

PHYS414 Final Project

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Newton

Part A)

We have:

$$(*) \frac{dm(r)}{dr} = 4\pi r^2 \rho(r) \quad (1)$$

$$(*) \frac{dp(r)}{dr} = -\frac{G m(r) \rho(r)}{r^2} \quad (2)$$

$$(*) p = K \rho^\gamma = K \rho^{(1+\frac{1}{n})} \quad (3)$$

Substitute from (3) into (2):

$$(*) \frac{d}{dr} [K \rho^{(1+\frac{1}{n})}] = -\frac{G m(r) \rho(r)}{r^2}$$

$$(*) K(1+\frac{1}{n}) \rho(r)^{\frac{1}{n}} \frac{d\rho(r)}{dr} = -\frac{G m(r) \rho(r)}{r^2}$$

$$(*) -m(r) = \left[K(1+\frac{1}{n}) \rho(r)^{\frac{1}{n}-1} \frac{d\rho(r)}{dr} r^2 \right] / G$$

$$(*) -m(r) = r^2 \frac{K(n+1)}{G} \frac{d}{dr} \rho(r)^{\frac{1}{n}} \quad ; \text{ use (1)}$$

$$(*) -\frac{dm(r)}{dr} = -4\pi r^2 \rho(r) = \frac{d}{dr} \left[r^2 \frac{K(n+1)}{G} \frac{d}{dr} \rho(r)^{\frac{1}{n}} \right]$$

$$(*) \frac{K(n+1)}{4\pi G} \frac{1}{r^2} \frac{d}{dr} \left[r^2 \frac{d}{dr} (\rho(r)^{\frac{1}{n}}) \right] + \rho(r) = 0$$

Let $\alpha \xi := r$ such that $\frac{K(n+1)}{4\pi G} \rho_c^{(1/n-1)} = \alpha^2$, then:

$$(*) \frac{1}{\xi^2} \frac{d}{d\xi} \left[\xi^2 \frac{d}{d\xi} \left(\frac{\rho(\xi)}{\rho_c} \right)^{\frac{1}{n}} \right] + \frac{\rho(\xi)}{\rho_c} = 0$$

Let $\theta(\xi) = \left(\frac{\rho(\xi)}{\rho_c} \right)^{\frac{1}{n}}$ such that $\theta(0) = 1$ since $\rho(0) = \rho_c$

Therefore; $\frac{1}{\xi^2} \frac{d}{d\xi} \left[\xi^2 \frac{d}{d\xi} \theta(\xi) \right] + \theta(\xi)^n = 0$; Q.E.D

We have:

$$(*) \frac{1}{z^2} \frac{d}{dz} \left[z^2 \frac{d}{dz} \theta(z) \right] + \theta(z)^n = 0$$

Want to find an expression for values around zero:

$$(*) \theta(z) = \theta(0) + \theta'(0)z + \frac{\theta''(0)}{2!} z^2 + \frac{\theta'''}{3!} z^3 + \dots$$

$$(*) \frac{d}{dz} \theta(z) = \theta'(0) + \theta''(0)z + \frac{\theta'''(0)}{2!} z^2 + \dots$$

$$(*) \frac{1}{z^2} \frac{d}{dz} \left[z^2 \frac{d}{dz} \theta(z) \right] = \frac{2}{z} \theta'(0) + 3 \theta''(0) + \frac{4}{2!} \theta'''(0)z + \dots \quad (1)$$

$$(*) \theta(z)^n = \theta(0)^n + (n) \theta'(0) \theta(0)^{n-1} z + (n) \frac{\theta''(0)}{2!} \theta(0)^{n-1} z^2 + \dots \quad (2)$$

We require that $(1) + (2) = 0$, coefficient of each term has to be 0

From these equations, we find that;

$$(*) \theta(0) = 1 \quad \& \quad \theta'(0) = 0 \quad \& \quad \theta''(0) = -\frac{1}{3} \\ \& \quad \theta'''(0) = 0 \quad \& \quad \theta^{(4)}(0) = \frac{n}{5}$$

Therefore, we have:

$$(*) \theta(z) = 1 - \frac{z^2}{6} + \frac{n z^4}{120} + \dots \quad ; \quad Q.E.D$$

For $n=1$;

$$(*) \theta(z) = 1 - \frac{z^2}{6} + \frac{z^4}{120} + \dots$$

We have:

$$(*) \quad \frac{d}{dr} m(r) = 4\pi r^2 \rho(r)$$

Want to find the total mass of the star.

