

# **Survey of Solution Techniques for Linear Systems from Finite Difference Methods in Numerical Radiative Transfer**

**Advanced Numerical Analysis II  
Final Project**

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**May 11, 2017**

# 1 Introduction

## 1.1 Kelp

## 1.2 Radiative Transfer

Let  $n$  be the number of spatial dimensions for the problem. (i.e., 2 or 3)

Let  $x \in \mathcal{R}^n$ .

Let  $\Omega$  be the unit sphere in  $\mathcal{R}^n$ . Let  $\omega \in \Omega$  be a unit vector in  $\mathcal{R}^n$ .

Let  $L(x, \omega)$  denote *radiance* - the intensity of light at position  $x$  in the direction  $\omega$ . Let  $I(x)$  denote *irradiance* - the total intensity of light at position  $x$ . Let  $a(x)$  and  $b(x)$  denote the absorption and scattering coefficients respectively of the medium. Then, the Radiative Transfer Equation is

$$\omega \cdot \nabla_x L(x, \omega) = -(a(x) + b(x))L(x, \omega) + \int_{\Omega} \beta(\omega \cdot \omega') L(x, \omega') d\omega' \quad (1)$$

## 2 Discretization

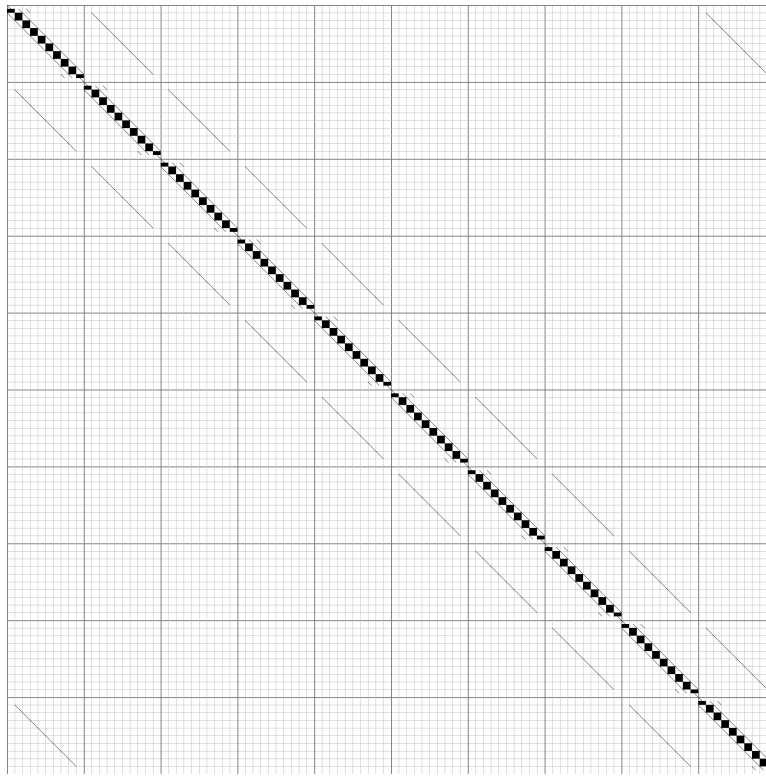


Figure 1: Sparsity plot: 10x10x16, ordering 012

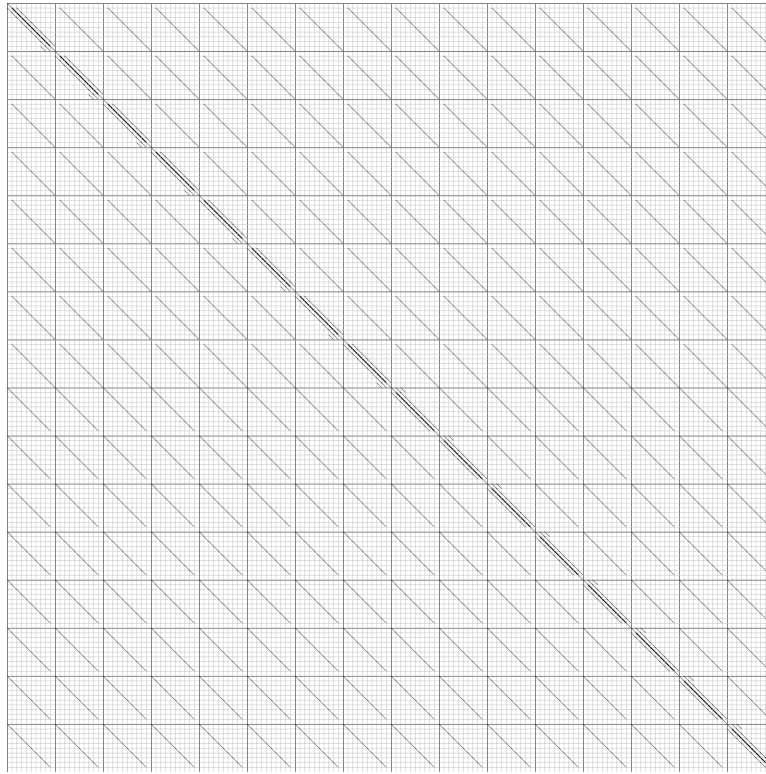


Figure 2: Sparsity plot: 10x10x16, ordering 210

## 2.1 Sparsity Plots

## 2.2 Matrix Properties

### 2.2.1 Diagonal Dominance

### 2.2.2 Spectral Radius

## 3 Direct Methods

### 3.1 Factorizations

### 3.2 Software Packages

## 4 Stationary Iterative Methods

### 4.1 Fixed-Point Iteration

### 4.2 Convergence and Preconditioning

## 5 Nonstationary Iterative Methods

### 5.1 Krylov Subspace Methods

### 5.2 Convergence and Preconditioning

## 6 Numerical Results

## 7 Conclusions