## Molecular Dynamics Simulation

Prabal K Maiti
Department of Physics

http://www.physics.iisc.ernet.in/~maiti

☐ Introduction (Basic facts and some history) ■ Molecular dynamics, Various schemes for integration, □ Inter and Intra molecular forces, Various ensemble (NVE, NVT, NPT, NPH), atomic charge derivation scheme ☐ How to make the simulation efficient (Cell List, Neighbor List), Periodic Boundary condition, computing long range interactions ☐ Free energy calculations, Umbrella sampling, Thermodynamic Integration, Entropy calculation ☐ Introduction to various solvation methods Application: Nanotube, Liquid Crystal, DNA simulations

### Ref:

Computer simulation of Liquids: M. P. Allen and D. J. Tildesley, Oxford (1987)

Understanding Molecular simulation: Daan Frenkel and B Smit (2<sup>nd</sup> ed)

Molecular Modelling Principles And Applications: Andrew Leach, Prentice Hall (2001)

The art of Molecular dynamics: D. C.Rappaport

Molecular Modeling and Simulation: Tamar Schlick

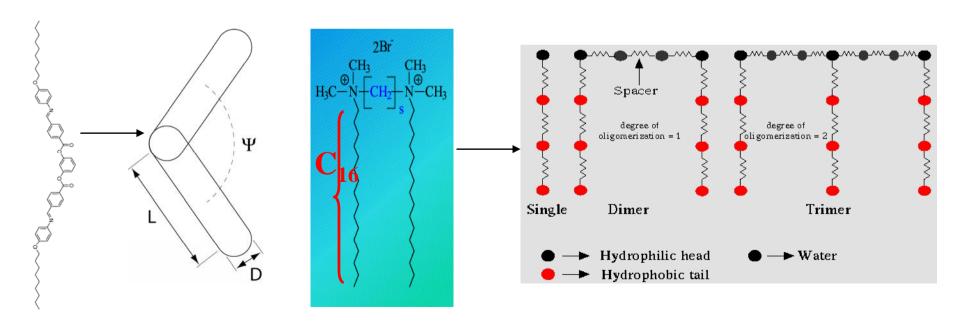
Web: CCP5 library http://www.ccp5.ac.uk

Acknowledgement: <a href="http://www.cheme.buffalo.edu/courses/ce530">http://www.cheme.buffalo.edu/courses/ce530</a>

### What is molecular simulation?

Molecular modeling is the science and art of studying molecular structure and function through model building and computation.

Model building could be as simple as representing molecule by hard/soft sphere (beads), rigid rods, or other geometrical shape, sphere/beds connected through springs or molecule with full chemical details.



### What is molecular simulation?

Computation can be carried out using following methods:

Molecular mechanics, molecular dynamics, Monte Carlo, Free energy and solvation methods, structure/activity relationship (SAR) and many other established procedure.

### Need for Molecular simulation

- There are no general method for a solving complex many-body problems. Hamiltonian is unknown, until we solve the quantum many-body problem! In fact in most cases not possible and requires lots of approximations.
- Molecular simulations are the only possible solution for such complex many body systems.
- In many cases experiments are limited and expensive. Simulations can complement the experiment.
- Simulation can give molecular level understanding even at a single molecule level

## Simulation Methodology

## Semi-empirical

Give us the phenomena and invent a model to mimic the problem.

### ab initio methods

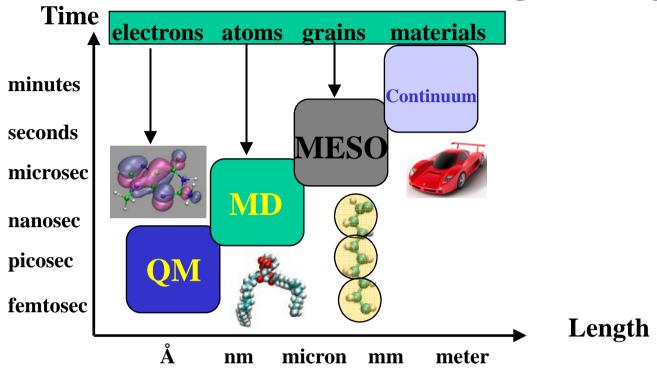
Maxwell, Boltzmann and Schrödinger gave us the model. All we must do is numerically solve the mathematical problem and determine the properties.

These two approaches can be combined to make what is termed as Multi-scale modeling strategies

"The general theory of quantum mechanics is now almost complete. The underlying physical laws necessary for the mathematical theory of a large part of physics and the whole of chemistry are thus completely known, and the difficulty is only that the exact application of these laws leads to equations much too complicated to be soluble."

Dirac, 1929

## **Multi-scale Modeling strategy**



### High quality multi-scale simulations

- Quantum Mechanical calculations
- First Principles force fields
- ➤ Large scale Molecular Dyanmics (MD) simulations
- Mesoscopic modeling (Coarse-grained MD, DPD, BD)
- Macroscopic modeling (finite elements, continuum simulations, Lattice Boltzmann)

## What size is too big? What times are too long?

• QM (ab initio molecular dynamics)

electrons / N basis sets, speed  $\sim N^3$  or  $N^4$  (may be linear with very high prefactor)

Time steps  $\sim 10^{-2}$  fs.

Example of big and long: 64-256 water molecules during 100 ps.

Classical atomistic MD

Atoms/ N atoms, speed ~N<sup>2</sup> (may be reduced with efficient algorithms, periodic coulomb is most expensive)

Time steps  $\sim 0.5$ -2 fs.

Example of big and long (1 processor): 50,000 atoms for 1 ns.

(parallel simulations, can improve this much. Parallel codes available free or almost free: LAMMPS, AMBER, GROMACS, NAMD)

### Short history of Molecular Simulations

- Metropolis, Rosenbluth, Teller (1953) Monte Carlo Simulation of hard disks.
- Fermi, Pasta Ulam (1954) experiment on ergodicity
- Alder & Wainwright (1958) liquid-solid transition in hard spheres. "long time tails" (1970)
- Vineyard (1960) Radiation damage using MD
- Rahman (1964) liquid argon, water(1971)
- Verlet (1967) Correlation functions, ...
- Andersen, Rahman, Parrinello (1980) constant pressure
   MD
- Nose, Hoover, (1983) constant temperature thermostats.
- Car, Parrinello (1985) ab initio MD.

The examples for each period are representative. The first five systems are modeled in vacuum and the others in solvent.

The 38  $\mu s$   $\beta$ -hairpin simulation in 2001 represents an ensemble (or aggregate dynamics) simulation, as accumulated over several short runs, rather than a long simulation.