Use the following change of variables:

$$(*) \quad r \rightarrow \alpha \xi \quad \& \quad \rho(r) \rightarrow \rho_c \theta(\xi)^n$$

We get:

$$(*) \quad \frac{d}{d\xi} m(\xi) = 4\pi \xi^2 \rho_c \theta(\xi)^n \alpha^3$$

Total mass M is given by;

$$(*) \quad M = \int_0^{\xi_f} 4\pi \rho_c \alpha^3 \xi^2 \theta(\xi)^n d\xi ; \text{ where } \alpha \xi_f = R$$

Using Lane-Emden equation to substitute for $\xi^2 \theta(\xi)^n$ gives:

$$(*) \quad M = 4\pi \rho_c \alpha^3 \int_0^{\xi_f} -\frac{d}{d\xi} \left[\xi^2 \frac{d}{d\xi} \theta(\xi) \right] d\xi ; \text{ use FTC}$$

$$(*) \quad M = 4\pi \rho_c \alpha^3 \xi_f^3 \left[-\frac{1}{\xi_f} \frac{d}{d\xi} \theta(\xi_f) \right] ; \text{ use } \alpha^3 \xi_f^3 = R^3$$

$$(*) \quad M = 4\pi \rho_c R^3 \left[-\frac{1}{\xi_f} \frac{d}{d\xi} \theta(\xi_f) \right] ; \text{ Q.E.D}$$

We have;

$$(*) \quad M = 4\pi \rho_c R^3 \left[\frac{-1}{z_f} \frac{d}{dz} \theta(z_f) \right] \quad (1)$$

Want to express mass as a function of radius.

$$(*) \quad \alpha^2 = \frac{K(n+1)}{4\pi G} \rho_c^{\left(\frac{1}{n}-1\right)} ; \text{ substitute to } (1)$$

$$(*) \quad M = 4\pi^{\left(\frac{1}{1-n}\right)} R^{\left(\frac{3-n}{1-n}\right)} \left[\frac{-d}{dz} \theta(z_f) \right] \left(\frac{K}{G} \right)^{\left(\frac{n}{n-1}\right)} (n+1)^{\left(\frac{n}{n-1}\right)} z_f^{\left(\frac{n+1}{n-1}\right)}$$

