Write your name here Surname	Oth	ner names
Pearson Edexcel GCE	Centre Number	Candidate Number
Physics Advanced Unit 5: Physics from	om Creation to	Collapse
Thursday 19 June 2014 - Time: 1 hour 35 minute	•	Paper Reference 6PH05/01R

### **Instructions**

- Use black ink or ball-point pen.
- **Fill in the boxes** at the top of this page with your name, centre number and candidate number.
- Answer all questions.
- Answer the questions in the spaces provided
  - there may be more space than you need.

#### Information

- The total mark for this paper is 80.
- The marks for **each** question are shown in brackets
  - use this as a guide as to how much time to spend on each question.
- Questions labelled with an asterisk (\*) are ones where the quality of your written communication will be assessed
  - you should take particular care with your spelling, punctuation and grammar, as well as the clarity of expression, on these questions.
- The list of data, formulae and relationships is printed at the end of this booklet.
- Candidates may use a scientific calculator.

#### **Advice**

- Read each question carefully before you start to answer it.
- Keep an eye on the time.
- Try to answer every question.
- Check your answers if you have time at the end.

P 4 3 2 7 2 A 0 1 2 8

Turn over ▶



# **SECTION A**

# **Answer ALL questions.**

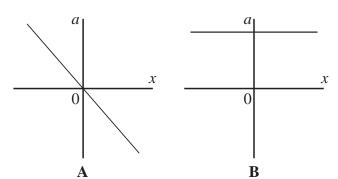
For questions 1–10, in Section A, select one answer from A to D and put a cross in the box  $\boxtimes$ . If you change your mind, put a line through the box  $\boxtimes$  and then mark your new answer with a cross  $\boxtimes$ .

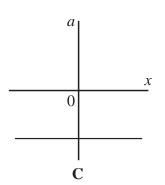
1	The av	verage kinetic energy of the molecules in an ideal gas is	
	⊠ A	directly proportional to the square root of the absolute temperature.	
	⊠ B	directly proportional to the absolute temperature.	
		independent of the absolute temperature.	
	<b>■ D</b>	inversely proportional to the absolute temperature.	
_		(Total for Question 1 = 1 mark)	
2	A sample of an ideal gas at 27 °C is placed in a sealed container. The gas is heated at constant volume to a temperature of 324 °C.  The ratio of the final pressure to the initial pressure exerted by the gas is approximately		
	⊠ A	1	
	⊠ B	2	
		4	
	<b>■ D</b>	12	
_		(Total for Question 2 = 1 mark)	
3	If an o	oscillating system is completely undamped, the system	
	A	exhibits simple harmonic motion.	
	<b>B</b> B	is said to be a free oscillation.	
		obeys Hooke's law.	
	<b>■ D</b>	oscillates indefinitely.	
_		(Total for Question 3 = 1 mark)	

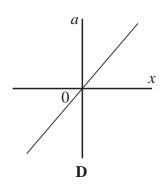
4		a in homeing wanting the faces a series. The series	4 into our oll our mliter il4! - 1
	A mas oscilla	s is hanging vertically from a spring. The mass is sections.	ei into smaii amplitude vertical
	The to	tal energy of the undamped oscillating system is	
	$\mathbf{X}$ A	a maximum at an extreme position of the mass.	
	$\boxtimes$ B	a maximum at the mean position of the mass.	
	<b>区</b> C	a minimum at the mean position of the mass.	
	<b>■</b> D	the same at all positions of the mass.	
			(Total for Question 4 = 1 mark
5		surface temperature of the Sun were to double, the rareceived on the Earth would increase by a factor of	ate at which energy from the
	$\mathbf{X}$ A	2	
	$\boxtimes$ B	4	
	<b>区</b> C	8	
	$\square$ D	16	
			(Total for Question 5 = 1 mark
6	are sho	light from the galaxy in Andromeda is analysed, it is orter than expected. ells us that the galaxy is	s found that the wavelengths
	$\boxtimes$ A	moving towards us.	
	$\blacksquare$ B	moving away from us.	
	<b>区</b> C	a very distant galaxy.	
	$\boxtimes$ <b>D</b>	rotating on an axis.	
			(Total for Question 6 = 1 mark

