



Physics 1: Mechanics and Waves

Week 7 – Circular motion

1. Acceleration in curvilinear Motion

$$\vec{a} = \frac{\mathrm{d}v}{\mathrm{d}t}\hat{t} + \frac{v^2}{\rho}\hat{n} = \vec{a}_t + \vec{a}_n$$

2. Nonuniform circular motion

$$\vec{a} = \frac{\mathrm{d}v}{\mathrm{d}t}\hat{t} + \frac{v^2}{r}\hat{n} = \vec{a}_t + \vec{a}_c$$

3. Acceleration in uniform circular motion

$$\vec{a} = \frac{\mathrm{d}v}{\mathrm{d}t}\hat{t} + \frac{v^2}{r}\hat{n} = \frac{v^2}{r}\hat{n}$$

Review

4. Description of the circular motion with angular coordinates

- ①Angular position (coordinate) $\theta(t)$
- ② Angular displacement $\Delta\theta(t)$
- 3 angular speed and angular velocity vector

$$\vec{\omega}(t) = \omega_z(t)\hat{k} = \frac{\mathrm{d}\theta(t)}{\mathrm{d}t}\hat{k}$$

rad/s

angular acceleration

$$\vec{\alpha}(t) = \frac{d\vec{\omega}(t)}{dt} = \alpha_z(t)\hat{k}$$
 SI: rad/s²

5. The relationship between angular and linear quantities in circular motion

The relationship between angular and linear quantities in circular otion
$$\vec{r}(t) = r \cos \theta \hat{i} + r \sin \theta \hat{j}$$

$$\Delta s = r \Delta \theta$$

$$v = \frac{\mathrm{d}s(t)}{\mathrm{d}t} = r\frac{\mathrm{d}\theta(t)}{\mathrm{d}t} = r\omega$$

$$\vec{v}(t) = \vec{\omega}(t) \times \vec{r}(t)$$

$$dt dt$$

$$\vec{a}(t) = \vec{\alpha}(t) \times \vec{r}(t) + \vec{\omega}(t) \times \vec{v}(t)$$

$$= r\alpha_z \hat{t} + r\omega^2 \hat{n}$$

$$= \vec{a}_t + \vec{a}_c$$

$$a_t = r\alpha_z$$

$$a_n = r\omega^2$$

Nonuniform circular motion - with a constant angular acceleration

Initial time:
$$\alpha_z = \text{constant}, \quad \omega(t_i) = \omega_{z0}, \quad \theta(t_i) = \theta_0$$

From
$$\alpha_z = \frac{d\omega_z}{dt} = \text{costant}$$

Integrate in both sides

From
$$\omega_z(t) = \frac{d\theta(t)}{dt}$$

Integrate in both sides

es
$$\int_{\theta_0}^{\theta} d\theta = \int_0^t \omega_z dt = \int_0^t (\omega_{z0} + \alpha_z t) dt$$
$$\theta(t) = \theta_0 + \omega_{z0} t + \frac{1}{2} \alpha_z t^2$$

 $\int_{\omega_z}^{\omega_z} d\omega_z = \int_0^t \alpha_z dt, \quad \omega_z(t) = \omega_{z0} + \alpha_z t$

les
$$\int_{\theta_0}^{\theta} d\theta = \int_0^t \omega_z dt = \int_0^t (\omega_{z0})^t d\theta$$
$$\theta(t) = \theta_0 + \omega_{z0}t + \frac{1}{2}\alpha_z t^2$$

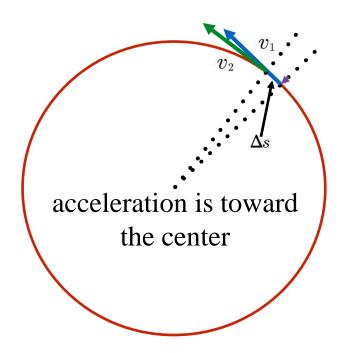
Nonuniform circular motion - with a constant angular acceleration

