Table 1: Sources included in our final sample, relevant physical parameters, and number of spectra of sufficient quality for our analysis.

| Name | mass ($M_\odot$) | distance (kpc) | inclination (°) | dips | $P_{orb}$ (h) | absorption lines reported in the iron band |
|---|---|---|---|---|---|---|
| 1E 1740.7-2942 | 8 | 8 | $31^{+29\,R}_{-18}$ (1.) | X | $303^{+1}_{-2}$(2.) | X |
| 4U 1543-475$^D$ (3.) | $8.4 \pm 1$(4.) | $7.5 \pm 0.5$ (5.) | $20.7 \pm 1.5^D$ (4.)/$67^{+7\,R}_{-8}$ (1.) | ✓ (6.) | 26.8(4.) | ✓ |
| 4U 1630-47 | 8 | $8.1 \pm 3.4$(7.) | $[60-75]^D$ (8.)/$55^{+8\;R}_{-11}$ (1.) | ✓ (8.) | X | ✓ |
| 4U 1755-338 | 8 | $6.5 \pm 2.5$(10.) | / | X | X | X |
| 4U 1957+115 | $3^{+2.5}_{-1}$ (9.) | 8 | $\sim 13^D$ (9.)/$52^{+12\,R}_{-13}$ (1.) | X | 9.33(9.) | X |
| A0620-00$^D$ (13.) | $6.6 \pm 0.3$(13.) | $1.06 \pm 0.1$(13.) | $52.6 \pm 2.5^D$ (14.) | ✓ (11.) | 7.75(12.) | X |
| A 1524-61$^D$ (15.) | $5.8^{+3}_{-2.4}$(15.) | $8 \pm 0.9$(15.) | $57 \pm 13^D$ (15.) | X | $6.2 \pm 2$(15.) | X |
| 1A 1742-289 | 8 | 8 | / | X(16.) | X(16.) | X |
| AT 2019wey | 8 | 8 | $14^{+12\,R}_{-10}$ (1.) | X | X | X |
| CXOGC J174540.0-290031 | 8 | 8 | X | ✓ (17.) | 7.8(17.) | X |
| EXO 1846-031 | 8 | $\sim 7$(18.) | $40 \pm 3^R$ (19.)/$62^{+9\;R}_{-10}$ (1.) | X | X | ✓ |
| GRO J0422+32 (21.) | $2.7^{+0.7}_{-0.5}$(14.) | $2.5 \pm 0.3$(21.) | $56 \pm 4^D$ (14.) | X | 5.09 (21.) | X |
| GRO J1655-40$^D$ (22.) | $5.4 \pm 0.3$(23.) | $3.2 \pm 0.2$(24.) | $69 \pm 2^D$ (23.) | ✓ (25.) | 62.9(26.) | ✓ |
| GRS 1009-45$^D$ (29.) | 8 | $3.8 + 0.3$(27.) | $59 \pm 22^D$ (28.) | X | 6.85 (29.) | X |
| GRS 1716-249 | $6.4^{+3.2}_{-2}$ (30.) | $6.9 \pm 1.1$(30.) | $61 \pm 15^D$ (30.)/$59^{+7\;R}_{-12}$ (1.) | X | 6.67(30.) | X |
| GRS 1739-278 | 8 | $7.3 \pm 1.3$(31.) | $70^{+5\;R}_{-11}$ (1.) | X | X | X |
| GRS 1730-312 | 8 | 8 | / | X | X | X |
| GRS 1737-31 | 8 | 8 | / | X | X | X |
| GRS 1758-258 | 8 | 8 | $67^{+8\;R}_{-13}$ (1.) | X | X | ✓ |
| GRS 1915+105$^D$ (32.) | $11.2^{+2}_{-1.8}$(33.) | $9.4^{+1.6}_{-1.6}$(33.) | $64 \pm 4^D$ (33.)/ $60 \pm 8^R$ (1.) | ✓ (35.) | $812 \pm 4$(34.) | ✓ |
| GS 1354-64$^D$ (36.) | 8 | $\sim 25$(36.) | $< 79^D$ (36.)/ $47^{+11\,R}_{-10}$ (1.) | X | 61.1(36.) | X |
| GS 1734-275 | 8 | 8 | / | X | X | X |
| GS 2000+251$^D$ (37.) | $7.2 \pm 1.7$(37.) | $2.7 \pm 0.7$(5.) | $68 \pm 6^D$ (14.) | X | 8.26 (37.) | X |
| GX 339-4$^D$ (38.) | $5.9 \pm 3.6$(38.) | 8 | $[37-78]^D$ (38.)/$49 \pm 14^R$ (1.) | X | 42.2(38.) | X |
| H 1705-250$^D$ (41.) | $5.4 \pm 1.5$ (40.) | $8.6 \pm 2.1$ (5.) | $64 \pm 16^D$ (39.) (41.) | X | 12.51 (41.) | X |
| H 1743-322 | 8 | $8.5 \pm 0.8$(42.) | $75 \pm 3^J$ (42.)/$54^{+12\,R}_{-13}$ (1.) | ✓ (43.) | X | ✓ |
| IGR J17091-3624 | 8 | 8 | $\sim 70^H$ (44.) (45.)/$47^{+10\,R}_{-11}$ (1.) | ✓ (35.) | X | ✓ |

