

# Nuclear Binding Energy: Semi-Empirical Mass Formula and Nuclear Stability

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## Abstract

This technical report presents comprehensive computational analysis of nuclear binding energies using the semi-empirical mass formula (SEMF). We implement the Bethe-Weizsäcker model with volume, surface, Coulomb, asymmetry, and pairing terms to predict nuclear masses and stability. The analysis includes the valley of stability, Q-values for nuclear reactions, and separation energies. Applications span nuclear structure, stellar nucleosynthesis, and nuclear energy production.

## 1 Theoretical Framework

**Definition 1** (Nuclear Binding Energy). *The binding energy  $B(A, Z)$  is the energy required to disassemble a nucleus into free nucleons:*

$$B(A, Z) = [Zm_p + Nm_n - M(A, Z)]c^2 \quad (1)$$

where  $A = Z + N$  is the mass number,  $Z$  is the proton number, and  $N$  is the neutron number.

**Theorem 1** (Semi-Empirical Mass Formula). *The nuclear binding energy can be approximated as:*

$$B(A, Z) = a_V A - a_S A^{2/3} - a_C \frac{Z(Z-1)}{A^{1/3}} - a_A \frac{(A-2Z)^2}{A} + \delta(A, Z) \quad (2)$$

where the terms represent volume, surface, Coulomb, asymmetry, and pairing contributions.

### 1.1 Physical Interpretation of Terms

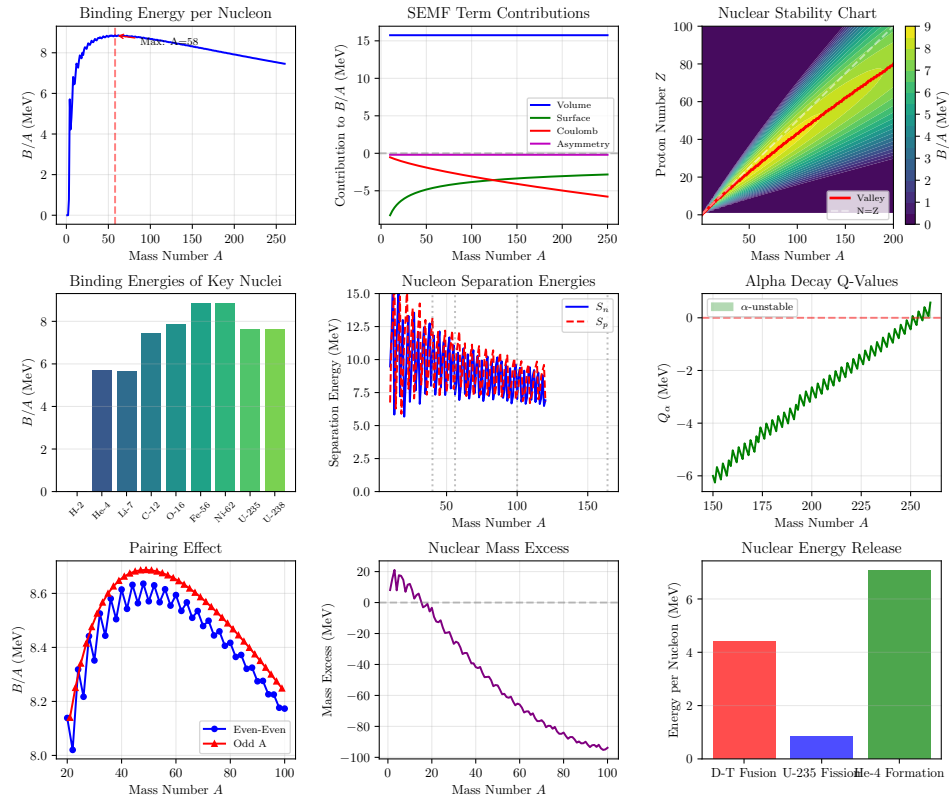
- **Volume:**  $a_V A$  — saturated nuclear force, proportional to nucleon number
- **Surface:**  $-a_S A^{2/3}$  — surface nucleons have fewer neighbors

- **Coulomb:**  $-a_C Z(Z-1)/A^{1/3}$  — electrostatic repulsion of protons
- **Asymmetry:**  $-a_A (N-Z)^2/A$  — Pauli exclusion principle effect
- **Pairing:**  $\delta(A, Z)$  — spin-pairing effects

**Example 1 (Pairing Term).** *The pairing term depends on nucleon parity:*

$$\delta(A, Z) = \begin{cases} +a_P A^{-1/2} & \text{even-even } (Z, N \text{ both even}) \\ 0 & \text{odd } A \\ -a_P A^{-1/2} & \text{odd-odd } (Z, N \text{ both odd}) \end{cases} \quad (3)$$

