

# JANUS Cosmological Model Constraints from Type Ia Supernovae: Reproduction and Pantheon+ Extension

Patrick Guerin\*

Independent Researcher, Brittany, France

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*Data:* <https://github.com/PGPLF/JANUS-S>

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## Abstract

We reproduce D'Agostini & Petit (2018) JANUS model constraints using JLA supernovae (740 SNe Ia), obtaining  $q_0 = -0.086 \pm 0.014$  with  $\chi^2/\text{dof} = 0.88$ , validating the original result ( $q_0 = -0.087$ ). Extension to Pantheon+ (1543 SNe Ia) yields  $q_0 = -0.035 \pm 0.014$ , revealing significant dataset dependence ( $\Delta q_0 = 0.05$ ). Sub-sample analysis shows  $q_0$  evolution:  $-0.26$  at  $z < 0.1$  to  $\sim 0$  at high- $z$ . Comparison with  $\Lambda\text{CDM}$  shows comparable fits ( $\Delta\chi^2/\text{dof} < 4\%$ ) with slight  $\Lambda\text{CDM}$  preference ( $\Delta\text{AIC} \approx -25$ ).

## 1 Introduction

The JANUS bimetric model (Petit and D'Agostini, 2014; D'Agostini and Petit, 2018; Petit et al., 2024) explains cosmic acceleration through positive/negative mass sector interactions, avoiding dark energy. D'Agostini & Petit (2018) constrained the deceleration parameter to  $q_0 = -0.087 \pm 0.015$  using JLA data (Betoule et al., 2014).

We aim to: (1) reproduce the 2018 analysis, (2) extend to Pantheon+ (Brout et al., 2022), and (3) compare with  $\Lambda\text{CDM}$ .

## 2 Methods

### 2.1 JANUS Model

The luminosity distance is:

$$d_L(z) = \frac{c}{H_0} \left[ z + \frac{z^2(1-q_0)}{1+q_0z+\sqrt{1+2q_0z}} \right] \quad (1)$$

with distance modulus  $\mu = 5 \log_{10}(d_L/\text{Mpc}) + 25$ .

### 2.2 Data

**JLA:** 740 SNe Ia,  $0.01 < z < 1.30$ , using  $\mu = m_B - M_B + \alpha z - \beta c$  with  $\alpha = 0.141$ ,  $\beta = 3.101$ ,  $M_B = -19.05$ .

**Pantheon+:** 1543 unique SNe Ia,  $0.001 < z < 2.26$ , using calibrated MU\_SH0ES distances.

### 2.3 Fitting

Minimize  $\chi^2 = \sum_i [(\mu_i^{\text{obs}} - \mu_i^{\text{th}} - \delta)/\sigma_i]^2$  via Nelder-Mead. Bootstrap (100 samples) for uncertainties.

## 3 Results

### 3.1 2018 Reproduction

Table 1: JLA results comparison

	This work	Ref. (2018)
$q_0$	$-0.086 \pm 0.014$	$-0.087 \pm 0.015$
$\chi^2/\text{dof}$	0.883	0.89

### 3.2 Pantheon+ Extension

Table 2: Dataset comparison

	JLA	Pantheon+
$N$	740	1543
$q_0$	$-0.086$	$-0.035$
$\chi^2/\text{dof}$	0.883	0.497

\*Corresponding author: pg@gfo.bzh

### 3.3 Redshift Dependence

Table 3: Pantheon+ subsamples			
Range	$N$	$q_0$	$\chi^2/\text{dof}$
$z < 0.1$	583	-0.260	0.58
$z < 0.5$	1333	-0.165	0.50
$z < 1.0$	1518	-0.070	0.49
Full	1543	-0.035	0.50

### 3.4 JANUS vs $\Lambda$ CDM

Table 4: Model comparison

Dataset	JANUS	$\Lambda$ CDM	$\Delta\text{AIC}$
JLA	0.883	0.852	-24
Pantheon+	0.497	0.481	-26

## 4 Discussion

The 2018 reproduction is excellent ( $\Delta q_0 = 0.001$ ). The JLA/Pantheon+ discrepancy ( $\Delta q_0 = 0.05$ ) reflects: (1) different  $z$  distributions, (2) calibration differences, (3) possible  $q_0(z)$  evolution.

