

# The Ovals of Cassini

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Page 1 of 11

Go Back

Full Screen

Close

#### Introduction

We will be looking at a curve that was developed by Giovanni Domenico Cassini in 1680. He had believed that the motion of the Earth and Sun followed a lobe of one of these ovals. If we are given two fixed points  $F_1$  and  $F_2$  and a constant c the Ovals of Cassini are defined as the locus of points that are  $|PF_1| \cdot |PF_2| = c$ . This is given to us in Cartesian coordinates as:

$$[(x-a)^2 + y^2][(x+a)^2 + y^2] = b^4$$

### Polar Form

This form can be quite cumbersome so we will convert the equation to one in polar coordinates.

Let  $x = r \cos \theta$  and  $y = r \sin \theta$  Then,



Home Page

Title Page





Page 2 of 11

Go Back

Full Screen

Close

$$[(x-a)^{2} + y^{2}][(x+a)^{2} + y^{2}] = b^{4}$$

$$[(r\cos\theta - a)^{2} + (r\sin\theta)^{2}][(r\cos\theta + a)^{2} + (r\sin\theta)^{2}] =$$

$$[r^{2}\cos^{2}\theta - 2ar\cos\theta + a^{2} + r^{2}\sin^{2}\theta][r^{2}\cos^{2}\theta + 2ar\cos\theta + a^{2} + r^{2}\sin^{2}\theta] =$$

$$r^{4}\sin^{4}\theta + r^{4}\cos^{4}\theta - 2a^{2}r^{2}\cos^{2}\theta + 2a^{2}r^{2}\sin^{2}\theta + 2r^{4}\sin^{2}\theta\cos^{2}\theta + a^{4} =$$

$$r^{4}\sin^{4}\theta + r^{4}\cos^{4}\theta - 2a^{2}r^{2}(\cos^{2}\theta - \sin^{2}\theta) + 2r^{4}\sin^{2}\theta\cos^{2}\theta + a^{4} =$$

$$r^{4}\sin^{4}\theta + r^{4}\cos^{4}\theta - 2a^{2}r^{2}\cos2\theta + 2r^{4}\sin^{2}\theta\cos^{2}\theta + a^{4} =$$

$$r^{4}(\sin^{4}\theta + 2\sin^{2}\theta\cos^{2}\theta + \cos^{4}\theta) - 2a^{2}r^{2}\cos2\theta + a^{4} =$$

$$r^{4}(\sin^{2}\theta + \cos^{2}\theta)^{2} - 2a^{2}r^{2}\cos2\theta + a^{4} =$$

$$r^{4}(\sin^{2}\theta + \cos^{2}\theta)^{2} - 2a^{2}r^{2}\cos2\theta + a^{4} =$$

$$r^{4}(\sin^{2}\theta + \cos^{2}\theta)^{2} - 2a^{2}r^{2}\cos2\theta + a^{4} =$$

### Parametric Form

Now that we have our equation in polar coordinates we can try to get a parametric form as well. To do this we will solve for r using the quadratic formula.



Home Page

Title Page





Page 3 of 11

Go Back

Full Screen

Close

$$r^{2} = \frac{2a^{2}\cos 2\theta \pm \sqrt{4a^{4}\cos^{2}2\theta - 4(a^{4} - b^{4})}}{2}$$

$$= a^{2}\cos 2\theta \pm \sqrt{a^{4}\cos^{2}2\theta - a^{4} + b^{4}}$$

$$= a^{2}\cos 2\theta \pm \sqrt{a^{4}(\cos^{2}2\theta - 1) + b^{4}}$$

$$= a^{2}\cos 2\theta \pm \sqrt{-a^{4}(\sin^{2}2\theta) + b^{4}}$$

$$= a^{2}\left[\cos 2\theta \pm \sqrt{\left(\frac{b}{a}\right)^{4} - \sin^{2}2\theta}\right]$$

$$r = \pm a\sqrt{\cos 2\theta \pm \sqrt{\left(\frac{b}{a}\right)^{4} - \sin^{2}2\theta}}$$

Now substitute this r into our earlier  $x = r \cos \theta$  and  $y = r \sin \theta$ ,



Home Page

Title Page





Page 4 of 11

Go Back

Full Screen

Close

$$x = \pm a \sqrt{\cos 2\theta} \pm \sqrt{\left(\frac{b}{a}\right)^4 - \sin^2 2\theta} \left[\cos \theta\right]$$
$$y = \pm a \sqrt{\cos 2\theta} \pm \sqrt{\left(\frac{b}{a}\right)^4 - \sin^2 2\theta} \left[\sin \theta\right]$$

While these may be our parametric equations they inherently have a problem, there will be times when we will receive complex numbers which are totally useless for us. To remedy this we will have to restrict  $\theta$  so that the radicand will be greater than or equal to zero. To find this we begin with:

$$\left[ \left( \frac{b}{a} \right)^4 - \sin^2 2\theta \right] = 0$$

$$\left( \frac{b}{a} \right)^2 = \sin 2\theta$$

$$t = \frac{1}{2} \arcsin \left( \frac{b}{a} \right)^2$$



Home Page

Title Page





Page 5 of 11

Go Back

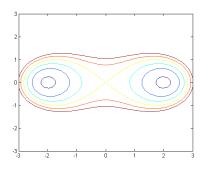
Full Screen

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Therefore our  $\theta$  values must run from  $\frac{1}{2}\arcsin\left(\frac{b}{a}\right)^2$  to  $-\frac{1}{2}\arcsin\left(\frac{b}{a}\right)^2$ . Now we have several useful versions for the Ovals of Cassini.

### Representations

The Cassinian Ovals exhibit peculiar behavior, at times becoming two separate curves! Let's take a look at this with  $Matlab^{\textcircled{\$}}$ .



As you may have noticed the yellow curve here is the Lemniscate of Bernoulli, a special case of the Cassini Ovals equation. Our Cartesian equation allows a and b to be modified, where a is the distance the foci are placed from the origin and  $b^2$  is our earlier stated constant c. Depending on the ratio of numbers chosen different behaviors are exhibited. If b = a



Home Page

Title Page





Page 6 of 11

Go Back

Full Screen

Close

then one can conceptually come to the conclusion that the Lemniscate is the result, as the distance from the foci to a point P is the same as the distance a focus is from the origin resulting in a common meeting point at the origin. There are still two more ratios of b and a that need to be covered. If  $\frac{b}{a} > 1$  one loop is created, but if  $\frac{b}{a} < 1$  two separate curves result. One can imagine this as the constant from b is small enough in comparison to a that a point P can't be found at the origin as with the Lemniscate case but rather are within a distance b < a from a focus, tracing out a region around each focus.

## **Experimenting With Variables**

Once again emphasizing the Cartesian equation discussion, the only variables we can modify are a and b. We will start with a varying a, take note that these images are of the same curves as earlier, we only modify a single variable. This is demonstrated in the first four images.

Understandably in this case the foci merely separate from each other with increasing a values. We know modifying b will increase the constant of where the curve is defined at.



Home Page

Title Page



Page 7 of 11

Go Back

Full Screen

Close

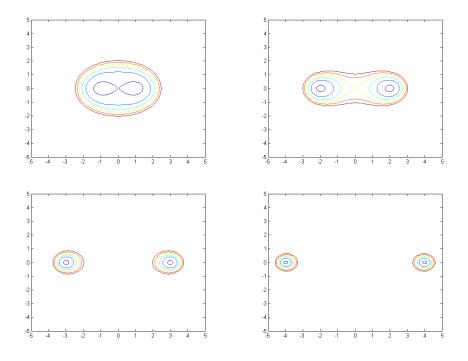


Figure 1: Varying a

### Conclusion

While it has been established that planetary bodies do not actually travel on these curves but rather on elliptical paths it is still an interesting curve



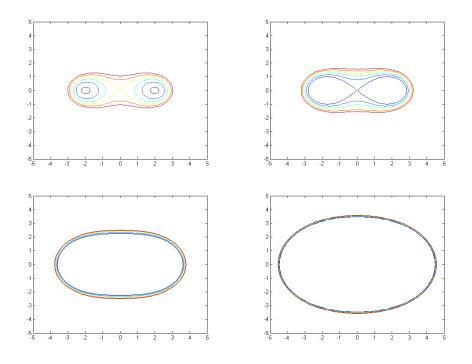
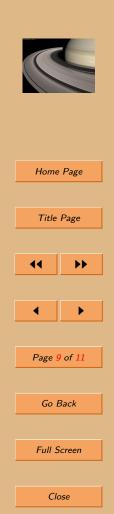
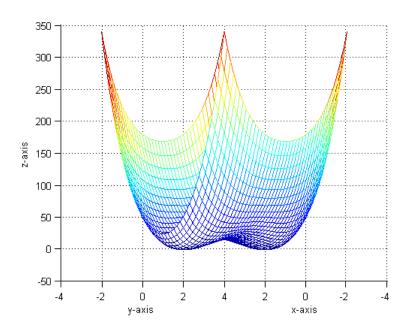


Figure 2: Varying b

to experiment with. For further consideration one can delve deeper into the curve and will undoubtedly discover that this curve is a toric section and



the three dimensional representation of the special case of the Lemniscate is the Möbius Strip. This will be an interesting thing for another time. For example I have included the Cassini Ovals plotted as a surface, the viewing angle is not standard but it emphasizes the features of the surface.





Home Page

Title Page

44

**→** 

Page 10 of 11

Go Back

Full Screen

Close

#### References

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Home Page

Title Page

**14** ▶

**◆** 

Page 11 of 11

Go Back

Full Screen

Close