Table 1: Continued.

| Name | mass ($M_\odot$) | distance (kpc) | inclination (°) | dips | $P_{orb}$ (h) | absorption lines reported in the iron band |
|---|---|---|---|---|---|---|
| IGR J17098-3628 | 8 | $\sim 10.5$ (47.) | / | X | X | X |
| IGR J17285-2922 | 8 | 8 | / | X | X | X |
| IGR J17451-3022 | 8 | 8 | $> 70^{D}$ (48.) | ✓ (48.) | 6.3 (49.) | ✓ |
| IGR J17454-2919 | 8 | 8 | $54^{+15\,R}_{-14}$ (1.) | X | X | X |
| IGR J17497-2821 | 8 | 8 | / | X | X | X |
| IGR J18175-1530 | 8 | 8 | / | X | X | X |
| IGR J18539+0727 | 8 | 8 | / | X | X | X |
| MAXI J0637-430 | 8 | 8 | $63^{+9\ R}_{-10}$ (1.) | X | $\approx 2.2 \pm 1$ (50.) | X |
| MAXI J1305-704$^{D}$ (51.) | $8.9^{+1.6}_{-1.}$ (51.) | $7.5^{+1.8}_{-1.4}$ (51.) | $72^{+5\,D}_{-8}$ (51.) | ✓ (52.) | $9.5 \pm 0.1$ (51.) | ✓ |
| MAXI J1348-630 | 8 | $3.4^{+0.4}_{-0.4}$ (53.) | $28 \pm 3^{J}$ (54.)/$65 \pm 7$ (55.)/ $52^{+8\ R}_{-11}$ (1.) | X | X | ✓ |
| MAXI J1535-571 | 8 | $4.1^{+0.6}_{-0.5}$ (56.) | $\leq 45^{J}$ (57.)/ $44^{+17\,R}_{-19}$ (1.) | X | X | ✓ |
| MAXI J1543-564 | 8 | 8 | / | X | X | X |
| MAXI J1631-479 | 8 | 8 | $22^{+10\,R}_{-12}$ (1.) | X | X | ✓ |
| MAXI J1659-152 | 8 | $8.6 \pm 3.7$ (58.) | $70 \pm 10^{D}$ (58.) | ✓ (58.) | 2.4 (59.) | X |
| MAXI J1727-203 | 8 | 8 | $65^{+11\,R}_{-14}$ (1.) | X | X | X |
| MAXI J1803-298 | 8 | 8 | $\sim 67 \pm 8^{D}$ (60.)/$72^{+6\,R}_{-9}$ (1.) | ✓ (60.) | $7 \pm 0.2$ (60.) | ✓ |
| MAXI J1810-222 | 8 | 8 | / | X | X | X |
| MAXI J1813-095 | 8 | 8 | $42^{+11\,R}_{-13}$ (1.) | X | X | X |
| MAXI J1820+070 $^{D}$ (61.) | $6.9 \pm 1.2$ (62.) | $2.96 \pm 0.33$ (63.) | $74 \pm 7^{D}$ (62.)/$64^{+8\,R}_{-9}$ (1.) | ✓ (64.) | 16.5 (61.) | ✓ |
| MAXI J1828-249 | 8 | 8 | / | X | X | X |
| MAXI J1836-194 | 8 | $7 \pm 3$ (65.) | $9^{+6\,D}_{-5}$ (65.) | X | X | X |
| MAXI J1848-015 | 8 | $3.4 \pm 0.3$ (66.) | $77 \pm 2^{D}$ (66.)/$29^{+13\,R}_{-10}$ (1.) | X | X | X |
| Nova Muscae 1991$^{D}$ (67.) | $11^{+2.1}_{-1.4}$ (67.) | $5 \pm 0.7$ (67.) | $43^{+2\,D}_{-3}$ (67.) | X | 10.4 (68.) | X |
| SAX J1711.6-3808 | 8 | 8 | / | X | X | X |
| SLX 1746-331 | 8 | 8 | / | X | X | X |
| Swift J1357.2-0933 | $11.6^{+2.5}_{-1.9}$ (14.) | 8 | $81^{+9\ D}_{-12}$ (14.) | ✓ (69.) | $2.8 \pm 0.3$ (70.) | X |
| Swift J151857.0-572147 | 8 | 8 | / | X | X | ✓ |

