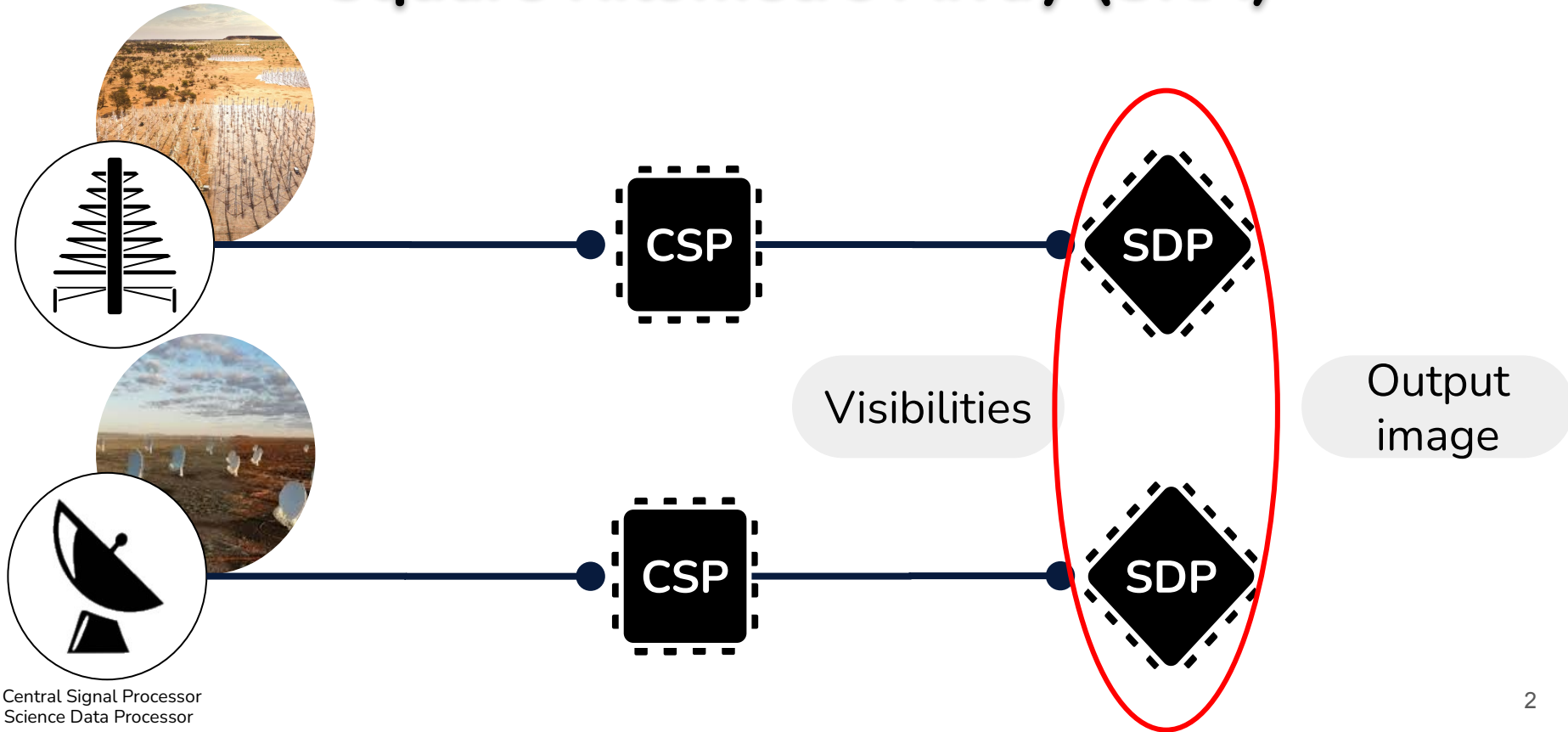




Simulating MeasurementSet for imaging pipeline validation

Hackathon Eclat Recap

Square Kilometre Array (SKA)



Near Real-time requirement



$\text{dataAcquisitionTime} \approx \text{pipelineProcessingTime}$

Note

$\text{Data} = \text{N_ANT} \times \text{N_CHANNEL} \times \text{N_POL} \times \text{BANDWIDTH} \times \text{OBS_TIME} \times \text{N_BYTE}$

