

Band-Pass Filters For Visual & Video Astronomy

by

Jim Thompson, P.Eng

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1.0 Introduction

Almost a year ago now I shared my astronomy filter research with the OAOG and Mallincam community. That personal project came about from my own selfish desire to know as much as I could about filters, both planetary (colour) and deep-sky (band-pass), so that I could make an informed purchase and later informed application of these filters. At the time I was devoted to observing visually, primarily from my backyard in Ottawa. Since that time I have been drawn into the world of video astronomy, and as a result I have been the proud owner of a Mallincam Xtreme since early January this year. With this recent shift in my observing methodology, some rework of my filter research was required.

2.0 Scope

A lot of technical data was compiled on band-pass type astronomical filters in my original research, but that data was compiled from the perspective of using the filters for visual observation only. Opening up the use of these filters to video astronomy as well results in some of the filters I ignored before now being of interest. Video astronomy also provides a very convenient method of comparing the performance of different filters not just using plots and % Luminous Transmissivity (%LT) numbers, but visually as well.

This whitepaper presents the revision I've made to my filter research specifically related to use with my Mallincam astrovideo camera. The following changes specifically are included:

- added 18 filters, mostly H-alpha which are not suitable for visual use but are suitable for video astronomy;
- calculated %LT for all filters based on spectral response of Sony CCD used in the Mallincam; and
- captured images using my Mallincam Xtreme of a fixed target (M42), using a range of band-pass filters, over a range of integration times.

The final task in my scope is an important one, as it provides some way of seeing firsthand how the use of a particular filter affects not only the view but technical issues like total integration time which in turn affects the importance of tracking accuracy (guided vs. unguided). The images collected are also useful for the visual observer since they give an idea what one can hope to achieve for contrast improvement, as will be discussed later.

3.0 Detector Spectral Response

In my filter research from a year ago I presented what is understood to be the sensitivity of the human eye to various wavelengths of light, both when light adapted (photopic) and dark adapted (scotopic). In order to consider the Mallincam or other video imaging device, a spectral sensitivity for this type of sensor was needed. The technical specification document for the Sony CCD sensor installed in the Mallincam has the spectral response of each of the four colour channels, CMYG. I took a straight average of these responses to get the average sensor response, as shown in Figure 1. The normalized average response is used later in the calculation of %LT. The spectral sensitivity of the Mallincam is compared to the human eye in Figure 2.

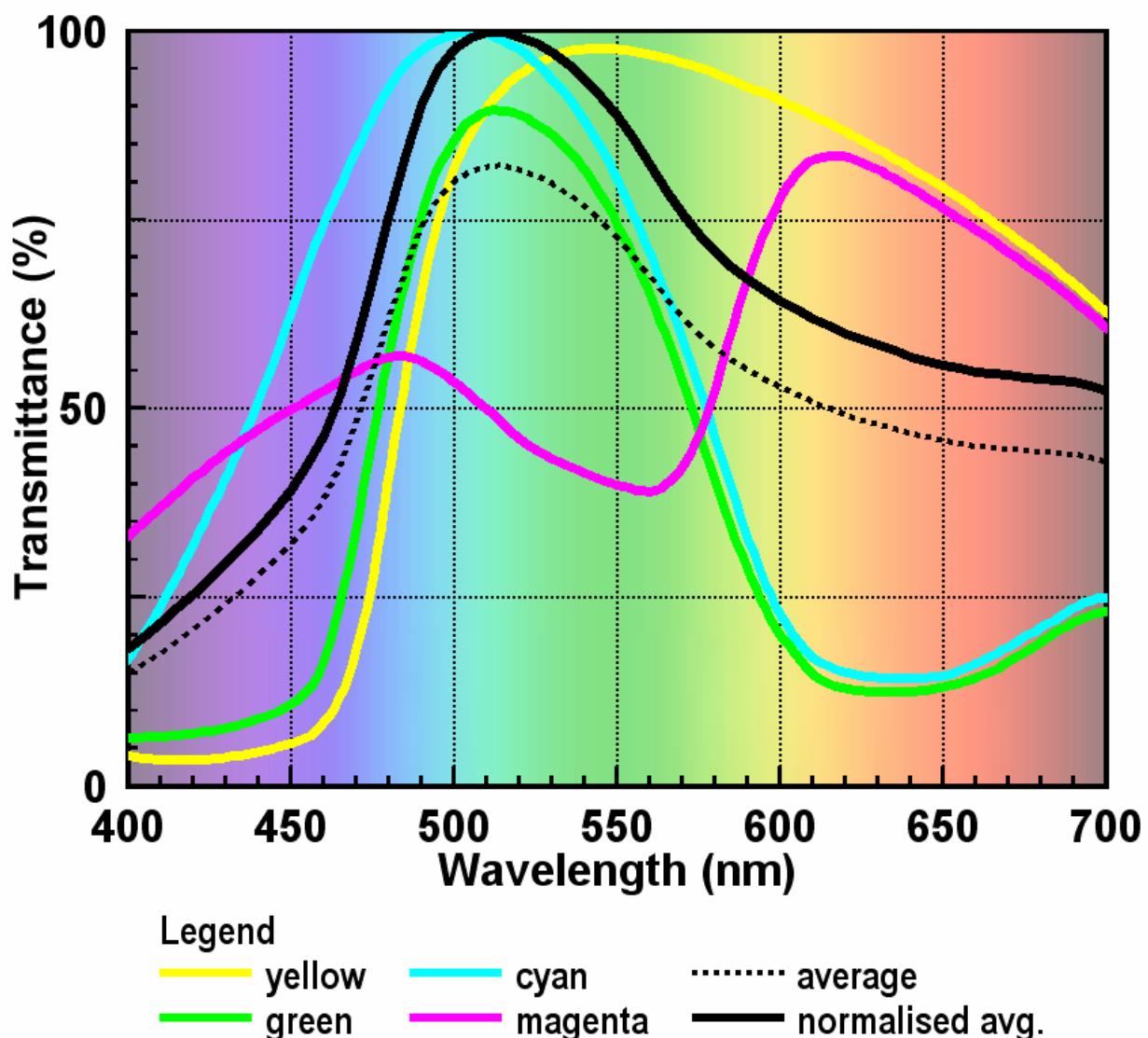


Figure 1 Spectral Response Of The Sony ICX418AKL-A CCD

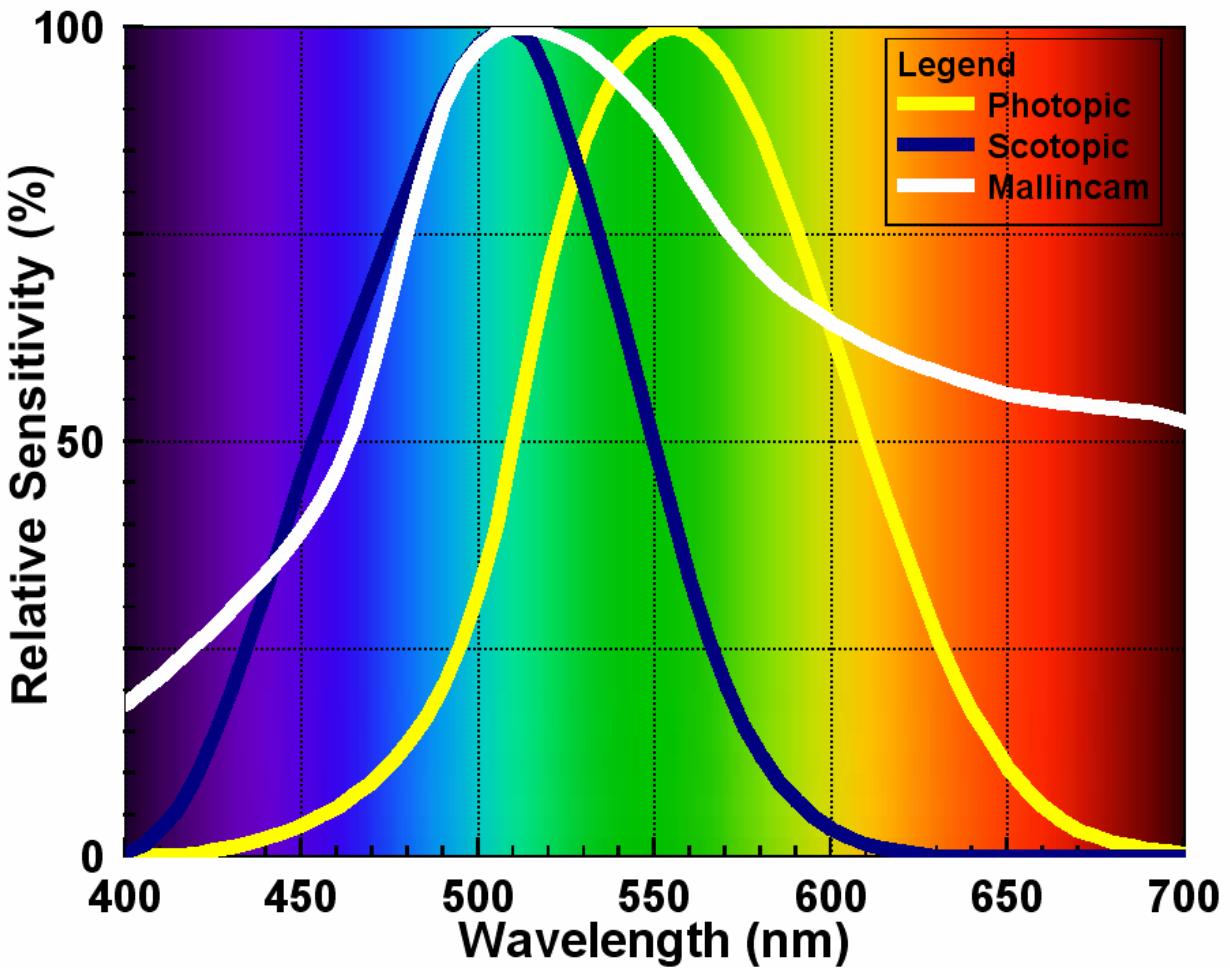


Figure 2 Relative Spectral Response - Human Eye vs. Mallincam

The Mallincam, as well as many other (unfiltered) CCD devices, has a much higher sensitivity in the red end of the spectrum, but is otherwise very similar to the human eye.

4.0 Filter Manufacturers, Types, & Categories

Using band-pass filters with CCD devices opens up a range of wavelengths that the human eye is not normally sensitive to, from 600 to 700nm and even into the infrared (700+). In my initial research I completely ignored H-alpha type filters since they have no application in visual astronomy. They do have application to video astronomy however, so I have now added them to my already exhaustive list of filters. Eighteen filters in total have been added, 13 H-alpha filters plus a couple of miscellaneous ones: Astronomik OIII CCD, Astronomik H-beta CCD,

Astronomik CLS CCD, Baader Planetarium Solar Continuum, and FLI OIII. This brings the total to 82 filters!

Appendix A contains a summary table of all 82 band-pass type filters. Appendix B contains the spectral transmissivity plots for these filters. The same filter categories defined in my first whitepaper are re-used again here, with the addition of the H-alpha filters. Table 1 below summarizes the 13 filter categories as well as the recommended use of each category of filter. Also listed in the table is the minimum recommended telescope aperture for each filter category when they are being used for visual astronomy. For those of you looking for a limit that is more filter specific, Figure 3 shows the relationship between filter % Scotopic Luminous Transmissivity (%SLT) found in Appendix A and telescope aperture. Table 1 also shows the amount by which one can expect the Mallincam integration time to increase for a fixed target when the particular category of filter is used. A more filter specific value of integration time increase can be calculated by simply taking the inverse of the % Mallincam Luminous Transmissivity (%MLT) which is also found in the table in Appendix A. For example a filter with %MLT = 50% would need an integration time that is approximately 2x longer than without a filter.

| Category | Prerequisite | Application | Min. Aperture for Visual Use | MC Integration Time |
|-----------------|---|--|---|---------------------|
| O-III Group A | Allow both doubly ionized Oxygen wavelengths to pass | View & image emission/planetary nebulae & supernova remnants under heavy light pollution | 5.5" (140mm) | 4-14x |
| O-III Group B | Allow only one doubly ionized Oxygen wavelength to pass | Image emission/planetary nebulae & supernova remnants under heavy light pollution | not suitable | 9-40x |
| H-beta Group A | Pass H-beta wavelength with >90% transmission | View & image faint emission nebulae, with or without light pollution | 8" (203mm) | 10-21x |
| H-beta Group B | Pass H-beta wavelength with <90% transmission | Image faint emission nebulae, with or without light pollution | not suitable | 10-48x |
| H-alpha Group A | H-alpha pass band is >10nm wide | Image emission/planetary nebulae & supernova remnants under heavy light pollution | not suitable | 5-40x |
| H-alpha Group B | H-alpha pass band is <10nm wide | Image emission/planetary nebulae & supernova remnants under heavy light pollution | not suitable | 42-71x |
| Narrow Band | H-beta + O-III pass band is <35nm wide | View & image emission/planetary nebulae & supernova remnants under moderate-to-no light pollution | 4.5" (114mm) | 3-9x |
| Medium Band | H-beta + O-III pass band is >35 but <50nm wide | View & image emission/planetary nebulae & supernova remnants under moderate-to-no light pollution | 3.5" (89mm) | 3-5x |
| Wide Band | H-beta + O-III pass band is >50 but <70nm wide | View emission/planetary nebulae & supernova remnants under mild-to-no light pollution; image all deep-sky objects | 2.5" (64mm) | 2-4x |
| Extra Wide Band | H-beta + O-III pass band is >70nm wide | View or image all objects under mild-to-no light pollution | no limit | 1.5-2x |
| Multi Band | More than two major pass bands in the visible range | View or image all objects under mild-to-no light pollution | no limit | 1.7-1.9x |
| Special A | Filters esp. designed for planets or other special object viewing | Baader Solar for Sun; Lumicon comet for comets; Orion Mars for Mars; all others for contrast improvement while viewing Moon or planets | Lumicon comet 5.5", all others no limit | 1.6-20x |
| Special B | Special filters for contrast enhancement based on Neodymium infused glass | Contrast improvement while viewing Moon or planets | no limit | 1.4-1.9x |

Table 1 Summary Of Band-Pass Filter Categories

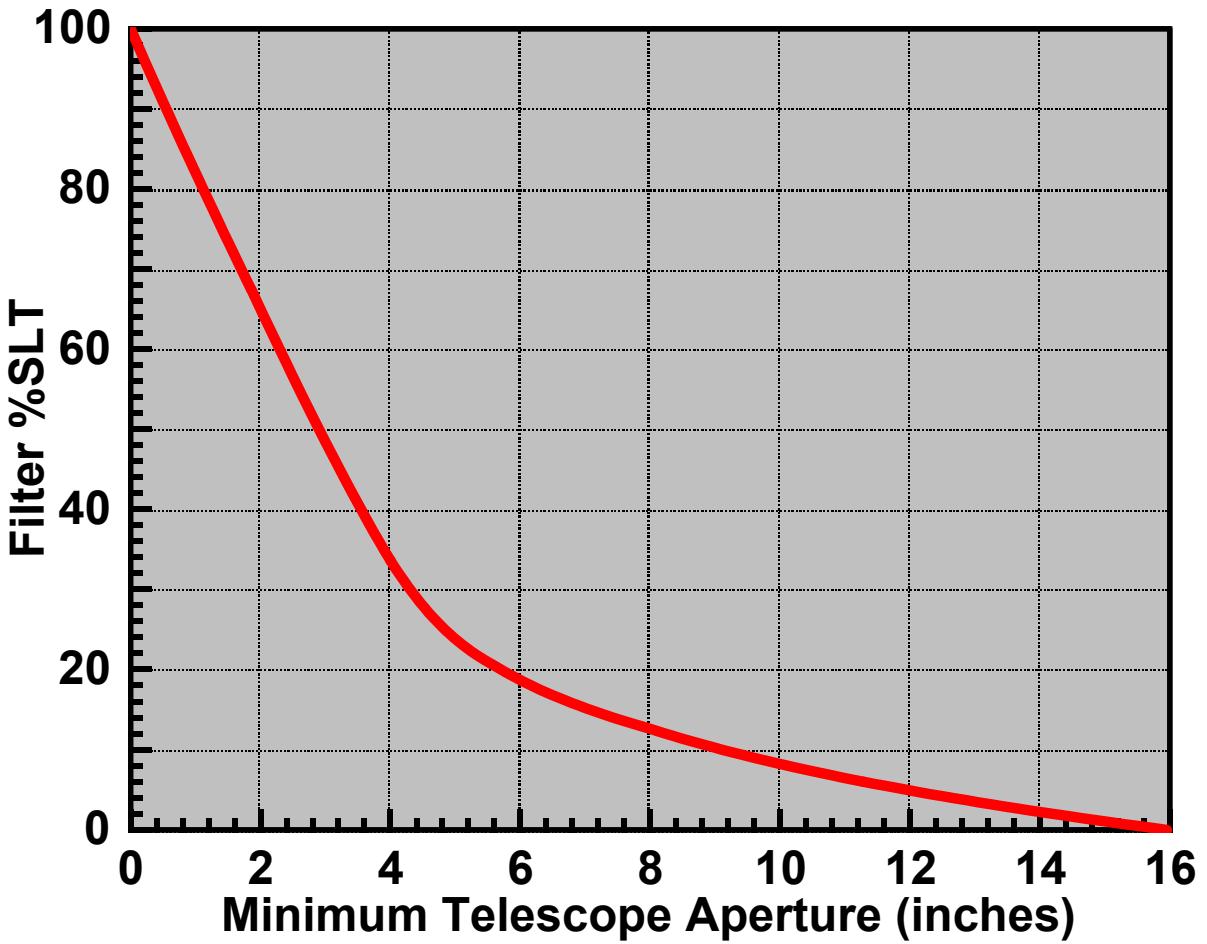


Figure 3 Minimum Recommended Telescope Aperture For Visual Observation

5.0 Visual Comparison Of Filters

Spectral transmissivity plots and %LT values are useful metrics for comparing different filters to each other. They are especially useful when deciding what filter to buy with a filter category. But how does one decide what filter category is best for their application? My suggested applications in Table 1 provide some guidance, but the real test for me is visually comparing the view that each filter produces. It has probably become obvious to everyone by now that I am a bit of a curious little monkey. In my fiddlings and experimental viewings I have compiled a fairly broad sampling of band-pass type filters:

- Baader Planetarium Moon & Skyglow
- Astro Hutech IDAS LPS-P2

- Orion Skyglow Broadband
- Baader Planetarium UHC-S
- Astronomik UHC
- Astronomik H-beta
- Astronomik OIII
- Lumicon #29 Dark Red + Baader Planetarium UV/IR Blocker

The last filter in my list was just for fun...boy was I pleasantly surprised by the results, as you will see later. Figures 4 and 5 show the spectral transmissivity curves for the eight filters.

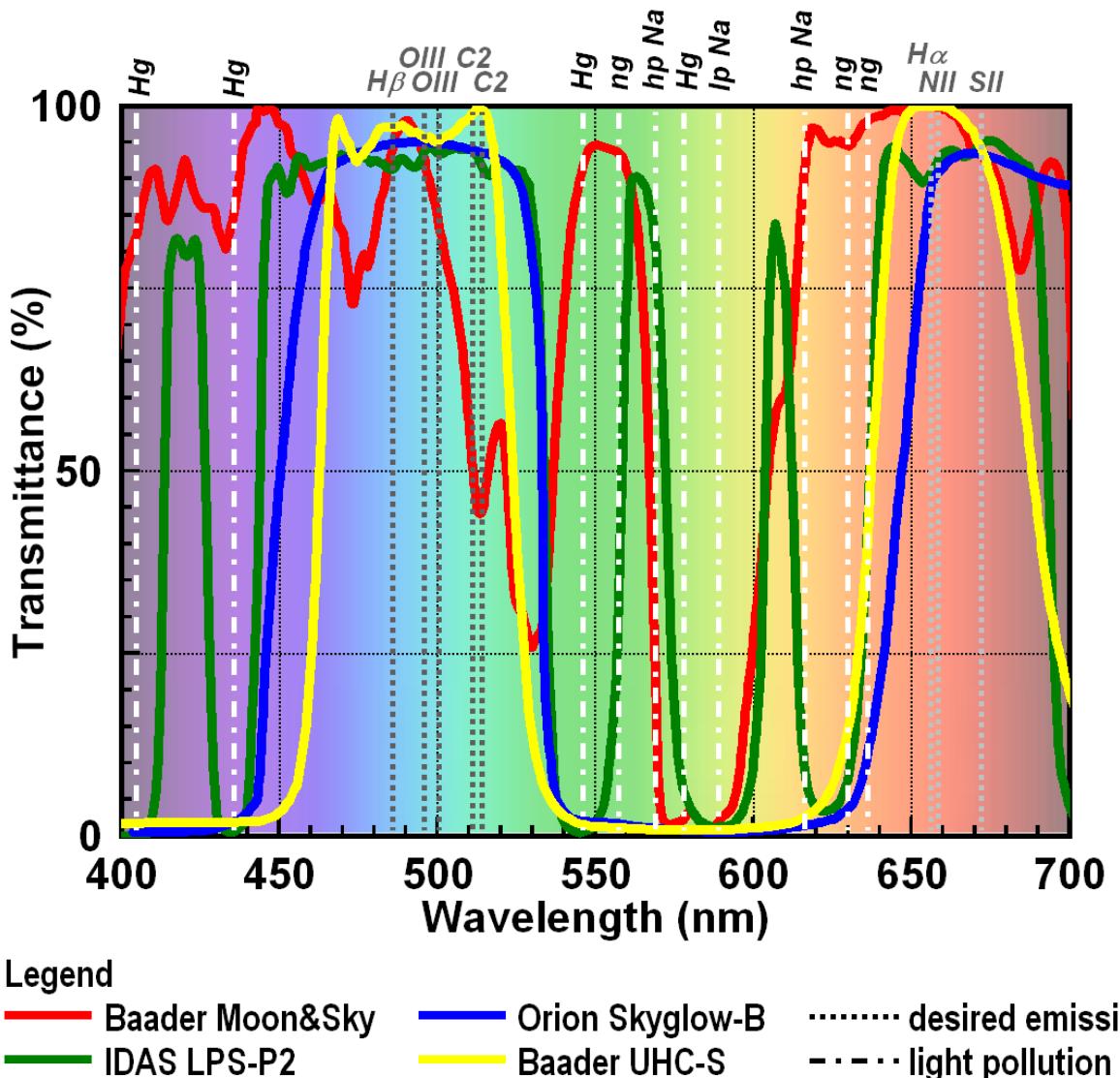


Figure 4 Band-Pass Filters Used In Test - Part 1

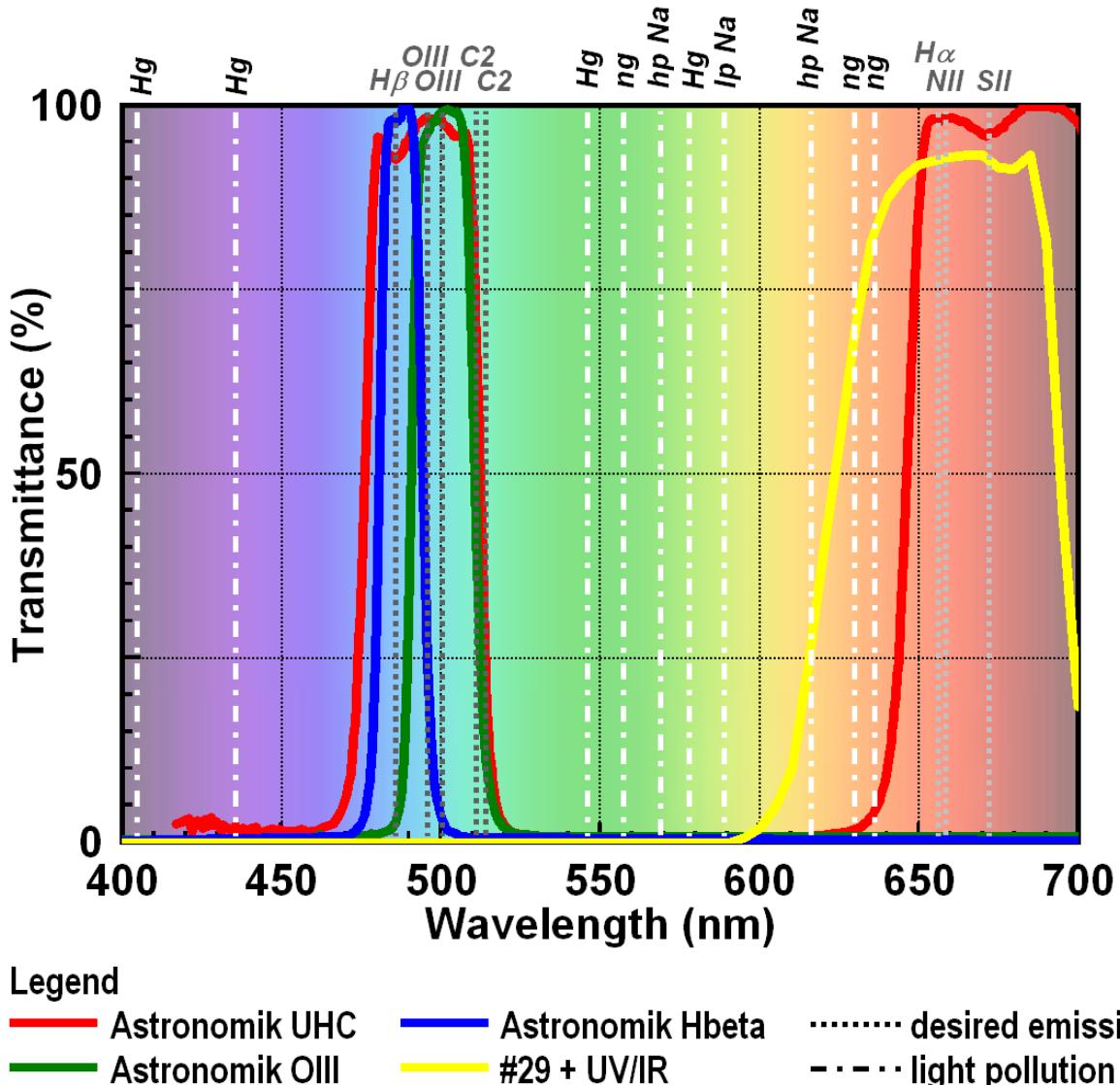


Figure 5 Band-Pass Filters Used In Test – Part 2

My images were collected in as consistent a manner possible, with the lovely M42 as the target. I used my Meade 8" LX10 telescope with Meade f/3.3 focal reducer on my Orion Atlas EQ/G mount, plus my Mallincam Xtreme astrovideo camera to collect the images. Each image was recorded at a fixed Hue (60), Saturation (75), Sharpness (2), AGC (5), and APC (8). Brightness was adjusted for each image to give a well balanced image with dark background, and Contrast was held mostly at 40 with some images varying up to 45 and 50. I recorded images at a fixed set of integration times: 2.1s (integration off), 4sec, 8sec, 12sec, 16sec, 24sec, 32sec, 48sec, 60sec, 75sec, 90sec. For each integration time I recorded images from three refreshes, which were then later hand stacked into one image for comparison purposes. No additional processing was performed on the per-integration time images. The images can be found in Appendix C. Note that the H-beta, OIII, and #29 Dark Red filters were recorded in B&W since the filters have such narrow wave bands they only let one colour through anyway, cyan, green, and red

respectively. Many of the filters did not make it to 90sec integration as the view was totally over exposed and/or I ran out of brightness adjustment. In addition to the per-integration time images, I made multi-exposure hand stacks using 4 different integration times, selected to have as close to the same relative exposure between each filter tested. The results from this can be found below in Figure 6. I did use tone balancing on these multi-exposure stacks in order to give the sexiest image possible with the selected image data.



No Filter



Baader Moon&Sky



IDAS LPS-P2



Orion Skyglow-B



Baader UHC-S



Astronomik UHC



Astronomik H-beta



Astronomik OIII



#29 Dark Red + UV/IR Blocker

6.0 Conclusions

In the end I was very happy with the results of my testing. Some of the filter images are slightly blurry, suggesting that I did not have the focus quite right on them, but all-in-all the images turned out nice. My conclusions on each of the filters tested are as follows:

None: The stacked image looks pretty good surprisingly, but there is an obvious hue to the background due to the light pollution. Integration times were as short as they can be, plus there are lots of stars, which is okay on the periphery, but in the trapezium area it results in the image being over exposed.

Baader M&S: Seems to provide a slight improvement over no filter. Hue due to light pollution seems to be mostly gone. Integration times still short.

LPS-P2: Provides nice colour balance and finer details in the nebula. Background is dark, and integration times are low. Stars still fairly bloated at longer integration times.

Orion Skyglow-B: More intense reds, but blues still visible. Contrast similar to slightly worse than LPS-P2. At longer integration times, stars had strange coma shape.

Baader UHC-S: Image very red, almost orange. Contrast similar to slightly less than Orion Skyglow-B.

Astro. UHC: The nicest colour image in my opinion. Very nice colour balance, and wispy blues much more refined. Finer contrast and detail in dark lanes, esp. in M43. Very nice dark background, and stars are small and round.

Astro. H-beta: Okay contrast, but very long integration time to get it. Seem to get the same from #29 but at shorter integration time.

Astro. OIII: Not very exciting, interesting how only the blues (OIII) regions are coming through and none of the H-alpha (red), but the resulting image is not very pleasing in my opinion. Very long integration times.

#29 + UV/IR: Wow, what a pleasant surprise! This very inexpensive, very wide band H-alpha filter produced what I feel are very impressive images. The best for contrast overall I think. I wonder what a proper H-alpha filter would do? Hmm....

In my opinion, the LPS-P2 and Astronomik UHC gave the best detailed and colour balanced images. I am however curious now what a filter from the Narrow Band category would produce.

There is a definite trade off between higher contrast in the image and longer integration times. I was lucky to get a good polar alignment for these images and had good weather to leave my setup outside for 3 days so that I could get all the images. I had good tracking up to 90sec with no guiding during my tests, but most of the time (with having to setup and teardown every night) I can only get 30-40sec.

For the visual observers out there, I hope you can get some useful information out of this work as well. For example, if you equate integration time to the light gathering capacity of a telescope (roughly equal to aperture²), you can get an idea what impact bigger apertures have on your ability to see detail. You can also visually see how the image gets more contrast but dimmer as a result of using each filter when you compare images of the same integration time.

One important behavior that bares mention is how band-pass filters respond in short focal length telescopes. The spectral transmissivity curve for these filters tends to shift back and forth with wavelength as the angle of the filter to the light beam is changed. In the case of short focal length scopes, the angle of the light passing through the middle of the filter is different enough from the light passing through the outer edge that there will be a visible difference in performance. This centre-to-edge variation is progressively more important as the filter's pass band gets narrower, and may become a serious issue with the narrowest filters (eg. H-beta, OIII, H-alpha, & Narrow Band). I don't have any observational data to know what the practical lower focal length limit is, so filter users should just be aware and keep their eyes open for this phenomenon, especially if they use small aperture fast focal ratio refractors or possibly even the Hyperstar lens system some have on their SCT. One way of mitigating this problem is to move the filter forward in the optical train, ahead of focal reducers, where the light rays are more parallel (ie. area of longer affective focal length).

I hope my little whitepaper is interesting and useful to someone out there. I am convinced that large improvements to the quality of the image can be achieved when combined with the proper telescope aperture (visual & video) or appropriate integration time (video only). This is definitely true for emission/planetary nebulae and super nova remnants. I assume the impact will be less so for clusters, galaxies, and reflection nebulae since they are broad spectrum targets (filter cuts emission from target along with cutting out light pollution). If you have any questions, feel free to contact me.

Cheers!