Table 1: Continued.

| Name | mass ($M_\odot$) | distance (kpc) | inclination (°) | dips | $P_{orb}$ (h) | absorption lines reported in the iron band |
|---|---|---|---|---|---|---|
| Swift J1539.2-6227 | 8 | 8 | / | X | X | X |
| Swift J1658.2-4242 | 8 | 8 | $50^{+9}_{-10}{}^{R}$ [1.] | ✓ [71.] | X | ✓ |
| Swift J1713.4-4219 | 8 | 8 | / | X | X | X |
| Swift J1727.8-1613 | 8 | $2.7 \pm 0.3$ [72.] | / | X | $7.6 \pm 0.2$ [72.] | X |
| Swift J1728.9-3613 | 8 | $8.4 \pm 0.8$ [73.] | $7^{+8}_{-3}{}^{R}$ [1.] | X | X | X |
| Swift J174510.8-262411 | 8 | $\sim 3.7 \pm 1.1$ [74.] | / | X | $< 11.3$ [74.] | X |
| Swift J174540.2-290005 | 8 | 8 | / | X | X | X |
| Swift J174540.2-290037 (T37) | 8 | 8 | $31^{+8}_{-9}{}^{R}$ [1.] | X | X | X |
| Swift J174540.7-290015 (T15) | 8 | 8 | $63^{+10}_{-8}{}^{R}$ [1.] | X | X | ✓ |
| Swift J1753.5-0127 | 8 | $5.6^{+1.8}_{-2.8}$ [75.] | $73 \pm 8^{R}$ [1.] | X | 3.2 [77.] | X |
| Swift J1753.7-2544 | 8 | 8 | / | X | X | X |
| Swift J1842.5-1124 | 8 | 8 | / | X | X | X |
| Swift J1910.2-0546 | 8 | 8 | / | X | $2.4 \pm 0.1$ [78.] | X |
| V404 Cyg$^{D}$ [79.] | $9^{+0.2}_{-0.6}$ [79.] | $2.4 \pm 0.2$ [80.] | $67^{+3}_{-1}{}^{D}$ [79.] / $37^{+9}_{-8}{}^{R}$ [1.] | X | 155.3 [81.] | ✓ |
| V4641 Sgr$^{D}$ [82.] | $6.4 \pm 0.6$ [83.] | $6.2 \pm 0.7$ [83.] | $72 \pm 4^{D}$ [83.] / $66^{+7}_{-11}{}^{R}$ [1.] | X | 67.6 [82.] | X |
| XMMSL1 J171900.4-353217 | 8 | 8 | / | X | X | X |
| XTE J1118+480$^{D}$ [86.] | $7.1 \pm 0.1$ [86.] | $1.7 \pm 0.1$ [84.] | $72 \pm 2^{D}$ [14.] | X | 4.1 [85.] | X |
| XTE J1550-564$^{D}$ [87.] | $11.7 \pm 3.9$ [87.] | $4.4^{+0.6}_{-0.4}$ [87.] | $75 \pm 4^{D}$ [87.] / $40 \pm 10^{R}$ [88.] | X | 37.0 [87.] | ✓ |
| XTE J1637-498 | 8 | 8 | / | X | X | X |
| XTE J1650-500$^{D}$ [89.] | 8 | $2.6 \pm 0.7$ [90.] | $\geq 47^{D}$ [89.] | X | 7.7 [89.] | X |
| XTE J1652-453 | 8 | 8 | $\leq 32^{R}$ [91.] | X | X | ✓ |
| XTE J1720-318 | 8 | $6.5 \pm 3.5$ [92.] | / | X | X | X |
| XTE J1719-291 | 8 | 8 | / | X | X | X |
| XTE J1726-476 | 8 | 8 | / | X | X | X |
| XTE J1748-288 | 8 | 8 | / | X | X | X |
| XTE J1752-223 | 8 | $\sim 6 \pm 2$ [93.] | $< 49^{J}$ [94.] / $35 \pm 4^{R}$ [95.] | X | $\lesssim 7$ [93.] | X |
| XTE J1755-324 | 8 | 8 | / | X | X | X |