Comparison: nonuniform circular motion with a constant angular acceleration and the rectilinear motion with a constant acceleration

$$\begin{cases} \alpha_z = \frac{d\omega_z}{dt} = costant \\ \omega_z(t) = \omega_{z0} + \alpha_z t \\ \theta(t) = \theta_0 + \omega_{z0} t + \frac{1}{2}\alpha_z t^2 \\ \omega_z^2 - \omega_{z0}^2 = 2\alpha_z \Delta \theta \end{cases}$$

$$\begin{cases} a_{x} = \frac{dv_{x}}{dt} = costant \\ v_{x}(t) = v_{x0} + a_{x}t \\ x(t) = x_{0} + v_{x0}t + \frac{1}{2}a_{x}t^{2} \\ v^{2} - v_{0}^{2} = 2a\Delta x \end{cases}$$

Uniform Circular motion - Centripetal acceleration



Direction and magnitude

$$a_{c} = \frac{v^{2}}{r}$$

$$v = \frac{2\pi r}{T}$$

Number of revolutions per second f = 1/T Time for one complete revolution

Correspondence

	Linear	Circular	Relationship
Displacement	$\boldsymbol{\mathcal{X}}$	θ	$\theta = \frac{x}{r}$
Velocity	v	ω	$\omega = \frac{v}{r}$
Acceleration	$a_{ au}$	α	$\alpha = \frac{a}{r}$

Why



How to make sure the speed limit when a car rounds a curve?



Why a people in a satellite orbit close to the Earth will experience "weightlessness"
There is no gravity in space?
What is the net force on an astronaut at rest inside the space station?

Dynamics of Uniform Circular Motion

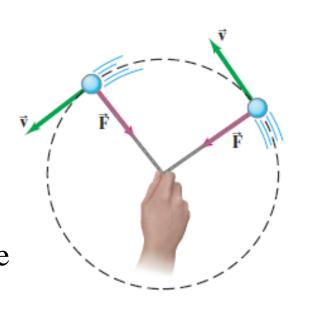
Centripetal force:

If the particle has a mass *m* and is being accelerated, it must be subject to a force.

$$F = ma$$

This force is acting towards the center of the circle – it is centripetal (meaning = center seeking)

$$F = ma_{cent} = \frac{mv^2}{r}$$



Dynamics of Uniform Circular Motion

What is the normal force at the top and at the bottom?

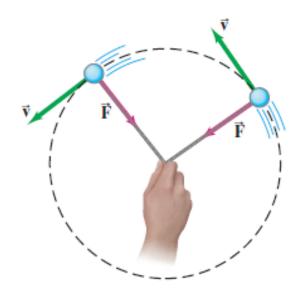


speed = 15 m/s diameter = 40 m $total\ mass = 1200 \text{ kg}$

Roller Coaster

Centripetal Force

What will happen without this force?





Centripetal Force

EXERCISE D To negotiate a flat (unbanked) curve at a *faster* speed, a driver puts a couple of sand bags in his van aiming to increase the force of friction between the tires and the road. Will the sand bags help?



Increasing the friction force to increase *a* and *v*

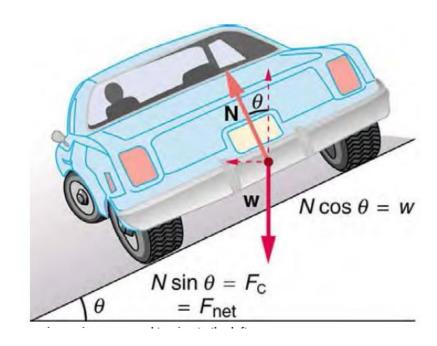
For a faster speed, will the sand bags help?

What will help?

