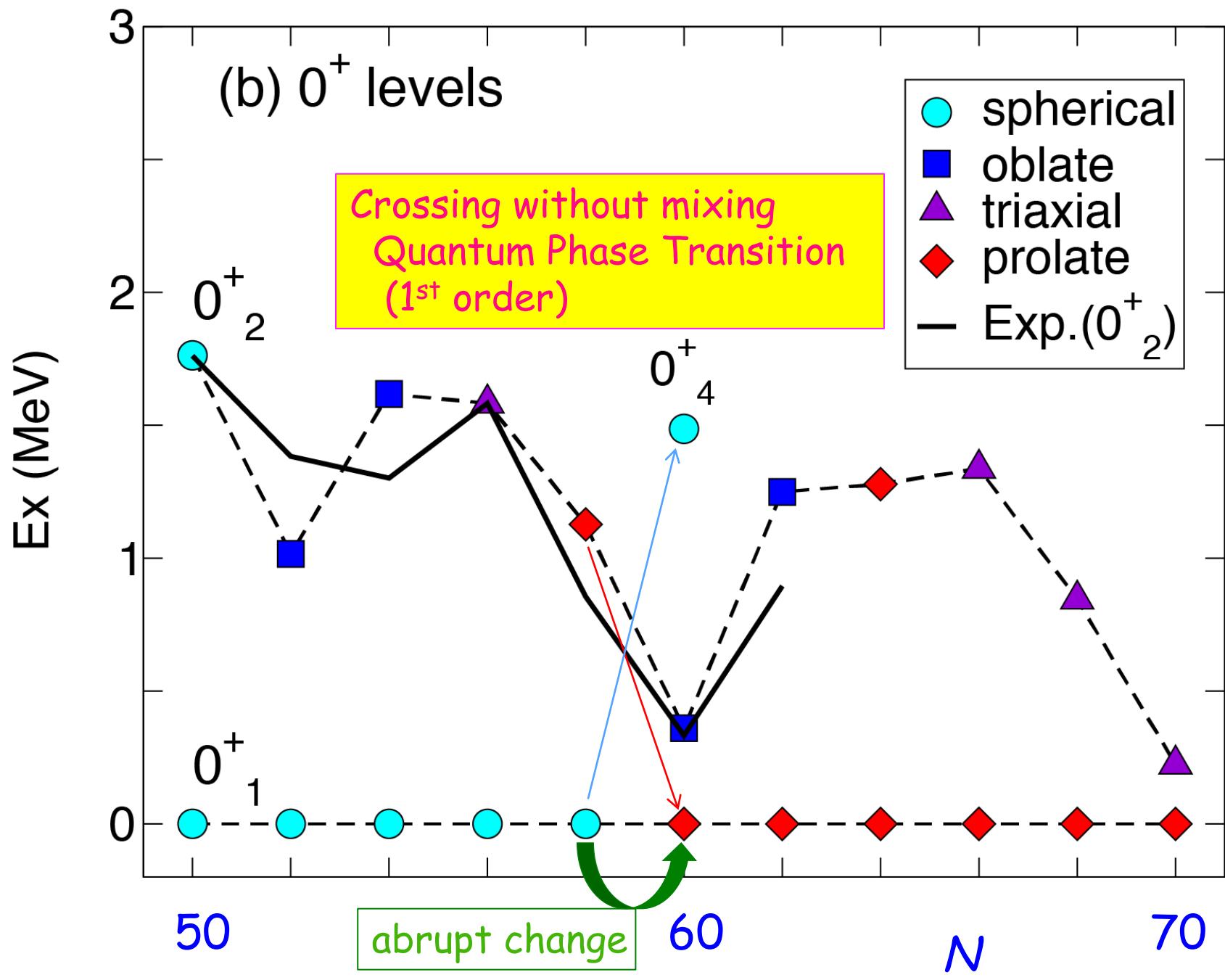
Tomoaki Togashi,¹ Yusuke Tsunoda,¹ Takaharu Otsuka,^{1,2,3,4} and Noritaka Shimizu¹



FRDM: S. Moeller et al. At. Data Nucl. Data Tables 59, 185 (1995).

IBM: M. Boyukata et al. J. Phys. G 37, 105102 (2010).