Table 1: Continued.

| Name | mass ($M_\odot$) | distance (kpc) | inclination (°) | dips | $P_{orb}$ (h) | absorption lines reported in the iron band |
|---|---|---|---|---|---|---|
| XTE J1817-330 | 8 | $5.5 \pm 4.5_{(96.)}$ | / | ✓$_{(97.)}$ | X | X |
| XTE J1818-245 | 8 | $3.6 \pm 0.8_{(98.)}$ | / | X | X | X |
| XTE J1856+053 | 8 | 8 | / | X | X | X |
| XTE J1859+226 $^{D}{}_{(100.)}$ | $8 \pm 2_{(99.)}$ | $12.5 \pm 1.5_{(100.)}$ | $67 \pm 4^{D}{}_{(99.)}/71 \pm 1^{R}{}_{(101.)}$ | ✓$_{(102.)(103.)}$ | 6.6 $_{(99.)}$ | X |
| XTE J1901+014 | 8 | 8 | / | X | X | X |
| XTE J1908+094 | 8 | $6.5 \pm 3.5_{(104.)}$ | $28 \pm 11^{R}{}_{(1.)}$ | X | X | X |
| XTE J2012+381 | 8 | 8 | $46 \pm 4^{R}{}_{(105.)}/68^{+6}_{-11}{}^{R}{}_{(1.)}$ | X | X | X |

**Notes:** The letter $D$ in the object name column identifies dynamically confirmed BHs. A fiducial mass of 8 $M_\odot$ and distance of 8 kpc are used when not reliably known, including when dynamical constraints are only lower limits, according to the properties of the bulk of the Galactic BHLMXB population (see e.g. Corral-Santana et al. 2016). For inclination measurements letters $D$, $J$, $H$, $R$ refer respectively to dynamical inclination measurements (dips/eclipses/modulations), jets, heartbeats, and reflection fits. Sources analyzed in (Parra et al. 2024) are highlighted in blue. Details and references for line detection reports are provided in Tab. 2.