The table is taken from the book by Tamar Schlick

Period	System and Size <sup>a</sup>	Trajectory Length <sup>b</sup> [ns]	CPU Time/Computer <sup>a</sup>
1973	Dinucleoside (GpC) in vacuum	_	_
	(8 flexible dihedral angles)	0.01	
1977	BPTI, vacuum	0.01	
	(58 residues, 885 atoms)	0.00	
1983	DNA, vacuum, 12 & 24 bp	0.09	several weeks each
1984	(754/1530 atoms) GnRH, vacuum	0.15	Vax 780
1984	(decapeptide, 161 atoms)	0.15	
1985	Myoglobin, vacuum	0.30	50 days
1905	(1423 atoms)	0.30	VAX 11/780
1985	DNA, 5 bp	0.50	20 hrs
1505	(2800 atoms)	0.50	Cray X-MP
1989	Phospholipid Micelle	0.10	City It III
1505	(≈ 7,000 atoms)	0.10	
1992	HIV protease	0.10	100 hrs.
	(25,000 atoms)		Cray Y-MP
1997	Estrogen/DNA	0.10	22 days
	(36,000 atoms, multipoles)		HP-735 (8)
1998	DNA, 24 bp	0.50	1 year, SGI
	(21,000 atoms, PME)		Challenge
1998	$\beta$ -heptapeptide in methanol	200	8 months, SGI-
	$(\approx 5000/9000 \text{ atoms})$		Challenge (3)
1998	Villin headpiece (36 residues,	1000	4 months, 256-proc.
	12,000 atoms, cutoffs)		Cray T3D/E
1999	$bc_1$ complex in phospholipid	1	75 days, 64 450-MHz-
	bilayer (91,061 atoms, cutoffs)		proc. Cray T3E
2001	C-terminal β-hairpin of protein-	38000	~ 8 days, 5000 proc.
****	G (177 atoms, implicit solvent)		Folding@home megacluster
2002	channel protein in lipid mem-	5	30 hrs, 500 proc. LeMieux terascale
	brane (106,189 atoms, PME)		system; 50 days, 32 proc. Linux (Athlon)

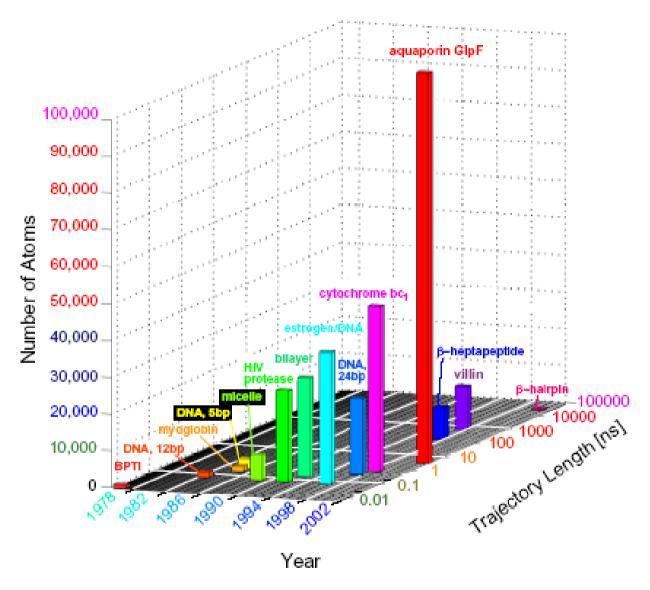


Figure 1.1. The evolution of molecular dynamics simulations with respect to system sizes and simulation lengths (see also Table 1.2).

### NSF Peta scale machine

A molecular dynamics (MD) simulation of curvature-inducing protein BAR domains binding to a charged phospholipid vesicle over 10 ns simulation time under periodic boundary conditions. The vesicle, 100 nm in diameter, should consist of a mixture of dioleoylphosphatidylcholine (DOPC) and dioleoylphosphatidylserine (DOPS) at a ratio of 2:1. The entire system should consist of 100,000 lipids and 1000 BAR domains solvated in 30 million water molecules, with NaCl also included at a concentration of 0.15 M, for a total system size of 100 million atoms. All system components should be modeled using the CHARMM27 all-atom empirical force field. The target wall-clock time for completion of the model problem using the NAMD MD package with the velocity Verlet time-stepping algorithm, Langevin dynamics temperature coupling, Nose-Hoover Langevin piston pressure control, the Particle Mesh Ewald algorithm with a tolerance of 1.0e-6 for calculation of electrostatics, a short-range (van der Waals) cut-off of 12 Angstroms, and a time step of 0.002 ps, with 64-bit floating point (or similar) arithmetic, is 25 hours. The positions, velocities, and forces of all the atoms should be saved to disk every 500 timesteps.

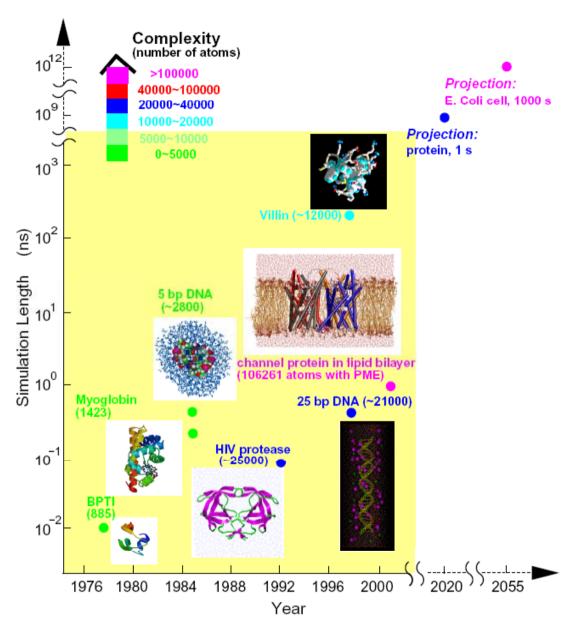


Figure 1.2. The evolution of molecular dynamics simulations with respect to simulation lengths (see also Table 1.2 and Figure 1.1). The data points for 2020 and 2055 represent extrapolations from the 1977 BPTI [136] and 1998 villin [57, 55] simulations, assuming a computational power increase by a factor of 10 every 3–4 years, as reported in [56].

# Molecular Dynamics (MD)

- Pick particles, masses and potential.
- Initialize positions and momentum. (boundary conditions in space and time)
- Solve  $\mathbf{F} = \mathbf{m} \mathbf{a}$  to determine  $\mathbf{r}(t)$ ,  $\mathbf{v}(t)$ .

## Newton (1667-87)

- Compute properties along the trajectory
- Estimate errors.
- Try to use the simulation to answer physical questions.

## What are the forces?

$$\frac{d^2\vec{r}}{dt^2} = -\nabla V(\vec{r})$$

- Crucial since V(r) determines the quality of result.
- Semi-empirical potentials: potential is constructed on theoretical grounds but using some experimental data.
- Common examples are Lennard-Jones, Coulomb, embedded atom potentials. They are only good for simple materials. The *ab initio* philosophy is that potentials are to be determined directly from quantum mechanics as needed.
- But computer power is not yet adequate in general.
- A powerful approach is to use simulations at one level to determine parameters at the next level.

## Statistical Ensembles

- Classical phase space is 6N variables ( $\mathbf{p}_i$ ,  $\mathbf{q}_i$ ) and a Hamiltonian function H( $\mathbf{q}$ , $\mathbf{p}$ ,t).
- We may know a few constants of motion such as energy, number of particles, volume...
- Ergodic hypothesis: each state consistent with our knowledge is equally "likely"; the *microcanonical* ensemble.
- Implies the average value does not depend on initial conditions.
- A system in contact with a heat bath at temperature T will be distributed according to the canonical ensemble:

$$\exp(-H(q,p)/k_BT)/Z$$

- The momentum integrals can be performed.
- Are systems in nature really ergodic? Not always!

# Criteria for an Integrator

- ➤ Newton's equation of motion are time reversible and so should be our algorithm.
- ➤ Hamiltonian dynamics preserve the magnitude of volume element in phase space and so our algorithm should have this area preserving property
- > simplicity (How long does it take to write and debug?)
- > efficiency (How fast to advance a given system?)
- > stability (what is the long-term energy conservation?)
- reliability (Can it handle a variety of temperatures, densities, potentials?)