Jim Thompson

jimmythepuker2@yahoo.ca

Appendix A Band-Pass Filter Summary Table

Deep-Sky Filter Summary Table

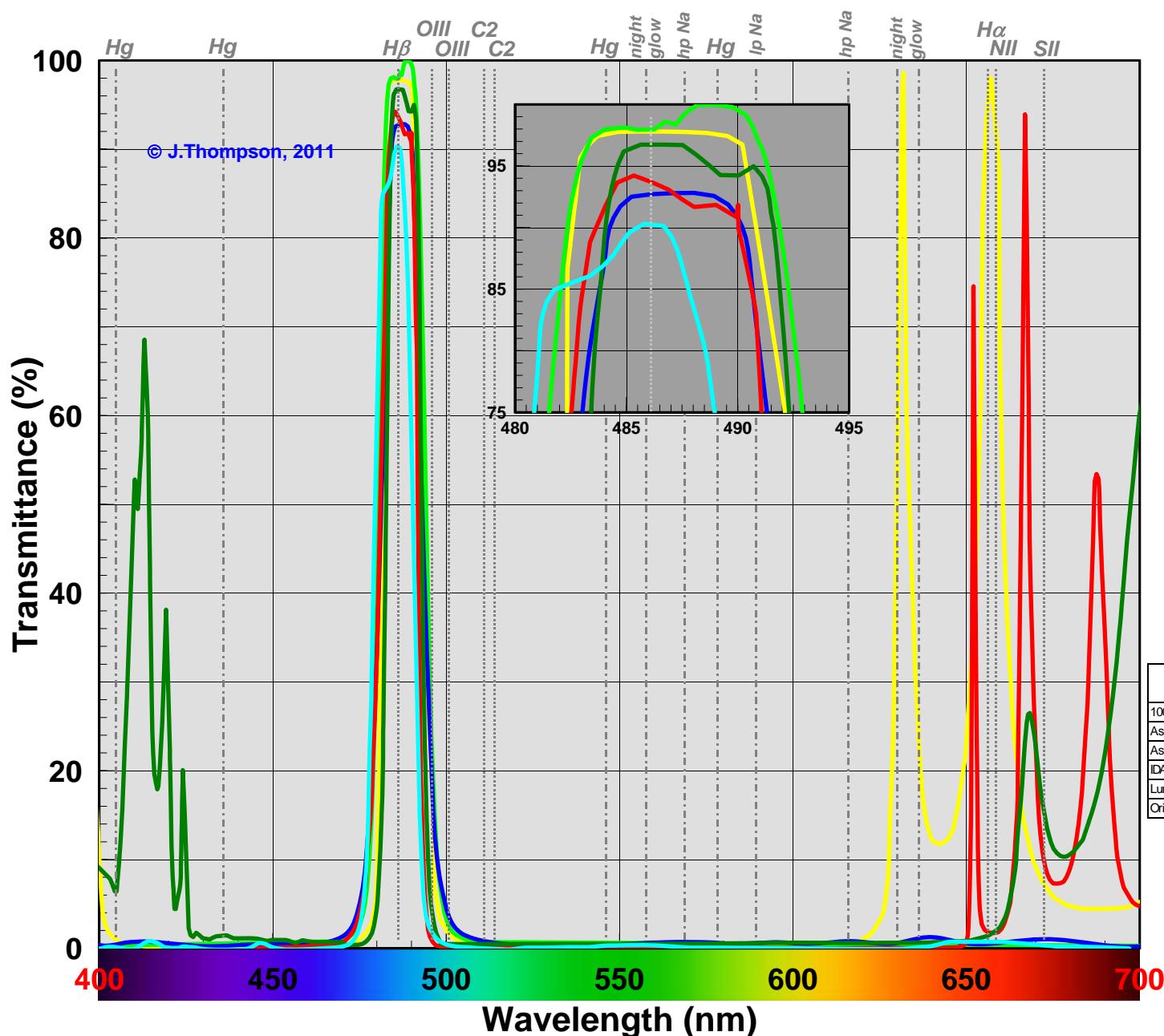
| # | Manufacturer | Model | Full Name | Manuf.'s Recommended Use | Contact | Category | % Luminous Transmissivity | | |
|----|---------------------|------------|-------------------------------------|---|--|-----------------|---------------------------|----------|-----------|
| | | | | | | | Photopic | Scotopic | Mallincam |
| 1 | 1000 Oaks | LP-1 | LP-1 Broadband | Slight to moderately light polluted areas | www.thousandoakoptical.com | wide band | 16.8 | 50.6 | 39.0 |
| 2 | | LP-2 | LP-2 Narrowband | Heavily light polluted areas | | narrow band | 9.3 | 26.5 | 18.7 |
| 3 | | LP-3 | LP-3 O-III | Diffuse & planetary nebulae | | O-III A | 4.5 | 11.8 | 10.9 |
| 4 | | LP-4 | LP-4 H-beta | Horsehead, California, & other faint nebulae | | H-beta A | 4.2 | 10.8 | 10.4 |
| 5 | Andover Corporation | O-III | O-III | Heavily light polluted areas, planetary nebulae | www.andovercorp.com | O-III A | 4.4 | 14.2 | 7.2 |
| 6 | | 3ch Nebula | 3 Channel | Light polluted areas, planetary & emission nebulae | | narrow band | 12.5 | 32.9 | 30.9 |
| 7 | Antares | ALP | Anti-Light Pollution, Broadband | block light pollution, improve view of emission nebulae | www.antaresoptical.com | wide band | 26.0 | 59.6 | 46.6 |
| 8 | Astrodon | H-alpha1 | H-alpha 3nm | CCD imaging | www.astrodon.com | H-alpha B | 0.4 | 0.0 | 1.4 |
| 9 | | H-alpha2 | H-alpha 5nm | CCD imaging | | H-alpha B | 0.4 | 0.0 | 1.6 |
| 10 | Astro Hutech | O-III | IDAS O-III | narrow-band nebular + O-III passband modifier filter | www.sciencecenter.net/hutech/ | O-III A | 7.3 | 16.2 | 20.7 |
| 11 | | H-beta | IDAS H-beta | narrow-band nebular + Hbeta passband modifier filter | | H-beta A | 2.6 | 10.2 | 9.1 |
| 12 | | LPS-P1 | IDAS Light Pollution Suppression v1 | Light polluted areas, nebulae viewing, color balanced photography | | multi band | 46.3 | 73.5 | 56.0 |
| 13 | | LPS-P2 | IDAS Light Pollution Suppression v2 | Light polluted areas, nebulae viewing, color balanced photography | | multi band | 44.9 | 72.7 | 59.2 |
| 14 | | LPS-V3 | IDAS LPS v3, Narrow Band Nebular | Provide maximum contrast between key emission lines and light polluted sky | | wide band | 22.5 | 54.3 | 38.0 |
| 15 | | LPS-V4 | IDAS LPS v4, Narrow Band Nebular | Provide maximum contrast between key emission lines and light polluted sky | | wide band | 20.6 | 54.1 | 33.5 |
| 16 | | O-III | O-III | f3 to f15 & >6" aperture, substantial contrast gain on emission & planetary nebulae & supernova remnants | | O-III A | 6.6 | 20.5 | 10.4 |
| 17 | | O-III CCD | O-III CCD | CCD imaging | | O-III A | 6.2 | 20.9 | 10.8 |
| 18 | | H-beta | H-beta | f4.5 to f6 optimal, f3.5 to f15 possible, >8" aperture, dim hydrogen emission nebulae | | H-beta A | 2.5 | 12.6 | 6.3 |
| 19 | | H-beta CCD | H-beta CCD | CCD imaging | | H-beta A | 2.8 | 11.8 | 6.2 |
| 20 | Astronomik | UHC | Ultra High Contrast | LPR, better views of deep-sky-objects, f4 to f15 & >4" aperture, CCD and DSLR photography | www.astronomik.com | medium band | 11.8 | 33.6 | 33.0 |
| 21 | | UHC-E | Ultra High Contrast - Economy | increase contrast between target and night sky, well suited for smaller telescopes <5", comets, jupiter's clouds, double stars, LPR for photography | | medium band | 23.9 | 42.5 | 38.2 |
| 22 | | CLS | "Clear Sky"? | budget filter for visual LPR, very good colour balance for photography (film or digital) | | extra wide band | 31.1 | 67.5 | 52.0 |
| 23 | | CLS CCD | "Clear Sky"? CCD | CCD imaging | | extra wide band | 24.3 | 65.8 | 47.6 |
| 24 | | H-alpha1 | H-alpha 6nm | CCD imaging | | H-alpha B | 0.8 | 0.5 | 2.4 |
| 25 | | H-alpha2 | H-alpha 13nm | CCD imaging | | H-alpha A | 1.9 | 0.5 | 5.9 |
| 26 | | O-III | O-III Visual Nebula | maximum contrast and image sharpness, emission & planetary nebulae | | O-III B | 2.0 | 6.5 | 3.3 |
| 27 | | H-beta | H-beta Narrowband CCD | maximum contrast and image sharpness | | H-beta B | 1.1 | 5.4 | 2.7 |

| | | | | | | | | | |
|----|-----------------------|------------------|--|--|--|-----------------|------|------|------|
| 28 | | UHC-S | UHC-S Nebula | improved contrast over typical broadband filters without sacrificing stars like other UHC filters, great for smaller scopes, good for imaging | | wide band | 22.5 | 54.7 | 42.4 |
| 29 | | Moon&Sky | Moon & Sky Glow | neodymium infused glass, enhances both planetary & deep sky contrasts by reducing skyglow from LP & Moon, RGB intensifier | | special B | 55.3 | 72.3 | 70.7 |
| 30 | Baader Planetarium | Contrast | Contrast Booster | neodymium infused glass + minus violet filter, boosts lunar & planetary contrast, cuts skyglow, totally eliminates de-focused blue halo in achromats, natural colour balance, great for Mars | | special B | 53.1 | 48.9 | 56.0 |
| 31 | | Solar | Solar Continuum | enhance solar granulation & sunspot detail, boost contrast & reduce atmos. Turbulence, also good for star testing telescopes | | special A | 8.4 | 7.0 | 4.9 |
| 32 | | H-alpha1 | H-alpha 7nm | CCD imaging | | H-alpha B | 0.5 | 0.0 | 2.2 |
| 33 | | H-alpha2 | H-alpha 35nm | CCD imaging | | H-alpha A | 4.7 | 0.0 | 9.7 |
| 34 | Burgess Optical | LPR | Broadband Nebula - Light Pollution Reduction | a fine light pollution filter that passes a very high percentage of light originating from stellar sources, blocks light at wavelengths typically found in outdoor lighting | www.burgessoptical.com | wide band | 26.9 | 47.7 | 26.8 |
| 35 | O-III | Narrowband O-III | one of most commonly used filters by researchers and serious amateurs, useful for capturing high resolution images with high light pollution | | O-III B | 1.5 | 4.8 | 2.5 | |
| 36 | Custom Scientific | H-beta | Narrowband H-beta | one of most commonly used filters by researchers and serious amateurs, useful for capturing high resolution images with high light pollution | | H-beta B | 0.8 | 4.3 | 2.1 |
| 37 | | Multiband | Multiband H-beta / O-III / H-alpha | one of most commonly used filters by researchers and serious amateurs, useful for capturing high resolution images with high light pollution | | narrow band | 5.5 | 21.5 | 12.9 |
| 38 | | H-alpha1 | H-alpha 4nm | CCD imaging | | H-alpha B | 0.4 | 0.0 | 1.5 |
| 39 | | H-alpha2 | H-alpha 10nm | CCD imaging | | H-alpha A | 0.6 | 0.0 | 2.5 |
| 40 | DGM | O-III | High Performance O-III | maximum enhancement of emission and reflective nebula | | O-III B | 10.4 | 10.7 | 11.4 |
| 41 | | NPB | Narrow Pass Band Nebula | UHC type, small & fainter emission & planetary nebula, retains natural star colours | | narrow band | 12.3 | 22.6 | 25.7 |
| 42 | | VHT | Very High Throughput Nebula | smaller scopes (4-6"), compromise between UHC and broadband, enhance view of emission & reflective nebula with minimum star dimming | | medium band | 14.1 | 33.3 | 34.5 |
| 43 | | GCE | Galaxy Contrast Enhancement | aids visual observation of galaxies & Milky Way dust clouds & dark lanes, general purpose LPR, most of visible spectrum passed | users.erols.com/dgmoptics/ | extra wide band | 33.4 | 67.7 | 62.8 |
| 44 | | O-III | Hi Def O-III | the square transmission curves mean only photons in the desired emission bandpasses of the observed object are viewed, red halos around stars will not be present, a more natural and contrasty view results | | O-III B | 2.0 | 7.8 | 7.9 |

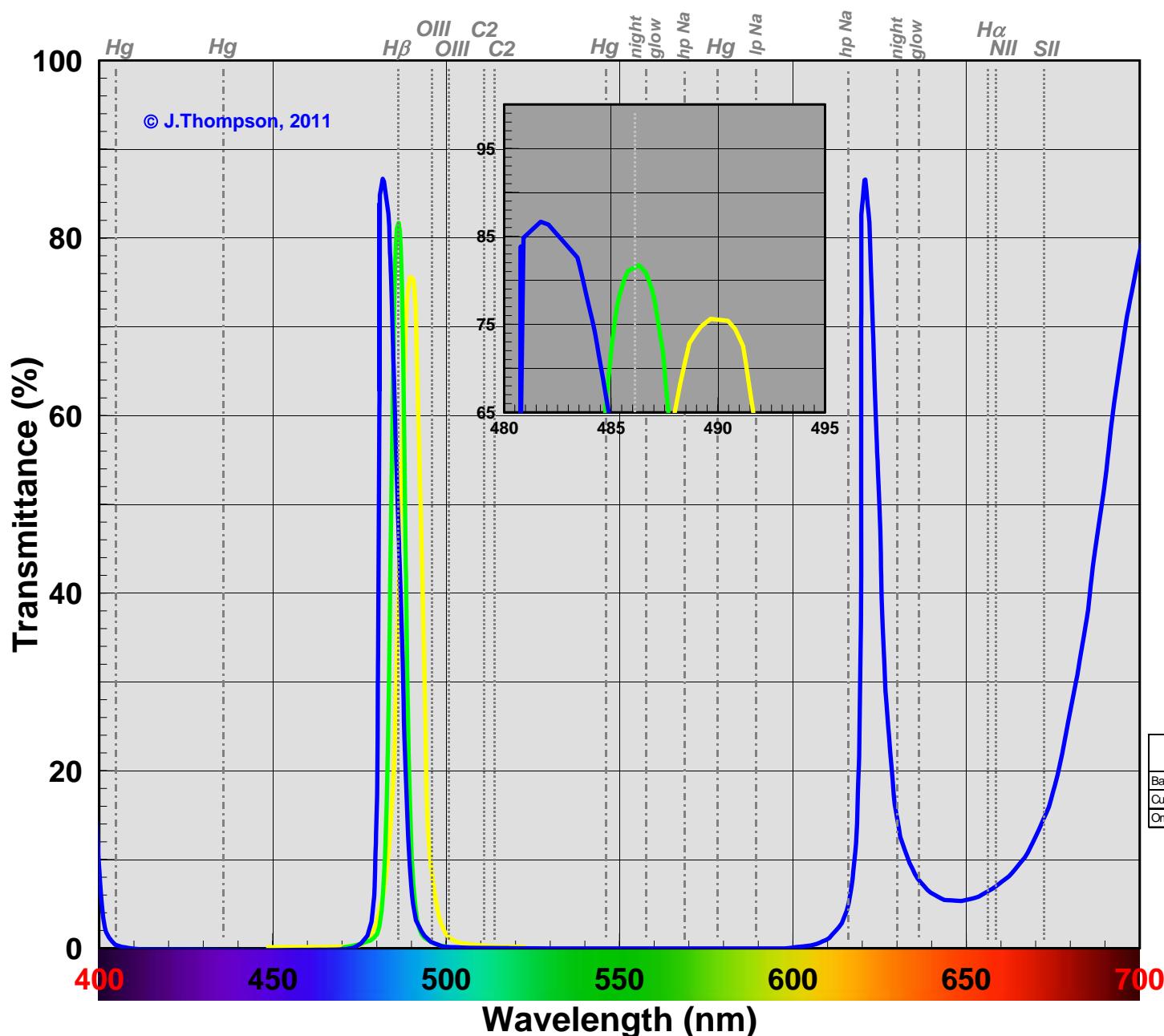
| | | | | | | | |
|----|-------------------|-----------|----------------------------|--|--|------|------|
| | | | | | | | |
| 45 | | UHC | Hi Def Ultra High Contrast | the square transmission curves mean only photons in the desired emission bandpasses of the observed object are viewed, red halos around stars will not be present, a more natural and contrasty view results | | | |
| 46 | Denkmeier Optical | Planetary | Hi Def Planetary | very unique contrast enhancement filter, see brighter greens and reds, greatly improve contrast of Mars, Jupiter, & Saturn | www.denkmeier.com | | |
| 47 | | O-III | O-III 8nm | CCD imaging | | | |
| 48 | FLI | H-alpha | H-alpha 8nm | CCD imaging | www.fliscamera.com | | |
| 49 | | O-III | O-III | narrow band pass, near-photographic views of Veil, Ring, Dumbbell, Orion, use on diffuse/planetary/faint nebulae, optimum exit pupil L2-5mm/D3-7mm | | | |
| 50 | | H-beta | H-beta | extremely faint nebulae like California, Cocoon & Horsehead, used best under clear skies & large aperture, optimum exit pupil L3-7mm/D4-7mm | | | |
| 51 | | UHC | Ultra High Contrast | superb views of Orion, Lagoon, Swan and other extended nebulae, best all-around dark-sky nebular filter, optimum exit pupil L1-4mm/D2-6mm | | | |
| 52 | | Deepsky | Deepsky | LPR, imaging of all types of deepsky objects, high contrast views of Martian polar caps, optimum exit pupil L0.5-2mm/D1-4mm | | | |
| 53 | | Comet | SWAN | enhances cyanogen wavelength in comet tails, narrow pass band allows OII and C2 | | | |
| 54 | | H-alpha | H-alpha Pass | CCD imaging | | | |
| 55 | Lumicon | #29 | #29 Dark Red | planetary viewing | www.lumicon.com | | |
| 56 | | O-III | O-III | Diffuse & planetary nebulae | | | |
| 57 | | Narrow | Narrowband Nebular | striking contrast between nebula and background, best with >25mm ep , not intended for photography, useful on fewer objects but those objects are greatly enhanced | | | |
| 58 | Meade | Wide | Wideband Nebular | LPR, photography, enhances nebula mostly but does improve contrast on galaxies | www.meade.com | | |
| 59 | | O-III | OIII Narrow CCD | precision interference filter, narrow band pass, best for CCD imaging | | | |
| 60 | | H-beta | Hb Narrow | precision interference filter, narrow band pass, best for CCD imaging | | | |
| 61 | | Wide | Hb&OIII Nebula II | precision interference filter, relatively broad pass band allwos Hbeta & OIII wavelengths | | | |
| 62 | | Narrow | Hb&OIII Nebula | precision interference filter, narrow pass band allows Hbeta & OIII wavelengths | | | |
| 63 | | Hg&Na | Hg&Na Skylight Reject | multi band interference filter, cuts out prominent light pollution wavelengths but maintains high overall transmission | | | |
| 64 | Omega Optical | Imaging | Colour Enhancing LPF | enhances recording of colour images, especially on bright objects like moon and planets, reduces light pollution, removes cyan and yellow that mute Hue | www.omegafilters.com | | |
| | | | | medium band | 10.7 | 38.8 | 24.6 |
| | | | | special A | 52.8 | 54.3 | 62.1 |
| | | | | O-III B | 1.9 | 6.1 | 3.1 |
| | | | | H-alpha B | 0.6 | 0.0 | 2.1 |
| | | | | O-III A | 3.8 | 12.6 | 9.9 |
| | | | | H-beta A | 2.4 | 10.1 | 8.6 |
| | | | | narrow band | 7.0 | 24.8 | 16.5 |
| | | | | wide band | 23.8 | 60.6 | 44.0 |
| | | | | special A | 9.3 | 21.3 | 11.0 |
| | | | | H-alpha A | 3.2 | 0.0 | 17.4 |
| | | | | H-alpha A | 7.3 | 0.1 | 21.6 |
| | | | | O-III A | 6.7 | 16.7 | 23.7 |
| | | | | narrow band | 8.9 | 28.1 | 25.2 |
| | | | | medium band | 10.2 | 37.8 | 29.2 |
| | | | | O-III B | 1.7 | 5.6 | 2.8 |
| | | | | H-beta B | 4.1 | 6.8 | 10.5 |
| | | | | wide band | 20.2 | 49.4 | 25.5 |
| | | | | medium band | 12.0 | 34.3 | 25.1 |
| | | | | extra wide band | 52.5 | 72.8 | 65.0 |
| | | | | multi band | 50.4 | 50.9 | 51.5 |

| | | | | | | | | | |
|----|----------------|------------|------------------------------|---|--|-----------------|------|------|------|
| 65 | Optec Inc. | O-III | O-III | Diffuse & planetary nebulae | www.optecinc.com/astronomy | O-III B | 3.2 | 9.7 | 4.9 |
| 66 | | Deepsky | Deepsky | blocks UV, violet, & sodium light completely | | extra wide band | 33.6 | 59.4 | 55.6 |
| 67 | | H-alpha | H-alpha | CCD imaging | | H-alpha B | 0.6 | 0.0 | 2.3 |
| 68 | | O-III | O-III | reveal more wondrous details when viewing nebulae, completely blocks all other visible wavelengths, for >8" aperture, visual use only | | O-III A | 4.2 | 13.8 | 7.2 |
| 69 | | H-beta | H-beta | visually capture elusive faint nebulae, perhaps only way to see Horsehead, California, & Cocoon, best with a moderate to large aperture scope & clear dark skies | | H-beta A | 1.8 | 9.5 | 4.7 |
| 70 | | Ultrablock | Ultrablock | for deep-sky observers at highly light polluted sites, enhances the sky presence of a significant number of fainter deep-sky objects | | narrow band | 8.7 | 26.5 | 13.6 |
| 71 | | Skyglow-B | Skyglow Broadband | enhances deep-sky observing in moderately light polluted skies, improves the view of nebulae, galaxies, & clusters | | extra wide band | 26.5 | 64.8 | 47.6 |
| 72 | | Skyglow-I | Skyglow Imaging | deep-sky CCD or DSLR imaging from light polluted skies, enhances all types of deep-sky objects (galaxies, nebulae, clusters), preserves neutral colour balance | | multi band | 61.9 | 68.5 | 60.2 |
| 73 | | Mars | Mars | high performance visual filter for improving views of Mars, improves view of polar ice caps, subtle mare shadings, cloud activity, dramatic improvements even in smaller telescopes | | special A | 29.2 | 52.6 | 53.7 |
| 74 | Orion | H-alpha | H-alpha Narrow Band | CCD imaging | www.telescope.com | H-alpha B | 0.5 | 0.0 | 1.8 |
| 75 | Sirius Optics | NEB1 | Nebula | - | | medium band | 17.7 | 41.1 | 23.3 |
| 76 | | CE1 | Contrast Enhance | - | | special A | 60.9 | 50.1 | 60.4 |
| 77 | | PC1 | Planetary Contrast | - | | special A | 40.7 | 35.9 | 36.5 |
| 78 | | NPC | Neodymium Planetary Contrast | - | | special B | 49.0 | 55.1 | 53.3 |
| 79 | TS Optics | O-III | O-III | narrowband nebular filter for observing and imaging | www.telescope-service.com | O-III B | 3.0 | 9.9 | 5.0 |
| 80 | | UHC | Ultra High Contrast | contrast boosting filter for deep sky observing | | narrow band | 5.9 | 22.2 | 11.2 |
| 81 | | O-III | Bandmate O-III | enhances planetary nebulae in larger scopes | | O-III A | 9.7 | 27.3 | 14.7 |
| 82 | Televue Optics | Nebustar | Bandmate Nebustar | high-performance dielectric squarewave "UHC" type general purpose narrow-band filter, great for all nebulae types and instrument sizes | www.televue.com | medium band | 14.9 | 42.5 | 22.1 |

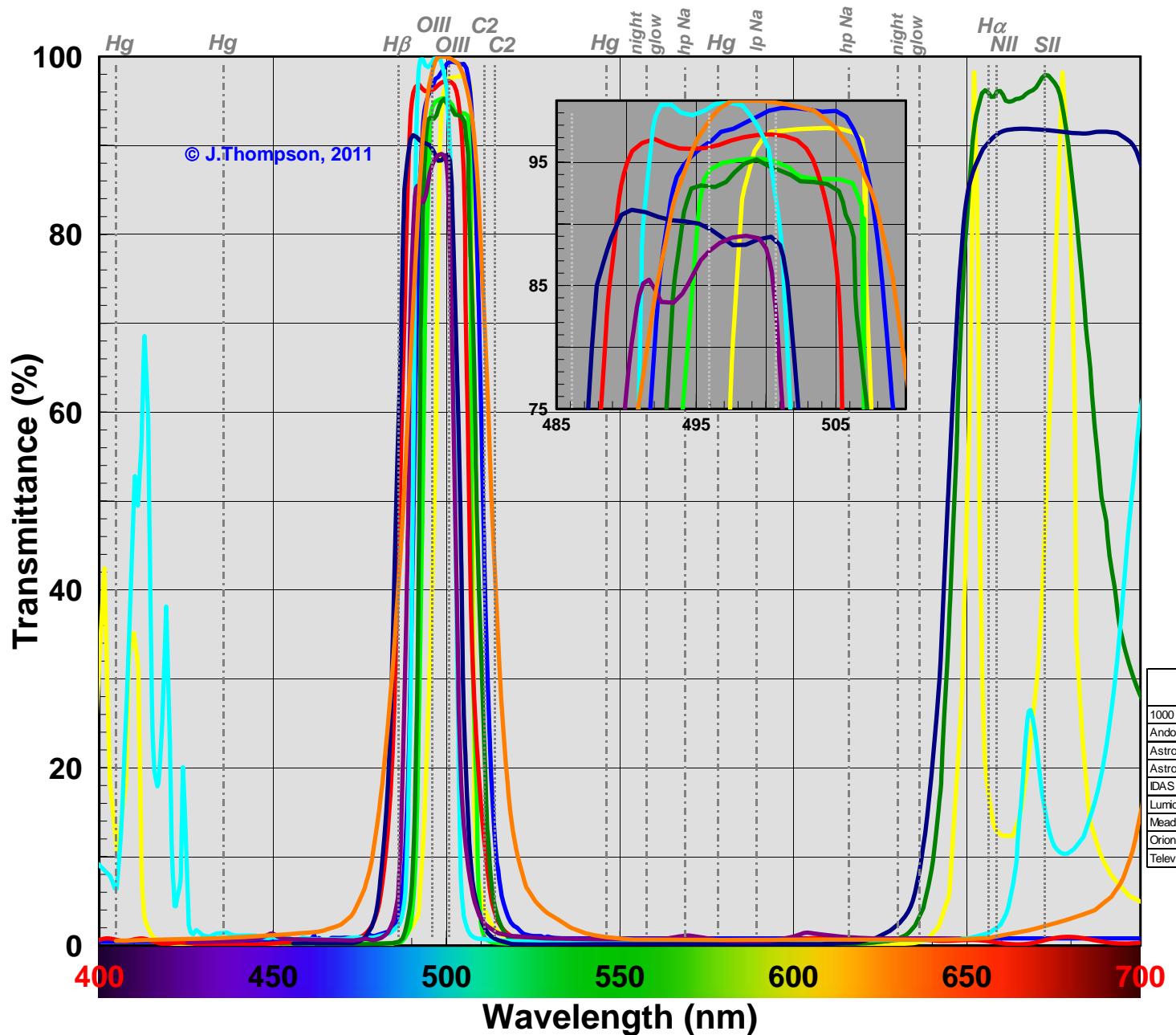
Appendix B Band-Pass Filter Spectral Transmissivities



H-beta Group A



**H-beta
Group B**

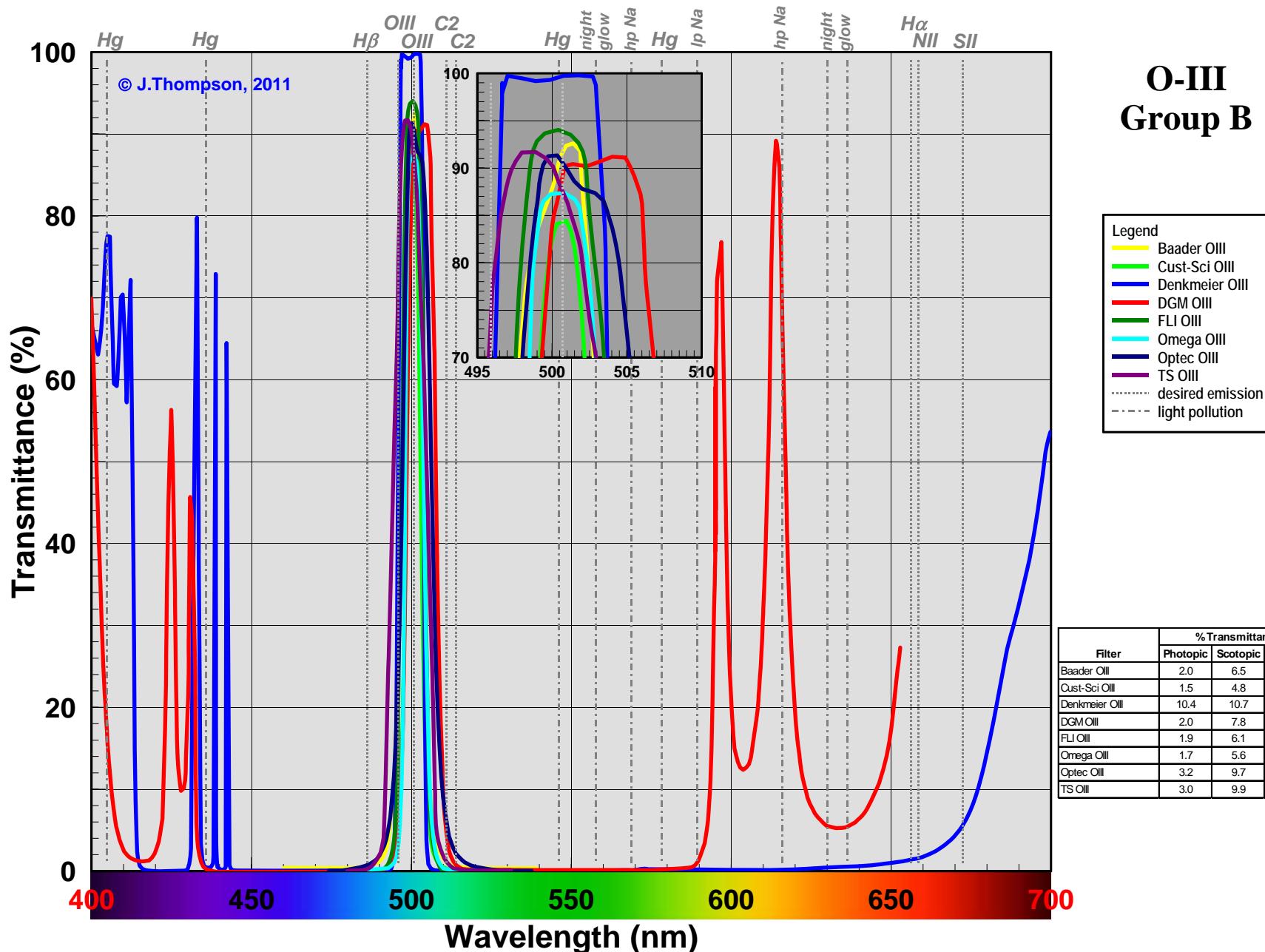


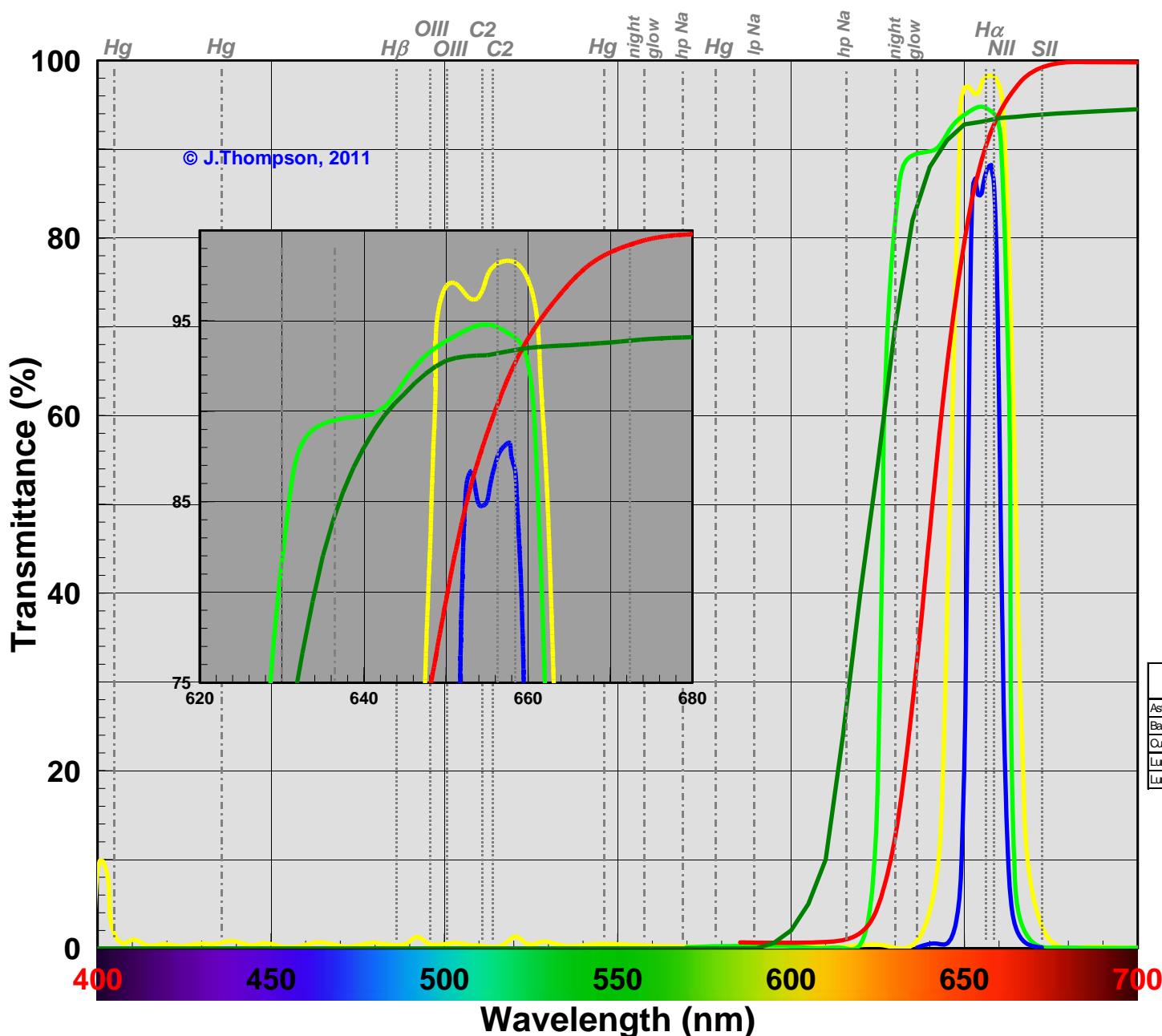
O-III Group A

| Legend | |
|---------------------|-------------|
| 1000 Oaks LP3 | Yellow |
| Andover OIII | Green |
| Astronomik OIII | Blue |
| Astronomik OIII CCD | Red |
| IDAS OIII | Dark Green |
| Lumicon OIII | Cyan |
| Meade OIII | Dark Blue |
| Orion OIII | Purple |
| Televue OIII | Orange |
| desired emission | Dotted line |
| light pollution | Dashed line |

| Filter | % Transmittance | | |
|---------------------|-----------------|----------|------|
| | Photopic | Scotopic | MC |
| 1000 Oaks LP3 | 4.5 | 11.8 | 10.9 |
| Andover OIII | 4.4 | 14.2 | 7.2 |
| Astronomik OIII | 6.6 | 20.5 | 10.4 |
| Astronomik OIII CCD | 6.2 | 20.9 | 10.8 |
| IDAS OIII | 7.3 | 16.2 | 20.7 |
| Lumicon OIII | 3.8 | 12.6 | 9.9 |
| Meade OIII | 6.7 | 16.7 | 23.7 |
| Orion OIII | 4.2 | 13.8 | 7.2 |
| Televue OIII | 9.7 | 27.3 | 14.7 |

O-III Group B



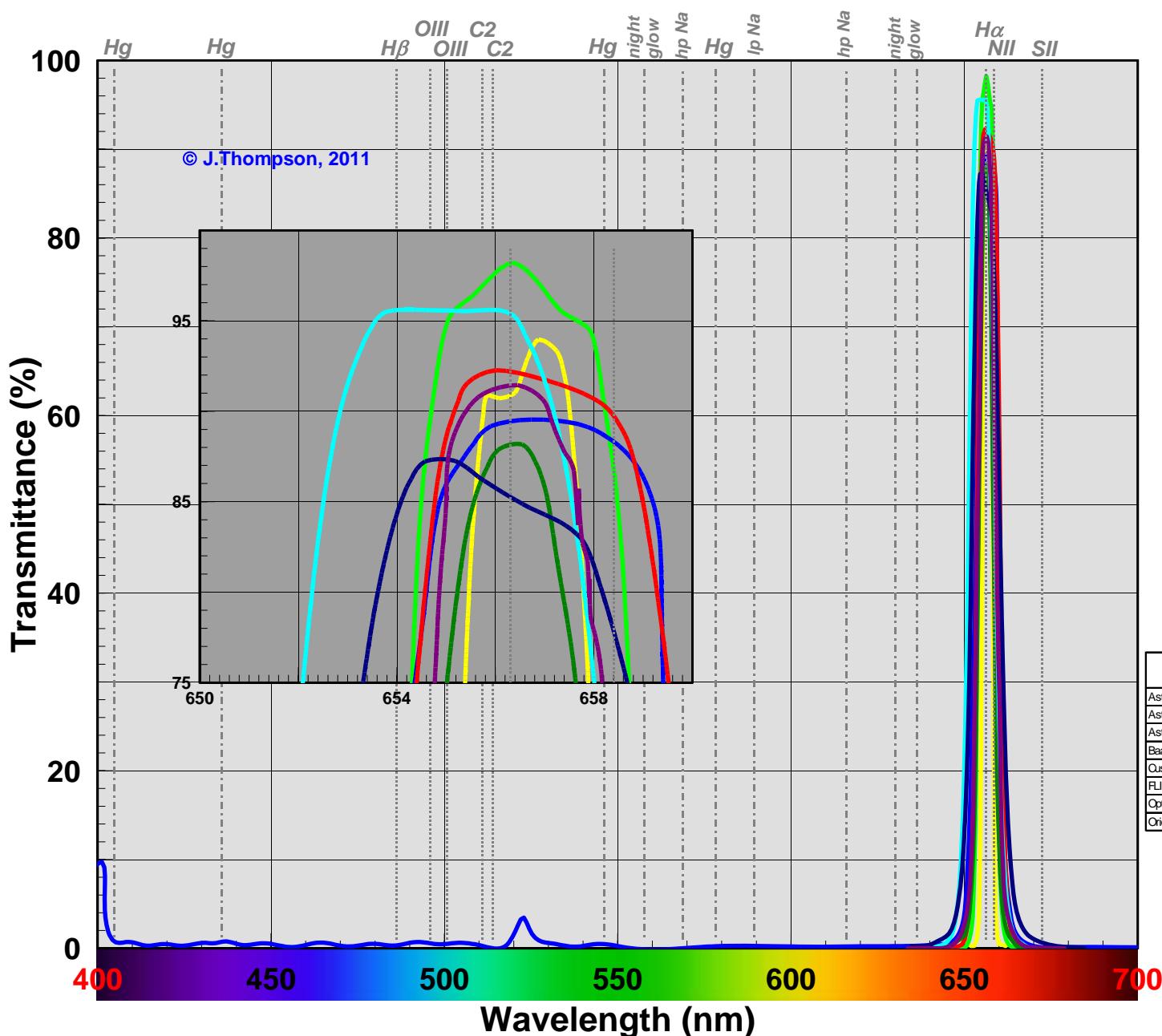


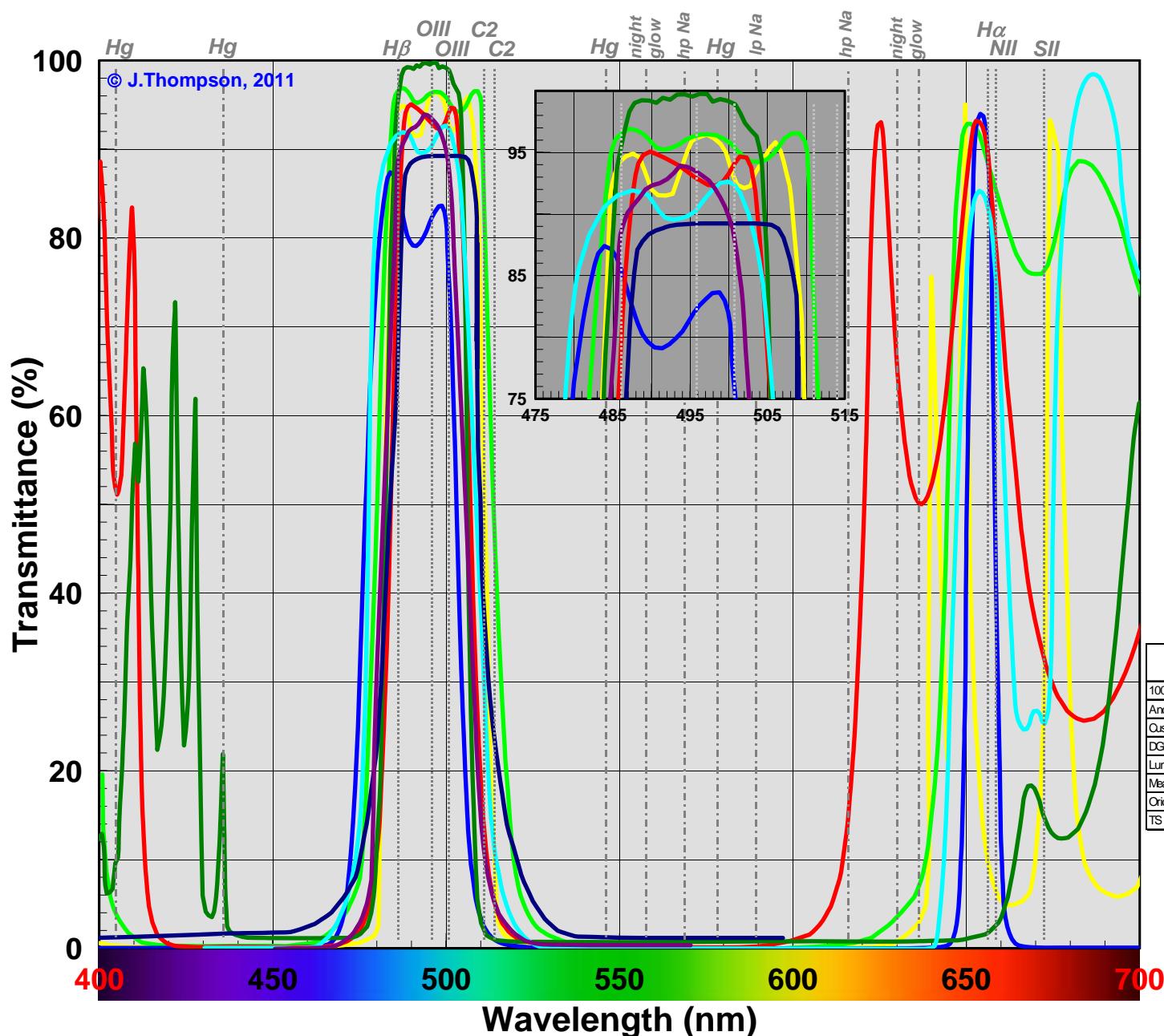
H-alpha Group A

Legend

- Astronomik Halphalpha2
- Baader Halphalpha2
- Cust-Sci Halphalpha2
- Lumicon H-alpha
- Lumicon #29
- desired emission
- - - light pollution

| Filter | % Transmittance | | |
|------------------------|-----------------|----------|------|
| | Photopic | Scotopic | MC |
| Astronomik Halphalpha2 | 1.9 | 0.5 | 5.9 |
| Baader Halphalpha2 | 4.7 | 0.0 | 97 |
| Cust-Sci Halphalpha2 | 0.6 | 0.0 | 2.5 |
| Lumicon H-alpha | 3.2 | 0.0 | 17.4 |
| Lumicon #29 | 7.3 | 0.1 | 21.6 |

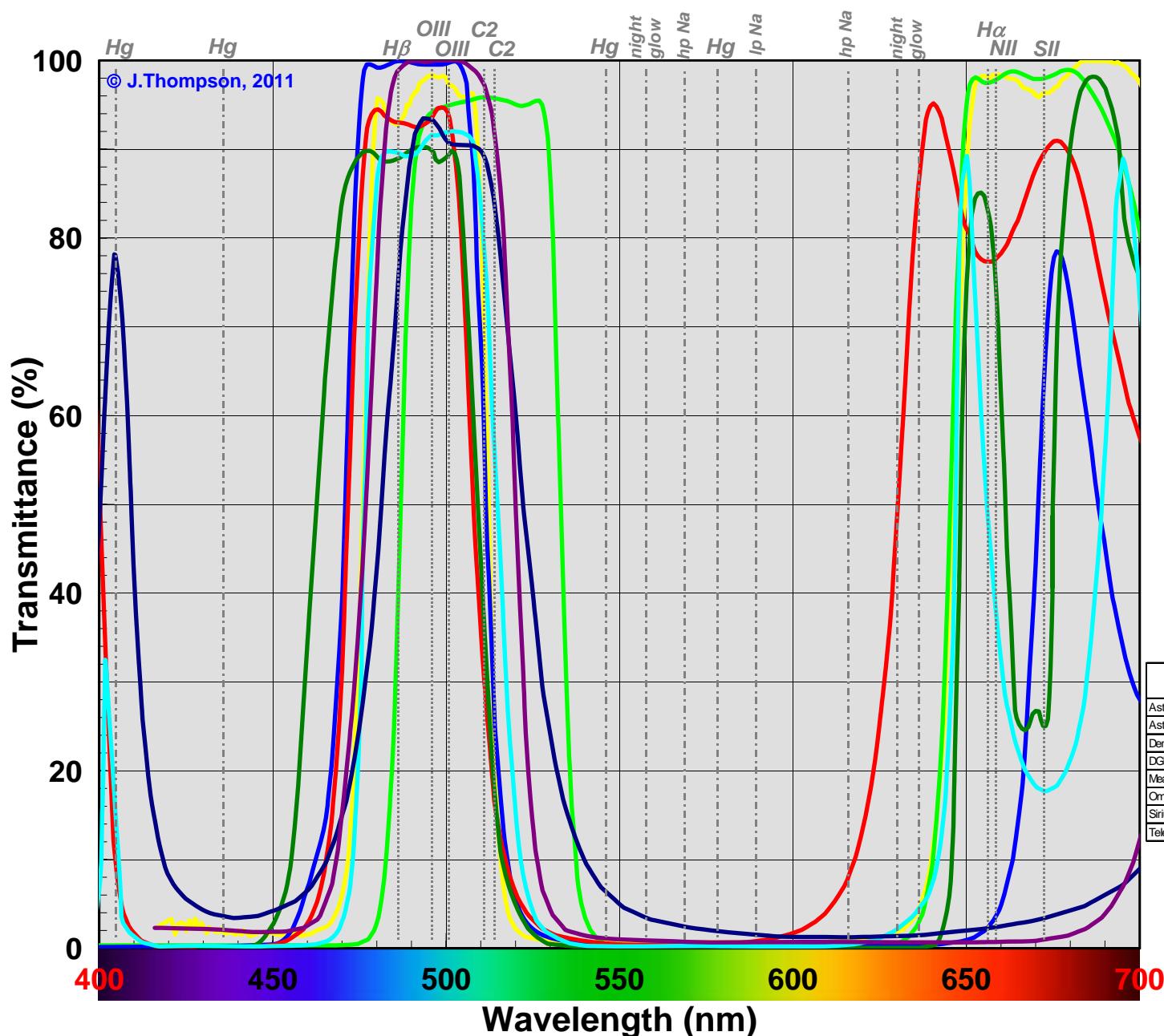


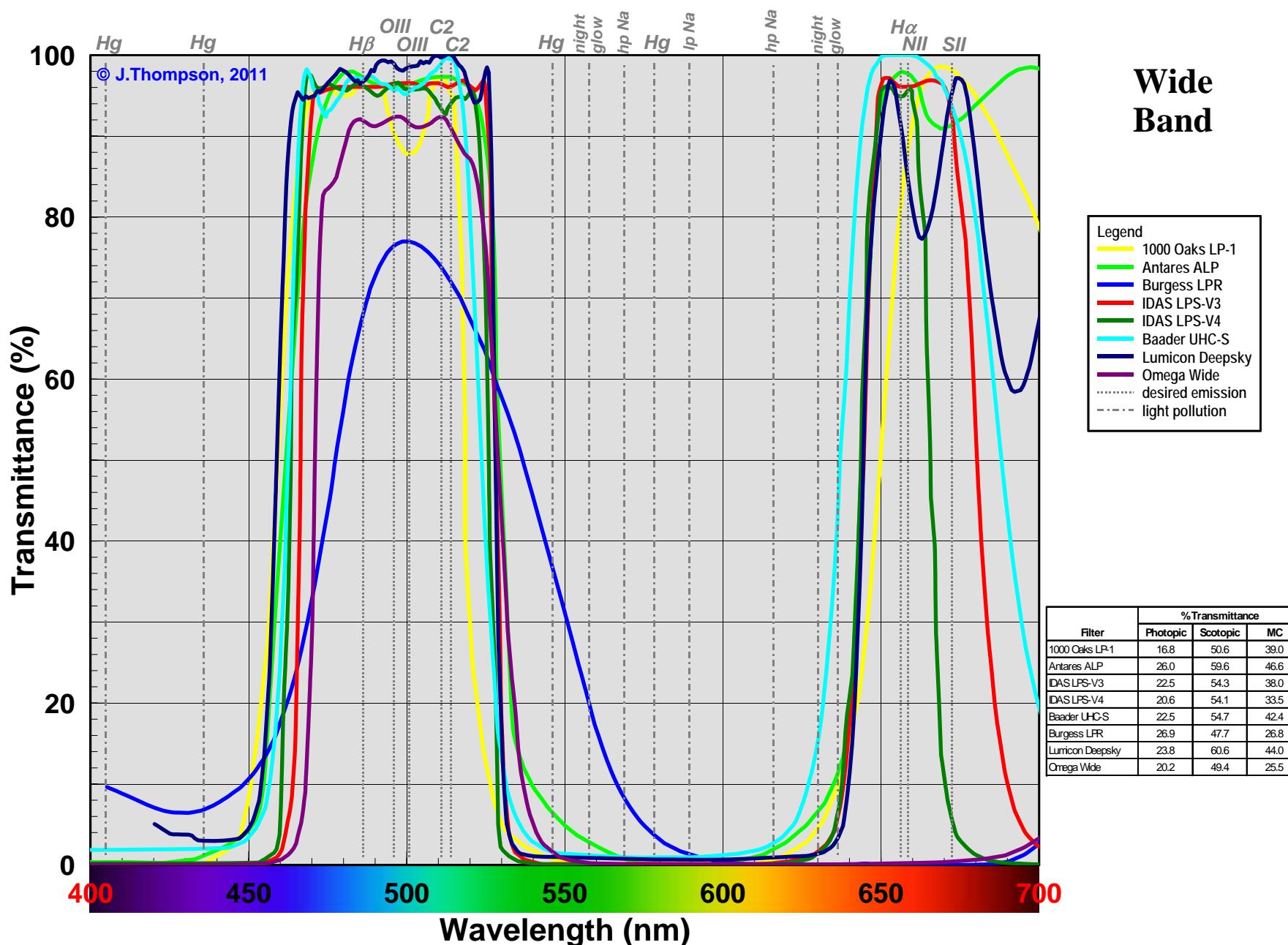


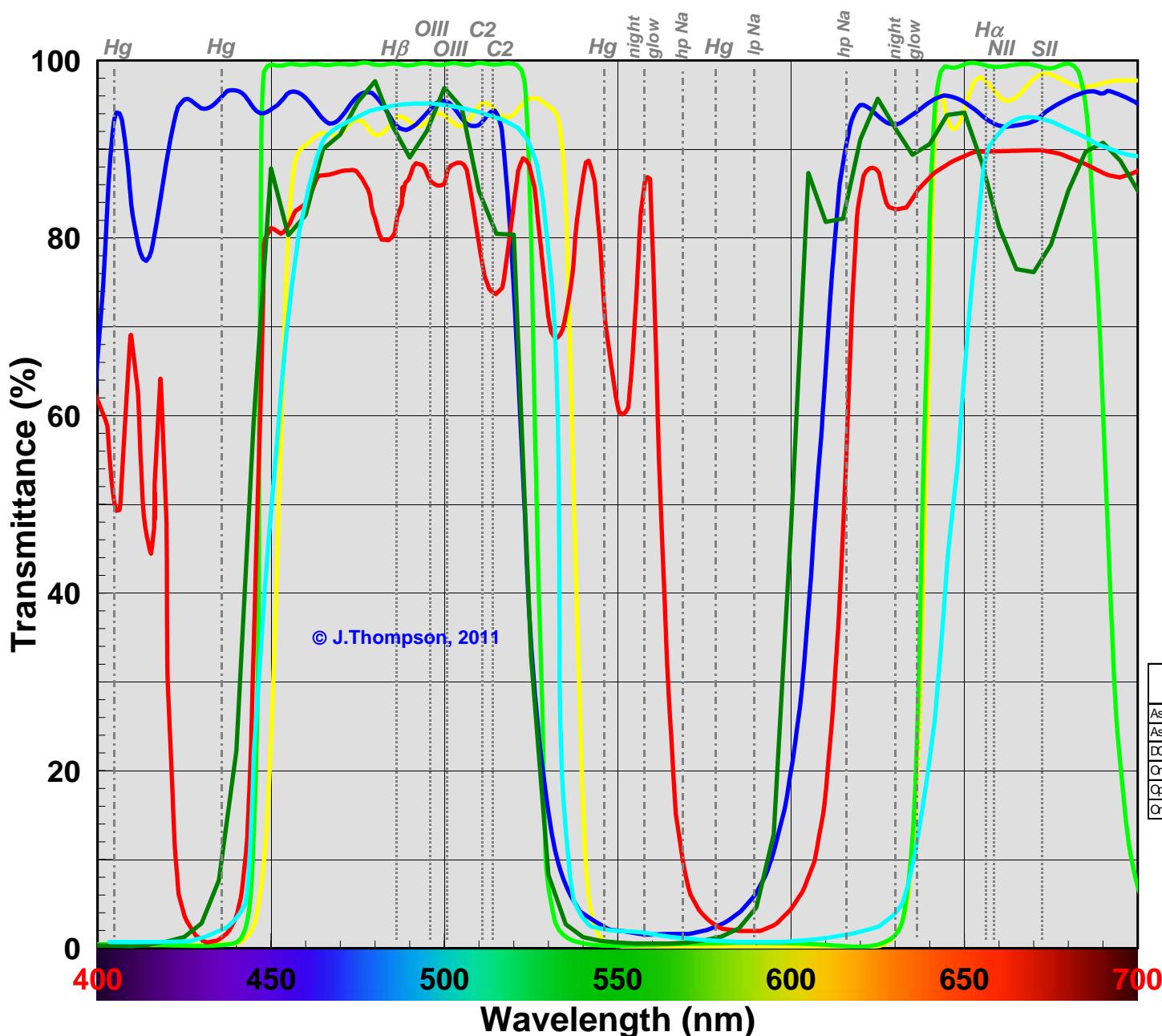
Narrow Band

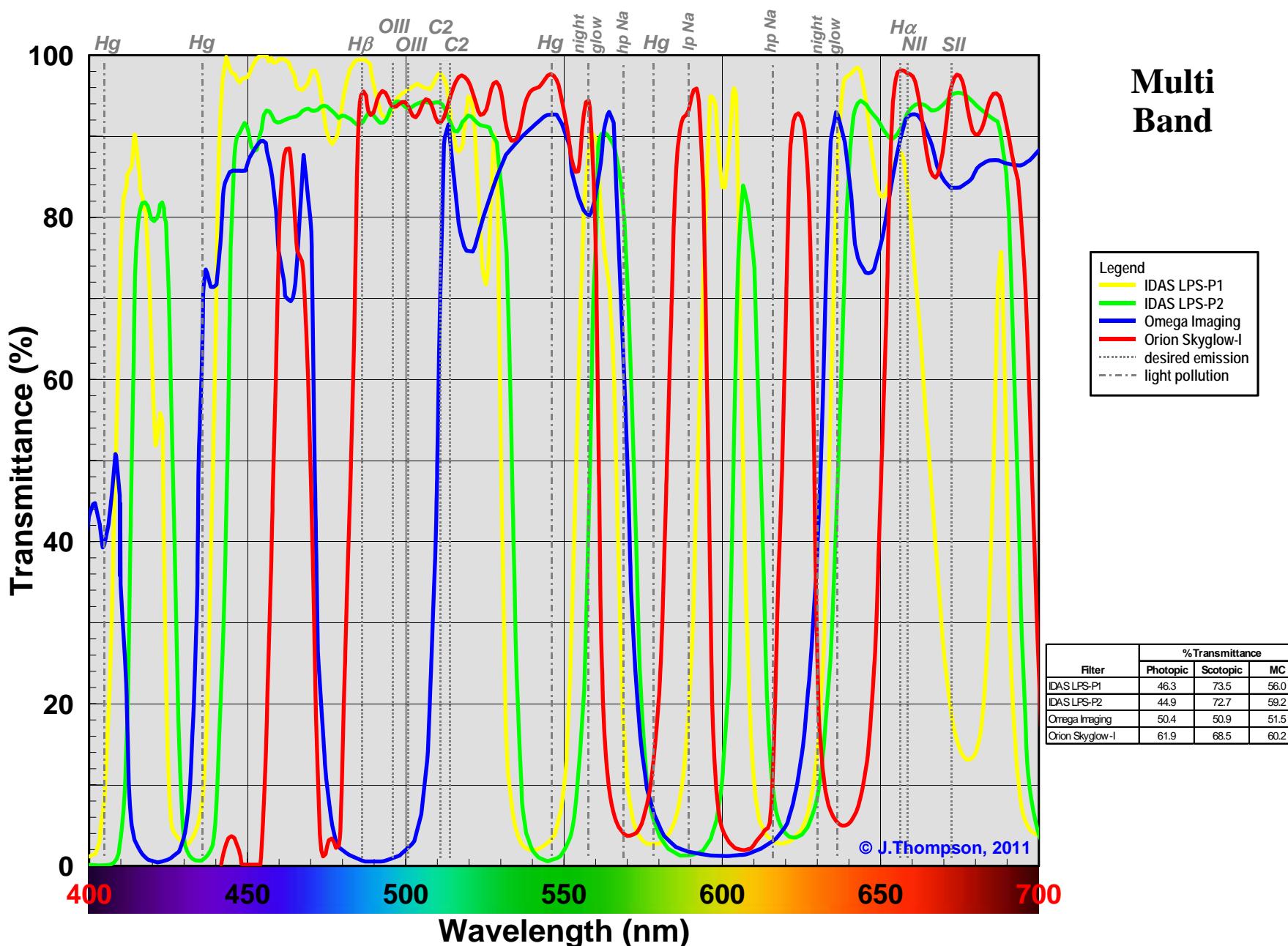
| Legend | 1000 Oaks LP-2 |
|--------------------|----------------|
| Andover 3ch Neb | 95 |
| Cust-Sci Multiband | 85 |
| DGM NPB | 80 |
| Lumicon UHC | 75 |
| Meade Narrow | 70 |
| Orion Ultrablock | 65 |
| TS UHC | 60 |
| desired emission | 55 |
| light pollution | 50 |

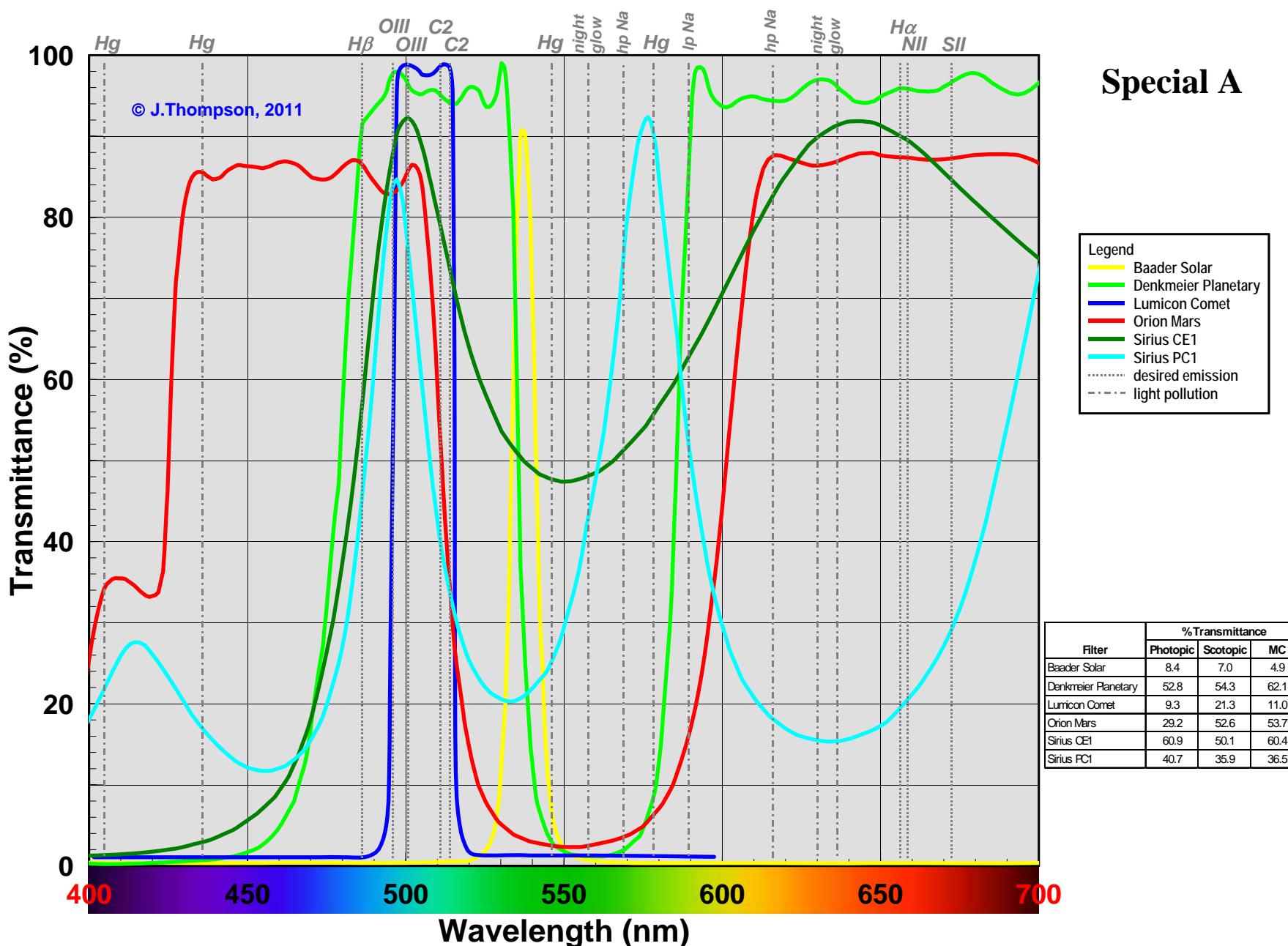
| Filter | %Transmittance | | |
|--------------------|----------------|----------|------|
| | Photopic | Scotopic | MC |
| 1000 Oaks LP-2 | 9.3 | 26.5 | 18.7 |
| Andover 3ch Neb | 12.5 | 32.9 | 30.9 |
| Cust-Sci Multiband | 5.5 | 21.5 | 12.9 |
| DGM/NPB | 12.3 | 22.6 | 25.7 |
| Lumicon UHC | 7.0 | 24.8 | 16.5 |
| Meade Narrow | 8.9 | 28.1 | 25.2 |
| Orion Ultrablock | 8.7 | 26.5 | 13.6 |
| TS UHC | 5.9 | 22.2 | 11.2 |

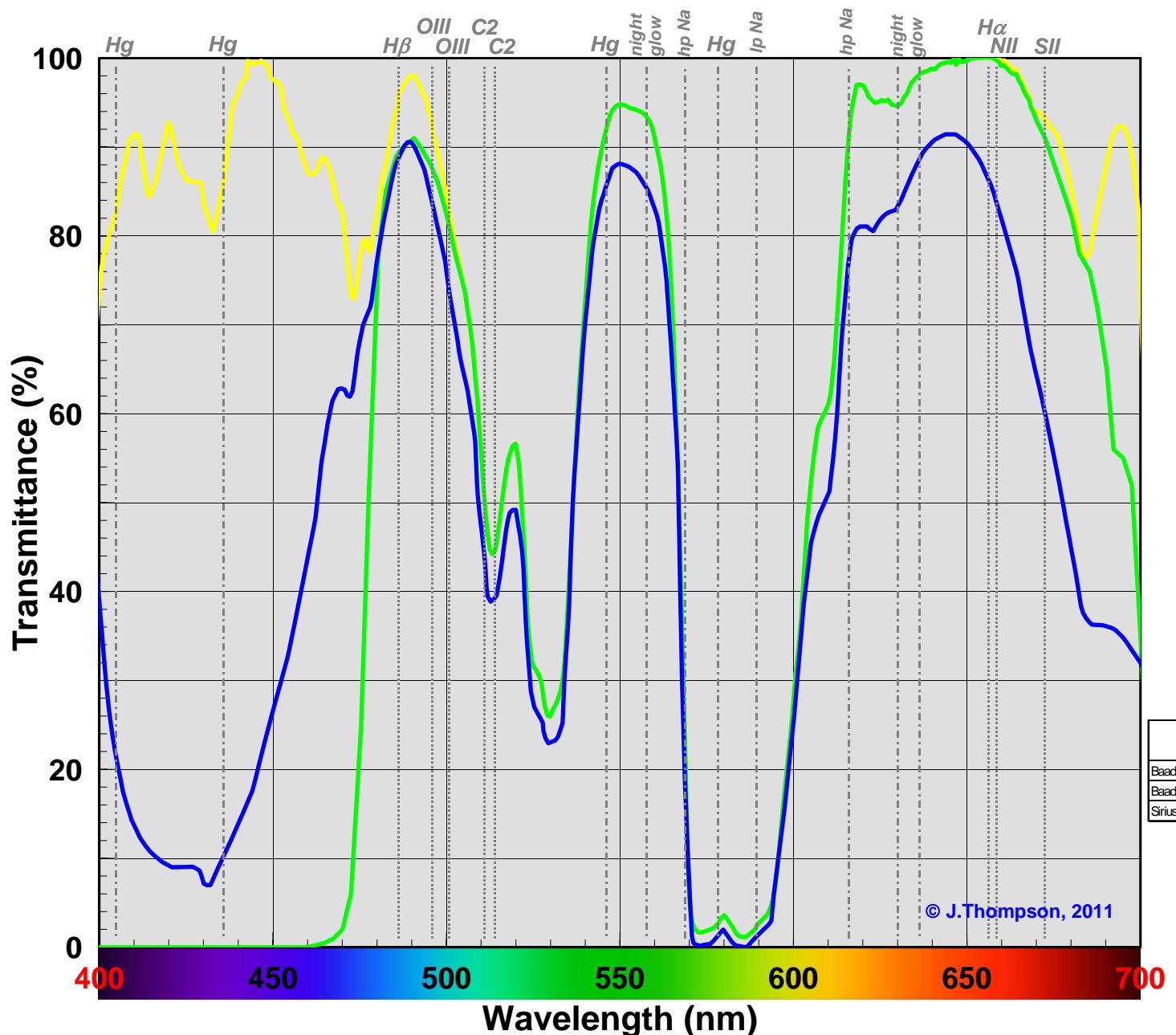












Special B

Appendix C Per-Integration Time Images

none-2sec-stack.tif



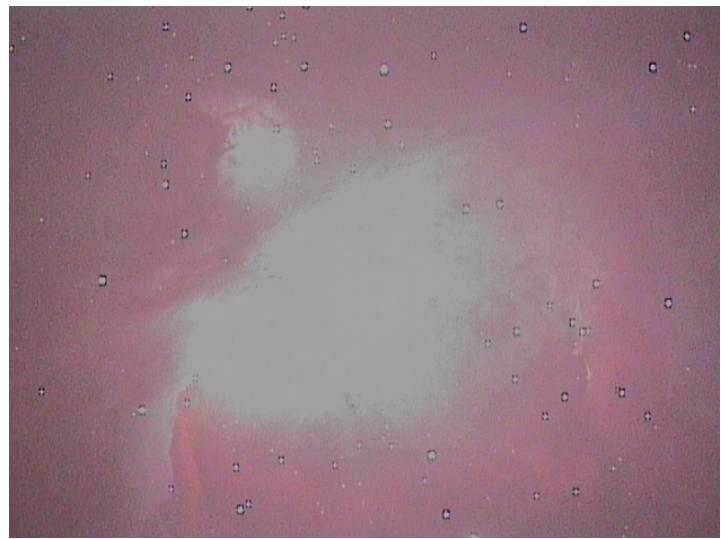
none-4sec-stack.tif



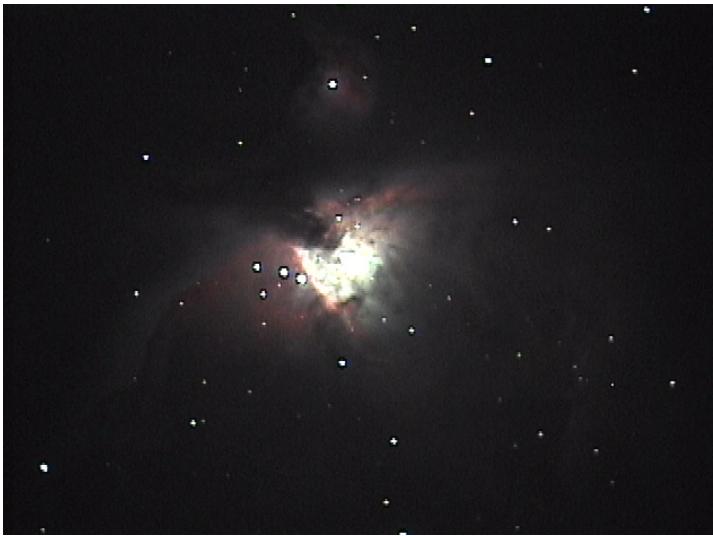
none-8sec-stack.tif



none-12sec-stack.tif



moonsky_2sec-stack.tif



moonsky_4sec-stacked.tif



moonsky_8sec-stacked.tif



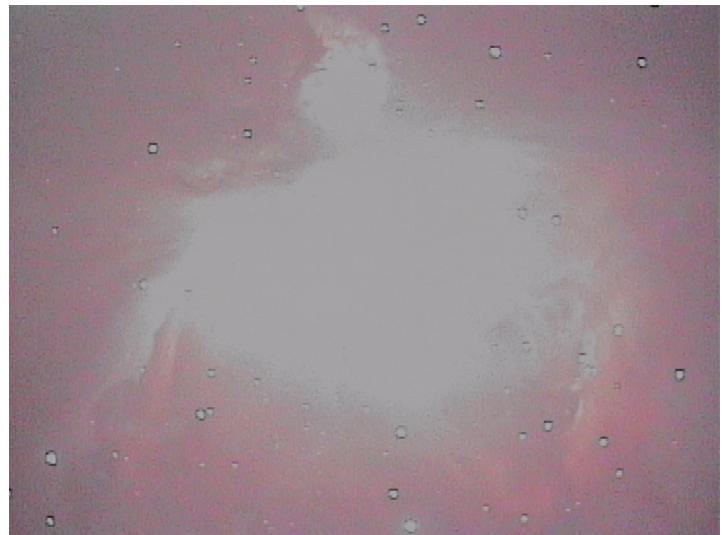
moonsky_12sec-stacked.tif



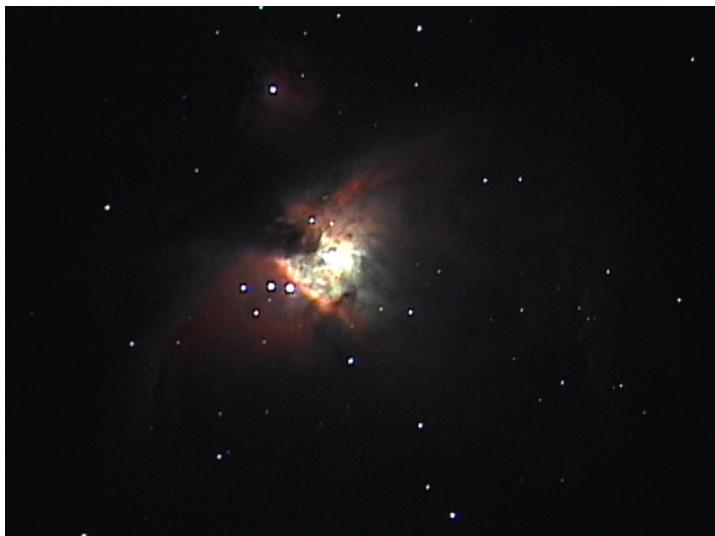
moonsky_16sec-stacked.tif



moonsky_24sec-stacked.tif



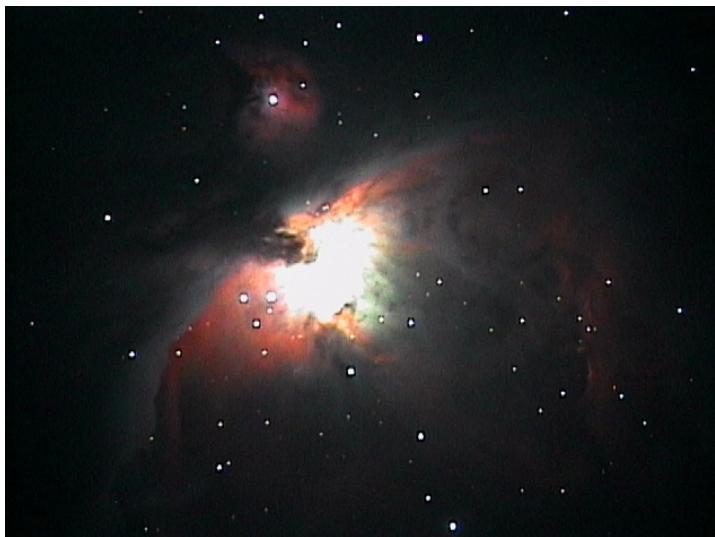
lps-2sec-stack.tif



lps-4sec-stack.tif



lps-8sec-stack.tif



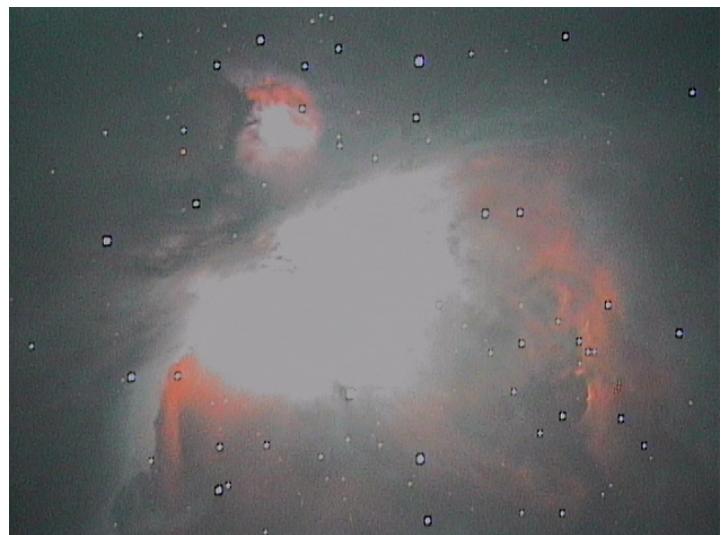
lps-12sec-stack.tif



lps-16sec-stack.tif



lps-24sec-stack.tif



lps-32sec-stack.tif



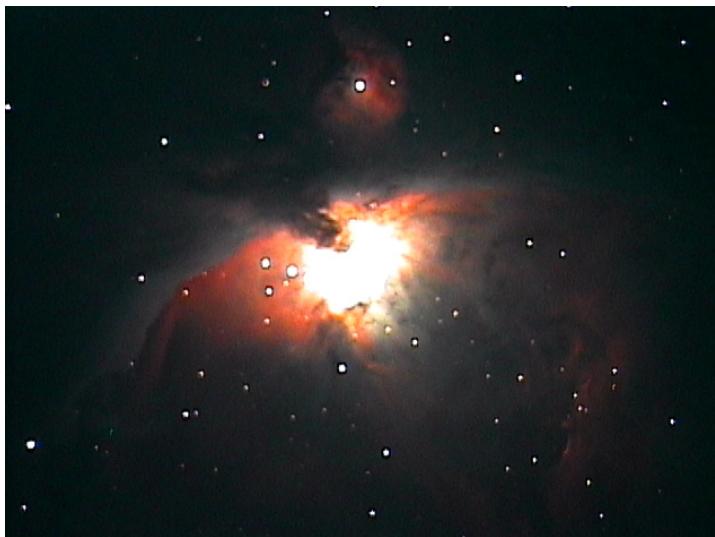
orion-2sec-stack.tif



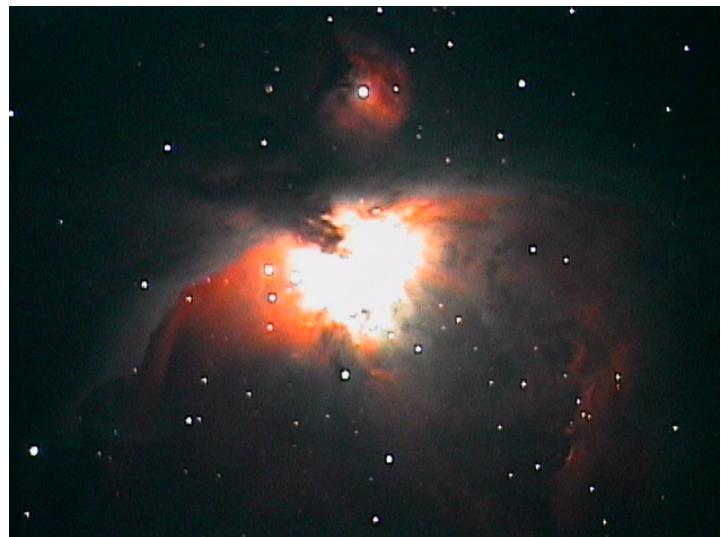
orion-4sec-stack.tif



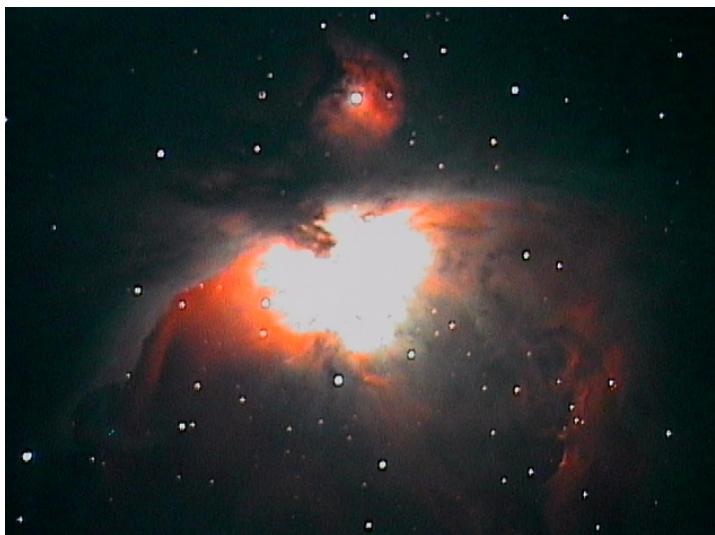
orion-8sec-stack.tif



orion-12sec-stack.tif



orion-16sec-stack.tif



orion-24sec-stack.tif



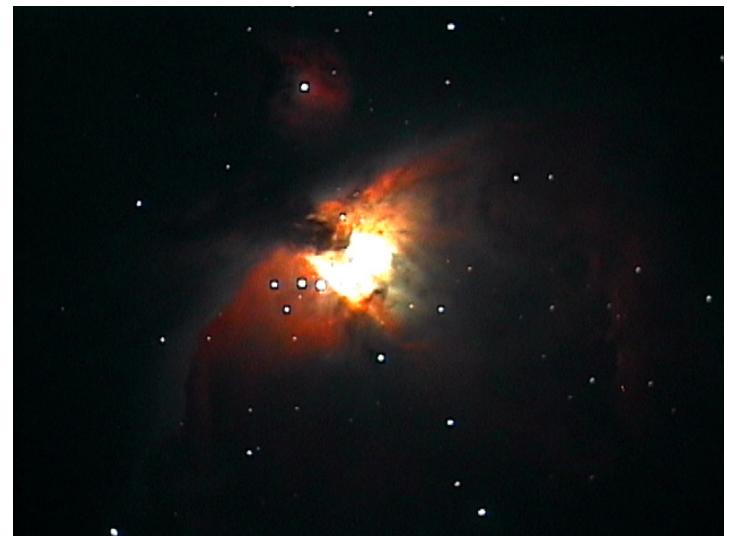
orion-32sec-stack.tif



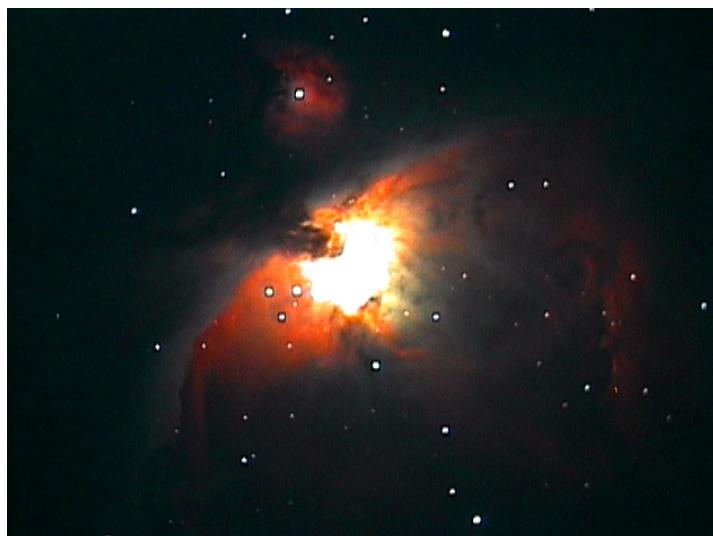
uhcs-2sec-stack.tif



uhcs-4sec-stack.tif



uhcs-8sec-stack.tif



uhcs-12sec-stack.tif



uhcs-16sec-stack.tif



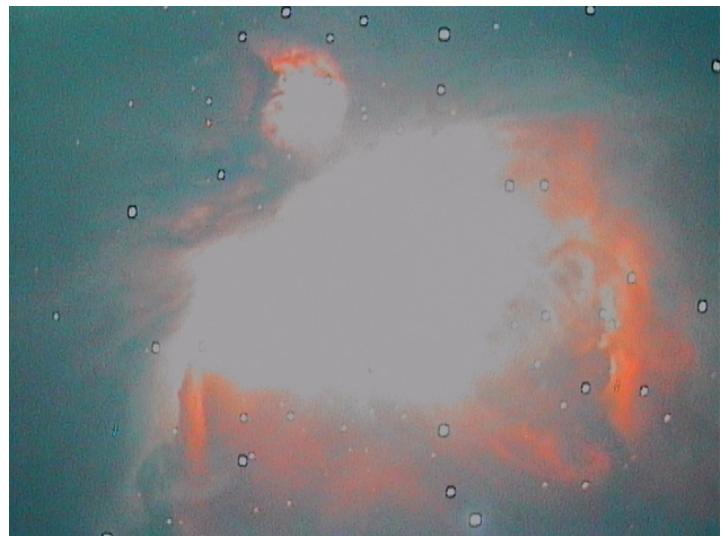
uhcs-24sec-stack.tif



uhcs-32sec-stack.tif