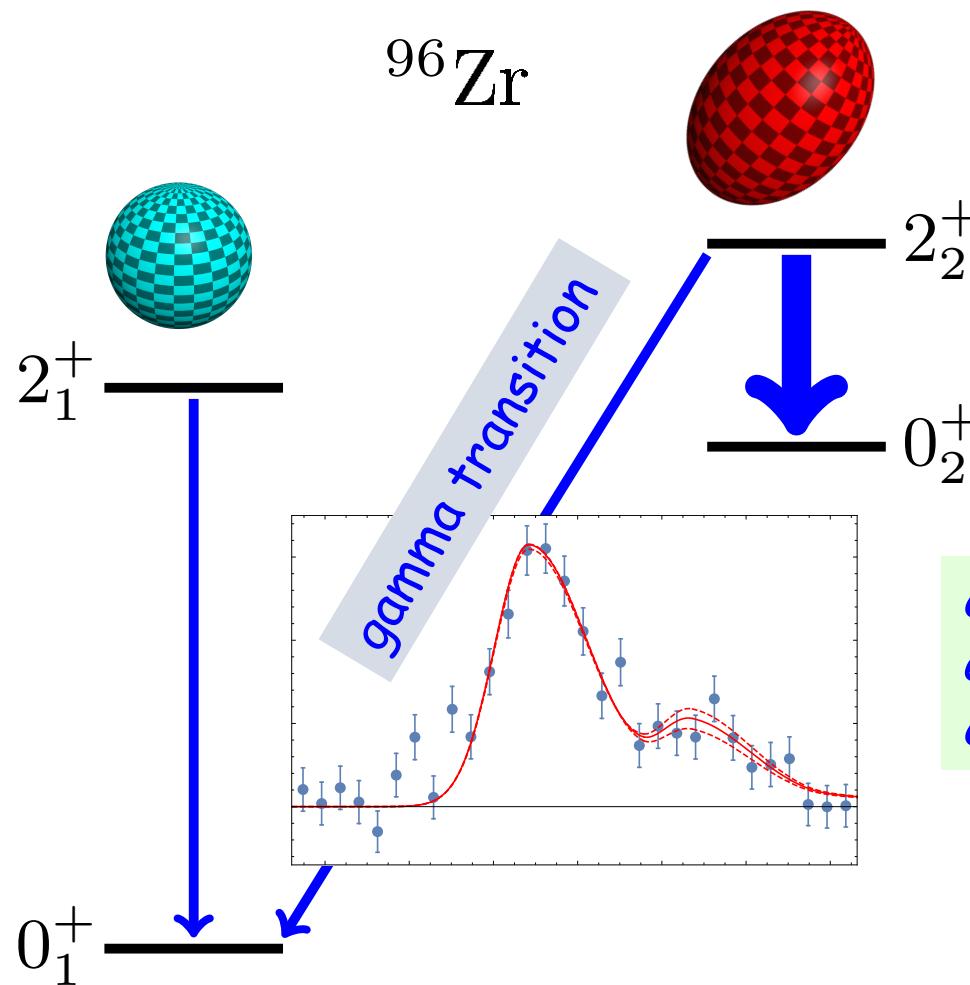
HFB: R. Rodriguez-Guzman et al. Phys. Lett. B 691, 202 (2010).⁶⁸