**References:** 1 (Draghis et al. 2024) 2 (Stecchini et al. 2017) 3 (Orosz et al. 1998) 4 (Orosz 2003) 5 (Jonker & Nelemans 2004) 6 (Park et al. 2004) 7 (Kalemci et al. 2018) 8 (Tomsick et al. 1998) 9 (Gomez et al. 2015) 10 (Angelini & White 2003) 11 (Haswell et al. 1993) 12 (Hernández et al. 2013) 13 (Cantrell et al. 2010) 14 (Casares et al. 2022) 15 (Yanes-Rizo et al. 2024) 16 (Kennea et al. 1996) 17 (Porquet et al. 2005) 18 (Parmar et al. 1993) 19 (Wang et al. 2020) 20 (Gelino & Harrison 2003) 21 (Webb et al. 2000) 22 (Van Der Hooft et al. 1998) 23 (Beer & Podsiadlowski 2002) 24 (Hjellming & Rupen 1995) 25 (Kuulkers et al. 1998) 26 (Petretti et al. 2023) 27 (Gelino & M. 2002) 28 (Shahbaz et al. 1996) 29 (Filippenko et al. 1999) 30 (Casares et al. 2023) 31 (Greiner et al. 1996) 32 (Reid et al. 2014) 33 (Reid & Miller-Jones 2023) 34 (Steeghs et al. 2013) 35 (Pahari et al. 2013) 36 (Casares et al. 2009) 37 (Ioannou et al. 2004) 38 (Heida et al. 2017) 39 (Martin et al. 1995) 40 (Harlaftis et al. 1997) 41 (Remillard et al. 1996) 42 (Steiner et al. 2012) 43 (Miller et al. 2006b) 44 (Capitanio et al. 2012) 45 (Rao & Vadawale 2012) 46 (Xu et al. 2017) 47 (Grebenev et al. 2006) 48 (Jaisawal et al. 2015) 49 (Bozzo et al. 2016) 50 (Soria et al. 2022) 51 (Sánchez et al. 2021) 52 (Shidatsu et al. 2013) 53 (Lamer et al. 2021) 54 (Carotenuto et al. 2022) 55 (Titarchuk & Seifina 2023) 56 (Chauhan et al. 2019) 57 (Russell et al. 2019) 58 (Kuulkers et al. 2013) 59 (Corral-Santana et al. 2018) 60 (Jana et al. 2022) 61 (Torres et al. 2019) 62 (Torres et al. 2020) 63 (Atri et al. 2020) 64 (Homan et al. 2018) 65 (Russell et al. 2014) 66 (Bahramian et al. 2023) 67 (Wu et al. 2016) 68 (González Hernández et al. 2017) 69 (Corral-Santana et al. 2013) 70 (Sánchez et al. 2015) 71 (Xu et al. 2018a) 72 (Mata Sánchez et al. 2024) 73 (Balakrishnan et al. 2023) 74 (Chaty et al. 2020) 75 (Arnason et al. 2021) 76 (Reis et al. 2009) 77 (Zurita et al. 2008) 78 (Saikia et al. 2023) 79 (Khargharia et al. 2010) 80 (Miller-Jones et al. 2009) 81 (Casares et al. 2019) 82 (Orosz et al. 2001) 83 (Macdonald et al. 2014) 84 (Gelino et al. 2006) 85 (González Hernández et al. 2014) 86 (Cherepashchuk et al. 2019) 87 (Orosz et al. 2011) 88 (Connors et al. 2020) 89 (Orosz et al. 2004) 90 (Homan et al. 2006) 91 (Chiang et al. 2012) 92 (Chaty & Bessolaz 2006) 93 (Ratti et al. 2012) 94 (Miller-Jones et al. 2011) 95 (García et al. 2018) 96 (Sala et al. 2007) 97 (Sriram et al. 2012) 98 (Bel et al. 2009) 99 (Yanes-Rizo et al. 2022) 100 (Corral-Santana et al. 2011) 101 (Mall et al. 2024) 102 (Rodriguez & Varnière 2011) 103 (Sriram et al. 2013) 104 (Chaty et al. 2006) 105 (Kumar 2024)