(10min)= $130000 \times 4000 \times 2 \times 10^6 \times 600 \times 4 = 2 \text{ Po}$

(12h)= $130000 \times 4000 \times 2 \times 10^6 \times 43200 \times 4 = 179 \text{ Po}$

LOFAR (LOW Frequency ARray) comparison:

(10min)= $4992 \times 256 \times 2 \times 195000 \times 600 \times 2 = 300 \text{ To}$

(12h)= $4992 \times 256 \times 2 \times 195000 \times 43200 \times 2 = 21 \text{ Po}$

NenuFAR (New Extension in Nançay Upgrading LOFAR) comparison:

(10min)= $1936 \times 768 \times 2 \times 195000 \times 600 \times 2 = 348 \text{ To}$

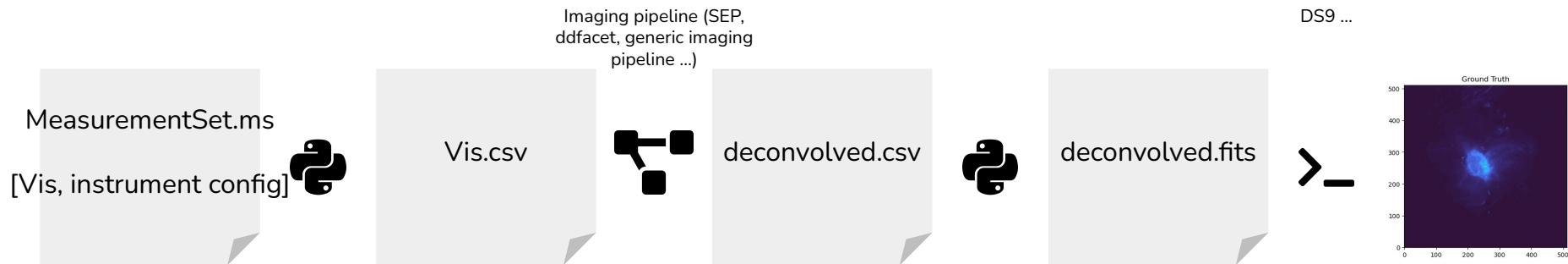
(12h)= $1936 \times 768 \times 2 \times 195000 \times 43200 \times 2 = 25 \text{ Po}$

Latency(pipeline)

Final goal

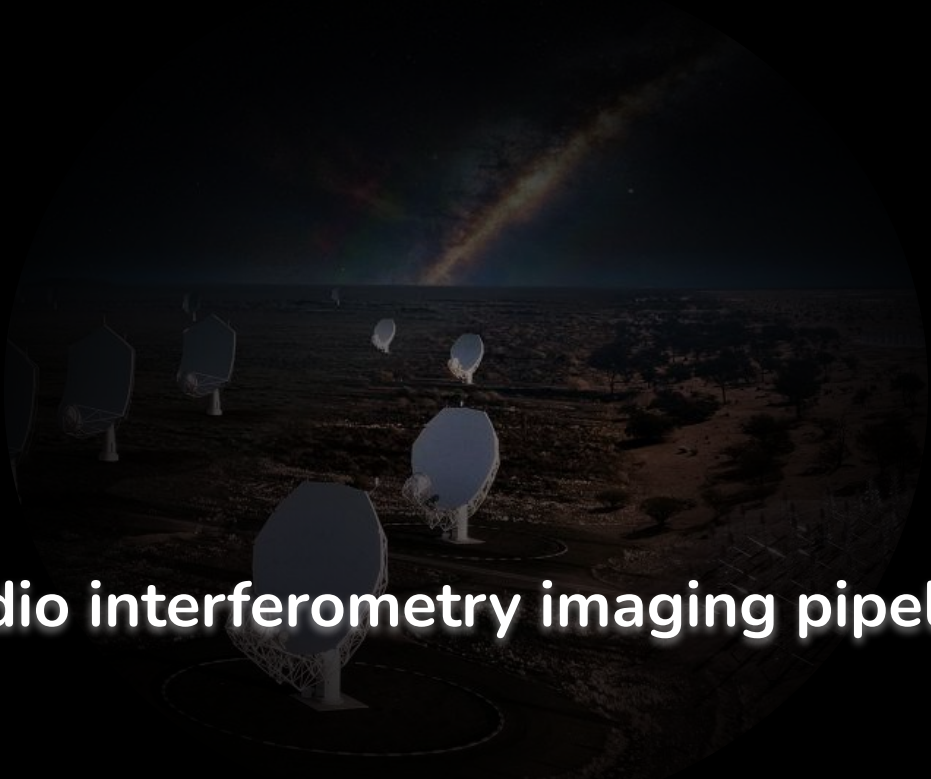
Generating MeasurementSet:

- 👉 To benchmark pipeline implementations (latency, memory requirement, energy, output quality ...)
- 👉 To control the “true sky”
- 👉 To simulate SKA data size (without requiring instrument)



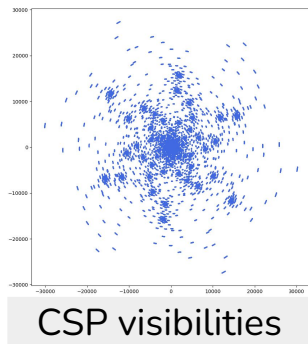
Common pipeline direction

This project

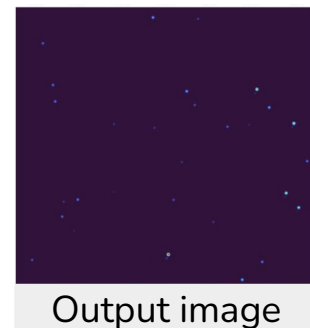
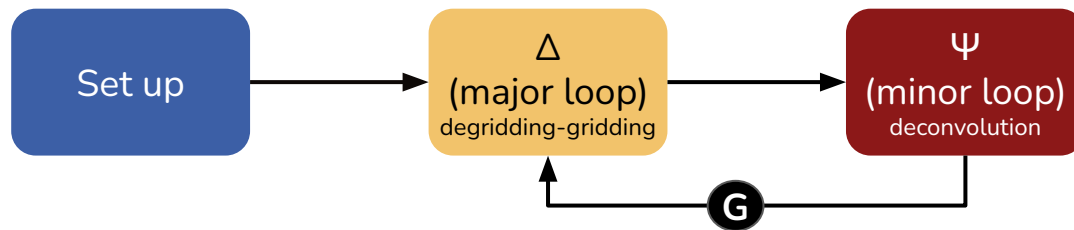
A circular inset image showing a radio telescope array at night. Several large parabolic dish antennas are visible on a dark, flat landscape. In the background, the Milky Way galaxy is visible in the dark sky, stretching from the horizon towards the top of the frame. The text "Radio interferometry imaging pipeline" is overlaid in white at the bottom of the circular image.

Radio interferometry imaging pipeline

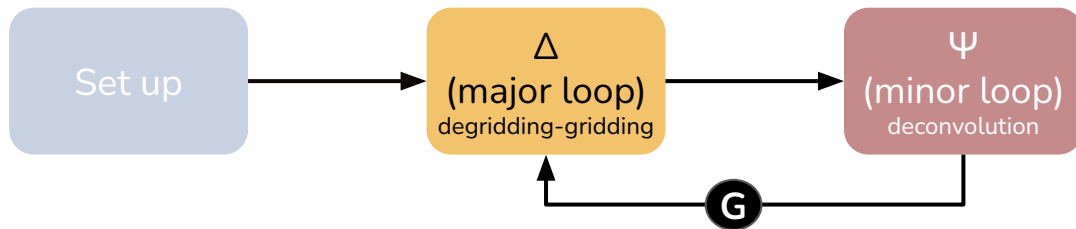
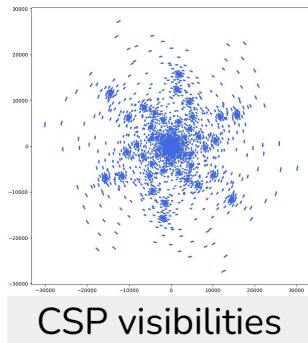
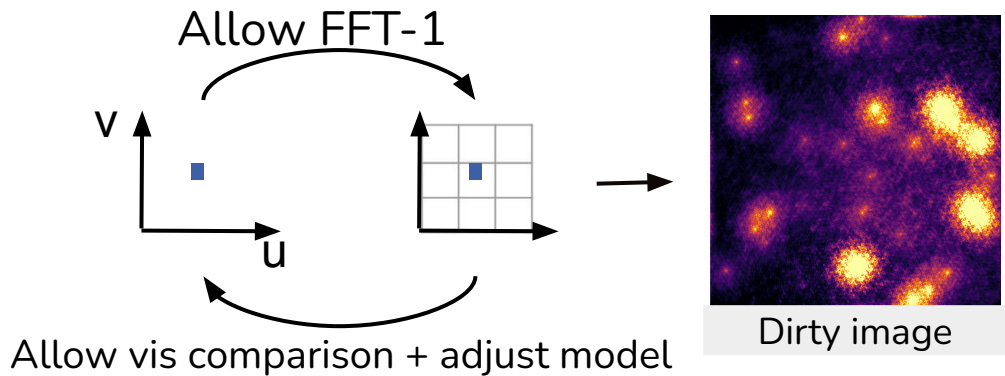
Generic radio-interferometric imaging pipeline



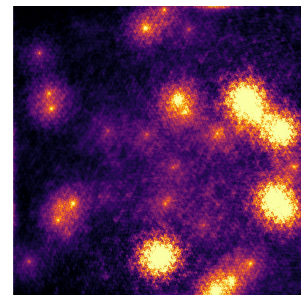
↑
correlation point of a pair of
antenna



$\Delta: v \rightarrow \text{dirty}$



Ψ : clean dirty



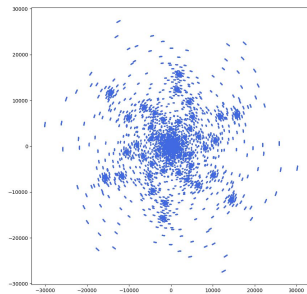
Dirty image

=

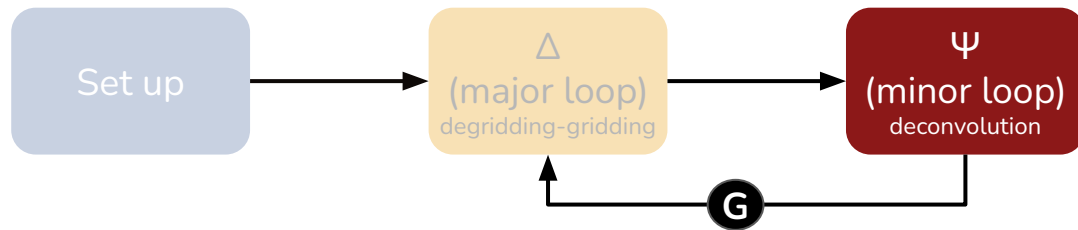
Sky Image

*

PSF



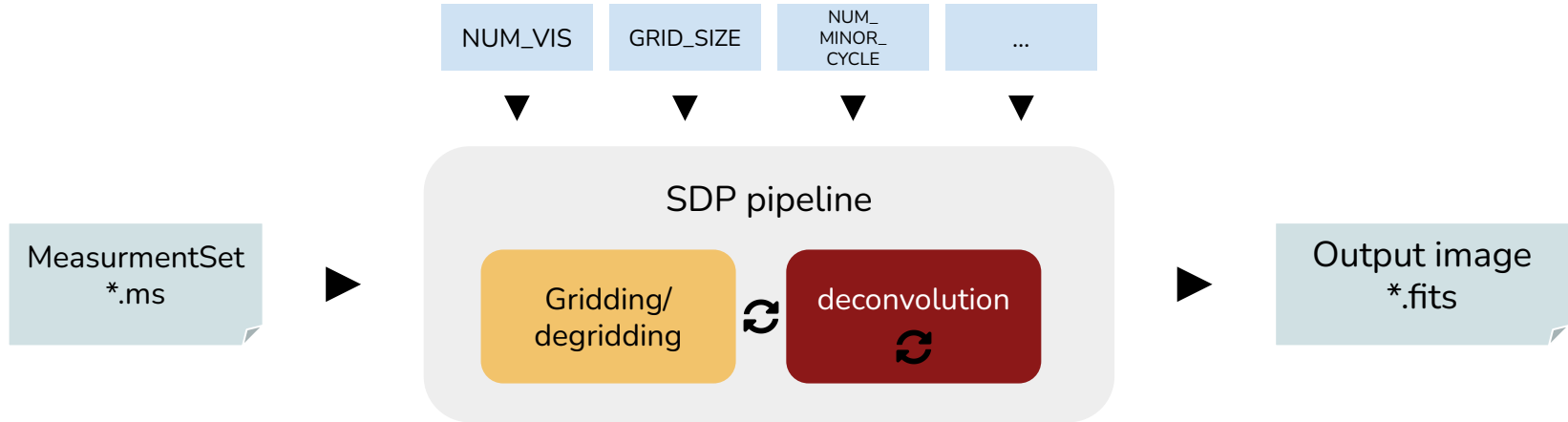
CSP visibilities

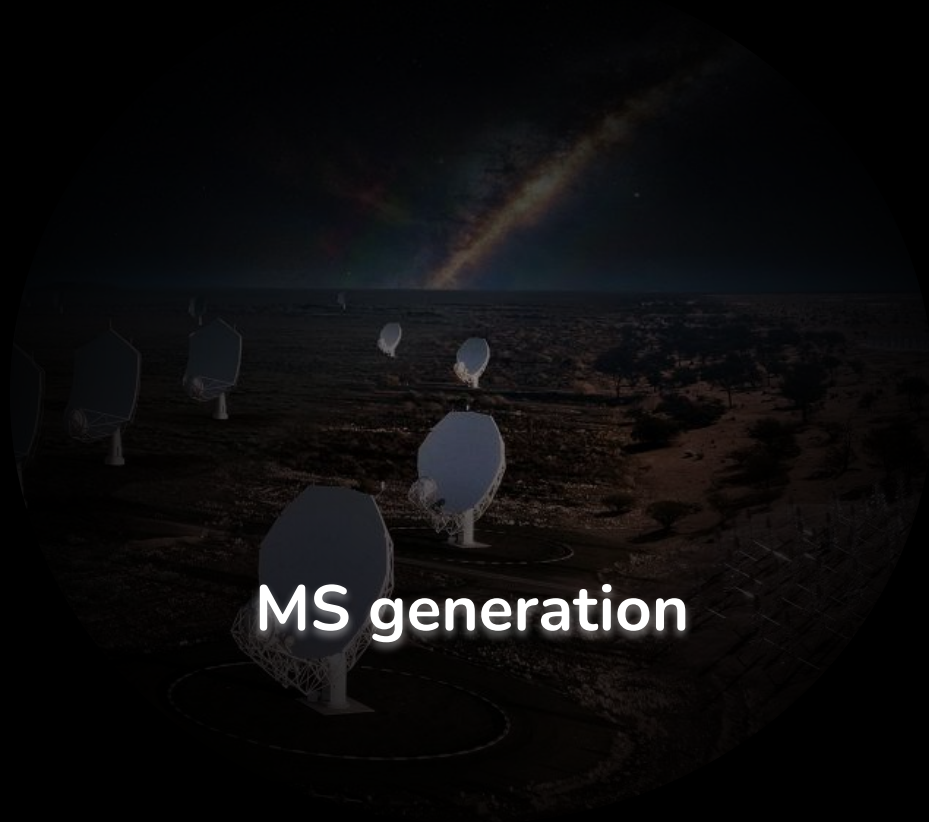


Existing imaging pipelines

	Dataflow implem	Deployment
SDP Evolutionary Pipeline (SEP) (insa gitlab here)	✓	out of date
Generic Imaging Pipeline <ul style="list-style-type: none">- DFT + Hogbom clean- FFT + Hogbom clean- G2G + Hogbom clean (I m working here)	✓	Under discussion for SKA
DDFacet (Cyril github here)	✗ (python)	LOFAR
RASCILL (SKAO github here)	✗ (python)	??
...		