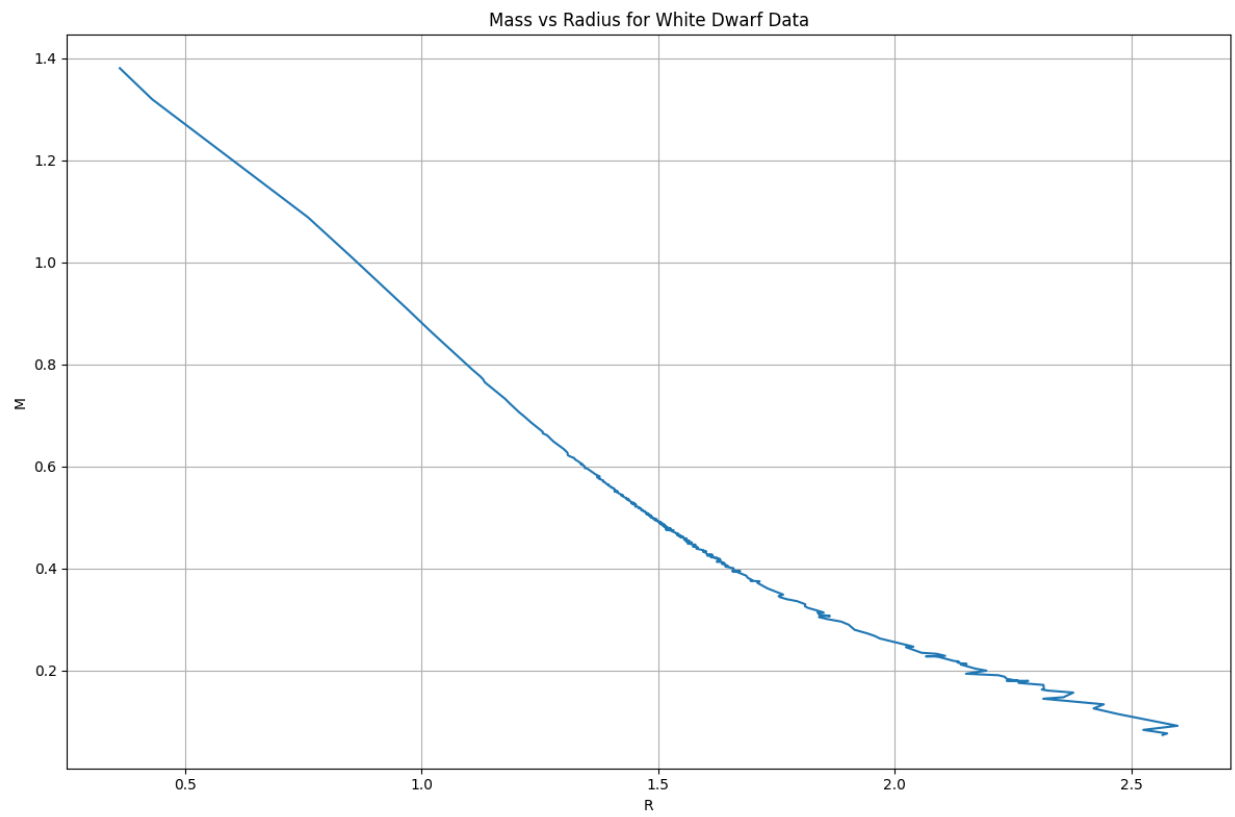
So, we can see that;

$$(*) \quad M \propto R^{\left(\frac{3-n}{1-n}\right)} ; \text{ Q.E.D}$$

with a proportionality constant of;

$$(*) \quad 4\pi^{\left(\frac{1}{1-n}\right)} \left[\frac{-d}{dz} \theta(z_f) \right] \left(\frac{K}{G} \right)^{\left(\frac{n}{n-1}\right)} (n+1)^{\left(\frac{n}{n-1}\right)} z_f^{\left(\frac{n+1}{n-1}\right)}$$

Part B)



Plot 1 Mass vs Radius for White Dwarf Data

Part C)

We have;

$$(*) \quad \rho = C [x(2x^2-3)(x^2+1)^{\frac{1}{2}} + 3 \sinh^{-1}(x)] \quad (1)$$

$$(*) \quad x = \left(\frac{\rho}{D}\right)^{\frac{1}{q}} \quad (2)$$

$$(*) \quad \rho = K^* \rho^{1+\frac{1}{n^*}} \quad (3)$$

Considering $x \ll 1$;

$$(*) \quad (x^2+1)^{\frac{1}{2}} \approx 1 + \frac{1}{2}x^2 - \frac{1}{8}x^4 + \dots$$