Table 2: Details of accretion states with reports of absorption line detection in both our work and the literature.

| Source | accretion states with absorption lines reported | | |
|---|---|---|---|
| | P24 | other works | |
| | iron band | iron band | other energies |
| **4U 1543-47** | X | soft$^{\dagger}$ $_{(1.)}$ | soft$^{X}$ $_{(2.)}$ |
| **4U 1630-47** | soft | soft $_{(3.)}$ , SPL $_{(4.)}$ | soft$^{X}$ $_{(5.)}$ |
| **A 0620-00** | X | X | outburst$^{V}$ $_{(6.)}$ |
| EXO 1846-031 | X | hard$^{\circledast}$ $_{(7.)\,(54.)}$ | X |
| GRO J0422+32 | X | X | outburst$^{V}$ $_{(8.)}$ |
| **GRO J1655-40** | soft | soft$_{(9.)}$ | soft$^{X}$ $_{(10.)}$, soft$^{V}$ $_{(11.)}$ |
| GRS 1009-45 | X | X | outburst$^{V}$ $_{(12.)}$ |
| GRS 1716-249 | X | X | hard$^{V}$ $_{(13.)}$ |
| GRS 1758-258 | X | hard$_{(14.)}$$^{\dagger}$ | X |
| **GRS 1915+105** | soft,hard | soft:$\phi,\gamma,\rho,\beta_{(15.)},\theta_{(16.)},\kappa_{(17.)},\lambda_{(18.)},\psi_{(19.)},\delta_{(23.)}$<br>hard:$\chi_{\mathrm{loud}\,(20.)}$,obscured*$_{(22.)}$, obscured* $\chi_{\mathrm{quiet}\,(23.)}$ | soft$^{X}$:$\phi_{(21.)}$<br>hard*:obscured$^{IR}$ $_{(24.)}$ |
| GX 339-4 | X | X | hard $^{X}$ $_{(26.)}$,soft$^{V}$ $_{(25.)}$,hard$^{V}$ $_{(25.)}$ |
| **H 1743-322** | soft | soft$_{(27.)}$ | soft$_{(28.)}$ |
| **IGR J17091-3624** | X | soft $_{(29.)}$,class-V$_{(30.)}$,class-X$_{(30.)}$, hard$^{\dagger}$ $_{(31.)}$ | hard$^{X}$ $_{(32.)}$ |
| **IGR J17451-3022** | soft | soft$_{(33.)}$ | soft$^{X}$ $_{(33.)}$ |
| **MAXI J1305-704** | X | soft$_{(34.)\,(35.)}$ | soft$^{X}$ $_{(34.)\,(35.)}$,hard$^{X}$ $_{(35.)}$, soft$^{V}$ $_{(36.)}$ |
| MAXI J1348-630 | X | soft$^{\circledast}$ $_{(37.)}$, hard$^{\circledast}$ $_{(37.)}$ | hard$^{X\dagger}$ $_{(38.)}$,hard$^{V,IR}$ $_{(39.)}$ |
| MAXI J1535-571 | X | soft$^{\circledast}$ $_{(54.)}$, hard$^{\circledast}$ $_{(54.)}$ | X |
| MAXI J1631-479 | X | soft$^{\circledast}$ $_{(40.)\,(1.)}$ | X |
| **MAXI J1659-152** | | X | quiescence$_{(41.)}$ |
| **MAXI J1803-298** | X | soft$_{(42.)\,(43.)}$ | soft$^{V}$ $_{(44.)}$,hard$^{V}$ $_{(44.)}$ |
| **MAXI J1810-222** | | X | soft$^{X}$ $_{(45.)}$, hard$^{X}$ $_{(45.)}$ |
| **MAXI J1820+070** | X | soft$_{(46.)}$ | hard$^{V,IR}$ $_{(47.)}$, quiescence$^{V}$ $_{(48.)}$ |
| **Nova Muscae 1991** | X | X | outburst$^{V}$ $_{(49.)}$ |
| Swift J1727.8-1613 | | X | soft$^{V\dagger}$ $_{(56.)}$, hard$^{V}$ $_{(56.)}$ |