The nearly universal choice for an integrator is the Verlet (leapfrog) algorithm

$$r(t+\delta t) = r(t) + v(t) \, \delta t + 1/2 \, a(t) \, \delta t^2 + b(t) \, \delta t^3 + O(\delta t^4)$$
 Taylor expand  $r(t-\delta t) = r(t) - v(t) \, \delta t + 1/2 \, a(t) \, \delta t^2 - b(t) \, \delta t^3 + O(\delta t^4)$  Reverse time  $r(t+\delta t) = 2 \, r(t) - r(t-\delta t) + a(t) \, \delta t^2 + O(\delta t^4)$  Add  $v(t) = (r(t+\delta t) - r(t-\delta t))/(2 \, \delta t) + O(\delta t^2)$  estimate velocities

Time reversal invariance is built in the energy does not drift.

Velocity is not required to compute the new position.

Once the new position is computed using position at t-  $\delta t$ , discard the old

Position. The current position become the old positions and the new position become the current position.

Note that velocity is used only to compute the kinetic energy and hence the temperature of the system

### Time reversible and area preserving

Tuckerman, Berne, Martyna, JCP, 97, 1990 1992

# Review of Hamiltonian Dynamics and Operators in Classical Mechanics

For a classical system, specifying the instantaneous positions and momenta of all the particles constituting the system can specify the microstate at any time t. For N particles there are 3N coordinates  $q_1,q_2...q_{3N}$  and 3N conjugate momenta  $p_1,p_2...p_{3N}$ . The equations of motion are *first order* differential equations

$$\dot{q}_{i} = \frac{\partial H(q_{i}, p_{i})}{\partial p_{i}} \qquad \dot{p}_{i} = -\frac{\partial H(q_{i}, p_{i})}{\partial q_{i}}$$

Let us consider a simple one-particle system in one dimension with a Hamiltonian

$$H = \frac{p^2}{2m} + U(x)$$

The equations of motion are

$$\dot{q} = \frac{p}{m}$$
  $\dot{p} = -\frac{dU}{dx} = F(x)$ 

Now Liouville's theorem says that any phase space function A(x, p) evolves as

$$\frac{dA}{dt} = \{A, H\}$$

where the {A, H} is the Poisson bracket given by

$$\{A, H\} = \frac{\partial H}{\partial p} \frac{\partial A}{\partial x} - \frac{\partial H}{\partial x} \frac{\partial A}{\partial p}$$

The evolution equation gives back Hamilton's equation of motion: To see that take A(x,p) = x. Then

$$\frac{dx}{dt} = \dot{x} = \{x, H\}$$

$$\{x, H\} = \frac{p}{m} \frac{\partial x}{\partial x} - \frac{dU}{dx} \frac{\partial x}{\partial p} = \frac{p}{m}$$

$$\dot{x} = p/m \quad \text{Since} \quad \frac{\partial x}{\partial p} = \mathbf{0}$$

So we have

Similarly if we take A(x,p) = p

$$\frac{dp}{dt} = \dot{p} = \{p, H\}$$

$$\{p,H\} = \frac{p}{m} \frac{\partial p}{\partial x} - \frac{dU}{dx} \frac{\partial p}{\partial p} = -\frac{dU}{dx} = F(x)$$

So we have  $\dot{p} = F(x)$ 

# As expected evolution equation gives back Hamilton's equation of motion

Now define a two-dimensional phase space vector  $\Gamma$ =(x,p). Hamilton's equation of motion for this  $\Gamma$  is

$$\frac{d\Gamma}{dt} = \{\Gamma, H\}$$

Now we define Liouville operator L such that  $iL\Gamma = \{\Gamma, H\}$ 

### iL can be expressed as differential operator using Hamilton's equation

$$iL = \frac{\partial H}{\partial p} \frac{\partial}{\partial x} - \frac{\partial H}{\partial x} \frac{\partial}{\partial p}$$

$$= \frac{p}{m} \frac{\partial}{\partial x} - \frac{dU}{dx} \frac{\partial}{\partial p}$$

$$= \frac{p}{m} \frac{\partial}{\partial x} - F(x) \frac{\partial}{\partial p}$$

$$= \dot{x} \frac{\partial}{\partial x} + \dot{p} \frac{\partial}{\partial p}$$

The equation of motion in operator form is given by

$$\frac{d\Gamma}{dt} = iL\Gamma$$

which can be solved to give

$$\Gamma(t) = e^{iLt}\Gamma(0)$$

The operator  $\exp(iLt)$  is called the classical propagator and the presence of i gives a nice analogy with the QM propagator  $\exp(-iHt/\hbar)$ 

### Properties of Liouville Operator and propagator

It is Hermitian:  $L^{\dagger} = L$  Prove this? (Home work)

The propagator  $U(t) \equiv exp(iLt)$  is a unitary operator

$$U^{\dagger}(t)U(t) = I$$
 Prove this? (Home work)

The unitarity of the propagator implies time reversal symmetry in the equations of motion. If the system is propagated forward in time up to a time t and then the clock is allowed to run backwards for a time -t, the system will evolve according to the same equations of motion but the direction of the velocities will be reversed, so that the system will simply return to its initial condition.

$$U(-t) = exp(-iLt)$$

Now apply U(t) on  $\Gamma(0)$  to get  $\Gamma(t)$  followed by U(-t):  $\Gamma(t) = U(t) \Gamma(0)$ 

$$U(-t) U(t) \Gamma(0) = e^{-iLt} e^{iLt}\Gamma(0) = \Gamma(0)$$

So U(-t) U(t) = I 
$$\Rightarrow$$
 U(-t) = U<sup>†</sup>(t) since U(t) is unitary

Since  $U^{\dagger}(t)$  is equivalent to backward propagation in time, it implies time reversibility since  $U^{\dagger}(t)U(t)=I$ 

Another important property of the unitary operator U(t) is that its determinant is 1 (Homework)

Unitarity of the propagator is consistent with the fact the volume in phase space is preserved under Hamilton's equation (Homework)

### **Trotter Theorem**

We have the evolution of the phase space vector

$$\Gamma(t) = e^{iLt}\Gamma(0)$$

In general it is difficult to evaluate exp(iLt) the reason for which will be clear from the following discussion

Remember iL can be written as

$$iL = \frac{p}{m} \frac{\partial}{\partial x} + F(x) \frac{\partial}{\partial p} = iL_1 + iL_2$$

$$iL_1 = \frac{p}{m} \frac{\partial}{\partial x}$$
  $iL_2 = F(x) \frac{\partial}{\partial p}$ 

The difficulty in the any computation arises from the fact that  $iL_1$  and  $iL_2$  do not commute :  $[iL_1,iL_2] \neq 0$ 

Show that  $iL_1$  and  $iL_2$  do not commute :  $[iL_1, iL_2] \neq 0$  (Homework)

### Since they don't commute

$$\exp(iL_1t+iL_2t) \neq \exp(iL_1t)\exp(iL_2t)$$

We can see this easily if we expand both side by Taylor expansion

$$\begin{split} e^{iLt} = &1 + (iL_1 + iL_2)t + \frac{1}{2}(iL_1 + iL_2)^2t^2 + \dots \\ = &1 + iLt + \frac{1}{2}\Big[(iL_1)^2 + (iL_2)^2 + (iL_1)(iL_2) + (iL_2)(iL_1)\Big]t^2 + \dots \end{split}$$

Now 
$$e^{iL_{1}t} = 1 + (iL_{1})t + \frac{1}{2}(iL_{1})^{2}t^{2} + \dots \qquad e^{iL_{2}t} = 1 + (iL_{2})t + \frac{1}{2}(iL_{2})^{2}t^{2} + \dots$$

$$e^{iL_{1}t}e^{iL_{2}t} = [1 + (iL_{1})t + \frac{1}{2}(iL_{1})^{2}t^{2} + \dots][1 + (iL_{2})t + \frac{1}{2}(iL_{2})^{2}t^{2} + \dots]$$

$$= 1 + iL_{1}t + iL_{2}t + [(iL_{1})(iL_{2}) + \frac{1}{2}(iL_{1})^{2} + \frac{1}{2}(iL_{2})^{2}]t^{2} + \dots$$

$$= (iL_{1} + iL_{2})t$$

$$\neq e^{(iL_{1} + iL_{2})t}$$

### Trotter theorem comes to our rescue

$$e^{(iL_1+iL_2)t} = \lim_{M \to \infty} \left[ e^{iL_2t/2M} e^{iL_1t/M} e^{iL_2t/2M} \right]^M$$