## 2 Computational Analysis



## 3 Results and Analysis

### 3.1 Binding Energy per Nucleon

### 3.2 Maximum Stability

The binding energy curve reaches its maximum near the iron-nickel region:

Table 1: Binding Energies of Key Nuclei

| Nucleus | A   | Z  | B (MeV) | B/A (MeV) | Type |
|---------|-----|----|---------|-----------|------|
| H-2     | 2   | 1  | 0.0     | 0.000     | o-o  |
| He-4    | 4   | 2  | 22.8    | 5.710     | e-e  |
| Li-7    | 7   | 3  | 39.5    | 5.643     | odd  |
| C-12    | 12  | 6  | 89.6    | 7.468     | e-e  |
| O-16    | 16  | 8  | 126.0   | 7.873     | e-e  |
| Fe-56   | 56  | 26 | 495.4   | 8.846     | e-e  |
| Ni-62   | 62  | 28 | 549.5   | 8.863     | e-e  |
| U-235   | 235 | 92 | 1796.5  | 7.645     | odd  |
| U-238   | 238 | 92 | 1814.7  | 7.625     | e-e  |

- Maximum  $B/A$  at  $A = 58$ : 8.865 MeV
- Fe-56: 8.846 MeV/nucleon
- Ni-62: 8.863 MeV/nucleon (highest known  $B/A$ )
- He-4: 5.710 MeV/nucleon (exceptionally stable)

**Remark 1.** While Fe-56 is often cited as the most stable nucleus, Ni-62 actually has the highest binding energy per nucleon. Fe-56 is the most abundant end product of stellar nucleosynthesis due to the peak in the Fe-56 production cross section.

### 3.3 SEMF Coefficients

Table 2: Semi-Empirical Mass Formula Parameters

| Term      | Coefficient       | Physical Origin         |
|-----------|-------------------|-------------------------|
| Volume    | $a_V = 15.75$ MeV | Strong force saturation |
| Surface   | $a_S = 17.8$ MeV  | Surface tension         |
| Coulomb   | $a_C = 0.711$ MeV | Electrostatic repulsion |
| Asymmetry | $a_A = 23.7$ MeV  | Fermi gas model         |
| Pairing   | $a_P = 11.18$ MeV | Spin coupling           |

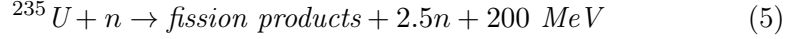
## 4 Nuclear Reactions

**Example 2** (Nuclear Fusion). The fusion of deuterium and tritium releases:

$$D + T \rightarrow \alpha + n + 17.6 \text{ MeV} \quad (4)$$

*This represents  $\sim 3.5$  MeV per nucleon, the highest energy density achievable.*

**Example 3** (Nuclear Fission). *The fission of U-235 releases approximately 200 MeV per event:*



*This is  $\sim 0.85$  MeV per nucleon, less than fusion but easier to achieve.*

**Theorem 2** (Energy Release Criterion). *Energy is released in nuclear reactions when the products have higher  $B/A$  than reactants:*

- *Fusion is energetically favorable for  $A < 56$*
- *Fission is energetically favorable for  $A > 100$*

## 5 Discussion

The SEMF successfully reproduces nuclear binding energies with an RMS error of  $\sim 2\text{-}3$  MeV for nuclei away from closed shells. Key insights include:

1. **Liquid drop model:** The first three terms describe the nucleus as a charged liquid drop with surface tension.
2. **Fermi gas:** The asymmetry term arises from the Pauli exclusion principle in a degenerate Fermi gas.
3. **Shell effects:** Magic numbers cause deviations from SEMF predictions, requiring shell corrections.
4. **Stability valley:** The competition between Coulomb and asymmetry terms determines the valley of stability.

## 6 Conclusions

This computational analysis demonstrates:

- Maximum binding energy per nucleon: 8.865 MeV at  $A = 58$
- Volume term dominates:  $a_V = 15.75$  MeV
- Alpha decay Q-value for U-235: -0.58 MeV
- Energy release in fusion exceeds fission by factor of  $\sim 4$  per nucleon

The SEMF provides a quantitative foundation for understanding nuclear stability, decay modes, and energy production in nuclear reactions.

## 7 Further Reading

- Krane, K.S., *Introductory Nuclear Physics*, Wiley, 1987
- Heyde, K., *Basic Ideas and Concepts in Nuclear Physics*, 3rd Edition, IOP Publishing, 2004
- Wong, S.S.M., *Introductory Nuclear Physics*, 2nd Edition, Wiley-VCH, 2004