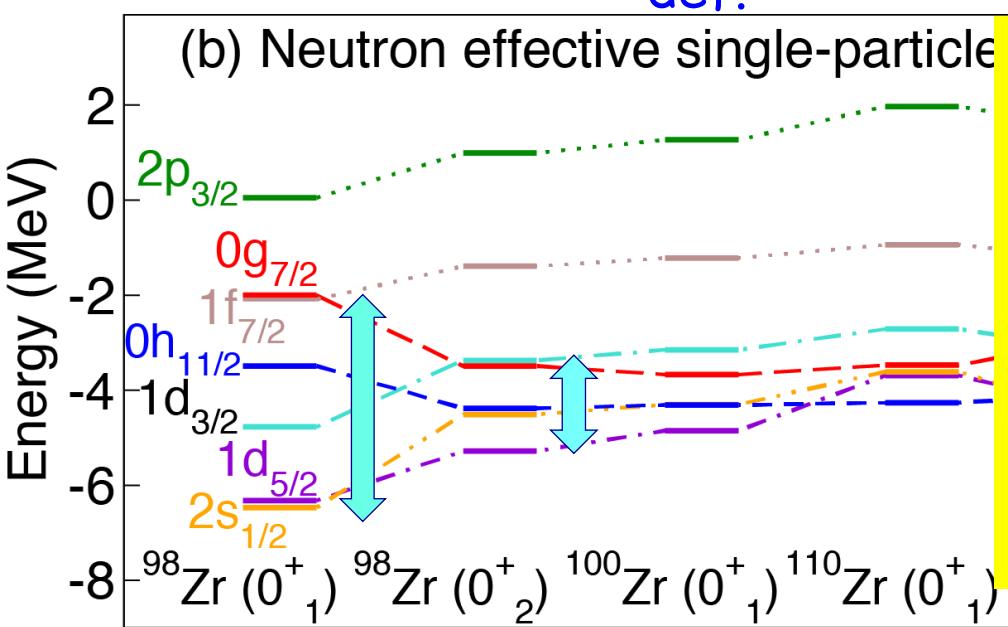
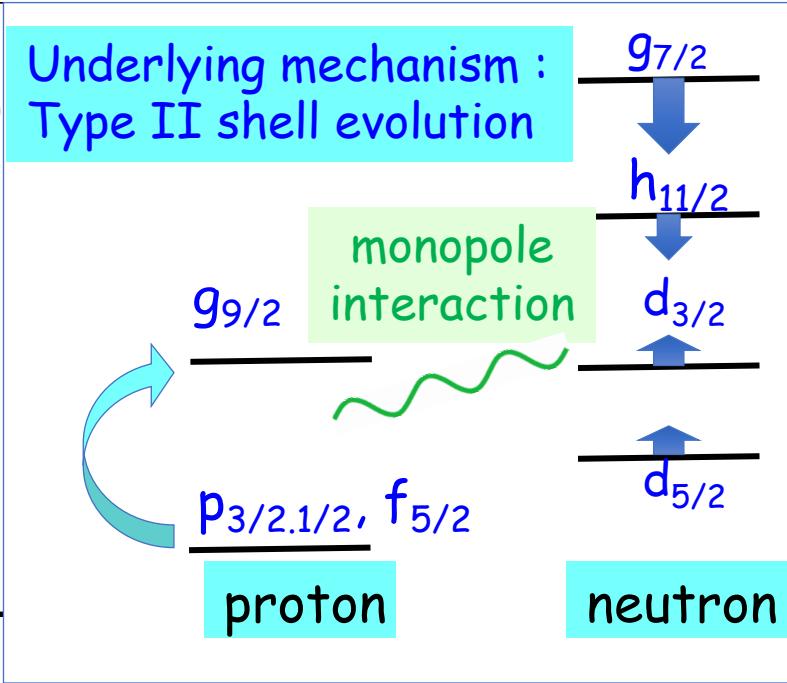
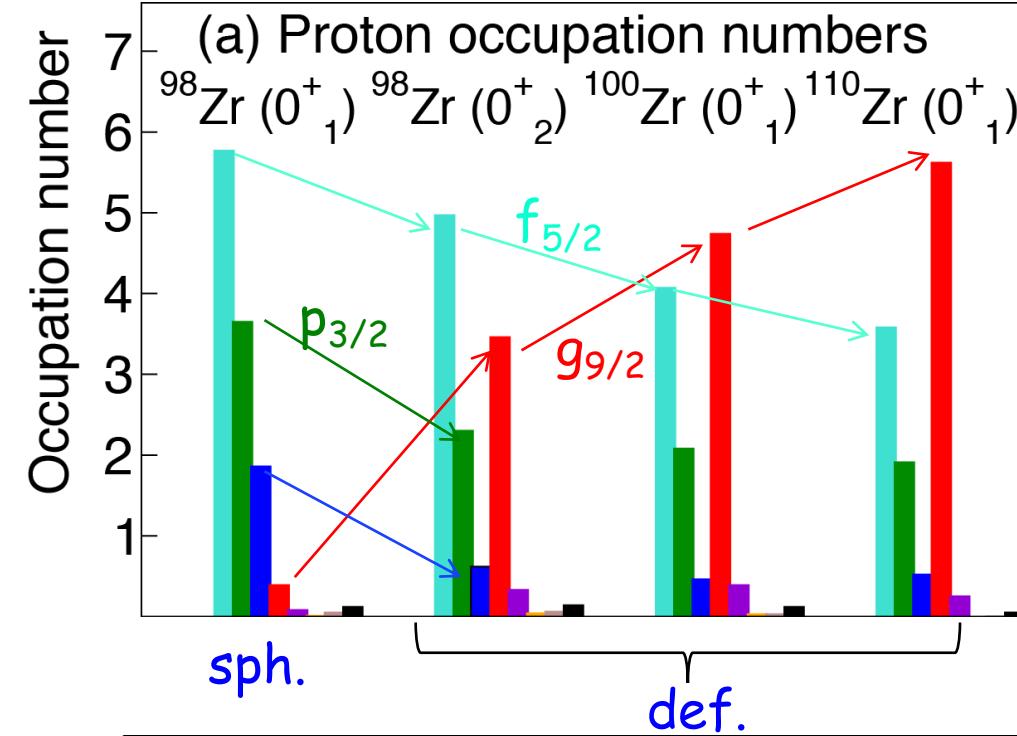