For a proof see Techniques and Applications of Path Integral by L. S. Schulman

For large but finite M above equation can be approximated as

$$e^{(iL_{1}+iL_{2})t} = \begin{bmatrix} iL_{2}t/2M & iL_{1}t/M & iL_{2}t/2M \\ e^{(iL_{1}+iL_{2})t/M} & e^{iL_{2}t/2M} & e^{iL_{1}t/M} & e^{iL_{2}t/2M} \end{bmatrix}^{M}$$

$$e^{(iL_{1}+iL_{2})t/M} \approx e^{iL_{2}t/2M} e^{iL_{1}t/M} e^{iL_{2}t/2M}$$

The expression on the left looks like approximate propagation of the system up to time t by M application of the operator in the bracket. If we interpret t/M as single time step,  $\delta t$ , then we have

$$e^{iL\delta t} = e^{(iL_1 + iL_2)\delta t} \approx e^{iL_2\delta t/2} e^{iL_1\delta t} e^{iL_2\delta t/2}$$

This is the propagator  $U(\delta t)$  for the time step  $\delta t$ . Like U(t),  $U(\delta t)$  is unitary and preserve the time reversibility of the dynamics

Show 
$$U^{\dagger}(\delta t) = U(-\delta t) = U^{-1}(\delta t)$$
  
and  $U(-\delta t)$   $U(\delta t) = I$  (Homework)

Now let us see what is the effect of the propagator  $U(\delta t)$  on the coordinates and momenta of the particles

Some useful identity  $c \frac{\partial}{\partial x} = x + c$   $[1 + c \frac{\partial}{\partial x} + \frac{1}{2}c^2 \frac{\partial^2}{\partial x^2} + \dots]x = x + c$ 

$$e^{c\frac{\partial}{\partial x}}g(x) = \sum_{k=0}^{\infty} \frac{1}{k!} \left[c\frac{\partial}{\partial x}\right]^{k} g(x)$$

$$= \sum_{k=0}^{\infty} \frac{1}{k!} c^{k} g^{(k)} x$$

$$= \sum_{k=0}^{\infty} \frac{1}{k!} c^{k} g^{(k)} x$$
This is just the Taylor series of  $g(x+c)$  so  $e^{c\frac{\partial}{\partial x}} g(x) = g(x+c)$ 

# Note that the action of $\exp(a\partial/\partial p)$ on x or g(x) has no effect: it acts like identity operator

$$U(\delta t)x = e^{\frac{\delta t}{2}F(x)} \frac{\partial}{\partial p} \int_{e}^{\delta t} \frac{p}{m} \frac{\partial}{\partial x} \int_{e}^{\delta t} F(x) \frac{\partial}{\partial p} \int_{x}^{\delta t} F(x) \frac{\partial}{\partial p} \int_{x}$$

$$\frac{\delta t}{2}F(x)\frac{\partial}{\partial p}\delta t\frac{p}{m}\frac{\partial}{\partial x}x$$

$$= e^{\frac{\delta t}{2}}F(x)\frac{\partial}{\partial p}(x+\frac{\delta t}{m}p)$$

$$= x+\frac{\delta t}{m}(p+\frac{\delta t}{2}F(x))$$

$$x(\delta t) = x+\frac{\delta t}{m}p+\frac{\delta t^{2}}{2m}F(x)$$

 $\frac{\partial t}{\partial p} F(x) \frac{\partial}{\partial p} \frac{\partial t}{\partial p} \frac{p}{e} \frac{\partial}{m} \frac{\partial}{\partial x} x$ First operator acting on x has no effect, it involves only momentum derivative

Second operator changes x to  $x+\delta t p/m$ 

 $= x + \frac{\delta t}{m} (p + \frac{\delta t}{2} F(x))$  Last operator acting on x has no effect. It acts on p and changes p to p+\delta t F(x)/2

### Similarly we can apply $U(\delta t)$ on p to get

$$U(\delta t)p = e^{\frac{\delta t}{2}F(x)}\frac{\partial}{\partial p} \delta t \frac{p}{m}\frac{\partial}{\partial x} \frac{\delta t}{2}F(x)\frac{\partial}{\partial p} p$$

$$= e^{\frac{\delta t}{2}F(x)}\frac{\partial}{\partial p} \delta t \frac{p}{m}\frac{\partial}{\partial x} (p + \frac{\delta t}{2}F(x)) \quad \text{First operator acting on p changes p to } \\ = e^{\frac{\delta t}{2}F(x)}\frac{\partial}{\partial p} (p + \frac{\delta t}{2}F(x + \frac{\delta t}{m}p) \quad \text{Next operator acting on } F(x) \text{ changes it to } \\ = e^{\frac{\delta t}{2}F(x)}\frac{\partial}{\partial p} (p + \frac{\delta t}{2}F(x + \frac{\delta t}{m}p) \quad \text{Next operator acting on } F(x) \text{ changes it to } \\ = p + \frac{\delta t}{2}F(x) + \frac{\delta t}{2}F(x + \frac{\delta t}{m}(p + \frac{\delta t}{2}F(x))) \quad \text{Last operator acting on p changes } p \\ = p(\delta t) = p + \frac{\delta t}{2}[F(x) + F(x + \frac{\delta t}{m}p + \frac{\delta t^2}{2m}F(x))] \\ = p(0) + \frac{\delta t}{2}[F(x(0)) + F(x(\delta t))] \quad \text{Argument of the second force in the above expression is just } x(\delta t)$$

#### THE GLOBAL MD ALGORITHM

### 1. Input initial conditions

Potential interaction V as a function of atom positions Positions r of all atoms in the system Velocities v of all atoms in the system



### repeat 2,3,4 for the required number of steps:

#### 2. Compute forces

The force on any atom

$$m{F}_i = -rac{\partial V}{\partial m{r}_i}$$

is computed by calculating the force between non-bonded atom

pairs:

$$F_i = \sum_j F_{ij}$$

plus the forces due to bonded interactions (which may depend on 1, 2, 3, or 4 atoms), plus restraining and/or external forces.

The potential and kinetic energies and the pressure tensor are

computed.



### 3. Update configuration

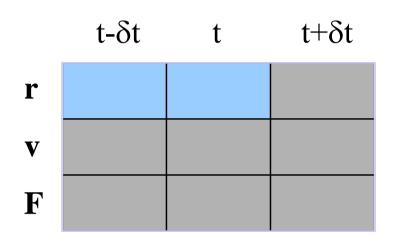
The movement of the atoms is simulated by numerically solving Newton's equations of motion

$$\frac{\mathrm{d}^2 r_i}{\mathrm{d}t^2} = \frac{F_i}{m_i}$$
or
$$\frac{\mathrm{d}r_i}{\mathrm{d}t} = v_i; \quad \frac{\mathrm{d}v_i}{\mathrm{d}t} = \frac{F}{m_i}$$

### 4. if required: Output step

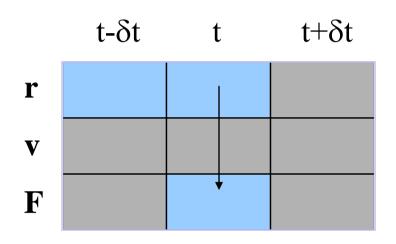
write positions, velocities, energies, temperature, pressure, etc.