7			
,			at 100 °C turns into steam at 100 °C. of the following statements is true?
	X	A	The internal energy is unchanged, but the kinetic energy of the molecules increases.
	X	В	The internal energy is unchanged, but the potential energy of the molecules increases.
	×	C	Both the internal energy and the kinetic energy of the molecules increase
	X	D	Both the internal energy and the potential energy of the molecules increase
			(Total for Question 7 = 1 mark)
8			unit for mass is the kilogram. However, particle physicists often use the tive unit
	×	A	MeV
	X	В	MeV/c
	×	C	$MeV/c^2$
	X	D	$MeV^2/c^2$
			(Total for Question 8 = 1 mark)
9		nich	el used in a nuclear fission reactor is uranium. of the following is required for fission to proceed?  Neutrons must be removed from the reactor core.
	X	В	The reactor core must be very hot.
	X	C	The uranium nuclei must absorb neutrons.
	X	D	The uranium nuclei must absorb protons.
			(Total for Question 9 = 1 mark)

**10** An object is undergoing simple harmonic motion. Which graph shows how the acceleration *a* varies with displacement *x* from the equilibrium position?







- $\mathbf{X}$   $\mathbf{A}$
- $\mathbf{B}$
- $\square$  C
- $\mathbf{X}$  **D**

(Total for Question 10 = 1 mark)

**TOTAL FOR SECTION A = 10 MARKS** 

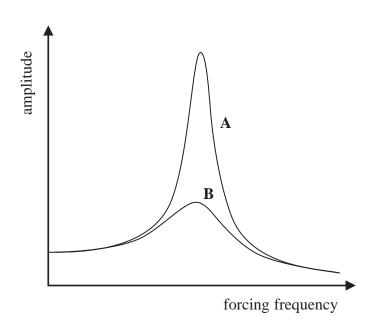
### **SECTION B**

# Answer ALL questions in the spaces provided.

\*11 A student uses the apparatus shown below to investigate the behaviour of a mass-spring system when it is forced into oscillation.



The graph shows how the amplitude of the oscillating mass varies over a range of forcing frequencies.



Curve A shows the results of the investigation using the apparatus as shown.

The student repeats the investigation with the oscillating mass in a beaker of water. Curve B shows these results.

ults.		(4)
	(Total for Question	11 = 4 marks)



a) (i) Describe the process of nuclear fusion.	
ty (1) Describe the process of national rusion.	(2)
(ii) Explain why it is difficult to maintain the conditions needed for nuclear	fusion in
a reactor.	
	(2)
b) Explain why the fusion of hydrogen nuclei should release energy.	
	(2)
(Total for Question 1	2 = 6 marks)

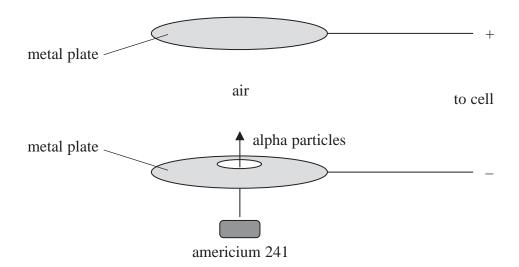
*13 Describe the similarities and differences between electric and gravitational fields.	(6)
(Total for Question 13 = 6 I	marks)



14	Some types of smoke detector contain a radioactive isotope of americium, <sup>241</sup> Am.	The
	nuclei of <sup>241</sup> Am decay by emitting an alpha particle.	

The diagram shows part of a smoke detector.

The detectors use a small amount of <sup>241</sup>Am to make the air between two metal plates conduct charge.





(ii) Suggest how smoke particles entering the space between the plates will cause the current to decrease.

**(1)** 

(b) (i) The decay of <sup>241</sup> Am is said to be random and spontaneous.  State what is meant by random and spontaneous.  Random	(2)
Spontaneous	
(ii) Complete the equation for the decay of <sup>241</sup> Am.	(2)
	arks)



(a) Su	ggest why a pre-warmed teapot may allow more flavour to be extracted.	
(a) Su	ggest why a pre-warmed teapor may allow more havour to be extracted.	(1)
(b) (i)	0.26 kg of water at 95 °C is added to a stainless steel teapot. In a very short time	
	the teapot and water both reach a temperature of 81 °C.	
	Show that the energy transferred from the water is about 15 kJ.	
	specific heat capacity of water = $4200 \text{ J kg}^{-1} \text{ K}^{-1}$	(2)
	Calculate the specific heat capacity of stainless steel, stating any assumption you make.  mass of teapot = 0.43 kg initial temperature of teapot = 22 °C	(3)
	Specific heat capacity =	J kg <sup>-1</sup> K

difference.	(2)
	(Total for Question 15 = 8 marks)



16	Radioactive isotopes are often used as markers, so that chemical substances can be traced around the body. In one medical procedure tritium is used as a means of studying protein absorption by the intestine.	
	A patient was given a sample containing the tritium to drink and then monitored. The initial activity of the sample was 3450 Bq.	
	Tritium is a beta-emitter with a half-life of $3.89 \times 10^8$ s.	
	(a) State what is meant by the activity of a radioactive source.	(1)
	(b) Show that the decay constant of the tritium is about $1.8 \times 10^{-9}  \mathrm{s^{-1}}$ and hence calculate the number of tritium nuclei in the initial sample.	(3)
	Number of nuclei =	

	(3)
(ii) Comment on the time given in (c) (i).	
(ii) Comment on the time given in (c) (i):	(1)
	(T. 10. 0. 1. 16. 0. 1.)
	(Total for Question 16 = 8 marks)

- 17 The first satellite weather picture was taken in 1960. Today more than 200 weather satellites are in use. Some of these satellites are in a geostationary orbit around the Earth, so that they remain at the same point above the Earth's surface all the time.
  - (a) (i) Show that the magnitude of the gravitational field strength g at a point outside of the Earth is given by

$$g = \frac{GM}{r^2}$$

where r is the distance of the point from the centre of the Earth and M is the mass of the Earth.

(2)

(ii) Use this expression together with an expression for the centripetal acceleration to show that the radius of a satellite's orbit is given by

$$r^3 = \frac{GMT^2}{4\pi^2}$$

where *T* is the time for one orbit of the satellite.

(3)

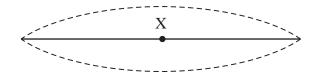
(iii) Hence calculate a value for the radius of the geostationary orbit.	
$M = 6.0 \times 10^{24} \text{ kg}$	
	(3)
Radius =	
(b) State why all geostationary satellites are in an orbit above the Earth's equator.	(1)
(Total for Question 17 = 9	) marks)
(Total for Question 17 = 9	marks)
(Total for Question 17 = 9	) marks)
(Total for Question 17 = 9	) marks)



18 Guitar strings can oscillate with simple harmonic motion.



Shortly after the string is plucked, a standing wave exists on the string. The simplified diagram below shows a string in three positions of the standing wave.



(a) State what is meant by simple harmonic motion.	(2)
(b) (i) Describe the acceleration of point X on the string as it moves between the extreme positions of its motion.	(2)

positions of its motion.	(3)
c) The oscillating string has a length of 0.53 m. Calculate emitted when the string oscillates as shown previously. speed of the wave on the string = $270 \text{ m s}^{-1}$	the frequency of the sound (3)
emitted when the string oscillates as shown previously.	
emitted when the string oscillates as shown previously.	
emitted when the string oscillates as shown previously.	

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The H-R diag	grams below are for a young star of	cluster and an old star cluster.
luminosity		luminosity
L	temperature	<b>←</b> temperati
	Young star cluster	Old star cluster



(b) Trigonometric parallax is one way in which stellar distances can be measured.

Astronomers measure the parallax angle for two nearby stars. The parallax angle for star A is  $3.74 \times 10^{-6}$  rad and that for star B is  $1.84 \times 10^{-7}$  rad.

(i) Without calculation, state what can be deduced from this data about the relative distances of the two stars.

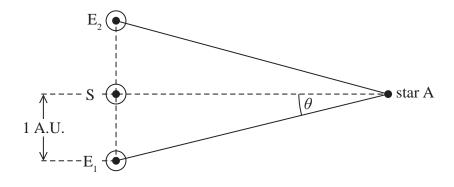
(1)

(ii) The diagram shows the parallax angle for star A.

Calculate the distance of star A from the Earth.

1 A.U. is  $1.50 \times 10^{11}$  m

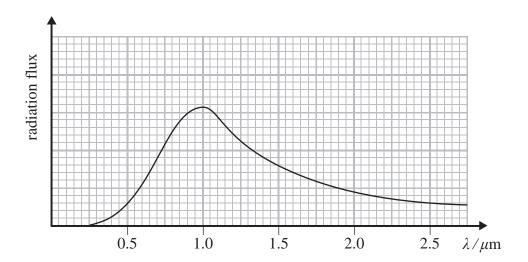
**(2)** 



Distance =

(c) In addition to finding the distances to stars astronomers are interested in determining the temperatures of stars.

The spectrum of star A is shown below.



Use data from the graph to determine the surface temperature of star A.

**(3)** 

Temperature =

(Total for Question 19 = 12 marks)

TOTAL FOR SECTION B = 70 MARKS

TOTAL FOR PAPER = 80 MARKS



## List of data, formulae and relationships

Acceleration of free fall  $g = 9.81 \text{ m s}^{-2}$  (close to Earth's surface)

Boltzmann constant  $k = 1.38 \times 10^{-23} \text{ J K}^{-1}$ 

Coulomb's law constant  $k = 1/4\pi\varepsilon_0$ 

 $= 8.99 \times 10^9 \text{ N m}^2 \text{ C}^{-2}$ 

Electron charge  $e = -1.60 \times 10^{-19} \text{ C}$ 

Electron mass  $m_{\rm e} = 9.11 \times 10^{-31} \,\mathrm{kg}$ 

Electronvolt  $1 \text{ eV} = 1.60 \times 10^{-19} \text{ J}$ 

Gravitational constant  $G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$ 

Gravitational field strength  $g = 9.81 \text{ N kg}^{-1}$  (close to Earth's surface)

Permittivity of free space  $\varepsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$ 

Planck constant  $h = 6.63 \times 10^{-34} \,\mathrm{J s}$ 

Proton mass  $m_p = 1.67 \times 10^{-27} \text{ kg}$ 

Speed of light in a vacuum  $c = 3.00 \times 10^8 \, \mathrm{m \ s^{-1}}$ 

Stefan-Boltzmann constant  $\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$ 

Unified atomic mass unit  $u = 1.66 \times 10^{-27} \text{ kg}$ 

### Unit 1

#### **Mechanics**

Kinematic equations of motion v = u + at

 $s = ut + \frac{1}{2}at^2$  $v^2 = u^2 + 2as$ 

Forces  $\Sigma F = ma$ 

g = F/mW = mg

Work and energy  $\Delta W = F \Delta s$ 

 $E_{k} = \frac{1}{2}mv^{2}$  $\Delta E_{\text{grav}} = mg\Delta h$ 

Materials

Stokes' law  $F = 6\pi \eta r v$ 

Hooke's law  $F = k\Delta x$ 

Density  $\rho = m/V$ 

Pressure p = F/A

Young modulus  $E = \sigma/\varepsilon$  where

Stress  $\sigma = F/A$ Strain  $\varepsilon = \Delta x/x$ 

Elastic strain energy  $E_{al} = \frac{1}{2}F\Delta x$ 

### Unit 2

Waves

Wave speed  $v = f\lambda$ 

Refractive index  $_{1}\mu_{2} = \sin i / \sin r = v_{1}/v_{2}$ 

Electricity

Potential difference V = W/Q

Resistance R = V/I

Electrical power, energy and P = VI efficiency  $P = I^2R$ 

 $P = I^{2}R$   $P = V^{2}/R$  W = VIt

% efficiency =  $\frac{\text{useful energy output}}{\text{total energy input}} \times 100$ 

% efficiency =  $\frac{\text{useful power output}}{\text{total power input}} \times 100$ 

Resistivity  $R = \rho l/A$ 

Current  $I = \Delta Q/\Delta t$ 

I = nqvA

Resistors in series  $R = R_1 + R_2 + R_3$ 

Resistors in parallel  $\frac{1}{R} = \frac{1}{R_1} + \frac{1}{R_2} + \frac{1}{R_3}$ 

Quantum physics

Photon model E = hf

Einstein's photoelectric  $hf = \phi + \frac{1}{2}mv_{\text{max}}^2$ 

equation

### Unit 4

#### Mechanics

Momentum p = mv

Kinetic energy of a

non-relativistic particle  $E_k = p^2/2m$ 

Motion in a circle  $v = \omega r$ 

 $T = 2\pi/\omega$ 

 $F = ma = mv^2/r$ 

 $a = v^2/r$ 

 $a = r\omega^2$ 

### Fields

Coulomb's law  $F = kQ_1Q_1/r^2$  where  $k = 1/4\pi\epsilon_0$ 

Electric field E = F/Q

 $E = kQ/r^2$ 

E = V/d

Capacitance C = Q/V

Energy stored in capacitor  $W = \frac{1}{2}QV$ 

Capacitor discharge  $Q = Q_0 e^{-t/RC}$ 

In a magnetic field  $F = BIl \sin \theta$ 

 $F = Bqv \sin \theta$ 

r = p/BQ

Faraday's and Lenz's Laws  $\varepsilon = -d(N\phi)/dt$ 

# Particle physics

Mass-energy  $\Delta E = c^2 \Delta m$ 

de Broglie wavelength  $\lambda = h/p$ 

### Unit 5

Energy and matter

Heating  $\Delta E = mc\Delta\theta$ 

Molecular kinetic theory  $\frac{1}{2}m\langle c^2\rangle = \frac{3}{2}kT$ 

Ideal gas equation pV = NkT

Nuclear Physics

Radioactive decay  $dN/dt = -\lambda N$ 

 $\lambda = \ln 2/t_{1/2}$ 

 $N = N_0 e^{-\lambda t}$ 

Mechanics

Simple harmonic motion  $a = -\omega^2 x$ 

 $a = -A\omega^2 \cos \omega t$   $v = -A\omega \sin \omega t$   $x = A \cos \omega t$   $T = 1/f = 2\pi/\omega$ 

Gravitational force  $F = Gm_1m_2/r^2$ 

Observing the universe

Radiant energy flux  $F = L/4\pi d^2$ 

Stefan-Boltzmann law  $L = \sigma T^4 A$ 

 $L=4\pi r^2\sigma T^4$ 

Wien's Law  $\lambda_{\max} T = 2.898 \times 10^{-3} \text{ m K}$ 

Redshift of electromagnetic

radiation  $z = \Delta \lambda / \lambda \approx \Delta f / f \approx v / c$ 

Cosmological expansion  $v = H_0 d$ 

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