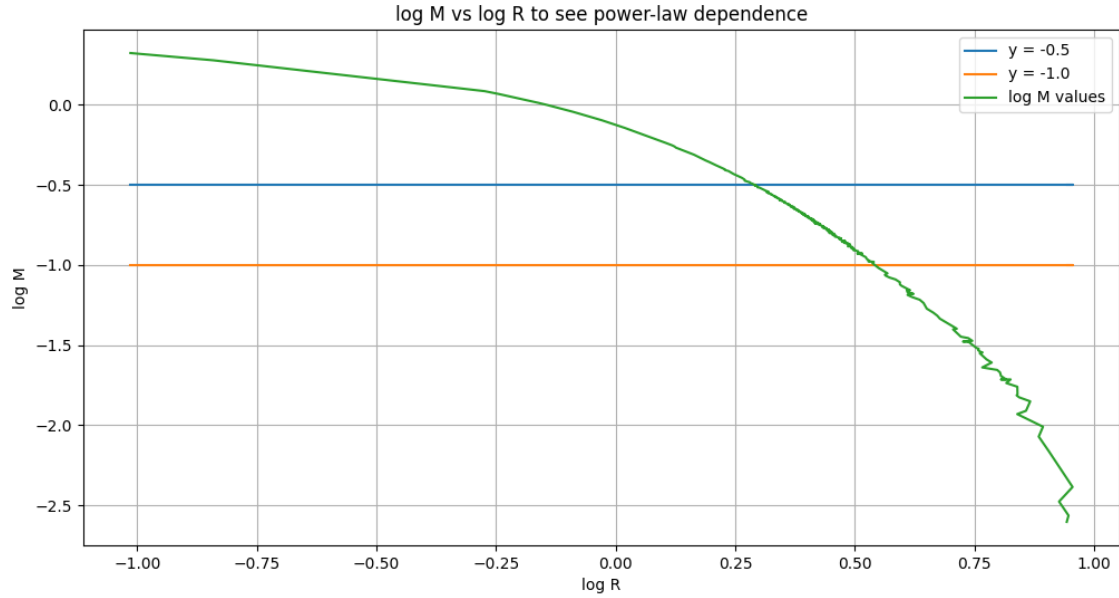
$$(*) \quad \sinh^{-1}(x) \approx x - \frac{1}{6}x^3 + \frac{3}{40}x^5 - \frac{5}{112}x^7 + \dots$$

$$(*) \quad \rho \approx C \left[\frac{11}{8}x^5 + \frac{1}{2}x^3 - 3x + 3\left(x - \frac{1}{6}x^3 + \frac{3}{40}x^5\right) + O(x^7) \right]$$

$$(*) \quad \rho \approx C \left[\frac{8}{5}x^5 + O(x^7) \right]$$

$$(*) \quad \rho = \frac{8}{5} \frac{C}{D^{\frac{5}{q}}} \rho^{\frac{5}{q}} \quad ; \text{ fitting for (3) gives:}$$

$$(*) \quad K^* = \frac{8C}{5D^{\frac{5}{q}}} \quad \& \quad n^* = \frac{q}{5-q} \quad ; \text{ Q.E.O}$$



Plot 2 Demonstration of power-law dependence

From the above graph, the cutoff value is chosen to be -1.1 for the calculations regarding the rest of this section.

The calculated value of q from the fit is: 2.9576226797905165 and the closest integer value 3 is used for q in the remaining part of this section.

Since $n = q / (5 - q)$ using $q = 3$ gives n_{star} as $3/2$

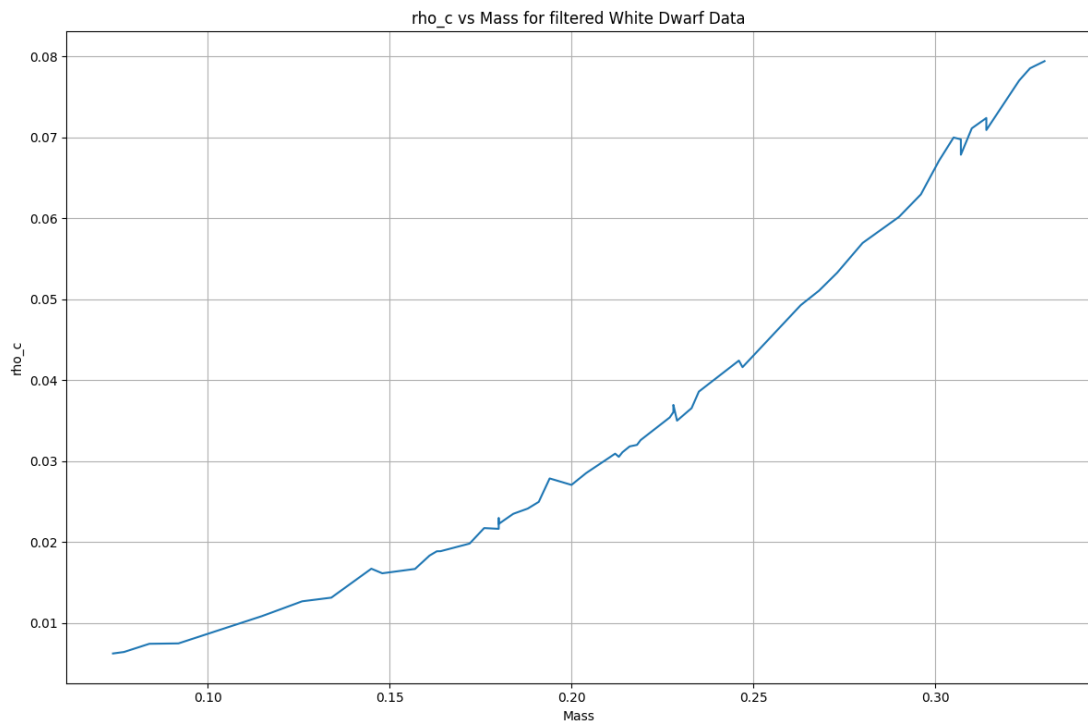
Then, the calculated value for K in the scaled units is: 0.3059168066552295