# Verlet Algorithm: Flow Diagram



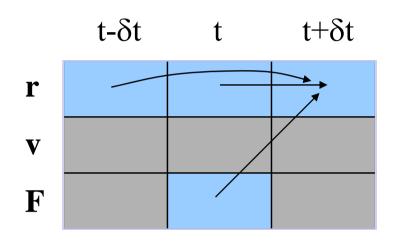
Given current position and position at end of previous time step

# Verlet Algorithm: Flow Diagram



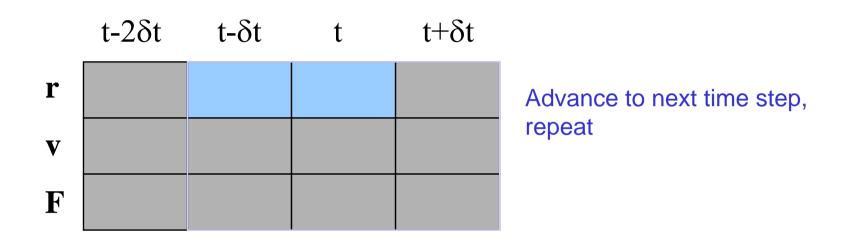
Compute the force at the current position

# Verlet Algorithm: Flow Diagram



Compute new position from present and previous positions, and present force

## Verlet Algorithm: Flow Diagram



## Verlet Algorithm: Loose Ends

#### Initialization

- how to get position at "previous time step" when starting out?
- simple approximation

$$\mathbf{r}(t_0 - \delta t) = \mathbf{r}(t_0) - \mathbf{v}(t_0) \delta t$$

#### Obtaining the velocities

- not evaluated during normal course of algorithm
- needed to compute some properties, e.g.
  - temperature
  - diffusion constant
- finite difference

$$\mathbf{v}(t) = \frac{1}{2\delta t} \left[ \mathbf{r}(t + \delta t) - \mathbf{r}(t - \delta t) \right] + O(\delta t^2)$$

### Verlet Algorithm Performance Issues

- Time reversible
  - forward time step

$$\mathbf{r}(t+\delta t) = 2\mathbf{r}(t) - \mathbf{r}(t-\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t^{2}$$

- replace  $\delta t$  with  $-\delta t$ 

$$\mathbf{r}(t + (-\delta t)) = 2\mathbf{r}(t) - \mathbf{r}(t - (-\delta t)) + \frac{1}{m}\mathbf{F}(t)(-\delta t)^{2}$$
$$\mathbf{r}(t - \delta t) = 2\mathbf{r}(t) - \mathbf{r}(t + \delta t) + \frac{1}{m}\mathbf{F}(t)\delta t^{2}$$

- same algorithm, with same positions and forces, moves system backward in time
- Numerical imprecision of adding large/small numbers

$$\mathbf{r}(t+\delta t) - \mathbf{r}(t) = \mathbf{r}(t) - \mathbf{r}(t-\delta t) + \frac{1}{m} \mathbf{F}(t) \delta t^{2}$$

$$O(\delta t^{0}) O(\delta t^{1}) O(\delta t^{2})$$

### Initial Velocity

- Random direction
  - randomize each component independently
  - randomize direction by choosing point on spherical surface
- Magnitude consistent with desired temperature. Choices:
  - Maxwell-Boltzmann:  $prob(v_x) \propto \exp(-\frac{1}{2}mv_x^2/kT)$
  - Same for y, z components
- Be sure to shift so center-of-mass momentum is zero  $P_x = \frac{1}{N} \sum p_{i,x}$

$$p_{i,x} \to p_{i,x} - P_x$$

- Eliminates addition of small numbers  $O(\delta t^2)$  to differences in large ones  $O(\delta t^0)$
- Algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t+\frac{1}{2}\delta t)\delta t$$
$$\mathbf{v}(t+\frac{1}{2}\delta t) = \mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t$$

- Eliminates addition of small numbers  $O(\delta t^2)$  to differences in large ones  $O(\delta t^0)$
- Algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t+\frac{1}{2}\delta t)\delta t$$

$$\mathbf{v}(t+\frac{1}{2}\delta t) = \mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t$$

• Mathematically equivalent to Verlet algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \left[\mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t\right]\delta t$$

- Eliminates addition of small numbers  $O(\delta t^2)$  to differences in large ones  $O(\delta t^0)$
- Algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t+\frac{1}{2}\delta t)\delta t$$

$$\mathbf{v}(t+\frac{1}{2}\delta t) = \mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t$$

Mathematically equivalent to Verlet algorithm

$$\mathbf{r}(t + \delta t) = \mathbf{r}(t) + \left[\mathbf{v}(t - \frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t\right]\delta t$$

$$\mathbf{r}(t) \text{ as evaluated from previous time step}$$

$$\mathbf{r}(t) = \mathbf{r}(t - \delta t) + \mathbf{v}(t - \frac{1}{2}\delta t)\delta t$$

previous time step

- Eliminates addition of small numbers  $O(\delta t^2)$ to differences in large ones  $O(\delta t^0)$
- Algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t+\frac{1}{2}\delta t)\delta t$$

$$\mathbf{v}(t+\frac{1}{2}\delta t) = \mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t$$

• Mathematically equivalent to Verlet algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \left[\mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t\right]\delta t$$

previous time step

$$\mathbf{r}(t)$$
 as evaluated from  $\mathbf{r}(t) = \mathbf{r}(t - \delta t) + \mathbf{v}(t - \frac{1}{2}\delta t)\delta t$ 

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \left[ \left( \mathbf{r}(t) - \mathbf{r}(t-\delta t) \right) + \frac{1}{m} \mathbf{F}(t) \delta t^2 \right]$$

- Eliminates addition of small numbers  $O(\delta t^2)$  to differences in large ones  $O(\delta t^0)$
- Algorithm

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t+\frac{1}{2}\delta t)\delta t$$

$$\mathbf{v}(t+\frac{1}{2}\delta t) = \mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t$$

• Mathematically equivalent to Verlet algorithm

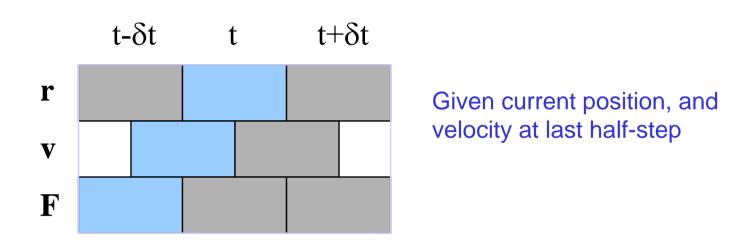
$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \left[\mathbf{v}(t-\frac{1}{2}\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t\right]\delta t$$

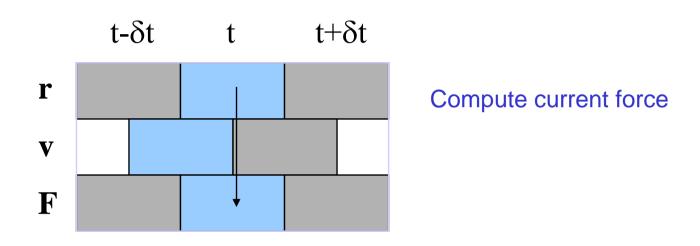
**r**(t) as evaluated from previous time step