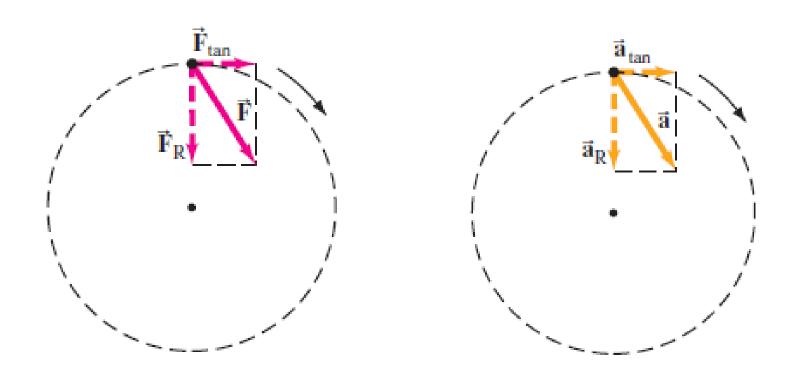
Banked curves

To reduce the chance of skidding, banking road

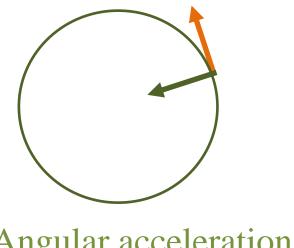
What's the ideal angle for *v*



Nonuniform Circular Motion



Nonuniform Circular Motion - Angular acceleration

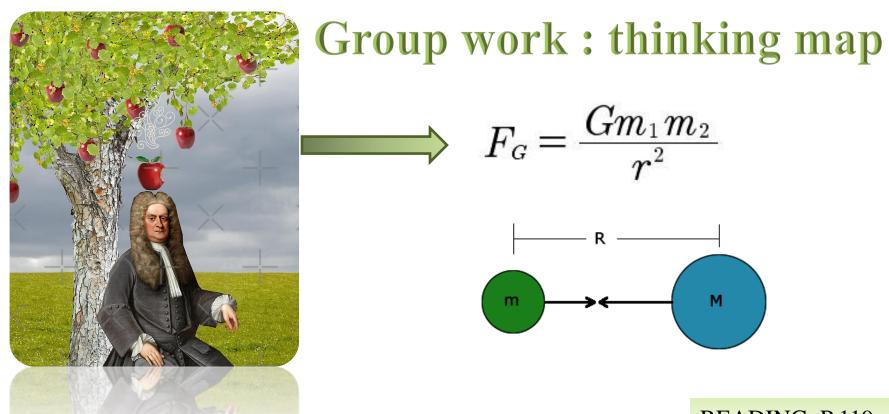


$$\vec{a} = \vec{a}_{tan} + \vec{a}_R$$

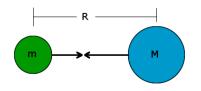
Angular acceleration

Tangential and centripetal (radial) acceleration

Newton's law of Universal Gravitation



Newton's law of Universal Gravitation



Every particle in the universe **attracts** every other particle with a force that is

proportional to the product of their masses and*inversely proportional* to the square of the distance between them.This force acts *along the line* joining the two particles.

$$F_{\scriptscriptstyle G} = rac{Gm_1m_2}{r^2}$$

$$G = 6.67 \times 10^{-11} \text{N m}^2/\text{kg}^2$$
Universal Gravitational Constant (m³/kg·s²)

Example

Estimate the gravitational force of attraction between two students sitting next to each other.

Gravitational constant: $G = 6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2 \cdot \text{kg}^{-2}$

Assume the mass of each student to be m = 60 kg ($m_1 = m_2$)

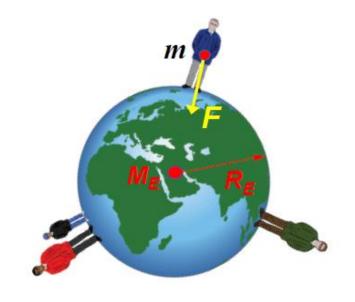
Assume two students are sitting 0.5 m apart in classroom

The gravity of the earth

For a body with *spherical symmetry*, such as the Earth, all the mass can be regarded to be at the *center of mass*, when calculating the gravitational force.

