

# Calculating the actual wave

## The Problem

In the **Fundamental Solution** jupyter notebook we have analytically approximated the fundamental solution. Here we are going to use that solution in order to solve the following Cauchy Global Problem.

$$\begin{cases} \Delta \left( p(\vec{x}, t) - \frac{1}{\omega_0} \frac{\partial}{\partial t} p(\vec{x}, t) \right) - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} p(\vec{x}, t) = -\frac{\beta}{C_p} \frac{\partial E}{\partial t} & \forall (\vec{x}, t) \in \mathbb{R}^3 \times \mathbb{R}^+ \\ p(\vec{x}, 0) = 0 & \forall \vec{x} \in \mathbb{R}^3 \\ \frac{\partial}{\partial t} p(\vec{x}, 0) = 0 & \forall \vec{x} \in \mathbb{R}^3 \end{cases}$$

## Notation

This notation, even though it is the most accurate is pretentious and cumbersome to work with. Therefore, from now on we will employ a different notation that is cleaner and simpler to read and write.

First we define the linear differential operator  $\mathcal{L}$  as such.

$$\mathcal{L}u(\vec{x}, t) = \Delta \left( u(\vec{x}, t) - \frac{1}{\omega_0} \frac{\partial}{\partial t} u(\vec{x}, t) \right) - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} u(\vec{x}, t)$$

Now we can rewrite the cauchy problem like so:

$$\begin{cases} \mathcal{L}p(\vec{x}, t) = -\frac{\beta}{C_p} \frac{\partial E}{\partial t} = \psi(\vec{x}, t) & \forall (\vec{x}, t) \in \mathbb{R}^3 \times \mathbb{R}^+ \\ p(\vec{x}, 0) = 0 & \forall \vec{x} \in \mathbb{R}^3 \\ p_t(\vec{x}, 0) = 0 & \forall \vec{x} \in \mathbb{R}^3 \end{cases}$$

This is undoubtedly better. Now let's try to find the solution.

## Duhamel's Method

Duhamel's method is used in this case to simplify and solve the problem by solving the following set of problems: Let  $F(\vec{x}, t)$  be the *Fundamental Solution* solution defined by the following global cauchy problem:

\$\$ \begin{cases}

$$\begin{aligned} \mathcal{L}F(\vec{x}, t) &= \delta(\vec{x})\delta(t) \quad \& \quad \forall (\vec{x}, t) \in \mathbb{R}^3 \times \mathbb{R}^+ \\ F(\vec{x}, 0) &= 0 \quad \& \quad \forall \vec{x} \in \mathbb{R}^3 \\ F_t(\vec{x}, 0) &= 0 \quad \& \quad \forall \vec{x} \in \mathbb{R}^3 \end{aligned}$$

$\end{cases}$  \$\$

Now we can use it to find the solutions  $u(\vec{x}, t; s)$  to the following family of problems that arise from decomposing  $\psi(\vec{x}, t)$  in time. So  $\forall s > 0$ :

$$\begin{cases} \mathcal{L}u(\vec{x}, t; s) = 0 & \forall (\vec{x}, t) \in \mathbb{R}^3 \times (s, +\infty) \\ u(\vec{x}, s; s) = 0 & \forall \vec{x} \in \mathbb{R}^3 \\ u_t(\vec{x}, s; s) = \psi(\vec{x}, s) & \forall \vec{x} \in \mathbb{R}^3 \end{cases}$$

Therefore the solution to our original PDE will be the following convolution

$$p(\vec{x}, t) = \int_0^t \int_{\mathbb{R}^3} \psi(\vec{y}, s) F(\vec{x} - \vec{y}, t - s) d\vec{y} ds$$

## Forcing Term

Now we need to find  $\psi(\vec{x}, t)$ . To do this we need to think of the analytical form of the energy deposition of the particle in the fluid. Hence we introduce the Bethe-Bloch formula that will predict the spatial rate of energy deposition of the particle through the medium along the distance travelled from its path.

$$-\frac{dE}{dx} \approx -\left\langle \frac{dE}{dx} \right\rangle = \frac{4\pi}{m_e c^2} \frac{n z^2}{\beta^2} \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \left[ \ln \left( \frac{2m_e c^2 \beta^2}{I \cdot (1 - \beta^2)} \right) - \beta^2 \right]$$

Now what we can do, is to assume that this energy is deposition will be Gaussian distributed over time and space in the medium. As a result the energy deposition in the medium is:

$$-\frac{dE}{dx} \cdot G(\vec{x}, t)$$

where  $G(\vec{x}, t)$  is a distribution in  $\vec{x}$  and  $t$ :

$$G(\vec{x}, t) = \frac{1}{4\pi^2\sigma_x^3\sigma_t} \exp\left(-\frac{|\vec{x}|^2}{2\sigma_x^2}\right) \exp\left(-\frac{t^2}{2\sigma_t^2}\right)$$

