Summary of Initial Papers

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Abstract

Summary of questions and findings from the eight initial papers from May 27, 2016.

1 Eight-Dimensional Mid-Infrared/Optical Bayesian Quasar Selection

Explored multi-dimensional, multiwavelength selection of quasars from the IRAC and SDSS. Selection traditionally in two-colour space, used a combination of 8-D and 4-D techniques. Used Bayesian selection techniques and completness and contamination to evaluate selection.

- 1. Converted between Vega and AB photometry
- 2. IRAC channels: 3.6, 4.5, 5.8, 8.0
- 3. Made 8 unique colours with ugriz magnitudes
- 4. Used all SDSS filters and two short-wave IRAC bands
- 5. Bayesian selection section 3
- 6. Used mean colours to classify types of quasars. Can we use that in our classification?
- 7. Set colour limits to reduce error and removed faint and saturated objects. sec 3.1. Can we do the same?
- 8. Can we use completeness and contamination? Need a training set, could use a set of points from the data?

2 Towards auto classification of all WISE sources

Applied support vector machines with a training sample to spectroscopic dataset to auto classify objects.

- 1. used four infrared bands 3.4 23 um
- 2. used signal to noise 2 in shorter wavelengths and deteriorates in longer.
- 3. significant work on colour colour space for WISE survey
- 4. used magnitude, color, and differential aperture mag space
- 5. computed completeness and contamination for training set.

3 Meaning of WISE colours

Colour magnitude criteria to select AGB stars with dust shells and seperate into classes.

- 1. colour plots showing distribution of object types in survey
- 2. set magnitude limits to isolate certain objects. sec 2.
- 3. heat map distributions of colors
- 4. chose 12 colours, only 3 independent. Used the 3 to classify objects
- 5. use two-sided Kolmogorov Smirnov test to test distribution hypothesis. Sec 3.3.4
- 6. created model to predict object type based on colour

4 CLaSPS: new method for knowledge extraction

Using unspurvised clustering to identify correlations among astronomical obersvations.

- 1. use combination of features and labels. We have colour features for objects, could we use labels as well?
- 2. use a score and fraction of objects similar to our summary.
- 3. use Kmeans and vary number of clusters applied to data set