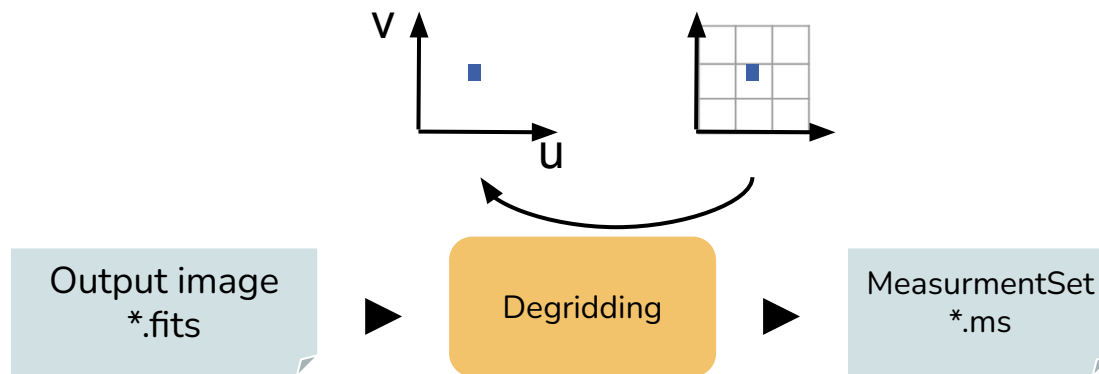
Pipelines operation





MS generation

The process



For the Hackathon, **RASCIL** degridder is used in order to avoid circular validation
Radio Astronomy Simulation, Calibration and Imaging Library (RASCIL)

Advantages

- ✓ integrate ska sdp library
- ✓ facilitate existing radio-telescope simulation

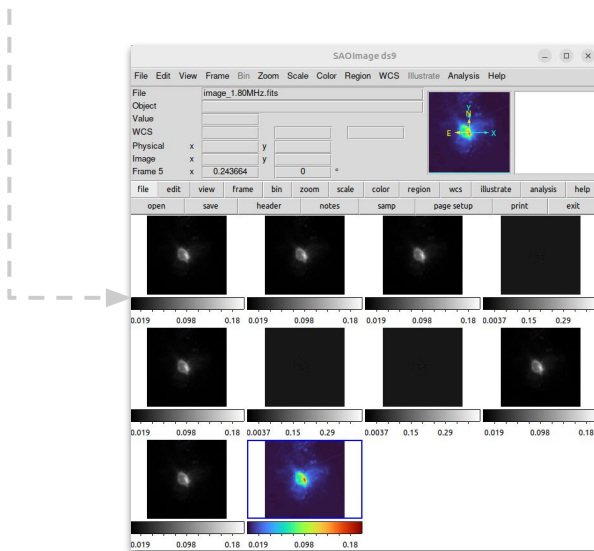
Limitations

- ✗ python
- ✗ sequential (super long for big files)
- ✗ ...

Visualize .fits

Install ds9: `sudo apt install saods9`

Run: `ds9 *.fits -lock frame wcs -zoom to fit`



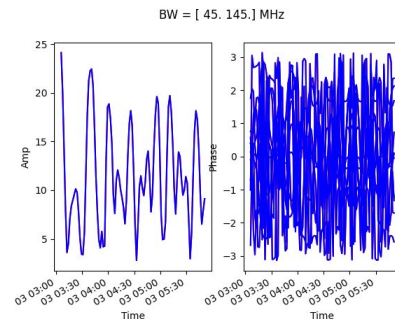
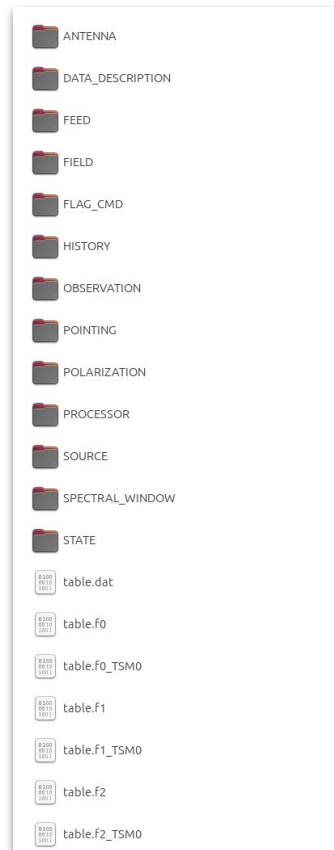
To reveal the contrasts:

- Color > Matplotlib > turbo (recommended by Sunrise)
- Color > Matplotlib > viridis / inferno (most popular in astro-papers)

Validating .ms



casacore based script



- Distributed MS simulation



Jupyter notebook to simplify the usage



Contain DDFacet tutorial



Contain MS validation (casacore)



Project available on github:

<https://github.com/Ophelie-Renaud/vis-generator>



Official repos: <https://framagit.org/eclat>

- Project in progress since we completed 1/2 hackathon



Configurable MS



SCAN ME