If we revert the value to SI units, we have $K_{\text{SI}} = 3187788.5335691073$

The analytical values for C , D , and K are calculated as:

- $C: 6.002334828410536 \times 10^{21}$
- $D: 1947864981.4433918$
- $K: 3161128.260832043$

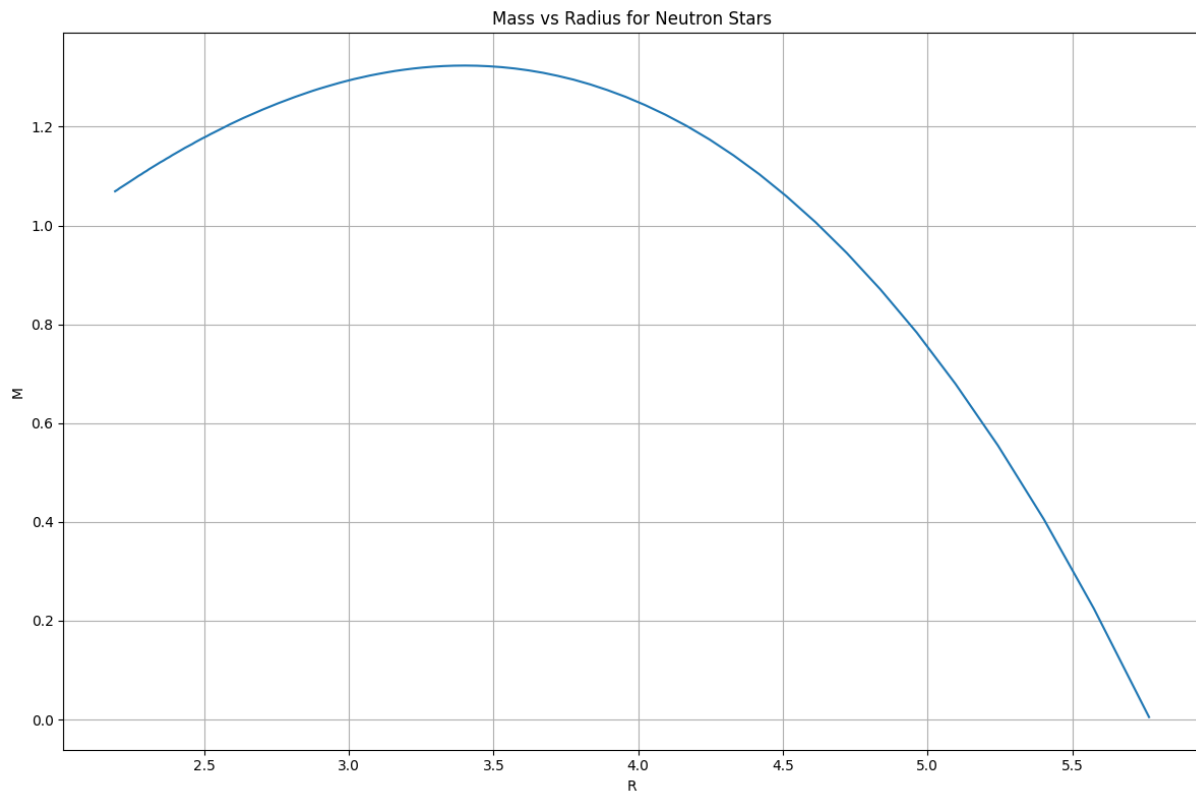
Therefore, the error for K is 0.84%



Plot 3 ρ_c vs Mass for Filtered White Dwarf Data

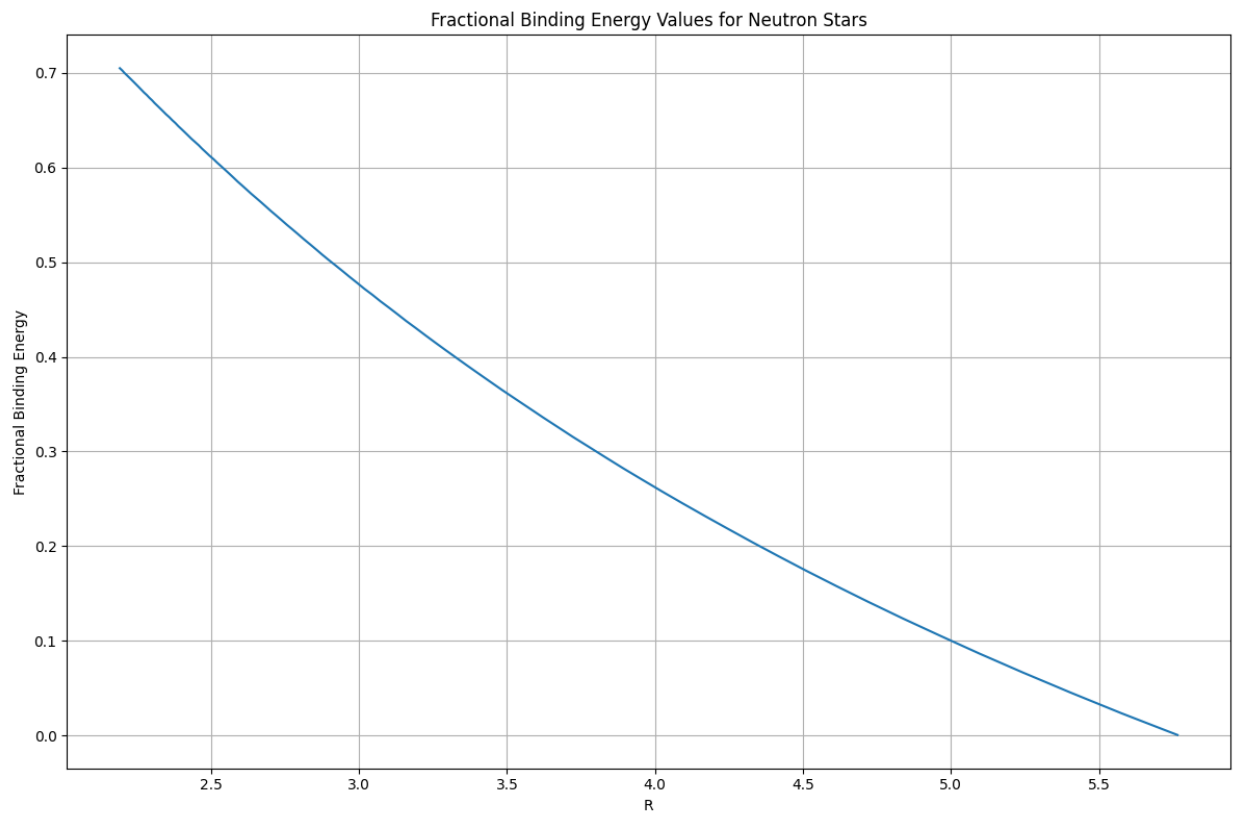
Einstein

Part A)