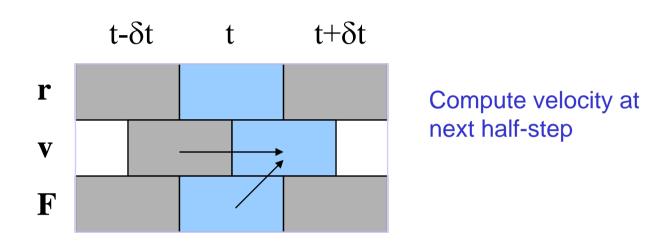
$$\mathbf{r}(t) = \mathbf{r}(t - \delta t) + \mathbf{v}(t - \frac{1}{2}\delta t)\delta t$$

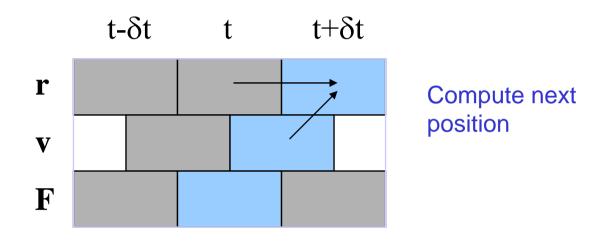
$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \left[ \left( \mathbf{r}(t) - \mathbf{r}(t-\delta t) \right) + \frac{1}{m} \mathbf{F}(t) \delta t^2 \right]$$

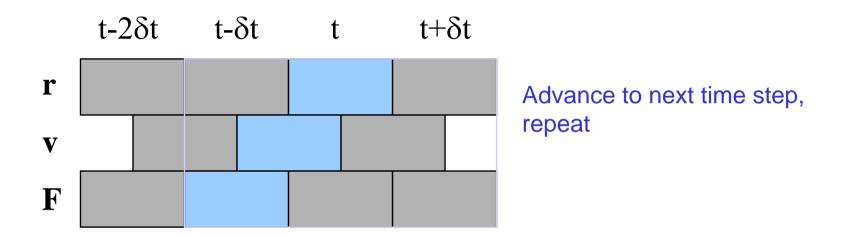
$$\mathbf{r}(t+\delta t) = 2\mathbf{r}(t) - \mathbf{r}(t-\delta t) + \frac{1}{m}\mathbf{F}(t)\delta t^2$$
 original Verlet algorithm











### Leapfrog Algorithm Loose Ends

- Initialization
  - how to get velocity at "previous time step" when starting out?
  - simple approximation

$$\mathbf{v}(t_0 - \delta t) = \mathbf{v}(t_0) - \frac{1}{m} \mathbf{F}(t_0) \frac{1}{2} \delta t$$

- Obtaining the velocities
  - interpolate

$$\mathbf{v}(t) = \frac{1}{2} \left[ \mathbf{v}(t + \frac{1}{2}\delta t) + \mathbf{v}(t - \frac{1}{2}\delta t) \right]$$

#### Velocity Verlet algorithm

 $r(\delta t) = r(0) + v(0) \delta t + 1/2 a(r(0)) \delta t^2 + b(t) \delta t^3 + O(\delta t^4)$  good to  $2^{nd}$  order in  $\delta t$  Do the Taylor expand for velocities

$$v(\delta t) = v(0) + \delta t \ a(r(0)) + O(\delta t^2) \ good \ to \ 1^{st} \ order \ in \ \delta t$$

To get V also accurate to the  $2^{nd}$  order in  $\delta t$  we consider starting from  $\delta t$  and applying the rule backward in time (I.e. for a time  $-\delta t$ ) so that we end up back at

$$r(0) = r(\delta t) - v(\delta t) \delta t + 1/2 a(r(\delta t)) \delta t^{2}$$
$$v(\delta t) = (r(\delta t) - r(0))/(\delta t) + \delta t/2 a(r(\delta t))$$

Now using the position equation we have

$$r(\delta t) - r(0) = v(0) \delta t + 1/2 a(r(0)) \delta t^2$$

So we have  $v(\delta t) = v(0) + \delta t / 2 [a(r(0)) + a(r(\delta t))]$ 

Combining we have

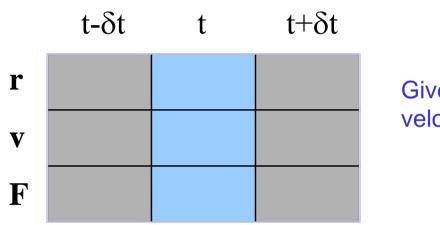
$$r(\delta t) = r(0) + v(0) \delta t + 1/2 a(r(0)) \delta t^2$$
  
 $v(\delta t) = v(0) + \delta t/2 [a(r(0)) + a(r(\delta t))]$  Velocity Verlet

#### Velocity Verlet Algorithm

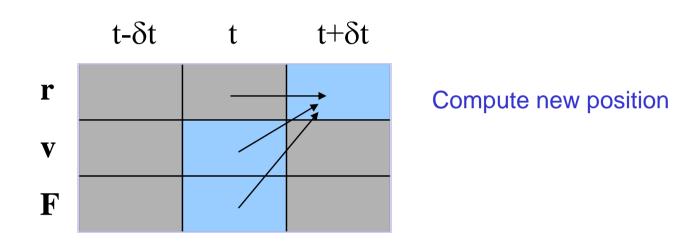
- Round off advantage of leapfrog, but better treatment of velocities
- Algorithm

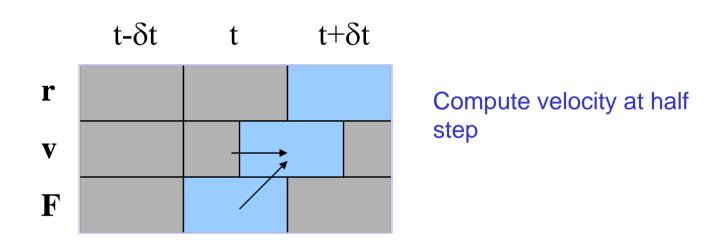
$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \mathbf{v}(t)\delta t + \frac{1}{2m}\mathbf{F}(t)\delta t^{2}$$
$$\mathbf{v}(t+\delta t) = \mathbf{v}(t) + \frac{1}{2m}[\mathbf{F}(t) + \mathbf{F}(t+\delta t)]\delta t$$

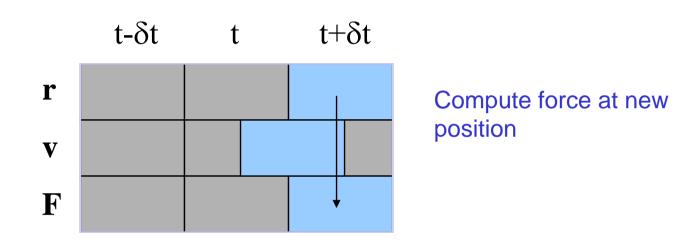
- Implemented in stages
  - evaluate current force
  - compute *r* at new time
  - add current-force term to velocity (gives v at half-time step)
  - compute new force
  - add new-force term to velocity
- Also mathematically equivalent to Verlet algorithm (in giving values of **r**)

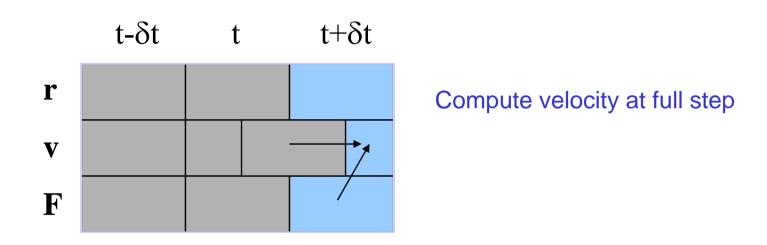


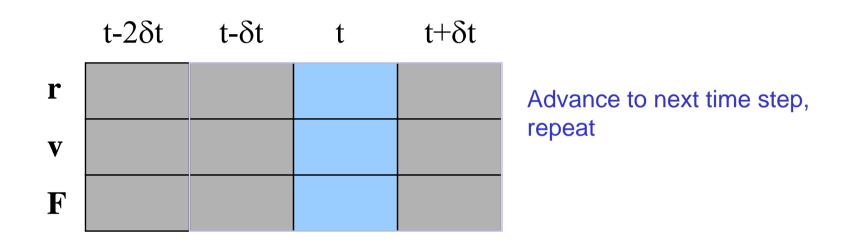
Given current position, velocity, and force