The  $q_0$  redshift trend (Fig. 3) suggests single-parameter JANUS may be insufficient for extended  $z$  ranges, motivating theoretical extensions.

Both models fit comparably;  $\Lambda$ CDM is slightly preferred statistically but JANUS remains competitive with one free parameter.

## 5 Conclusions

1. 2018 reproduction validated:  $q_0 = -0.086$ ,  $\chi^2/\text{dof} = 0.88$
2. Pantheon+ yields different  $q_0 = -0.035$
3. Evidence for  $q_0(z)$  evolution: -0.26 to  $\sim 0$
4. JANUS and  $\Lambda$ CDM comparably fit SNe Ia data

## Acknowledgments

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## References

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## Figures

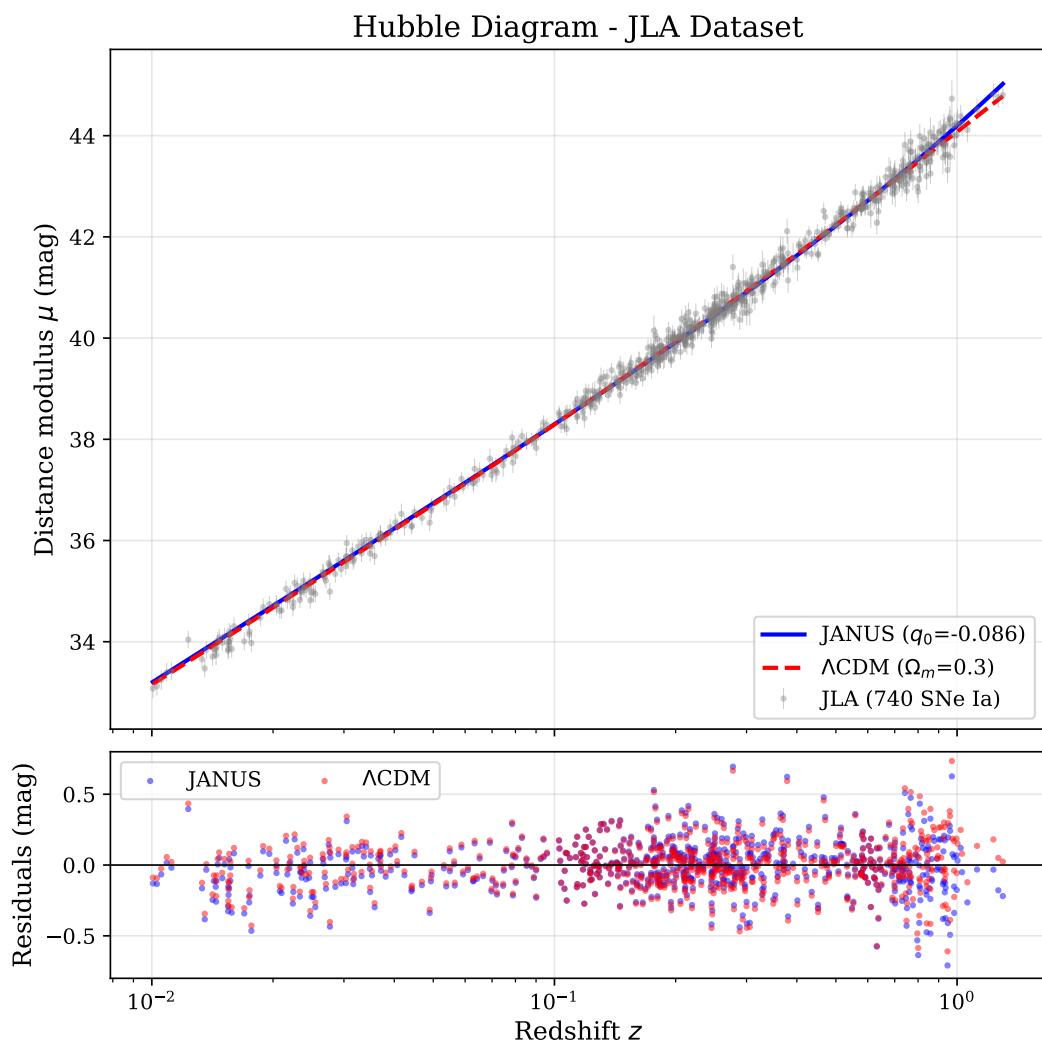


Figure 1: JLA Hubble diagram with JANUS ( $q_0 = -0.086$ ) and  $\Lambda$ CDM fits.