3

First Measurement of Collectivity of Coexisting Shapes Based on Type II Shell Evolution: The Case of ^{96}Zr

C. Kremer,¹ S. Aslanidou,¹ S. Bassauer,¹ M. Hilcker,¹ A. Krugmann,¹ P. von Neumann-Cosel,¹
T. Otsuka,^{2,3,4,5} N. Pietralla,¹ V. Yu. Ponomarev,¹ N. Shimizu,³ M. Singer,¹ G. Steinhilber,¹
T. Togashi,³ Y. Tsunoda,³ V. Werner,¹ and M. Zweidinger¹





Relevant neutron single-particle levels get closer as a combined effect of nuclear forces (tensor and central) and particular configurations. The resistance power against deformation is then reduced. Large difference in ESPEs and configurations \rightarrow crossing w/o mixing

Type II shell evolution is a simplest and visible case of

Quantum Self Organization

$$\text{deformation} = \frac{\text{quadrupole force}}{\text{resistance power}}$$

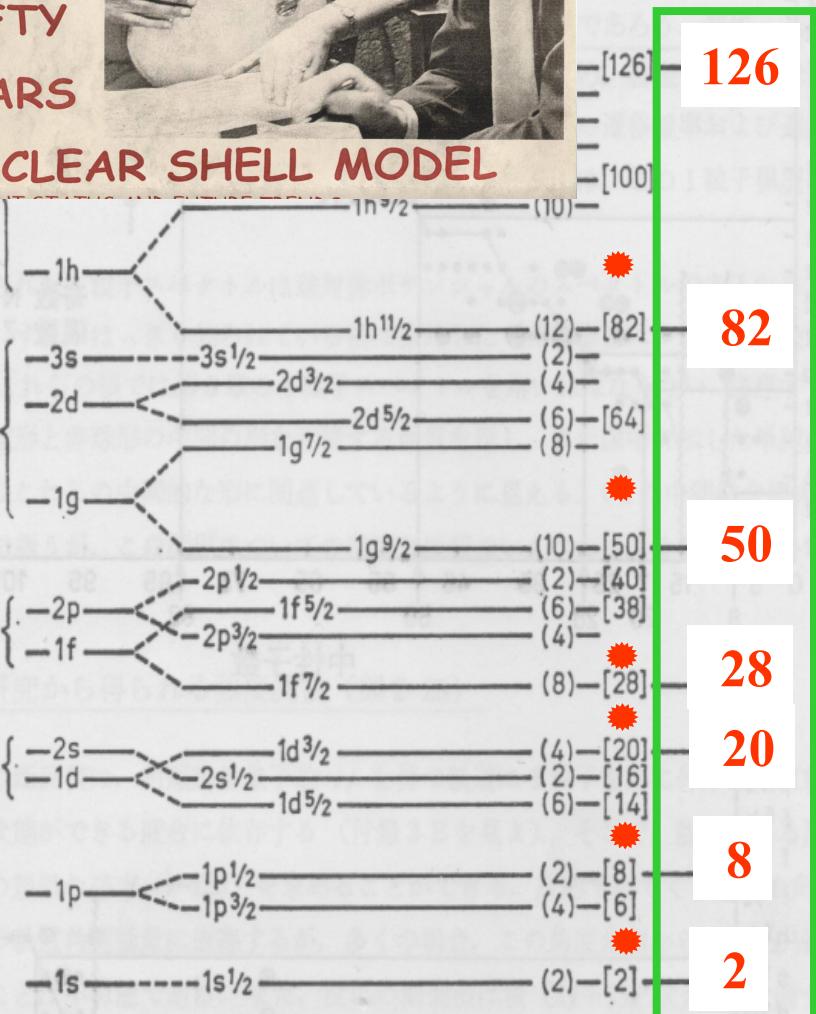
resistance power \leftarrow pairing force