uhcs-48sec-stack.tif



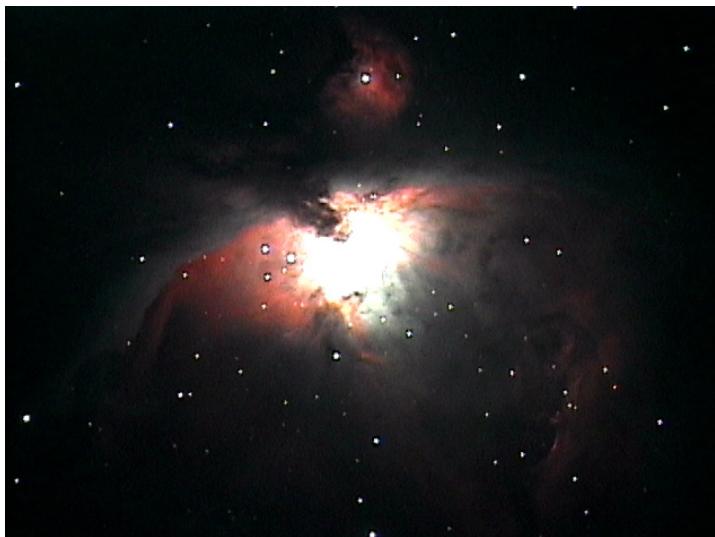
auhc_2sec-stacked.tif



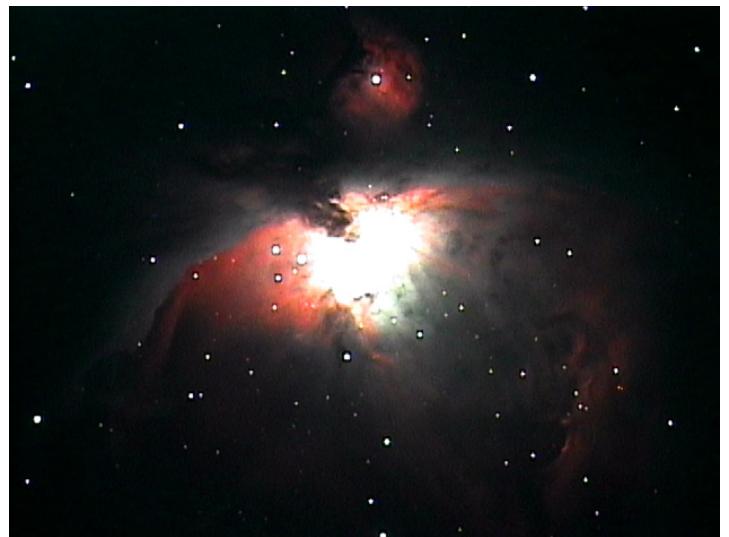
auhc_4sec-stacked.tif



auhc_8sec-stacked.tif



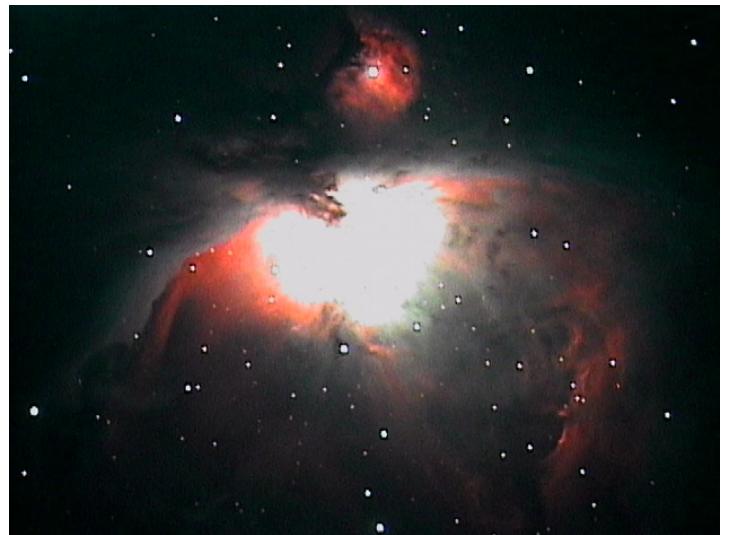
auhc_12sec-stacked.tif



auhc_16sec-stacked.tif



auhc_24sec-stacked.tif



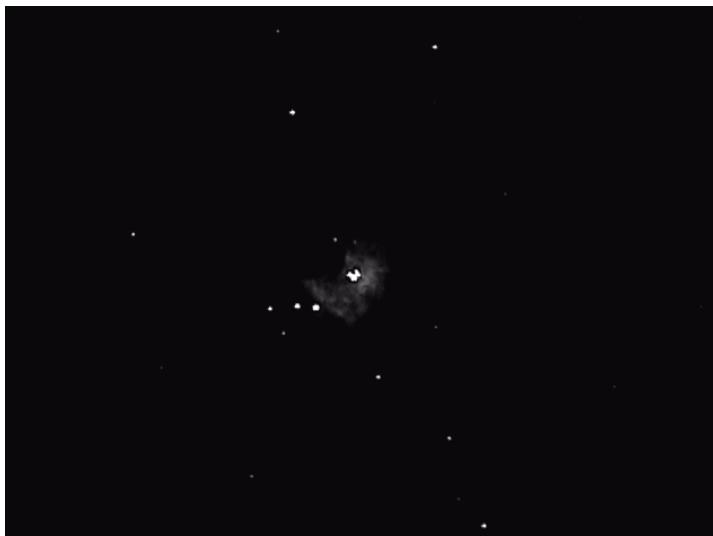
auhc_32sec-stacked.tif



auhc_48sec-stacked-badcoll.tif



abeta_2sec-stacked.tif



abeta_4sec3-stacked.tif



abeta_8sec-stacked.tif



abeta_12sec-stacked.tif



abeta_16sec-stacked.tif



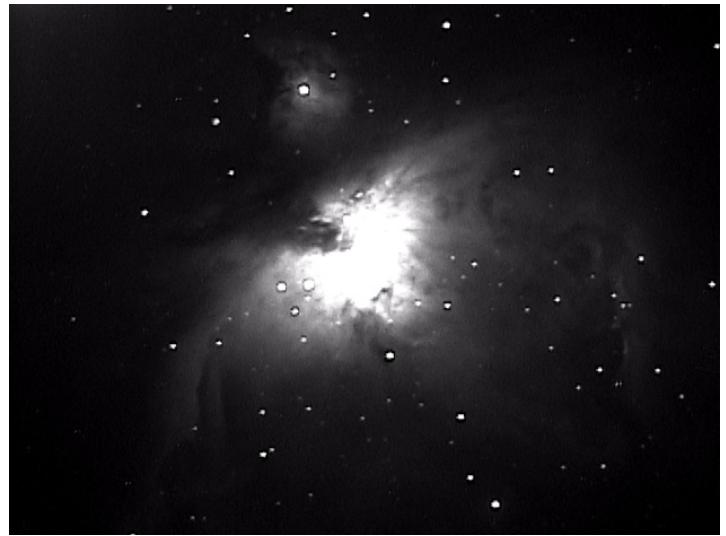
abeta_24sec-stacked.tif



abeta_32sec-stacked.tif



abeta_48sec-stacked.tif



abeta_60sec-stacked.tif



abeta_75sec-stacked.tif



abeta_90sec-stacked.tif



o3-2sec-stack.tif



o3-4sec-stack.tif



o3-8sec-stack.tif



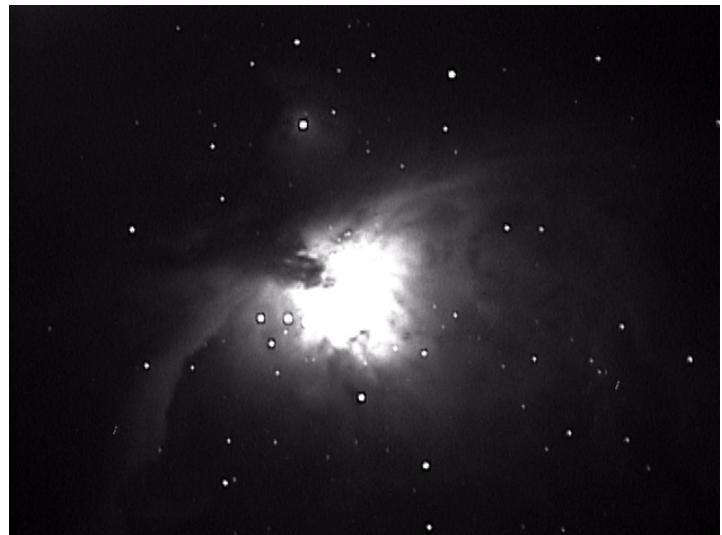
o3-12sec-stack.tif



o3-16sec-stack.tif



o3-24sec-stack.tif



o3-32sec-stack.tif



o3-48sec-stack.tif



o3-60sec-stack.tif



red_2sec-stacked.tif



red_4sec-stacked.tif



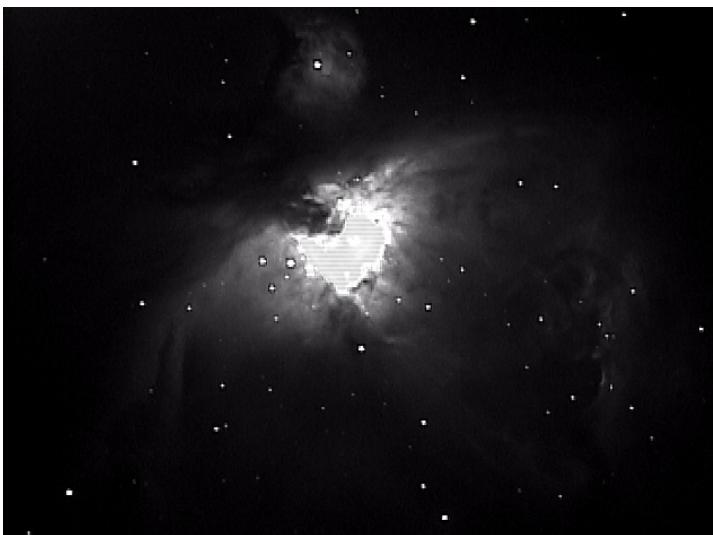
red_8sec-stacked.tif



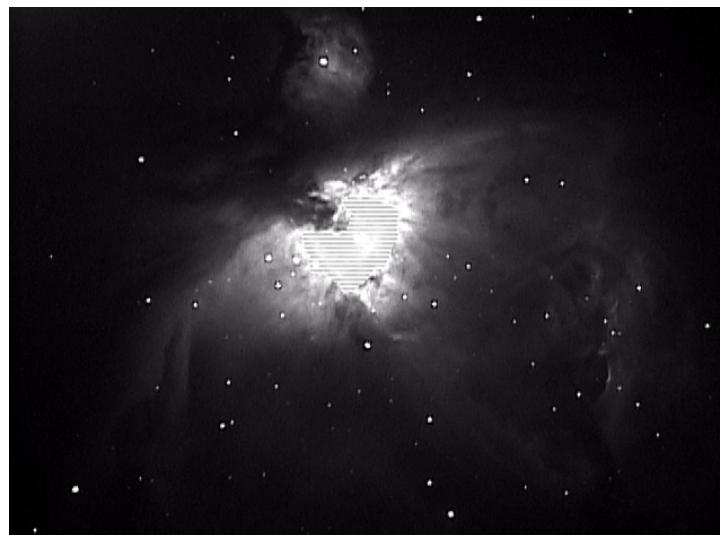
red_12sec-stacked.tif



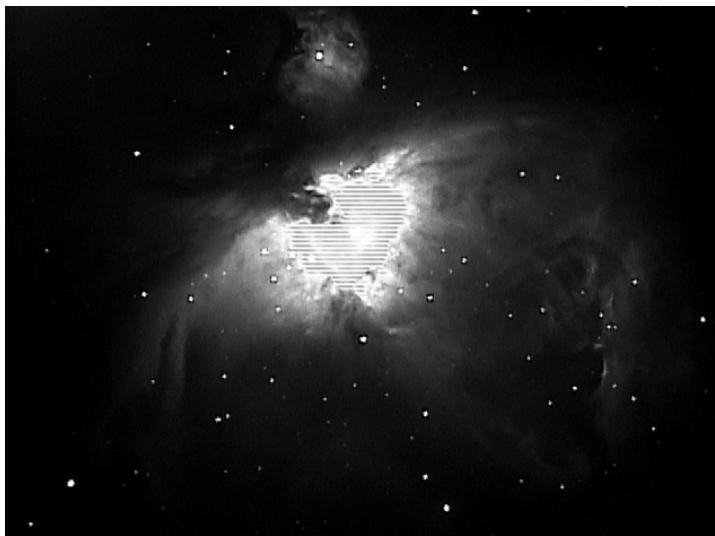
red_16sec-stacked.tif



red_24sec-stacked.tif



red_32sec-stacked.tif



red_48sec-stacked.tif



red_60sec-stacked.tif



red_75sec-stacked.tif



red_90sec-stacked.tif