#### Lines of code for Leap-frog Verlet algorithm

```
/* Carry out half-timestep update of atomic velocities using old forces. */
 velocity_step(n_atoms, atom_vels, atom_mass, delta, comp_forces);
 /* Carry out full-timestep update of atomic positions using half-timestep velocities. */
 position step(n atoms, atom coords, atom vels, atom move,
          scaled atom coords, h inv, h, delta, neigh switch);
/* Add short-range nonbonded forces to force accumulators. */
   for (i = 0; i < n \text{ atoms}; ++i)
     for (k = 0; k < NDIM; ++k)
     comp forces[i][k] += f vdw s[i][k] + f coul s[i][k];
/* Carry out half-timestep update of atomic velocities using new forces. */
 velocity step(n atoms, atom vels, atom mass, delta, comp forces);
```

```
void velocity_step(int n_atoms, double **atom_vels,
           double *atom mass, double delta, double **comp forces)
 int i, k;
 double delta_over_2m;
 /* Update velocities. */
 for (i = 0; i < n_atoms; ++i) {
   delta_over_2m = delta /(2*atom_mass[i]);
   for ( k = 0; k < NDIM; ++k)
   atom_vels[i][k] += delta_over_2m * comp_forces[i][k];
```

```
/* Carry out full-timestep update of atomic positions using half-timestep velocities. */
void position_step(int n_atoms, double **atom_coords,
           double **atom vels,
           double **atom move, double **scaled atom coords,
           double **h inv, double **h, double delta, int neigh switch)
 int i, k;
 double dr[NDIM];
/* Update positions. */
 for (i = 0; i < n \text{ atoms}; ++i) {
    for (k = 0; k < NDIM; ++k) {
   dr[k] = delta * atom vels[i][k];
   atom coords[i][k] += dr[k];
} }
/* If we are using periodic boundary conditions, calculate scaled atomic coordinates. */
   scaled_atomic_coords(n_atoms, h_inv, atom_coords, scaled_atom_coords);
   periodic boundary conditions(n atoms, h, scaled atom coords, atom coords);
```

#### Velocity sampling routine

```
/* Sample atomic velocities from a Maxwellian distribution. */
void sample_velocities(long *idum, double **atom_vels, int n_atoms,
             double *sqrt kT over m)
 int i, k;
 double sqrt kT by m;
 /* Choose Gaussian-distributed cartesian velocity components for each
   atom, with zero mean and standard deviation equal to sqrt(kT/m). */
 for (i = 0; i < n_atoms; ++i) {
   sqrt_kT_by_m = sqrt_kT_over_m[i];
    for (k = 0; k < NDIM; ++k)
   atom vels[i][k] = sqrt kT by m * gasdev(idum);
```

#### **Other Algorithms**

Leap Frog: This can be derived from Verlet algorithm. Velocities at half integer time step can be written as

$$v(t-\delta t/2) = [r(t) - r(t-\delta t)]/\delta t$$

$$v(t+\delta t/2) = [r(t+\delta t) - r(t)]/\delta t$$

From the above two equation we get the expression for the new position

$$r(t+\delta t) = r(t) + \delta t v(t+\delta t/2)$$

To update the velocity we use the expression from Verlet algorithm

$$v(t+\delta t/2) = v(t-\delta t/2) + \delta t a(t)$$

Note that velocities are not defined at the same time as the positions, so KE and PE are also not defined at the same time, and hence we can not directly compute the total energy in the Leap-Frog scheme

## Leap-Frog cont.

How ever it is possible to cast it in a way to look more Verlet like

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \delta t \left[ \mathbf{v}(t) + (\delta t/2) \mathbf{a}(t) \right]$$

Velocity update is done as follows

$$v(t+\delta t)=v(t)+(\delta t/2)[a(t+\delta t)+a(t)]$$

New velocity computation requires new force and hence new positions

In this form it is equivalent to velocity Verlet algorithm

#### **Higher Order algorithm: Predictor-corrector**

Consider the Taylor expansion of the  $r(t+\delta t)$ 

$$r(t+\delta t) = r(t) + v(t) \delta t + 1/2 a(t) \delta t^2 + b(t) \delta t^3 + O(\delta t^4)$$

## How to set the time step

- Adjust to get energy conservation to 1% of kinetic energy.
- Even if errors are large, you are close to the exact trajectory of a nearby physical system with a different potential.
- Since we don't really know the potential surface that accurately, this is satisfactory.
- Leapfrog algorithm has a problem with round-off error.
- Use the equivalent velocity Verlet instead:

$$\mathbf{r}(t+\delta t) = \mathbf{r}(t) + \delta t \left[ \mathbf{v}(t) + (\delta t/2) \mathbf{a}(t) \right]$$

$$\mathbf{v}(t+\delta t/2) = \mathbf{v}(t) + (\delta t/2) \mathbf{a}(t)$$

$$\mathbf{v}(t+\delta t) = \mathbf{v}(t+\delta t/2) + (\delta t/2) \mathbf{a}(t+\delta t)$$

#### Linear Stability analysis for Harmonic oscillator

$$F(x) = -\omega^2 x$$

#### Position Verlet scheme can be written as

$$\begin{pmatrix} \omega x(t+\delta t) \\ v(t+\delta t) \end{pmatrix} = S \begin{pmatrix} \omega x(t) \\ v(t) \end{pmatrix}$$

where 
$$S = \begin{vmatrix} 1 - \varepsilon^2 / 2 & \varepsilon (1 - \varepsilon^2 / 4) \\ -\varepsilon & 1 - \varepsilon^2 / 2 \end{vmatrix}$$
 and  $\varepsilon = \omega \delta t$ 

Powers of S is bounded if  $\varepsilon^2 < 4$ 

$$\delta t < 2/\omega \qquad T_p = 2\pi/\omega$$

$$< T_p/2\pi$$

#### How to increase time step?

#### Limiting factors: intra-molecular motions

Vibrational mode	Wave number $(1/\lambda)$ cm <sup>-1</sup>	Period $T_p$ ( $\lambda/c$ ) fs	$T_p/2\pi$ (fs)
O-H, N-H stretch	3200-3600	9.8	3.1
C-H stretch	3000	11.1	3.5
C≡C, C≡N stretch	2100	15.9	5.1
C=C stretch	1700	19.6	6.2
H-O-H bend	1600	20.8	6.4
O-C-O bend	700	47.6	15

Freeze or constrain the fast motion: make all bond, angle involving H rigid

#### Physical Quantities in Molecular Simulation

- State variables
  - each variable has an associated "conjugate" variable
    - temperature  $\Leftrightarrow$  energy (kT,E)
    - pressure  $\Leftrightarrow$  volume (P,V)
    - chemical potential  $\Leftrightarrow$  number of molecules  $(\mu,N)$
  - specification of state requires fixing one of each pair
  - the dependent variable can be measured by the simulation
- Configuration variables
  - position, orientation, momentum of each atom or molecule
  - energy, forces and torques
  - time
- Properties
  - transport coefficients, free energy, structural quantities, etc.
- Molecular model parameters
  - characteristic energy, size, charge