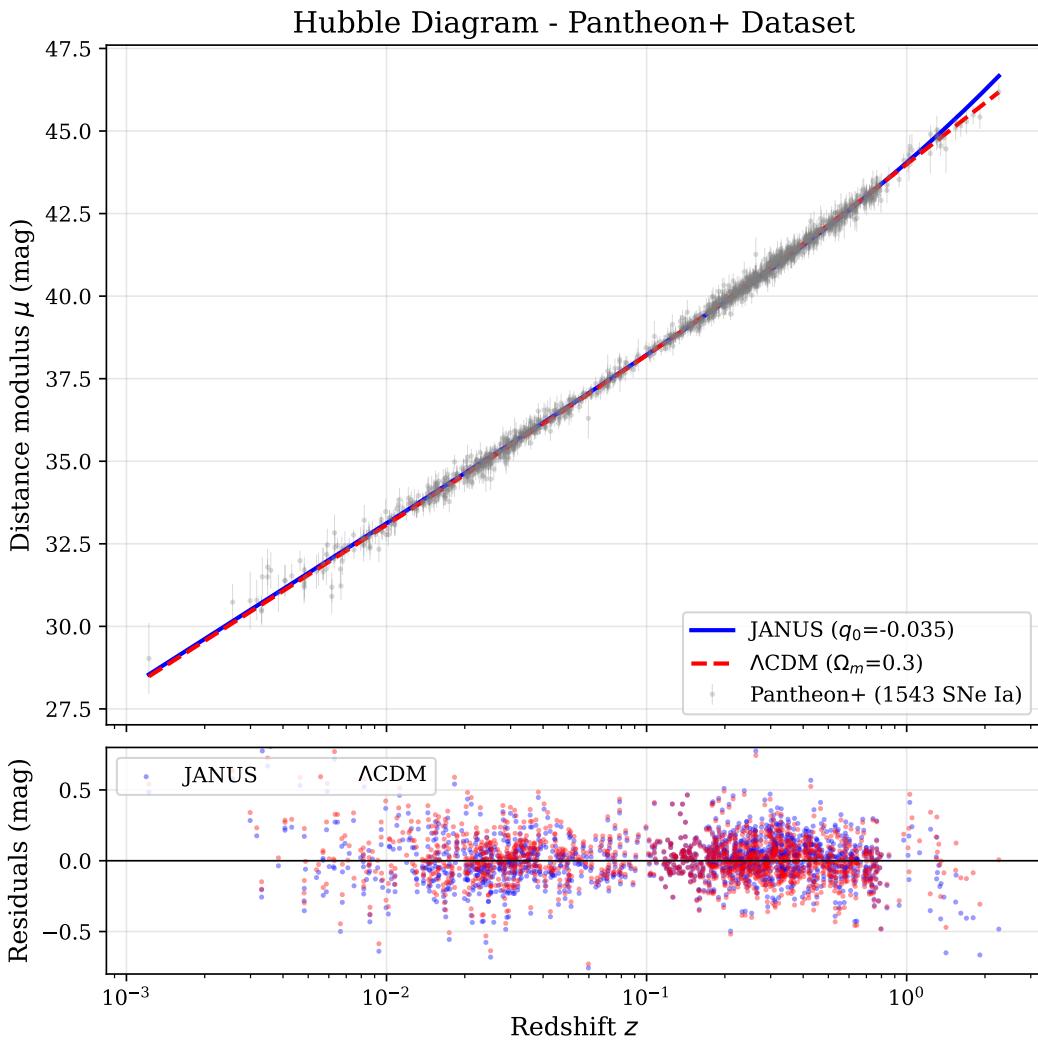


Figure 2: Pantheon+ Hubble diagram with JANUS ( $q_0 = -0.035$ ) and  $\Lambda$ CDM fits.

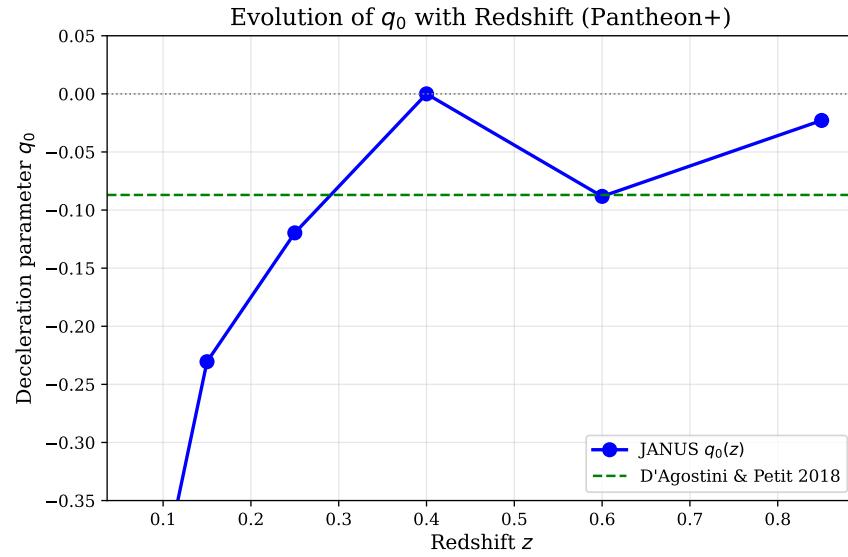