single-particle energies

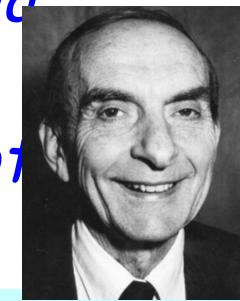
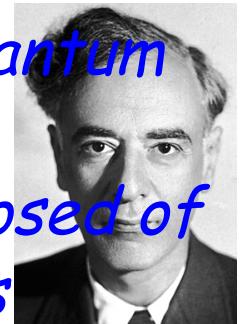
Atomic nuclei can “organize” their single-particle energies by taking particular configurations of protons and neutrons optimized for each eigenstate, thanks to orbit-dependences of monopole components of nuclear forces (e.g., tensor force).
→ an enhancement of Jahn-Teller effect.

CLEAR SHELL MODEL



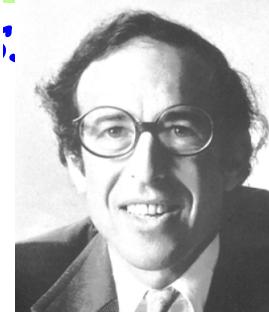
Magic numbers
by
Mayer and
Jensen (1949)

Atomic nucleus is a quantum
Fermi liquid :
The nucleus is composed of
almost free nucleons
interacting weakly via
residual forces
in a (solid) (mean) pot



“how single-particle states can coexist with collective modes” conceived also by Gerry Brown

Surface deformation produced.
by additional deformed mean
field



SVEN GÖSTA NILSSON

Nilsson model Hamiltonian

"Nuclear structure II" by Bohr and Mottelson

deformed nuclei, is obtained by a simple modification of the harmonic oscillator (Nilsson, 1955; Gustafson *et al.*, 1967),

$$H = \frac{\mathbf{p}^2}{2M} + \frac{1}{2}M(\omega_3^2 x_3^2 + \omega_{\perp}^2(x_1^2 + x_2^2)) + v_{ll}\hbar\omega_0(\mathbf{l}^2 - \langle \mathbf{l}^2 \rangle_N) + v_{ls}\hbar\omega_0(\mathbf{l} \cdot \mathbf{s})$$

quadrupole deformed field

spherical field (5-10)

$$\langle \mathbf{l}^2 \rangle_N = \frac{1}{2}N(N+3)$$

Figure	Region	$-v_{ls}$	$-v_{ll}$
5-1	N and $Z < 20$	0.16	0
5-2	$50 < Z < 82$	0.127	0.0382
5-3	$82 < N < 126$	0.127	0.0268
5-4	$82 < Z < 126$	0.115	0.0375
5-5	$126 < N$	0.127	0.0206

Table 5-1 Parameters used in the single-particle potentials of Figs. 5-1 to 5-5.

Spin-orbit force

$$\left. \begin{array}{ll} A=68 & 1.28 \\ A=100 & 1.12 \\ A=186 & 0.91 \end{array} \right\} (\mathbf{l} \cdot \mathbf{s})$$

Intuitively speaking,

$$\text{deformation} = \frac{\text{quadrupole force}}{\text{resistance power}}$$

resistance power \leftarrow pairing force

single-particle energies

Atomic nuclei can “organize” their single-particle energies by taking particular configurations of protons and neutrons, thanks to orbit-dependences of nuclear forces (e.g., tensor force).

Quantum Self Organization

Note : spherical single-particle energies are often treated being constant

After mastering the shell model,
three (possible) pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
dimensionality

massive
parallel
computers

Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
8+8 on ^{80}Zr core (Sn)
8+10 on ^{132}Sn core
(Sm)

...
island of stability
+
 χ EFT based
(multi-)shell int.

many-body dynamics

Shell evolution
(Type I & II)

Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization