

Aperture correction for Beamforming in radiometric detection of ultra-high energy cosmic rays

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For high-energy cosmic-ray physics, it is imperative to determine the mass and energy of the cosmic ray that initiated the air shower in the atmosphere. This information can be extracted from the longitudinal profile of the air shower. In radio-metric observations, this profile is customarily determined through an extensive fitting procedure where calculated radio intensity is fitted to data. Beamforming the measured signals offers a promising alternative to bypass the cumbersome fitting procedure and to determine the longitudinal profile directly. Finite aperture effects in beamforming hamper the resolution with which this profile can be determined. We present a comprehensive investigation of the beamforming resolution in radiometric observations of air showers. There are two, principally different, approaches possible in air-shower beamforming, one where the total beamforming intensity is determined and an alternative where the beamforming trace is cross-correlated with a known response function. The effects due to a finite aperture (size of antenna array and bandwidth) are large for both approaches. We argue that it is possible to correct for the aperture corrections using an unfolding procedure. We give an explicit expression for the folding function, the kernel. Being able to calculate the folding function allows for unfolding the finite aperture effects from the data. We show that, in a model-to-model comparison, this allows for an accurate reconstruction of the current profile as the shower develops in the atmosphere. We present also an example where we reconstruct the longitudinal current profile of a shower developing under thunderstorm conditions where the atmospheric electric fields greatly alter the orientation of the transverse current in the shower front.

I. INSTRUCTION TO THE USE OF THE KERNEL GENERATOR AND ITS APPLICATION TO BEAMFORMING

For the paper [1] use was made of the following construction. The aperture (or Kernel) is calculated in the code "BeamForm-v20.f90". This code also performs a fit of the Gaisser-Hillas profile longitudinal profile to reproduce the Beaming results for a given shower.

```
1 #!/bin/bash
2 #
3 #
4 Home="/home/olaf"
5 export RunFolder=$(pwd)
6 export WorkDir=${RunFolder}
7 export ProgDir=${Home}/MGMR3D/Program"
8 export GLEscr=${ProgDir}
9 export LIBRARY="-lm /home/olaf/NumLib/bin/libfftpack5.1d.a"
10 export FCFLAGS="-gdb -fbounds-check -fbacktrace -ffpe-trap=zero,overflow,underflow,invalid -finit-real=inf"
11 export AntennaFun="/home/olaf/LMA-fit/LMA2019/AntenFunct/"
12
13 #----- Generate radio footprint that will be fitted
14
15 cd ${ProgDir}
```

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```

16 gfortran ${FCFLAGS} -o MGMR3D_fit MGMR3D2_fit-v10.f90 ${LIBRARY}
17 rm *.mod
18
19 cd ${RunFolder}
20 ${ProgDir}/MGMR3D2_fit <MGMR3D2-SKA.in
21
22 #----- run Beaming
23
24 UseLib="/usr/lib/x86_64-linux-gnu"
25 export LIBRARY="-lm ${Home}/NumLib/bin/libfftpack5.ld.a ${UseLib}/lapack/liblapack.a ${UseLib}/blas/libblas.a"
26
27 cd ${ProgDir}
28 gfortran ${FCFLAGS} -o MGMR_Beam BeamForm-v20.f90 ${LIBRARY}
29 rm *.mod
30
31 cd ${RunFolder}
32 ${ProgDir}/MGMR_Beam <Beam-SKA.in
33

```

An example of the lines in the input for the run generating a radio footprint in MGMR3D, "MGMR3D2-SKA.in" for a fair-weather example using a 'SKA' bandwidth:

```

1  &ShPars IntegrateCurrent=-0.01
2  Intensity_Weight=.false.
3  OBSDIST_DIM=75
4  Fit_stl=.true
5  StParRange = 150 ! in down-sampled sample times for calculation of Stokes parameters= number of time-trace samples written to file
6  MOLIERERADIUS= 0 ! Used the default, varying for small xmax, 50 for large values
7  F_lim=1.
8  JOQ=0.22
9  JOT = 14.77, zen_B=22.19 , azi_B=-90. ! direction magnetic field at LOFAR (49.5 mu T)
10 RL_param=.true., R_0=0.3 , L_0=220,
11 tTrace_step=0.020
12 SAMPLINGTIME=1. ! in [ns]
13 nu_min=48. , nu_max= 300.
14 Zen_sh=43. , X_max=900 !
15 energy_sh=3624933134.61
16 PlotBase="SKA" ! 48 -- 300 MHz
17 Fitting=.false.
18 &end
19
20 24. -20 ! shift_x [m], shift_y [m], alpha_vB [deg] !30,10 or 10,20 or -10,30?
21
22 nstep
23
24 -4,9,10,-11, -12, -13, 15, 16, 17, -19, 20, 21, -15
25 "../runs/data/data_127944584xxx.dat" "-test"
26
27 -----
28 theta 0.
29 -----

```

The lines in the input for the run generating a a beam-forming fit to the just generated radio footprint "Beam-SKA.in"

:

```

1  &BeamPars
2  Base = "SKA/" ! 48 -- 300 MHz
3  !ReCreate = .true.
4  AntennaLayout = .false., DistMax = 500.
5  dN_tr=5
6  N_phir = 9 ! 5 ! 9 ! 11 !no effect
7  d_h = 0.1 ! 0.2 ! no real effect
8  N_h = 400 ! 100 !
9  d_crMax = 10.
10 zeta_K= 0.02, 0.1, 0.2, .500, 1.500, 2.500, 3.500, 4.500, 5.500, 7.000, 8.500, 10.000, 12., 14., 17., 20., -1.
11 GoodQual=5.
12 FracFold=0.4
13 &end
14

```

Reading file for specific antenna locations, real data, where the core position is fitted (no example of this case ready, files are produced by the routine "Observab_List" in "MGMR3D2_analyse-v10.f90") :

```

3  OPEN(UNIT=4,STATUS='old',FILE=trim(Base)//trim(label)//'.dat',iostat=nxx) ! E13.4
4  Read(4,*) N_antennas, N_timesamples, GroundLevel_sh, ZenithAngle_deg, Azimuth_deg, EstCore_N, EstCore_E, nu_min, nu_max
5  Do i_ant=1,N_antennas
6    Read(4,*) ant_file(i_ant), ant_pos(1:2,i_ant), i_min, i_max, SamplingTime, tTrace_Offset_dwn
7    ! tTrace_Offset_dwn = time corresponding to the first sample that is read-in
8    OPEN(UNIT=14,STATUS='old',FILE=trim(ant_file(i_ant))//'.csv',iostat=nxx) ! E13.4
9    read(14,*) labela
10   Do i=1,N_timesamples
11     Read(14,*,iostat=nxx) a, ant_xtrace(i, i_ant), b, ant_ytrace(i, i_ant), c
12

```

With the example file name "????.dat" :

```

2  ?
3  ?
4  .
5  .
6  .

```

- OR -

presently the only working option, assume antennas on a grid using vxB polarization only:

```

1  Do j_ant=1,N_antennas
2    Write(label,"(I2.2)") j_ant
3    OPEN(UNIT=4,STATUS='unknown',FILE=trim(Base)//'th_000'//trim(label)//'.csv',iostat=nxx) ! E13.4
4    If(j_ant.eq.1) Then
5      Read(4,*) label, Ant_d, N_timesamples, SamplingTime, t_offset, nu_min, nu_max, GroundLevel_sh, Zen_sh, Azi_sh

```

```

6       write(2,"(A,F7.2,A, I5,A,F6.3,A,F7.1,A)") 'Antenna separation=', Ant_d, '[m], trace length=', &
7         N_timesamples,'[samples], sampling time=', Samplingtime,'[m], offset=', t_offset,'[m]'
8       write(2,"(A,F7.2,A,F7.2,A, A,F6.3,A,F8.3,A,F8.3)") 'Frequency bandwidth=', nu_min,' --', nu_max,' [MHz], ' &
9         ', ground level @', GroundLevel_sh,'[m], Zenith angle=', Zen_sh,'[deg], azimuth=', Azi_sh
10      Else
11        If(nxx.ne.0) exit
12        Read(4,*) label, Ant_d
13      EndIf
14      Do i_t=1,N_timesamples
15        Read(4,*) i, ant_xtrace(i_t,N), ant_ytrace(i_t,N)
16      End Do ! i_t=1,N_t
17      Close(Unit=4)
18      End Do ! i_d=1,N_d

```

With the example file name "th_00001.csv" :

```

1  !    10.00 1094    0.2998    100.0    48.00    300.0    0.000    43.00    0.000
2    1    4.5419E+04    0.000    3361.
3    2    6.2863E+04    0.000    4089.
4    3    7.1446E+04    0.000    3440.
5    .
6    .
7    .

```

A. The renovated approach

To improve transparency the beamforming code has been split in two. One program "BeamForm-GenerateKernel" is constructing the Kernel for a particular cosmic-ray zenith angle (independent of azimuth angle) and for a given (irregular) grid of focal heights. The second program "BeamForm-ExtractCurrent" uses the Kernel to extract the longitudinal shower profile from a given radio footprint. To run both programs (and MGMR as well to generate a footprint) one uses the script "NewBeamForm-SKA.sh"

```

1  #!/bin/bash
2  #
3  #
4  source ${MGMR_Base}/MGMR_SC.sh
5
6  #----- Run MGMR3D to generate a sample shower profile -----
7  export LIBRARY=${FFTLib}
8  source ${ProgDir}/RunMGMR.sh MGMR3D_fit-v10 MGMR3D2-SKA.in
9
10 #----- Run Beaming Kernel generator -----
11 export LIBRARY=${FFTLib} ${BLASlib}
12 source ${ProgDir}/RunMGMR.sh BeamForm-GenerateKernel Kernel-SKA.in
13
14 #----- Run Beaming to extract source currents-----
15 source ${ProgDir}/RunMGMR.sh BeamForm-ExtractCurrent BeamCurr-SKA.in
16 exit

```

Here line 4 runs the initiations for the ShortCuts that are defined for a smooth installation. The are defined in "MGMR_SC.sh" to be edited to suit your particular installation In the file ".bashrc" residing in your home directory the following lines should be added,

```

export Home="/Users/users/scholten"
export MGMR_Base=${Home}/MGMR3D

```

where "MGMR_Base" should point to the folder containing "MGMR_SC.sh" . The files containing the input lines needed for running the various programs are denoted by the ".in" extension. An example for the one to run the MGMR program is was given earlier.

An example of the input for "BeamForm-GenerateKernel" is given by "Kernel-SKA.in" :

```

1  &BeamPars
2  Base = "SKA/" ! 48 -- 300 MHz
3  AntennaLayout = .false., DistMax = 500.
4  dN_tr=5
5  N_phir = 5 ! 9 ! 11 ! 9 ! no effect
6  d_h= 0.2 ! 0.1 ! no real effect
7  N_h= 100 ! 400 !
8  d_crMax = 10.
9  zeta_K= 0.02, 0.1, 0.2, .500, 1.500, 2.500, 3.500, 4.500, 5.500, 7.000, 8.500, 10.000, 12., 14., 17., 20., -1.
10 &end

```

Here the mandatory lines "&BeamPars" and "&end" specify the beginning and the end of the namelist input. The lines in between can be put in arbitrary order or omitted, in which case a default value is used (when so defined). All character on a line after an exclamation mark, !, are treated as comments. The output file "BeamForm-GenK.out" contains the complete listing of the possible namelist variables and their values for this particular run. Capitalization of the input parameters is not important,

- "Base=" The directory where all output will be written. It may be followed by some initial characters to distinguish different runs.

- **"DataBase="** The directory where the antenna info was written.
- **"AntennaLayout="** To switch between specifying all antenna locations (.true.) or to assume they lie on a regular grid in distance to the shower core (.false.). When antenna locations are specified the core location for the shower will be calculated. The option **"AntennaLayout=.t."** has not (yet) been tested well.
- **"del_z="** Step size for grid along the shower axis, in [m].
- **"del_tr="** Step size for distance to shower-axis grid, in [m].
- **"zeta_K="** The (irregular) grid heights (along the shower axis) for the beamforming calculations, in [km].
- **"N_h="** Number of steps probing the pancake
- **"d_h="** Step length [m] for probing pancake. It is important that $d_h=50./N_h$ is considerably smaller (factor 5) than the grid spacing imposed by the frequency filtering. For 30-80 MHz a value $d_h=0.25[m]$ appears sufficient. If this is not small enough the gradual move of the peak in pancake function with increasing t_r is not properly reflected, resulting in jumps in the weight function after filtering. Going to an even smaller d_h appears to be useless.
- **"dN_tr="** The resolution in height (dN_trx 10m) should be at least as good as the required precision for the currents to be extracted. 10 or 20 may thus be time-saving with the same results.
- **"N_phir="** If taken smaller than 5 it will result in a very large fluctuations in the results, in particular for those kinematic conditions where the Cherenkov angle just starts to becomes visible.
- **"d_crMax="** Maximal grid spacing for grid in radial extent.
- **"DistMax="** Maximal distance for which antennas will be included in the calculation. Will be set automatically is antenna layout is read-in (**"AntennaLayout=.t."**)

An example of the input for **"BeamForm-ExtractCurrent"** is given by **"BeamCurr-SKA.in"** :

```

1  &BeamPars
2  Base = "SKA/Crs-" ! 48 -- 300 MHz
3  DataBase = "SKA/"
4  KernelBase = "SKX/Coarse" ! 48 -- 300 MHz
5  AntennaLayout = .false., DistMax = 500.
6  zeta_K= 0.02, 0.1, 0.2, .500, 1.500, 2.500, 3.500, 4.500, 5.500, 7.000, 8.500, 10.000, 12., 14., 17., 20., -1.
7  GoodQual=5.
8  FracFold=0.4
9  &end

```

Here the mandatory lines **" &BeamPars"** and **" &end"** specify the beginning and the end of the namelist input. The lines in between can be put in arbitrary order or omitted, in which case a default value is used (when so defined). All character on a line after an exclamation mark, !, are treated as comments. The output file **"BeamForm-GenK.out"** contains the complete listing of the possible namelist variables and their values for this particular run. Capitalization of the input parameters is not important.

- **"Base="** The directory where all output will be written. It may be followed by some initial characters to distinguish different runs.
- **"DataBase="** The directory where the antenna info was written.
- **"KernelBase="** The directory where the Kernel info was written.
- **"AntennaLayout="** To switch between specifying all antenna locations (.true.) or to assume they lie on a regular grid in distance to the shower core (.false.). When antenna locations are specified the core location for the shower will be calculated. The option **"AntennaLayout=.t."** has not (yet) been tested well.
- **"zeta_K="** The (irregular) grid heights (along the shower axis) for the beamforming calculations, in [km], ending with a negative number.
- **"GoodQual="** If the quality indicator is worse than **"GoodQual"** a fit to the current profile is attempted with some more parameters (see subroutine **"MultFit_ShSrucl"** for more details).
- **"FracFold="** Parameter used in defining the grid for the piece-wise linear fit for the current (see subroutine **"UnFoldJ"** for more details).

- **"DistMax="** Maximal distance for which antennas will be included in the calculation. Will be set automatically is antenna layout is read-in (**"AntennaLayout=.t."**)

For more details on the used formulas, definitions and the info on the generated plot you are advised to look at the paper, Ref. [1] which is following. For all plots the GLE package [2] is used, which -I think- is part of the Ubuntu Linux package.

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II. BASIC FORMULATION USING MGMR3D

In MGMR3D, a cosmic-ray air shower is modeled as a dynamic three-dimensional charge-current cloud moving towards the ground with the speed of light. The radio emission of this cloud is calculated through the application of Maxwell's equations, see Ref. [3] for an in-depth discussion. This guarantees that all relativistic effects and all coherence effects, linked to the finite extent of the charge-current, are properly accounted for. Much effort has been devoted to the parametrization of this charge current cloud [3, 4], where of particular interest for this work is the transverse four-current density, where the space component is written as

$$\rho_x(D_c, h, r_s) = J_x(D_c) f(h) w(r_s) . \quad (\text{II.1})$$

In Eq. (II.1), \mathbf{x} denotes the direction of the current transverse to the shower axis, given by \mathbf{v} , i.e. each of the two directions $\mathbf{x} = \mathbf{v} \times \mathbf{B}$ and $\mathbf{y} = \mathbf{v} \times (\mathbf{v} \times \mathbf{B})$, where \mathbf{B} is the orientation of the geomagnetic field; D_c denotes the geometric distance of the shower front to the impact point on the ground and is measured along the shower axis, as is indicated in Fig. 1b; the shower reaches ground at $t = 0$ and $J_x(D_c)$ denotes the total current in the two transverse directions when the shower front is at D_c , i.e., at time $t = -D_c/c$; the distance behind the shower front is given by h , where $f(h)$ denotes the charged-particle density in the air shower with $f(h < 0) = 0$; the radial distribution of the current density is given by $w(r_s)$, where $r_s = 0$ is on the shower axis and cylindrical symmetry is assumed. The speed of light in vacuum is denoted by c , and the index of refraction of air is height-dependent and denoted by n with derivative n' . In this work, we assume that the antennas (observers) are in the shower plane, taken perpendicular to the shower axis. Note that for air showers developing under fair-weather conditions, we generally have that the current in the $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$ direction is vanishingly small, and it is thus sufficient to concentrate on $\mathbf{x} = \mathbf{v} \times \mathbf{B}$.

The dominant process for radio emission from air showers is driven by electric currents that are concentrated at the air shower front and are transverse to the shower axis. For fair-weather showers, these currents are induced by the geomagnetic field with an orientation given by the Lorentz force, $\mathbf{v} \times \mathbf{B}$. Under thunderstorm conditions, the strong atmospheric electric fields induce transverse currents that are not aligned with $\mathbf{v} \times \mathbf{B}$ and may change direction with distance along the shower axis. Additionally, there is radio emission due to the fact that there is a net charge excess in the air shower inducing Askaryan radiation [5] which is sub-dominant. As discussed in Section VII, we ignore the charge excess radiation in this work. Thus, only the projection of the electric field on the shower plane contributes to the transverse current.

Limiting to transverse-current emission, the field in an antenna at distance r_a from the core is given by

$$E^x(t_a; r_a) = \int dh dr_s^2 \frac{\mathcal{S}^x(D_c, h, r_s)}{\mathcal{D}} , \quad (\text{II.2})$$

with

$$\mathcal{D} = (t_a c + D_c) \left| \frac{dt_a}{dD_c} \right| = n |R - n\zeta - n' R^2| , \quad (\text{II.3})$$

where we have introduced $\zeta = D_c + h$ as the emission distance of a signal. Using Eq. (19) from Ref. [6], transcribed in the notation used in this work,

$$E_x(t, d) = \int dh dr_s^2 \left(J_x(D_c) \frac{df(h)}{dh} - f(h) \frac{dJ_x(D_c)}{dD_c} \right) \frac{w(r_s)}{\mathcal{D}} , \quad (\text{II.4})$$

we see that the source term can be written as

$$\begin{aligned} \mathcal{S}^x(D_c, h, r_s) &= \frac{d\rho_x(D_c, h, r_s)}{dh} \\ &= \left(J_x(D_c) \frac{df(h)}{dh} - f(h) \frac{dJ_x(D_c)}{dD_c} \right) w(r_s) . \end{aligned} \quad (\text{II.5})$$

The arrival time in the antenna t_a , shower-front distance D_c , and emission distance ζ are related by causality which can be written as

$$0 = t_a c - D_c + n_\zeta R_\zeta , \quad (\text{II.6})$$

with $n_\zeta R_\zeta = n(\zeta) \sqrt{\zeta^2 + (\vec{r}_a - \vec{r}_s)^2}$, the optical distance from the source point to the antenna. The denominator, \mathcal{D} , in Eq. (II.2) results from reducing the Dirac delta function $\delta(L^\mu L_\mu)$ where L_μ is the optical path [3, 7].

For the purpose of this work, it is essential to rewrite the integration in Eq. (II.2) as

$$\begin{aligned} E^x(t_a; \vec{r}_a) &= \int d\zeta \int dr_s^2 \mathcal{S}^x(D_c, h, r_s) \frac{1}{n_\zeta R_\zeta} \\ &= \int dD_c J_x(D_c) \int dr_s^2 \frac{\frac{df(h)}{dh} w(r_s)}{n_\zeta^2(h + D_c) - n_\zeta n'_\zeta R_\zeta^2}, \end{aligned} \quad (\text{II.7})$$

where we used $\int dh = \int |dh/d\zeta| d\zeta$ with $|dh/d\zeta| = \mathcal{D}/(n_\zeta R_\zeta)$, and assume the causality relation Eq. (II.6), changed integration variable in the second step, and have omitted the second term in Eq. (II.5) since the first is dominant by four orders of magnitude.

III. BEAMFORMING

By reverse propagation of the signals in the antennas, Eq. (II.2), to a focal point at distance D_b on the shower axis, we construct the beamforming trace. The trace depends on t_b ,

$$\begin{aligned} \mathcal{R}^x(t_b; D_b) &= \iint E^x(t'; r_a) \frac{\mathcal{F} \delta(G)}{n_b R_b} dt' d^2 r_a \\ &= \int E^x(t_a; r_a) \frac{\mathcal{F}}{n_b R_b} d^2 r_a, \end{aligned} \quad (\text{III.1})$$

where the Dirac delta-function factor, $\delta(G)$, determines the relation between t_a and the antenna position \vec{r}_a ,

$$0 = G(D_b, t_b; t_a, \vec{r}_a) = t_b c - t_a c - D_b + n_b R_b, \quad (\text{III.2})$$

with $n_b R_b = n(D_b) \sqrt{D_b^2 + \vec{r}_a^2}$. In principle, an arbitrary function $\mathcal{F}(D_b, t_b, t_a, r_a)$ can be added inside the integral but for ease of writing we put $\mathcal{F} = 1$. $t_b = 0$ corresponds to the time when the shower front passes distance D_b while $t_a = 0$ corresponds to the time when the shower front reaches ground.

In actual calculations, \mathcal{F} is written as a sum of delta-functions, changing the integral over antenna positions to a sum over discrete antenna locations. In principle, a weighting factor could be included to give more weight to antennas with a large signal-over-background ratio. See also the remarks at the end of Section VII.

In the examples presented here, we have limited ourselves to beamforming on the shower axis. Apart from simplifying the expressions, this yields the most sensitive measure of the longitudinal shower profile we are interested in. Off-axis beamforming (not reported on in this work) yields traces that are considerably different, which can be used to determine the position of the shower axis, as shown in [8].

Substituting Eq. (II.7) in Eq. (III.1) we obtain the central equation of the work,

$$\mathcal{R}^x(t_b; D_b) = \int J_x(D_c) \mathcal{K}(t_b; D_b, D_c) dD_c, \quad (\text{III.3})$$

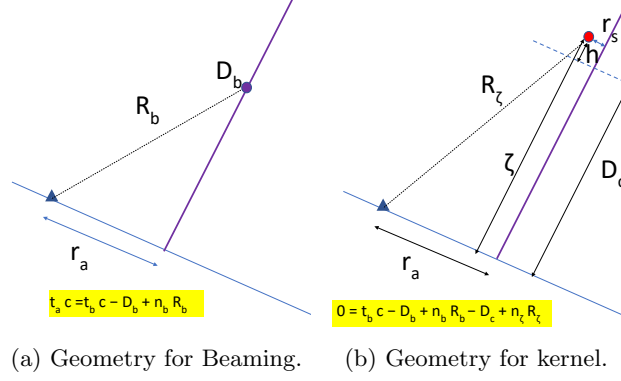
with the kernel

$$\begin{aligned} \mathcal{K}(t_b; D_b, D_c) &= \int \frac{w(r_s) df(h)}{n_b R_b dh} \times \\ &\quad \frac{1}{n_\zeta^2(h + D_c) - n_\zeta n'_\zeta R_\zeta^2} d^2 r_a d^2 r_s, \end{aligned} \quad (\text{III.4})$$

where the use of the causality relation, Eq. (III.2), is understood.

The variables used in the integrations are schematically indicated in Fig. 1. To evaluate the Kernel, we have to find the distance h relative to the shower front located at position D_c . For this, we use the focal point location parameters D_b , n_b , and R_b and time t_b to obtain the electric field evaluation time t_a from Eq. (III.2). Having obtained t_a , the distance h relative to the shower front, positioned at D_c , is found through the causality condition presented in Eq. (II.6).

The kernel depends explicitly on the antenna positions as expressed by $\int d^2 r_a$, which may be replaced by a sum over antenna position as noted earlier. The kernel also depends on the shape of the charge cloud, through the functions $f(h)$ and $w(r_s)$ as it moves to ground. As shown in Ref. [4], these functions are generic for a large variety of showers. These functions will probably not apply to very inclined showers, where also the assumption of cylindrical symmetry is broken. Cylindrical symmetry is assumed for calculational simplicity, but the approach can easily be generalized.



(a) Geometry for Beamforming. (b) Geometry for kernel.

FIG. 1: The geometry used for calculating the beamforming trace $\mathcal{R}^x(t_b; D_b)$ and the kernel $\mathcal{K}(t_b; D_b, D_c)$. Note that not all indicated parameters are independent since they have to obey the causality relations Eq. (III.2) and Eq. (II.6), as marked in yellow in the figure.

In future work, it needs to be investigated to what extent these assumptions hamper the applicability of the unfolding procedure as discussed in Section V. First indications are that these assumptions do not seriously affect the results since in the case of the thunderstorm examples the structure of $f(h)$ used in the model simulation of the ‘data’ differs from that used in the calculation of the kernel used to reconstruct the current profiles.

In principle, the function $w(r_s)$ could, after a re-ordering of the integration order in Eq. (III.3) and Eq. (III.4), be determined following a similar unfolding procedure as discussed in Section V. First results (not discussed in this work) indicate that it will be difficult to reach a sufficient accuracy as the emission very close to the shower axis tends to dominate strongly. This can be taken as an argument that thus the detailed form of $w(r_s)$ is not essential. This will be the subject of a future work.

IV. NUMERICAL CALCULATION KERNEL

To get some insight in the structure of the kernel, we investigate it through numerical calculations. We show the effects of the limited range of antennas as well as the effects of frequency filtering. In all our calculations, we have taken the shower at a rather arbitrarily chosen angle of 43° from the zenith. The precise value is not important for any aspect of the discussions presented in this work and neither the strength and relative angle of the magnetic field, as all fields and currents are expressed in arbitrary units.

At a fixed beamforming distance D_b , the structure of the kernel depends on the layout of the antennas through the explicit integral over antenna position \vec{r}_a as given in Eq. (III.4). Fig. 2 shows the structure of the kernel for the case where $D_b = 5.5$ km and the antennas cover a distance range of 250 m or 500 m from the shower axis in the shower plane (the shower core), i.e taking $\mathcal{F}(\vec{r}_a)$ equal to a sum of delta functions in $|\vec{r}_a|$ while keeping the azimuthal integration. A very sharp peak is shown for $t_b \approx 0$ at $D_c \approx D_b$ that extends over a range of about 2 km. For this study, the antennas are placed in continuous concentric rings at distances of 10 m. Except for small values of $D_b < 1$ km, the details of the layout are not important as long as the antennas are spread evenly covering the same area. The structure of the kernel is calculated using a time-step of $0.1 \text{ m} = 0.03 \text{ ns}$ in t_b .

The structure of the kernel is -at first sight- very similar for the two distance ranges of 250 m and 500 m as can be seen from Fig. 2. Both show a rather sharp positive-valued spike at $t_b = 0$ and $D_c = D_b = 5.5$ km with large wing-like structures extending to larger and smaller values of D_c . This tells immediately that for both cases the beamforming signal at distance D_b is influenced by the current in the shower from a large range of distances, D_c , extending well below and above D_b . The spike at $D_c = D_b$ is positive and followed by a negative (dark blue) tail extending to large values of t_b because it is formed by the coherent addition of pulses from all antennas that have the same basic structure. On closer inspection, one sees some important differences. One is that the peak in intensity has become much sharper for the 500 m case. The other is that for values of $D_c \neq D_b$ the ‘wings’ show differently in the two cases. For $D_c < D_b$ a negative-valued wing extends to larger positive values of t_b for a range of 500 m than is the case for 250 m. Thus, for antennas covering a smaller range, a current at lower altitudes gives rise to a kernel where the zero-crossing occurs for smaller t_b . This can be understood by the fact that the main positive part of the pulse emitted by a current at low altitudes, arrives relatively later in distant antennas than the pulse emitted from $D_c = D_b$. All these positive contributions from the more distant antennas will thus extend the positive part of the beamforming amplitude to larger t_b . Conversely, for $D_c > D_b$ the emission arrives earlier in the distant antennas

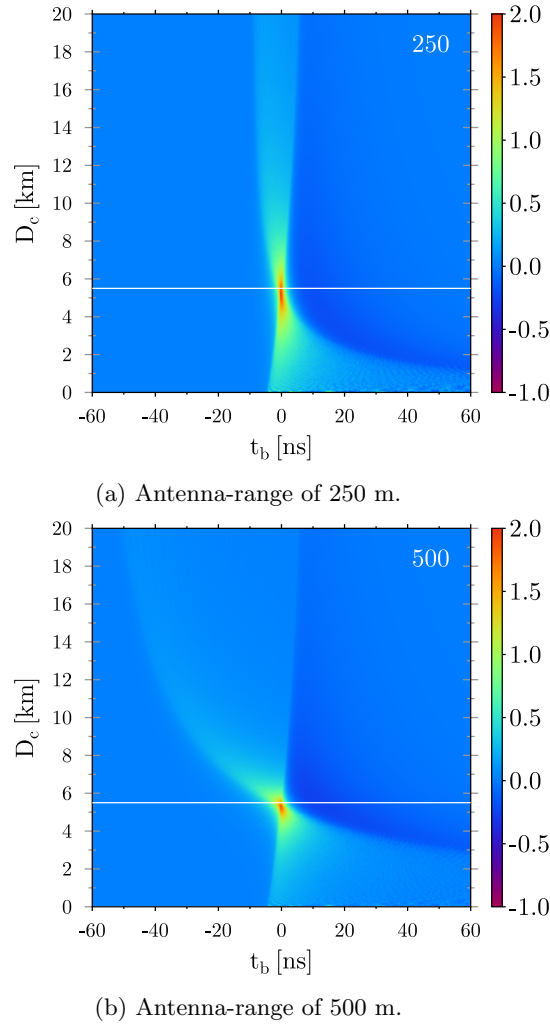
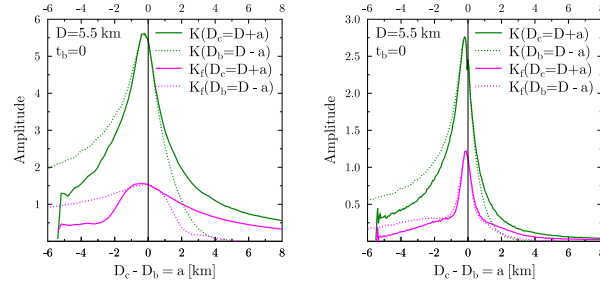


FIG. 2: The Kernel $\mathcal{K}(t_b; D_b, D_c)$, see Eq. (III.4), as calculated for the two different ranges of the maximal antenna distance to the shower core in the shower plane. The beaming distance is $D_b = 5.5$ km for both panes where $D_b = D_c$ is indicated by the horizontal white line. The color scale is in arbitrary units.

than the pulse emitted at $D_c = D_b$, thus giving rise to the much more pronounced wing extending to negative values of t_b , as seen by comparing Fig. 2b with Fig. 2a.

A more quantitative impression of the effects of the antenna range on the spreading width of the kernel is obtained from Fig. 3 where we show the dependence of the kernel $\mathcal{K}(t_b; D_b, D_c)$ at $t_b = 0$ around the point $D_c = D_b = D = 5.5$ km by varying $D_c - D_b = a$. The green curves show the results for the unfiltered kernel, $\mathcal{K}(t_b = 0; D_b = D - a, D_c = D)$ and $\mathcal{K}(t_b = 0; D_b = D, D_c = D + a)$. The magenta curves, labeled as \mathcal{K}_f show the filtered kernel using a 30-80 MHz block filter for the antenna range of 250 m (left panel) and 50-300 MHz for 500 m (right) and will be discussed in a following paragraph. This figure shows clearly that when the antenna array covers a larger range, Fig. 3b, the distribution is much more sharply peaked near $D_c = D_b$ than for a smaller antenna range, Fig. 3a, precisely as one would expect.

The distinct offset of the peak from $D_b = D_c$ is due to the fact that we have taken $t_b = 0$ in Fig. 3. When beaming at a distance $D_b = D = 5.5$ km, one ‘sees’ the maximum of the current which is some distance behind the shower front. The shower front thus has reached a slightly lower altitude $D_c < D_b$ (solid curve) due to the finite pancake thickness. Conversely, when the front of the shower just reached a distance of $D_c = D = 5.5$ km, the current at time $t = -Dc$ ($t_b = 0$) peaks at slightly larger altitudes and thus ‘seen’ when $D_b > D_c$ (dotted curve). For the quantitative relation between this shift and the pancake thickness, one should include relativistic beaming effects. For slightly larger values for t_b the peak is positioned at $D_b = D_c$ (not shown). Since the integrand for the calculation of the kernel is very non-smooth, due to the sharpness of the shower front near the core, a very fine integration grid was used to obtain Fig. 4. With a more coarse integration spurious structures may show that do not affect the final result



(a) Antenna-range of 250 m. (b) Antenna-range of 500 m.

FIG. 3: The value of the kernel $\mathcal{K}(t_b; D_b, D_c)$ (in arbitrary units) at $t_b = 0$ as function of $D_c = D + a$, keeping $D_b = D = 5.5$ km fixed (solid curves) and also as function of $D_b = D - a$, keeping $D_c = D = 5.5$ km fixed (dotted curves) for two different antenna layouts. The curves in green show the unfiltered kernels, while for the magenta curves in the left (right) panel a 30-80 MHz (50 – 300 MHz) block filter is applied.

when frequency filters are applied.

The magenta curves labeled by K_f in Fig. 3 display the kernel when semi-realistic frequency filters are applied spanning a frequency range one may encounter in realistic scenarios. Reminiscent of LOFAR is the 30 – 80 MHz block-filter [9] applied to the case where antennas up to a range of 250 m, case Fig. 3a, are included. Reminiscent of an SKA-like scenario is a block-filter of 50 – 300 MHz [10] for the case of Fig. 3b. It is seen by comparing the green and magenta curves in Fig. 3b that for the larger antenna range of 500 m the application of a 50 – 300 MHz block-filter does not greatly affect the peaked structure of the kernel. However, for the more limited antenna range of 250 m, applying the 30 – 80 MHz block-filter results in a considerable broadening of the peak in the kernel. The main reason for this difference is the effects of the antenna range as can also be deduced from Fig. 2, where the larger antenna range results in a very broad (in t_b) beaming trace for $D_b \neq D_c$ as compared to $D_b \approx D_c$ and thus a very fast drop in the value of the filtered kernel.

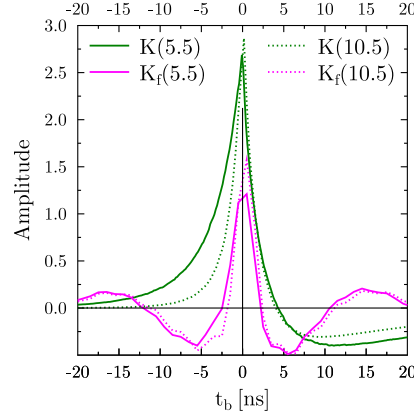


FIG. 4: The value of the kernel at $D_c = D_b = 5.5$ km (solid curves) and at 10.5 km (dotted curves) as function of $t_b = 0$. The curves labeled by K_f (magenta) display the structure of the kernel when a block frequency filter of 50 – 300 MHz is applied.

It can be seen already from Fig. 2 that the kernel as a function of t_b exhibits a sharp peak at $t_b \approx 0$ for the case that $D_b \approx D_c$ which gradually becomes less pronounced when the difference between D_b and D_c increases. As is discussed in following sections, this feature will allow to improve the sensitivity of the beaming signal to the shower profile. To this end we show in Fig. 4 the structure of the peak in the kernel at $D_b = D_c$ for two different distances of 5.5 km (same distance as in the previous figures) and 10.5 km, where we have limited ourselves to the case where antenna range is 500 m from the core. The unfiltered kernel (at a resolution of 0.3 ns) shows a clear peak at $t_b = 0$ that is sharper for higher altitudes. The peak is sufficiently sharp that when applying a 50 – 300 MHz block-filter the resulting time trace resembles closely a sinc-function (not shown) structure, the filtered response to a delta-function, very similar for the two altitudes. For future reference, we name the filtered peak response calculated for the highest

altitude the Pulse Kernel Response (PKR),

$$PKR(t_b) = N \mathcal{K}(t_b; D_b = D, D_c = D), \quad (\text{IV.1})$$

where D equals the largest distance considered in the calculation and where N is a normalisation constant such that, $\int |PKR(t_b)|^2 dt_b = 1$. As shown in Fig. 4, the structure of $\mathcal{K}(t_b; D_b = D, D_c = D)$ does not vary much with D , certainly after applying a frequency filter, and the precise value of D is not very relevant. For distances smaller than $D \approx 1$ km, the trace $\mathcal{K}(t_b; D_b = D, D_c = D)$ starts to differ significantly from the dependence shown in Fig. 4, where we have not investigated in detail why this is the case. For this reason the approach is simplified by selecting a single time trace, using the same for all heights.

V. EXTRACTION OF CURRENT PROFILE

In this section we discuss a few different approaches to extract the longitudinal current profile from the beamed antenna traces using the Kernel as discussed in the previous sections. The effectiveness of these approaches will be shown by using MGMR to generate signal traces for all antennas that are subsequently used in beamforming. For the different approaches the extracted longitudinal profile is compared to the one that was used to generate the antenna signals. For fair-weather events, where the current profile has a very simple structure, various extraction methods yield comparable results. A more severe test case is formed by an air shower where the current profile is strongly influenced by atmospheric electric fields since the direction of the transverse currents may vary greatly with altitude and may even be opposite to that for the fair weather case.

The first step in all approaches is to condense the beamforming trace $\mathcal{R}^x(t_b; D_b)$, calculated by beamforming the E-field traces of all antennas, into an amplitude. There are two basically different approaches possible. In the first approach, we extract the beamforming amplitude, $\mathcal{B}^x(D_b)$, as the square root of the beamforming power,

$$\mathcal{B}^x(D_b) = \sqrt{\sum_{t_b} |\mathcal{R}^x(t_b; D_b)|^2}. \quad (\text{V.1})$$

This approach is rather robust since it is not sensitive to an offset in t_b . In the second approach, we use the same beamforming trace $\mathcal{R}^x(t_b; D_b)$ and cross-correlate it with a pre-defined function, $PKR(t_b)$, that enhances the importance of the part near $t_b = 0$ and may be defined as in Eq. (IV.1). An alternative might be to take a filtered δ -function at $t_b = 0$. This yields the PKR-correlated beaming amplitude

$$\mathcal{P}^x(D_b) = \sum_{t_b} \mathcal{R}^x(t_b; D_b) PKR(t_b). \quad (\text{V.2})$$

In this approach, care should be taken to correct for any offsets in determining t_b which could be done by finding the offset in t_b that maximizes \mathcal{P}^x (not done in this work). By enhancing the part of the beamforming trace at $t_b = 0$, the contribution near $D_b = D_c$ is increased. An additional advantage is that the information on the sign of $\mathcal{R}^x(t_b; D_b)$ is kept, which is lost in the procedure of Eq. (V.1). An other advantage of the PKR-correlation approach is that it suppresses the effects of noise in the data (not considered in the present work), while for $\mathcal{B}^x(D_b)$ one would need an explicit correction. In principle, it is possible to keep the full beamforming trace $\mathcal{R}^x(t_b; D_b)$ and not reduce it to an amplitude, but in this exploratory work it is more insightful to explore simplifications.

The second step is the extraction of the longitudinal current profile from the beamforming amplitude. As discussed in the following two sections, this may be achieved following two different approaches, all based on the use of Eq. (III.3). In one an analytic parametrization is used, while in the other a more agnostic parametrization. Both approaches are discussed extensively in the following two sections.

To perform more detailed tests of the ‘data’, the E-field traces of all antennas, are generated from an MGMR3D simulation for this work. The beamforming is performed separately for the $\mathbf{x} = \mathbf{v} \times \mathbf{B}$ and $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$ polarization directions. For simplicity, it is assumed in the present study that the antennas are arranged in densely-packed concentric rings in the shower plane centered at the shower core at 10 m separation. The data-traces are time sampled as would be the case in a real measurement. It is essential that the kernel, $\mathcal{K}(t_b; D_b, D_c)$ see Eq. (III.4), is calculated for the same antenna layout as used in the ‘data’, using the same frequency filtering and the same sampling time for t_b . The distances D_b and D_c are calculated for a 50 m grid spacing. The kernel depends on the geometry, i.e. shower angle and antenna layout, while the simulated E-fields, the ‘data’, depend also on the shower profile. We use the fact that the structure of the charge-current cloud, in MGMR3D defined by the functions $f(h)$ and $w(r_s)$, are rather universal shower parameters. In the discussion in Section VII we will return to this point.

A. Using the beamforming amplitude

To extract the current profile from the beamforming amplitude, determined from the measured electric fields in the antennas, we use a very general parametrization of the current profile $J_x(D_c)$ and optimize the parameters by minimizing

$$\chi^2 = \sum_{D_b} \left(\frac{\mathcal{B}_K^x(D_b) - \mathcal{B}^x(D_b)}{\sigma_B} \right)^2, \quad (\text{V.3})$$

using a steepest descent method where $\mathcal{B}^x(D_b)$ is the beamforming amplitude constructed from the data as defined in Eq. (V.1), σ_B the estimate of the uncertainty and where

$$\mathcal{B}_K^x(D_b) = \sqrt{\sum_{t_b} |\mathcal{R}_K^x(t_b; D_b)|^2} \quad (\text{V.4})$$

is the modeled beamforming amplitude where the modeled beamforming trace is obtained from a simple folding procedure,

$$\mathcal{R}_K^x(t_b; D_b) = \sum_{D_c} J_x(D_c) \mathcal{K}(t_b; D_b, D_c). \quad (\text{V.5})$$

Since for the present examples we work with simulated data we have put $\sigma_B = 1$. Note that one should contract the current with the kernel before taking the the sum of the squares. Reversing this order would allow to simplify the procedure, however, introduces some approximations that make little difference for fair-weather showers but are severe for air showers under thunderstorm conditions. We will not pursue such an approach in this work.

In an alternative approach, one may minimize the root-mean-square differences between the traces instead of the amplitudes. This will probably improve the sensitivity of the approach, however, we have not pursued such an approach.

We have used a parameterization of the current profile based on the Gaisser-Hillas formula [11–13], where we have opted for a generalized parametrization in terms of the R and L parameters,

$$J^x(X) = \sum_{m=1}^{N_I} I_m^x \left(1 - R^x \frac{X_m^x - X}{L^x} \right)^{R^x - 2} e^{\frac{X_m^x - X}{L^x R}}, \quad (\text{V.6})$$

where I_m^x is the current strength at penetration depth $X = X_m^x$ and N_I typically equals 1 or 2 for fair weather showers. $N_I = 2$ allows for ‘double bump’ showers of the kind shown in Fig. 9 of Ref. [4], where the longitudinal shower profile may have more than one clear maximum. As a simplification, mainly to limit the number of parameters, we have taken the values for the R^x and L^x parameters independent of m . The parameters may differ for the two polarization directions, where for fair weather showers we take $\mathbf{x} = \mathbf{v} \times \mathbf{B}$ exclusively. The penetration depth and distance to ground are related by the structure of the atmosphere (taken equal to the US standard atmosphere parametrized by Linsley [14]) and the zenith angle of the cosmic ray taken equal to 43° , quite arbitrarily.

As an example, the results for the reconstructed currents, using Eq. (V.3) with Eq. (V.6) are shown in Fig. 5 using the more limited range of 250 m for the antennas and a 30 – 80 MHz block filter for calculating the fields. The top panel gives the beamforming amplitudes that show very limited structure towards the larger distances for all five values of X_{max} , the penetration depth at which the number of charged particles in the shower reached a maximum, ranging from 700 – 1100 g/cm^2 . For comparison, $X_{max} = 800 g/cm^2$ corresponds to $D = 6.2$ km and a Cherenkov radius of about 90 m in the antenna plane. The solid lines in the bottom panel show the longitudinal current profiles for the five different cases, while the dashed lines show the reconstructed profiles using $N_I = 1$ in Eq. (V.6) thus fitting the four parameters, I_1^x , X_1^x , R^x and L^x . Due to the fact that the filtered kernel is rather structure-less, as can be seen from Fig. 3a, this number of parameters is close to the limit that can be extracted, even for this rather idealized case where no realistic noise has been added. Increasing the number of parameters to 8 by taking $N_I = 3$ in Eq. (V.6) results in clear over fitting, as displayed by the dotted curve in the bottom panel for $X_{max} = 900 g/cm^2$. The fit to the beamforming amplitudes, a dotted curve in the top panel of Fig. 5 is indistinguishable from the one generated for $N_I = 1$ (dashed-dotted).

B. Parametrized current profile using the PKR

An alternative approach is to cross-correlate the beamforming trace, obtained from the data, with the pulse-shape function given in Eq. (IV.1) yielding the PKR-correlated beaming amplitude as was defined in Eq. (V.2),

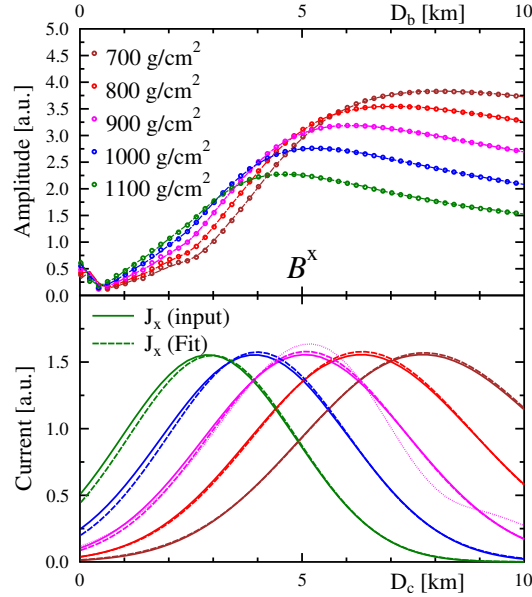


FIG. 5: The top panel gives the beamforming amplitude $\mathcal{B}^x(D_b)$ as defined in Eq. (V.1) for five showers with different values for X_{max} as indicated in the legend. The bottom panel displays the longitudinal current profile $J_x(D_c)$ for the five different showers (solid curves) as well as the extracted profile based on using Eq. (V.3) (dashed lines) using a single Gaisser-Hillas profile ($N_I = 1$ in Eq. (V.6)). The dotted curve shows for the case of $X_{max} = 900 \text{ g/cm}^2$ the result of the procedure using Eq. (V.6) with $N_I = 3$. Antennas up to a range of 250 m have been included as well as a 30 – 80 MHz block filter.

$\mathcal{P}^x(D_b) = \sum_{t_b} \mathcal{R}^x(t_b; D_b) PKR(t_b)$. Performing the same operation on the kernel, we obtain a weighting function

$$\mathcal{W}(D_b, D_c) = \sum_{t_b} \mathcal{K}(t_b; D_b, D_c) PKR(t_b). \quad (\text{V.7})$$

Following a similar approach as in the previous section, we obtain the current profile $J^x(D_c)$ by minimizing

$$\chi^2 = \sum_{D_b} \left| \frac{[\sum_{D_c} \mathcal{W}(D_b, D_c) J^x(D_c)] - \mathcal{P}^x(D_b)}{\sigma_P} \right|^2, \quad (\text{V.8})$$

where we have put $\sigma_P = 1$ as is more appropriate for a model-to-model comparison. This expression is numerically much simpler than the equivalent expression Eq. (V.3) since it is at most quadratic in $J^x(D_c)$.

To solve for $J^x(D_c)$, we have investigated two different parameterizations of the current profile, one based on the Gaisser-Hillas parametrization, Eq. (V.6). As was done in Section V A, we will use a steepest descent method to solve for the minimal χ^2 value.

A second solution method is using a piece-wise linear (PWL) parametrization, where the current is parametrized by its values on a grid with a linear interpolation for the in-between points. With this parametrization the expression for χ^2 , Eq. (V.8), reduces to Eq. (A.3) where the derivatives of the χ^2 are linear in the PWL-parameters. The minimization condition can thus be written as a matrix multiplication that can be solved analytically, see Appendix A, resulting in

$$J_i = A^{-1}B, \quad (\text{V.9})$$

where i labels the grid points, and the matrices A and B are defined in Eq. (A.6) and Eq. (A.5). The number of parameters for this case is equal to the grid points for defining the current.

Fig. 6 shows the results for the extracted current profiles using the PKR-correlation approach. Using the analytic parametrization for extracting the current profile gives results that closely match the the current profile used for calculating the E-field traces in the different antennas. The quality of the agreement is similar to the the beamforming-amplitude approach, shown in Fig. 5.

The currents can also be extracted using the PWL parametrization. This has the advantage that it contains no assumptions concerning the structure of the current profile, and additionally that the parameters can be extracted using

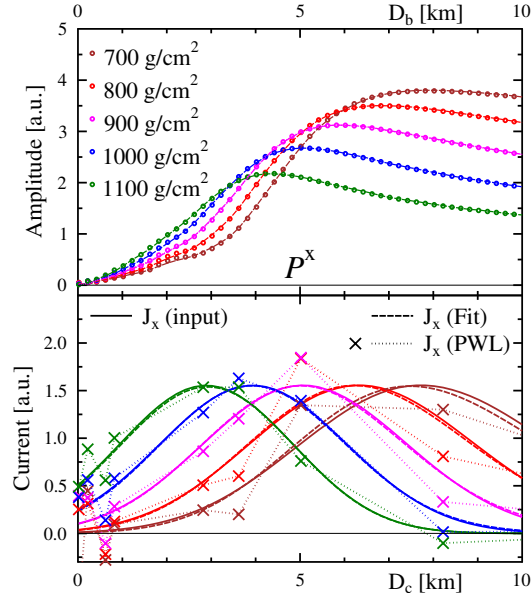


FIG. 6: The top panel gives the PKR-correlated beamforming amplitude $\mathcal{P}^x(D_b)$, as defined in Eq. (V.2), for five showers with different values of X_{max} as indicated in the legend. The bottom panel displays the longitudinal current profile $J_x(D_c)$ for the five different showers (solid curves) as well as the the extracted profile based on using Eq. (V.8) (dashed lines) using a single Gaisser-Hillas profile. The dotted lines show the results of the PWL calculation, using Eq. (V.9). Antennas up to a range of 250 m have been included as well as a 30 – 80 MHz block filter. The ordinate is in arbitrary units for both panels.

the analytic procedure of Eq. (V.9). The major disadvantage, as is shown clearly in Fig. 6, is that it is prone to over-fitting resulting in saw-tooth patterns. In the PWL results shown in Fig. 6, we have used a grid that is more dense at smaller distances where the weight functions Eq. (V.7) have a more pronounced structure. In all cases the fits of the Gaisser-Hillas as well as the PWL fits to the PKR-correlated beamforming amplitudes, $\mathcal{P}^x(D_b)$, shown in the top panel of Fig. 6, are almost indistinguishable. Care should be taken in choosing the PWL grid points; A too dense grid will give rise to results that show large zigzagging around the correct average. The PWL grid points in Fig. 6 are denoted by the crosses and have been taken more dense for values of D_c where the kernel shows more structure. The PWL fitting procedure can probably be improved by taking a more sophisticated procedure for selecting the grid and devising some kind of penalty procedure for a strongly fluctuating result. The deviations seen in Fig. 6 appear to be dependent on X_{max} and it thus might be that a more sophisticated procedure for the grid points, where the variation in the PKR-correlated beamforming amplitude is taken into account, is worth investigating.

Applying the two extraction procedures to the case where the antennas cover a larger range and a larger bandwidth, shown in Fig. 7, gives results that are qualitatively very similar to those shown in Fig. 6 although the results using the PWL-parametrization show considerably less scatter. This is a reflection of the fact that the spreading-width of the kernel for the larger antenna range is considerably smaller than for the small antenna range, as was shown in Fig. 2.

The most stringent test of the current profile extraction is for the case in which the shower develops in an atmosphere with strong electric fields as is the case under thunderstorm conditions. As mentioned in the second paragraph of Section II, we are mainly sensitive to the component of the electric field that is transverse to the shower axis. To investigate this more complicated shower profile, we have performed a simulation using a typical three-layered structure for the atmospheric electric field as found in Refs [15–17]. Specifically, we have used an atmospheric electric field in the top layer, between distances to the shower core of 8.1 and 6.2 km, with a projection in the shower plane of 103 kV/m at an angle of $\alpha = -77^\circ$ with respect to the $\mathbf{x} = \mathbf{v} \times \mathbf{B}$ direction, a middle layer between 6.2 and 5 km distance, with a projection in the shower plane of 100 kV/m at $\alpha = 66^\circ$, and a bottom layer from 5 km to the ground with a field of 108 kV/m at $\alpha = 118^\circ$ in the shower plane. For the cosmic ray, we have taken the same geometry as for the other examples in this work with $X_{max} = 900$ g/cm² (corresponding to $D = 5$ km). In order to extract the current profiles, we thus have used the same kernels as used in the previous examples.

The PKR-correlated beamforming amplitudes, as well as the extracted currents, are shown in Fig. 8. Since the atmospheric electric fields vary greatly in angle in the three different layers the currents are oriented in rather different directions transverse to the shower axis resulting in non-vanishing beaming traces extracted for the $\mathbf{v} \times \mathbf{B}$ as well as the $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$ polarization directions. Also notable is that \mathcal{P} is no longer positive definite as was the case for fair-

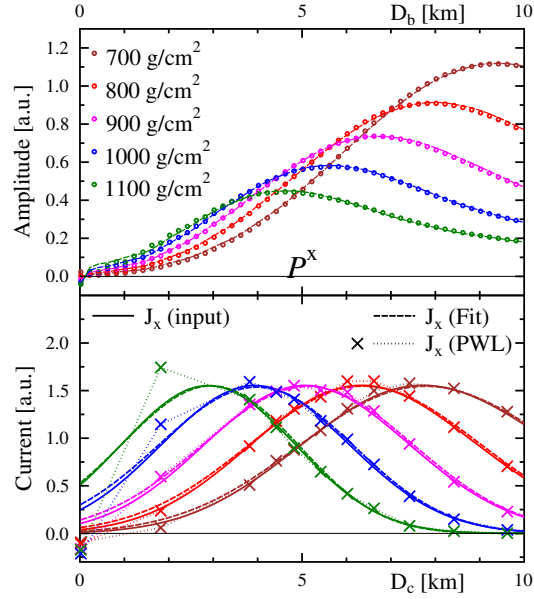


FIG. 7: Same as Fig. 6 but for the case where the antenna range is 500 m and a 50 – 300 MHz block filter is used.

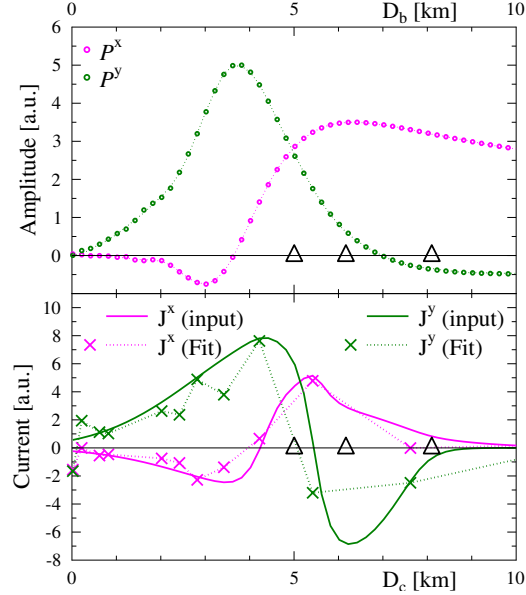


FIG. 8: The top panel gives the PKR-correlated beamforming amplitude, Eq. (V.2) for $\mathbf{x} = \mathbf{v} \times \mathbf{B}$ and $\mathbf{y} = \mathbf{v} \times (\mathbf{v} \times \mathbf{B})$ polarization directions for a cosmic-ray shower developing under influence of strong atmospheric electric fields as discussed in the text. The bottom panel displays the longitudinal profiles for the currents in the two directions. The solid curves show the profiles as used in the simulations while the crosses mark the extracted results obtained by using the PWL parametrization. The triangles mark the boundaries of the E-field layers. Antennas up to a distance of 250 m have been included as well as a 30 – 80 MHz block filter.

weather showers. This is due to the fact that the current at some distances may be opposite to the direction of the Lorentz force, $\mathbf{v} \times \mathbf{B}$. This is indicating the importance of using \mathcal{P} as \mathcal{B} , calculated from the absolute square of the beamforming trace, is not sensitive to the sign of the current.

We observed that using Eq. (V.6) to extract the current profile leads to rather unstable results and this method will thus not be explored any further for the case of atmospheric electric fields. This instability of the fit using the Gaisser-Hillas parametrization for the current profile is due to the fact that the current profile has considerably more structure than a typical fair-weather profile. This structure is difficult to capture in the parametrization of Eq. (V.6) even using $N_I = 3$ and the fit gets stuck in a local minimum. A reasonable result can be obtained only by carefully

choosing the starting values in the fit and this approach is thus not applicable in general. We have not explored if there are other analytic parametrizations possible that yield more stable results. The PWL-parametrization offers, in contrast, very reliable results for this case as is seen in Fig. 8. Even though the extracted current profile shows scatter, it traces the main features well.

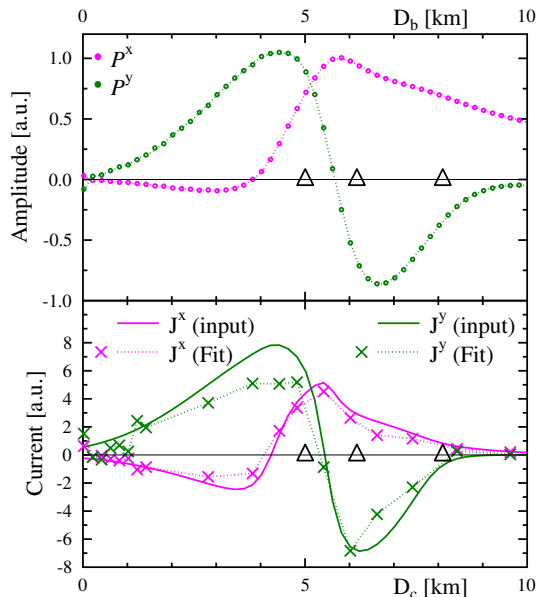


FIG. 9: Same as Fig. 8 but for the case where the antenna range is 500 m and a 50 – 300 MHz block filter is used.

Using a larger antenna range and a larger bandwidth gives results that are even more accurate as shown in Fig. 9. The current in the simulation (solid curves on the bottom panel) is the same as used in Fig. 8. With the increased range of antennas and the increased bandwidth, the sensitivity to the current profile at larger distances has greatly improved.

VI. SUMMARY AND DISCUSSION

One of the prime motives for radiometric observations of extensive air showers is to be able to extract the longitudinal intensity profile. Using beamforming, i.e. the coherent addition of the signals of all antennas, offers, potentially, a very sensitive method. However, due to the finite range of distances covered combined with the finite band-width of the antennas, the resolution of the procedure is hampered. We have investigated the effects of these aperture limitations. We have shown that these effects can be modeled using MGMR3D which gives an accurate description of the generic shower properties [4]. This resulted in the central equation of this paper, Eq. (II.7). Based on this, we presented different procedures that allow to unfold the effects of the limited aperture and reconstruct the longitudinal shower profiles in a macroscopic model-to-model comparison.

We have investigated two different methods for using beamforming, one where the total beamforming intensity is determined and an alternative where the beamforming trace is cross-correlated with a known response function. It is shown that both procedures allow for an accurate reconstruction of the longitudinal current profile using a generic Gaisser-Hillas parametrization that is known to give an accurate parametrization of the profile for fair-weather showers.

We have also explored a PWL parametrization of the current. This has the advantage that it is agnostic. For the case of fair weather showers, it appears more advantageous to implement in the fit that the current is varying rather smoothly with distance along the shower axis. For showers developing under the influence of (strong) atmospheric electric fields, as are present in thunderstorms, we find in contrast that the agnostic PWL approach gives more satisfactory results. It should be noted that the PWL parametrization is just one of the many possibilities. A different combination of a flexible parametrization combined with a well-chosen cost function will likely yield improvements.

In the present work, we have limited ourselves to transverse-current initiated radio emission thereby excluding charge-excess radiation. From observations it is known that the charge-excess radiation is sub-leading. The charge-excess radiation is polarized in the radial direction which interferes with the linearly polarized emission by the transverse current emission. This interference is the reason for the broken cylindrical symmetry in the fluency pattern. In spite of its marked effect on the structure of the radio footprint there are reasons for not considering charge-excess

radiation in this work. One reason is, that by summing the $\mathbf{v} \times \mathbf{B}$ or the $\mathbf{v} \times (\mathbf{v} \times \mathbf{B})$ polarization components for antennas in a ring around the core, the radial polarization component cancels and thus does not contribute to the analysis. To obtain an effect of charge excess radiation, one thus needs to introduce a particular asymmetric antenna layout, or integrate the radial polarization direction, which we defer to a discussion in a follow-up work where we plan to make a direct link with experiment. Another reason is that analyzing charge excess radiation requires a different expression for the kernel. Exploratory calculations (not reported here) show that the kernel for charge excess exhibits a much less pronounced peak near $D_c = D_b$ than the kernel for transverse current emission. The reason for this is probably that for charge excess radiation the intensity near the core is strongly suppressed. This results in a less accurate determination of the emitting current. Combined with the fact that charge excess radiation is sub-leading compared to transverse current emission does imply that for real experiments it is doubtful that beamforming the charge excess component will yield useful results.

In interpreting the longitudinal current profile, one should be careful that this will be different from the longitudinal charged-particle profile that is most often considered in air-shower physics. For fair-weather events, they are closely related as the drift velocity, defined as the ratio of the two, is a very smooth function of air density [4]. Under thunderstorm conditions, the two differ strongly as the current, in magnitude and direction, is dominated by the atmospheric electric fields.

From the onset, we have limited the paper to on-axis beamforming for air showers. The formalism can easily be extended to off-axis beamforming, see the discussion at the end of Section V.

In a future work, we plan to apply the proposed procedures to microscopic Monte-Carlo model predictions and real data such as measured at LOFAR. Attention points will be the investigation of the effects of charge-excess radiation, uncertainty in the core position, and arrival direction on the beamforming results. These will be particularly difficult to correct for with an uneven distribution of antennas over the area. One attention point for a realistic scenario will thus be the effect of the detailed antenna layout.

In this work, we have shown the importance of including larger distances, where we have used a relatively homogeneous coverage of the area and thus a large number of antennas at large distances. For a non-homogeneous coverage, as one will have in real measurements, there may be an excess antenna density at some radius which will thus dominate the beaming trace leading to worse resolution in D_c . It may be thus beneficial to compensate for this by adjusting the function $\mathcal{F}(D_b, t_b, t_a, r_a)$ introduced in Eq. (III.1). However, there is a balance as a larger antenna density improves the signal-to-noise ratio, where there exists an extensive discussion in the literature, see for example Refs. [18, 19], for the pros and cons of the different choices.

VII. CONCLUSION

Our investigation has demonstrated that it is possible to correct for finite aperture effects when beamforming the radio-frequency pulse emitted by air showers. This greatly improves the resolution with which the longitudinal profile can be extracted from measurements. As an example, one may compare the results shown in Fig. 5. The top panel shows the results from beamforming when no aperture correction is applied. The maximum of the beamforming amplitude for $X_{max} = 1100 \text{ g/cm}^2$ lies close to $D = 4.5 \text{ km}$ or at about 950 g/cm^2 with a profile shape that does not resemble that of a realistic shower profile. When applying the procedure proposed in this work the extracted X_{max} differs from the input value by only 9 g/cm^2 with a very reasonable result for the longitudinal profile function. In spite of the fact that much of the off-sets are systematic in nature the improvement of the procedure when applying aperture corrections is evident. The results show that beaming might be the ideal approach for air shower measurements at SKA, where a fairly homogeneous antenna coverage over a large area is combined with a large frequency band width.

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Appendix A: Piece-Wise Linear fit

In a piece-wise linear parametrisation of the current it is written as a sum of partially overlapping triangles where the base of the triangle with the peak at D_i stretches from D_{i-1} till D_{i+1} ,

$$F_i(D) = \begin{cases} \frac{D-D_{i-1}}{D_i-D_{i-1}} & \text{when } D_{i-1} < D < D_i , \\ \frac{D_{i+1}-D}{D_{i+1}-D_i} & \text{when } D_i < D < D_{i+1} , \\ 0 & \text{otherwise .} \end{cases} \quad (\text{A.1})$$

The current is subsequently written as

$$J(D) = \sum_{i=1}^N F_i(D) J_i . \quad (\text{A.2})$$

In terms of this current the expression for the quality factor becomes after some minor rewriting

$$\chi_{\text{PWL}}^2 = \int dD_b \left| \sum_{i=1}^N J_i \mathcal{W}_i(D_b) - \mathcal{P}(D_b) \right|^2 , \quad (\text{A.3})$$

with $\mathcal{W}_i(D_b) = \int_{D_c} dD_c \mathcal{W}(D_b, D_c) F_i(D_c)$. The minimum is found by setting the derivatives $\partial \chi_{\text{PWL}}^2 / \partial J_j = 0$ resulting in

$$B_j = \sum_i A_{j,i} J_i , \quad (\text{A.4})$$

with

$$B_j = \int dD_b \mathcal{P}(D_b) \mathcal{W}_j(D_b) , \quad (\text{A.5})$$

and

$$A_{j,i} = \int dD_b \mathcal{W}_j(D_b) \mathcal{W}_i(D_b) . \quad (\text{A.6})$$

Eq. (A.4) can be solved for J_i by a simple matrix inversion. For applying this procedure it is not necessary that the values D_i form a regular grid.

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