Figure 3:  $q_0(z)$  evolution from Pantheon+ subsamples. Dashed: 2018 value.

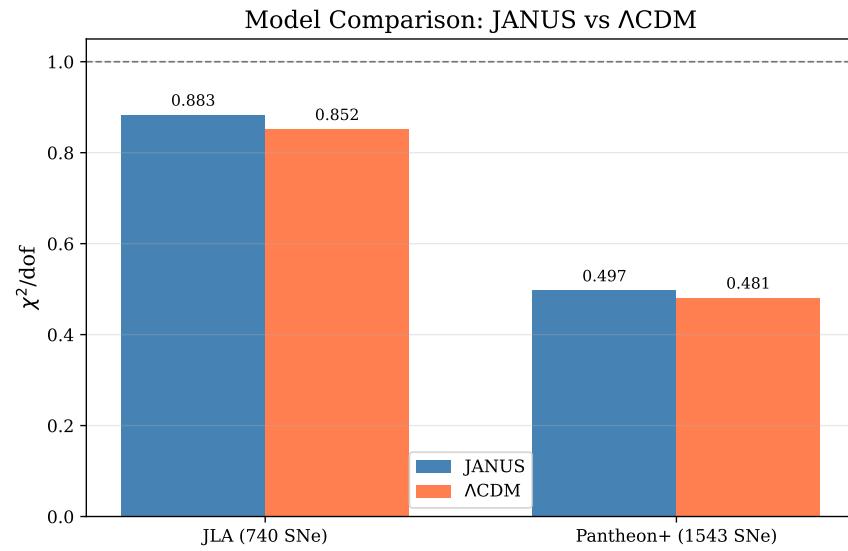


Figure 4:  $\chi^2/\text{dof}$  comparison: JANUS vs  $\Lambda$ CDM.