#### Separation of the Energy

• Total energy is sum of kinetic and potential parts

$$- E(p^N, r^N) = K(p^N) + U(r^N)$$

• Kinetic energy is quadratic in momenta

$$K(p^N) = \sum_i p_i^2 / 2m_i$$

• Kinetic contribution can be treated analytically in partition function

$$Q = \frac{1}{h^{3N}N!} \int dp^{N} e^{-\beta \sum p_{i}^{2}/2m} \int dr^{N} e^{-\beta U(r^{N})}$$

$$= \frac{1}{\Lambda^{3N}} \frac{1}{N!} \int dr^{N} e^{-\beta U(r^{N})}$$

$$= \frac{1}{\Lambda^{3N}} Z_{N}$$
 thermal de Broglie wavelength
$$= \frac{1}{\Lambda^{3N}} Z_{N}$$
 configuration integral
$$\Lambda = \frac{h}{\sqrt{2\pi mkT}}$$

And it drops out of position averages

$$\langle A \rangle = \frac{1}{Z_N} \frac{1}{N!} \int dr^N A(r^N) e^{-\beta U(r^N)}$$

### Simple Averages 1. Energy

Average energy

$$\langle E \rangle = \frac{1}{Q} \frac{1}{h^{3N}N!} \int dp^N \int dr^N E(p^N, r^N) e^{-\beta E(p^N, r^N)}$$

Note thermodynamic connection

$$\langle E \rangle = -\frac{\partial \ln Q}{\partial \beta} = \frac{\partial (A/kT)}{\partial (1/kT)} = E_{\text{internal}}$$

Average kinetic energy

$$\langle K \rangle = \frac{1}{h^{3N}} \int dp^N \sum \frac{p_i^2}{2m} e^{-\beta \sum p_i^2 / 2m}$$
  
=  $\frac{3}{2} NkT$  Equipartition of energy: kT/2 for each degree of freedom

Average potential energy

$$\langle U \rangle = \frac{1}{Z_N} \frac{1}{N!} \int dr^N U(r^N) e^{-\beta U(r^N)}$$

#### Simple Averages 2. Temperature

- Need to measure temperature in microcanonical ensemble (NVE) simulations
- Define instantaneous kinetic temperature

$$T = \frac{1}{3Nk} \sum p_i^2 / m$$
 More generally, divide by number of molecular degrees of freedom instead of 3N

- Thermodynamic temperature is then given as ensemble average
- Relies on equipartition as developed in canonical ensemble
- A better formulation has been developed recently (Thermostating)

#### Simple Averages 3a. Pressure

From thermodynamics and bridge equation

$$P = -\left(\frac{\partial A}{\partial V}\right)_{T,N} = kT \frac{\partial}{\partial V} \ln\left[\frac{1}{N!} \int dr^N e^{-\beta U(r^N)}\right]$$

$$P = \frac{NkT}{V} + \frac{1}{3V} \left\langle \sum_{\text{pairs i,j}} \vec{r}_{ij} \cdot \vec{f}_{ij} \right\rangle$$

### Simple Averages 4. Heat Capacity

• Example of a "2nd derivative" property

$$C_{v} = \left(\frac{\partial E}{\partial T}\right)_{V,N} = -k\beta^{2} \left(\frac{\partial^{2}(\beta A)}{\partial \beta^{2}}\right)_{V,N}$$
$$= -k\beta^{2} \frac{\partial}{\partial \beta} \frac{1}{Q(\beta)} \int dr^{N} dp^{N} E e^{-\beta E}$$

• Expressible in terms of fluctuations of the energy

$$C_v = k\beta^2 \left[ \left\langle E^2 \right\rangle - \left\langle E \right\rangle^2 \right]$$
 Note: difference between two O(N²) quantities to give a quantity of O(N)

- Other 2nd-derivative or "fluctuation" properties
  - isothermal compressibility

$$\kappa_T = -\frac{1}{V} \left( \frac{\partial V}{\partial P} \right)_{T,N}$$

#### Dimensions and Units 1. Magnitudes

- Important extensive quantities small in magnitude
  - when expressed in macroscopic units
- Small numbers are inconvenient
- Two ways to magnify them
  - work with atomic-scale units
    - ps, amu, nm or Å
  - make dimensionless with characteristic values
    - model values of size, energy, mass

Symbol	Definition	Value		
1. Constants				
k	Boltzmann's constant	1.3806×10 <sup>-28</sup> J/(molec·K)		
$N_{\theta}$	Ava gadro's number	$6.022 \times 10^{23}$		
2. Simule	ution Variables			
N	Number of molecules	$\sim 10^{3}$		
V	Simulation cell volume	$\sim 10^{-24}  \mathrm{m}^3$		
m	Molecular mass	$\sim 10^{-25}\mathrm{kg/molec}$		
ρ	Number density	$\sim 10^{27}  \mathrm{molec/m^3}$		
E	Energy (total)	$\sim 10^{-20}\mathrm{J/molec}$		
ţ	time	$\sim 10^{-12}  \mathrm{s}$		
3. Model Variables				
$\sigma$	Size variable	$\sim 5\times10^{-10}~\text{m}$		
E	Energy variable	$\sim 10^{-21}\mathrm{J/mo1ec}$		
$\gamma_{\delta}$	Bond distance	$\sim 10^{-10}  \text{m}$		
$k_{\bullet}$	Vibrational spring constant	$\sim 10^3 \text{ J/m}^2$		

#### Dimensions and Units 2. Scaling

• In simulations it is often convenient to express quantities such as temperature, density, pressure and the like in reduced units. This means that we choose a convenient unit of energy, length and mass and then express all the other quantities in terms of these basic units. A natural choice of our basic units is the following

- size σ
- energy ε
- mass m

In terms of these basic units, all other units follow

Symbol	Meaning	Definition
y*	dimensionless distance	r/σ
$\mathcal{E}^*$	dimensionless energy	E/∈
$T^*$	dimensionless temperature	$kT/\epsilon$
$U^*$	dimensionless internal energy	U/∈
<i>t</i> *	dimensionless time	$t/[\sigma(m/\epsilon)^{0.5}]$
v*	dimensionless velocity	$\forall l (\epsilon / m)^{0.5}$
₽*	dimensionless force	Fo∕e
$P^*$	dimensionless pressure	$P\sigma^3/\epsilon$
$D^*$	dimensionless self diffusion coefficient	$D/[\sigma(\epsilon/m)^{0.5}]$

#### Why reduced Units?

- Many combinations of  $\rho$ , T,  $\epsilon$  and  $\sigma$  all correspond to the same state in reduced units. This is the law of corresponding states: the same simulation can of a LJ model can be used to study the Argon at 60 K and density 840 kg/m3 and Xe at 112 K and a density at 1617 kg/m3. In reduced unit both simulations corresponds to the state point  $\rho = 0.5$  and T = 0.5. Scaling by model parameters
- In reduced units almost all quantities of interest are of order 1 (say between 10-3 and 103). Hence if we suddenly find very large (or very small) number in our simulations, suspect some error somewhere.
- Simulation results obtained in reduced units can be translated back into real units.

See the following table

# Conversion of reduced Units to real Units for LJ argon system: $\epsilon/k_B = 119.8$ K, $\sigma=3.405x10^{-10}$ m, m=0.03994 kg/mol

Quantity	Reduced Units	Real Units
Temperatur e	$T^* = 1$	$T = \varepsilon/k_B = 119.8 \text{ K}$
Density	$\rho^* = 1$	$\rho = 1680 \text{ kg/m}^3$
Time	$\delta t^* = 0.005$	$\delta t = 1.09 \times 10^{-14} s$
Pressure	P*=1	P = 41.9 MPa