$$M_{\rm E} = 5.97 \times 10^{24} \, \rm kg$$

 $m = 60 \, \rm kg$
 $R_{\rm E} = 6.37 \times 10^{6} \, \rm m$
 $G = 6.67 \times 10^{-11} \, \rm m^{3} \cdot kg^{-1} \cdot s^{-2}$



$$F_{\scriptscriptstyle G} = rac{Gm_1m_2}{r^2}$$

$$F = 588.8N$$

Gravitational force of a uniform sphere on a particle

1. Shell theorem #1

A uniformly dense spherical shell attracts an external particle as if all the mass of the shell were concentrated at its center.

2. Shell theorem #2

A uniformly dense spherical shell exerts no gravitational force on a particle located anywhere inside it.

Gravity Near the Earth's Surface

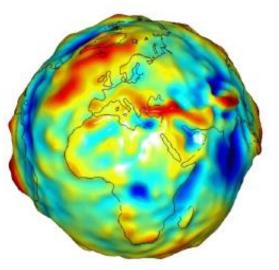
$$F_G = rac{Gm_1m_2}{r^2}$$

$$F_G = mg$$



$$g = G \frac{m_E}{r_E^2}$$

Map of Map of g



Gravity not exactly the same on the surface of Earth

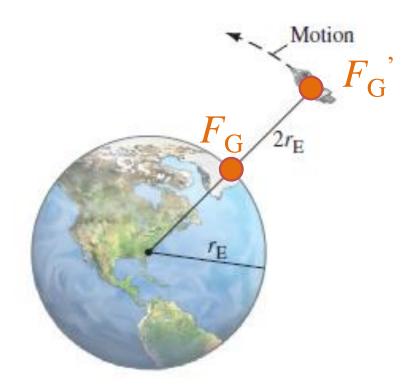
The gravity of the earth

For a body with *spherical symmetry*, such as the Earth, all the mass can be regarded to be at the *center of mass*, when calculating the gravitational force.

What is F_G for an object at the surface of the earth?

How about F_G for two Earth radii?

$$F_G = \frac{Gm_1m_2}{r^2}$$



Example

The mass of the Moon is 7.35×10^{22} kg and the radius of the Moon is 1740 km. What is \mathbf{g} on the surface of the Moon, \mathbf{g}_{Moon} ?

$$g_{Moon} = \frac{GM_{Moon}}{r_{Moon}^2} = \frac{6.67 \times 10^{-11} \times 7.35 \times 10^{22}}{\left(1740 \times 10^3\right)^2} = 1.62 \, \text{ms}^{-2}$$

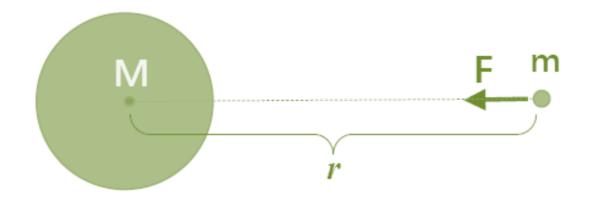
$$\frac{g_{Earth}}{g_{Moon}} = \frac{9.80}{1.62} = 6.0$$

- •How does your mass change on the moon?
- •How does your weight change on the moon?

Gravitational potential energy

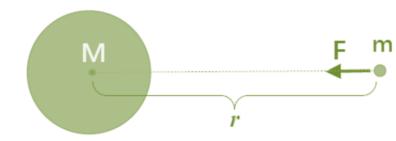
Suppose the Earth is stationary.

What is the potential energy of an object a distance *r* away from the center of Earth?



Gravitational potential energy

Let's move the object to



infinitely far away.