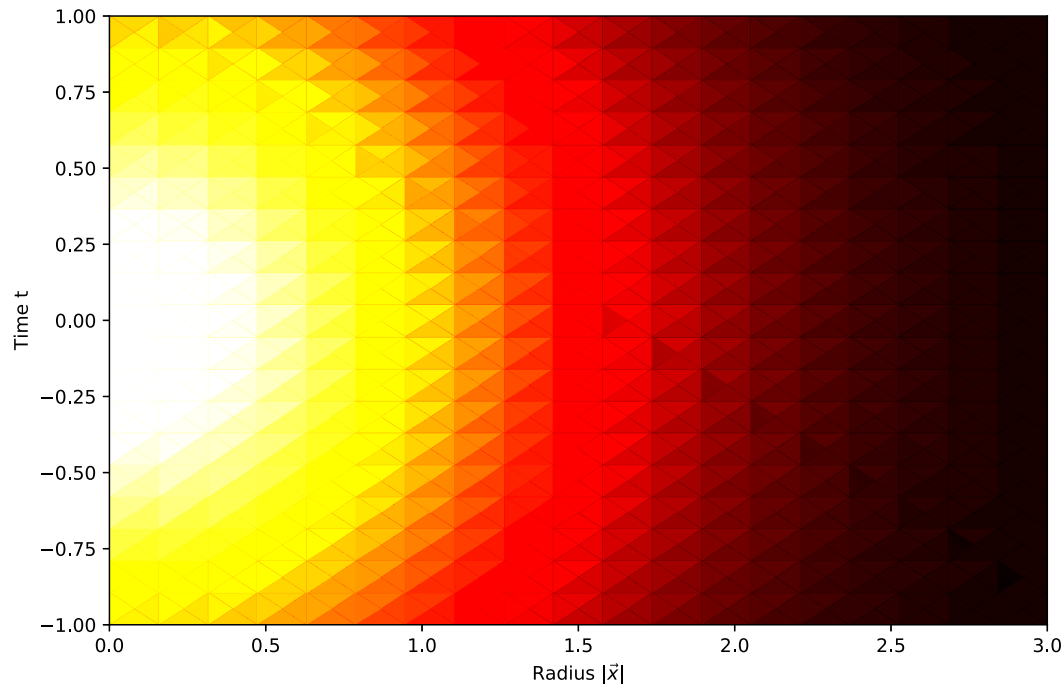
Therefore we can write the forcing function  $\psi(\vec{x}, t)$  like so.

$$\psi(\vec{x}, t) = -\frac{\beta_{Xe}}{C_p} \frac{dE}{dx} \frac{dx}{dt} G(\vec{x}, t)$$

Below we plot the energy envelope.

Out[2]: <matplotlib.collections.QuadMesh at 0x7f9583388710>

Distribution of heat in space and time



## Integrating

Now to do this integral we need to first express it analytically. This should be fun!

$$p(\vec{x}, t) = \int_0^t \int_{\mathbb{R}^3} -\frac{\beta_{Xe}}{C_p} \frac{4\pi}{m_e c^2} \frac{nz^2}{\beta^2} \left(\frac{e^2}{4\pi\epsilon_0}\right)^2 \left[ \ln\left(\frac{2m_e c^2 \beta^2}{I \cdot (1 - \beta^2)}\right) - \beta^2 \right] \frac{dz}{dt} \frac{1}{4\pi^2\sigma_x^3\sigma_t} e^{-\frac{|\vec{y}|^2}{2\sigma_x^2}} e^{-\frac{s^2}{2\sigma_t^2}} \Theta(t-s) \frac{8\pi^2 c}{|\vec{x} - \vec{y}|(t-s)} \sqrt{\frac{\pi\omega_0}{8c^2(t-s)}} \exp\left(-\frac{|\vec{x} - \vec{y}|^2 \omega_0}{c^2(t-s)} - (t-s)\omega_0\right) \left[ \exp\left(\frac{(|\vec{x} - \vec{y}| + c(t-s))^2 \omega_0}{2c^2(t-s)}\right) - \exp\left(\frac{(|\vec{x} - \vec{y}| - c(t-s))^2 \omega_0}{2c^2(t-s)}\right) \right] d\vec{y} ds$$

cool...

This keeps happening. But I think we can simplify the integral to a much better form. First of all, assuming the velocity of the particle to be constant we can rewrite everything like this.

$$p(\vec{x}, t) = \mathcal{A} \int_0^t \int_{\mathbb{R}^3} \frac{1}{|\vec{x} - \vec{y}|(t-s)^{\frac{3}{2}}} \exp\left(-\frac{|\vec{y}|^2}{2\sigma_x^2} - \frac{s^2}{2\sigma_t^2} - \frac{|\vec{x} - \vec{y}|^2 \omega_0}{c^2(t-s)} - (t-s)\omega_0\right) \left[ \exp\left(\frac{(|\vec{x} - \vec{y}| + c(t-s))^2 \omega_0}{2c^2(t-s)}\right) - \exp\left(\frac{(|\vec{x} - \vec{y}| - c(t-s))^2 \omega_0}{2c^2(t-s)}\right) \right] d\vec{y} ds$$