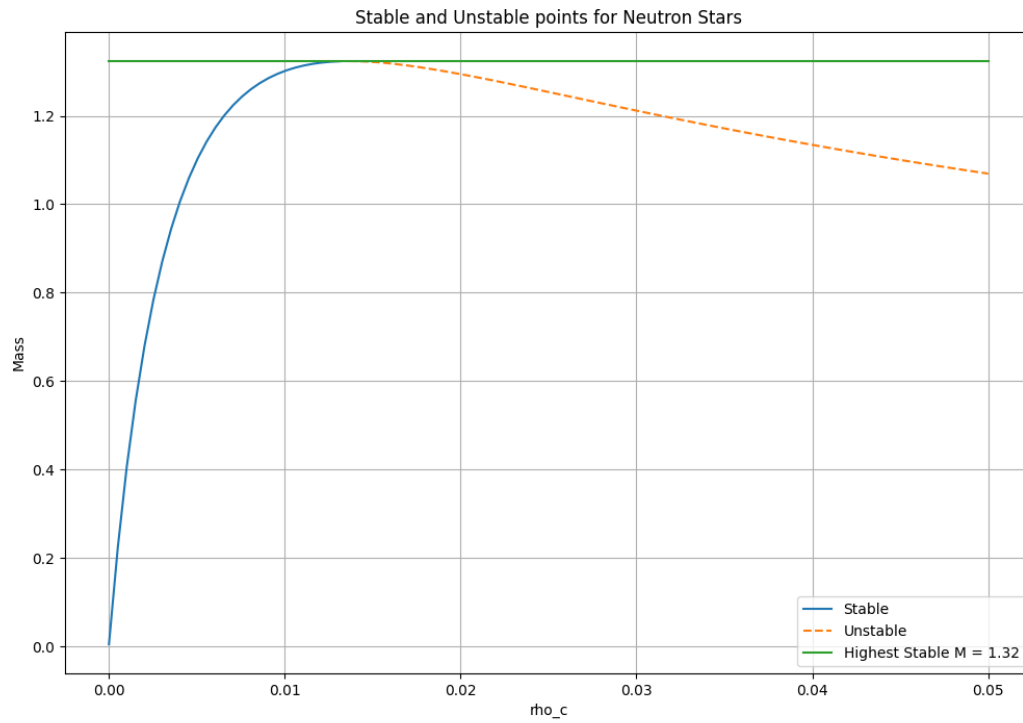
Plot 4 Mass vs Radius for Neutron Stars

Part B)



Plot 5 Fractional Binding Energy Values for Neutron Stars

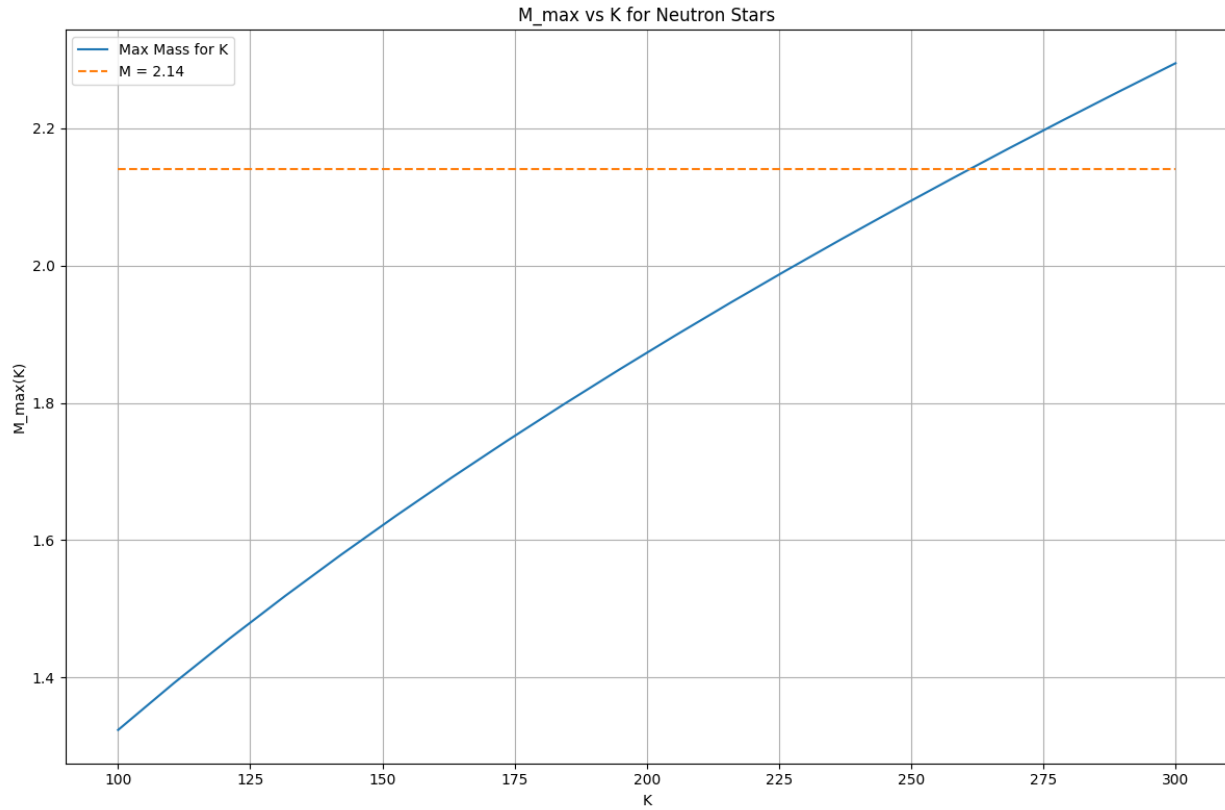
Part C)