The work done by the gravitational force is:

$$W = \int_{r}^{\infty} \vec{F} \cdot d\vec{r} = -\int_{r}^{\infty} F \, dr = -\int_{r}^{\infty} G \frac{Mm}{r^2} dr$$

$$W = \frac{-GMm}{r}$$
 Negative work = increase of potential energy

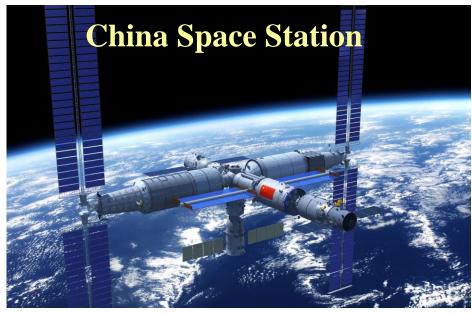
Gravity in orbit



$$r = 6751 \text{ km}$$

 $g = 8.75 \text{ ms } -2$

380 km above the ground



Satellites

Which needs a faster speed?

Close to the earth?

Far away from the earth?

r_A

Try to find speed and period.

near-Earth orbits geostationary orbit

Near-Earth orbits

Low near-Earth orbits are usually from 160 to 2000 km above earth.

With a very rough approximation, we assume **the orbit radius is the**radius of the Earth (which could be true for planets without atmosphere)

Orbital speed:
$$v = \sqrt{gr} = \sqrt{9.8 \times 6.37 \times 106} = 7910 \ m/s$$

The first cosmic speed

The orbital period is about 84 mins

Example

If a satellite is travelling at 7 km/s, what is the radius of its orbit?

Orbital Speed:
$$v = \sqrt{\frac{GM}{r}}$$
 $r = \frac{GM}{v^2}$
$$r = \frac{6.67 \times 10^{-11} \times 5.98 \times 10^{24}}{(7000)^2} = 8.14 \times 10^6 m \qquad \text{(1760 km above the Earth's surface)}$$

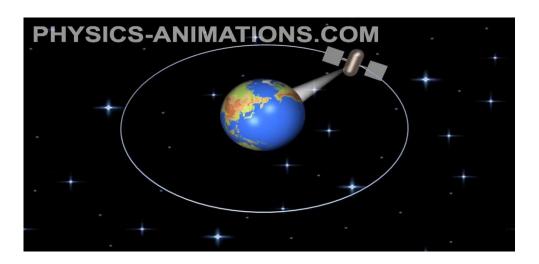
What is the total energy of the satellite?

$$E_k = \frac{1}{2}mv^2 \qquad E_p = -G\frac{Mm}{r} = -m\frac{GM}{r} = -mv^2$$

$$E_k + E_p = -\frac{1}{2}mv^2 = -\frac{GMm}{2r} \qquad \text{Negative and inversely proportional to the radius}$$

Geostationary satellite (Geosynchronous)

The orbit's period is equal to the Earth's rotation period



What is the height of this kind of satellite?

Orbit of the satellite



Weightlessness



Why a people in a satellite orbit close to the Earth will experience "weightlessness"?

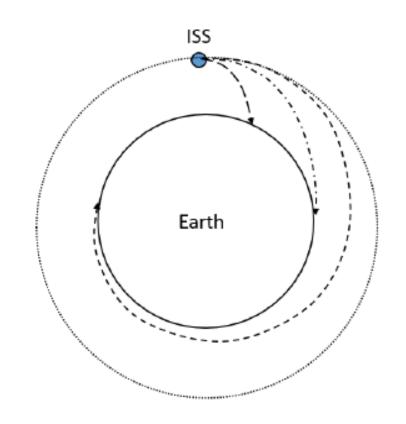
There is no gravity in space?

What is the net force on an astronaut at rest inside the space station?

Weightlessness

The station and astronauts are both in *free -fall*. But due to the horizontal speed, the projectile motion might be *a circular motion*.

Gravity provides the centripetal force to keep the space station orbiting around the Earth. For a uniform circular motion, the centripetal acceleration equals the gravitational acceleration.



Apparent weightlessness

A person on a scale in an elevator.

If the elevator fall freely, the *reading of the scale would be 0*

Apparent (not real)

- 1. Don't seem to have weight
- 2. Gravity doesn't disappear

Weightlessness

Apparent weightlessness?