Plot 6 Stable and Unstable points for Neutron Stars

The highest stable mass allowed by this Equation of State is 1.32 in scaled units.

Part D)



Plot 7 M_{max} vs K for Neutron Stars

Above is the graph for the highest stable mass for the Neutron Star corresponding to different K values.

Below are the values of the maximum masses for different K :

- Mass: 1.32344 for $k=100.00$
- Mass: 1.39154 for $k=110.53$
- Mass: 1.45647 for $k=121.05$
- Mass: 1.51863 for $k=131.58$
- Mass: 1.57834 for $k=142.11$
- Mass: 1.63589 for $k=152.63$
- Mass: 1.69148 for $k=163.16$
- Mass: 1.74530 for $k=173.68$
- Mass: 1.79752 for $k=184.21$
- Mass: 1.84826 for $k=194.74$
- Mass: 1.89765 for $k=205.26$
- Mass: 1.94579 for $k=215.79$
- Mass: 1.99277 for $k=226.32$
- Mass: 2.03867 for $k=236.84$
- Mass: 2.08355 for $k=247.37$
- Mass: 2.12750 for $k=257.89$
- Mass: 2.17055 for $k=268.42$
(first one above maximum)
- Mass: 2.21277 for $k=278.95$
- Mass: 2.25420 for $k=289.47$
- Mass: 2.29488 for $k=300.00$

Part E)

We have;

$$\textcircled{*} \quad v' = \frac{2M}{r(r-2M)}, \quad (r > R)$$

$$\textcircled{*} \quad \int_R^r v' dr' = \int_R^r \frac{2M}{r'(r'-2M)} dr'$$

$$\textcircled{*} \quad v(r) - v(R) = \int_R^r \frac{2M}{(r')^2} \frac{1}{1 - \frac{2M}{r'}} dr'; \text{ use change of variables}$$

$$\textcircled{*} \quad 1 - \frac{2M}{r'} = u \Rightarrow \frac{2M}{(r')^2} dr' = du$$

$$\textcircled{*} \quad v(r) - v(R) = \int_{1-\frac{2M}{R}}^{1-\frac{2M}{r}} \frac{1}{u} du$$

$$\textcircled{*} \quad v(r) - v(R) = \ln\left(1 - \frac{2M}{r}\right) - \ln\left(1 - \frac{2M}{R}\right)$$

$$\textcircled{*} \quad v(r) = \ln\left(1 - \frac{2M}{r}\right) - \ln\left(1 - \frac{2M}{R}\right) + v(R); \quad (r > R) \quad \text{Q.E.D.}$$