Table 2: Continued.

| Source | accretion states with absorption lines reported | | |
|---|---|---|---|
| | P24 | other works | |
| | iron band | iron band | other energies |
| **Swift J1357.2-0933** | X | X | hard$^{V}$ (50.) (51.), quiescence (52.) |
| Swift J151857.0-572147 | | soft⋆ | X |
| **Swift J1658.2-4242** | X | soft⊛ (54.) (53.), hard⊛ (55.) (54.) (53.) | X |
| Swift J174540.7-290015 (T15) | X | soft⊛ (54.) | X |
| V404 Cyg | X | hard*:obscured (57.) | hard*:obscured$^{X}$ (58.) $^{V}$ (59.) |
| V4641 Sgr | X | X | soft$^{X}$ (60.) (2.),hard*:obscured$^{V}$ (61.) |
| XTE J1118+480 | X | X | hard$^{V}$ (62.) |
| XTE J1550-564 | X | soft (63.) | X |
| XTE J1652-453 | X | hard† (64.) | X |
| **XTE J1859+226** | X | X | soft$^{V}$ (65.) (66.) |

Legend:
**Bold**: dippers.
blue arc: X-ray absorption lines embedded in reflection components.
red arc: broad absorption line in optical, unclear origin (see Miceli et al. 2024 and the main text for details).
⊛: confirmed (⊛), suspected (⊛), or possible (⊛) edge mismatchs in NuSTAR fits creating artificial absorption lines (see Bogensberger et al. 2020 and the main text for details).
$X/V/IR$: detection in soft X-rays / Visible / Infrared.
†: low significance or tentative detection.
∗: The observed HR value of the obscured state might not reflect the actual HR of the source.
⋆: Source: IXPE collaboration.
The list of reference papers is not exhaustive for objects with many wind detections. Last update: 06-24.

**References:** 1 (Prabhakar et al. 2023) 2 (Draghis et al. 2023) 3 (Kubota et al. 2007) 4 (Parra et al. 2024a, in prep) 5 (Trueba et al. 2019) 6 (Whelan et al. 1977) 7 (Wang et al. 2020) 8 (Callanan et al. 1995) 9 (Miller et al. 2006a) 10 (Miller et al. 2008) 11 (Della Valle et al. 1998) 12 (Della Valle et al. 1997) 13 (Cúneo et al. 2020) 14 (Reynolds & al. 2018) 15 (Neilsen & Lee 2009) 16 (Ueda et al. 2010) 17 (Liu et al. 2022) 18 (Neilsen et al. 2018) 19 (Shi et al. 2023) 20 (Lee et al. 2002) 21 (Ueda et al. 2009) 22 (Neilsen et al. 2020) 23 (Athulya & Nandi 2023) 24 (Sánchez-Sierras et al. 2023) 25 (Rahoui et al. 2014) 26 (Miller et al. 2004) 27 (Miller et al. 2006b) 28 (Parra et al. 2024b, in prep. - PhD manuscript) 29 (King et al. 2012) 30 (Wang et al. 2024) 31 (Wang et al. 2018) 32 (Gatuzz et al. 2020) 33 (Jaisawal et al. 2015) 34 (Miller et al. 2014) 35 (Shidatsu et al. 2013) 36 (Miceli et al. 2024) 37 (Chakraborty et al. 2021) 38 (Saha et al. 2021) 39 (Panizo-Espinar et al. 2022) 40 (Xu et al. 2020) 41 (Torres et al. 2021) 42 (Coughenour et al. 2023) 43 (Zhang et al. 2024) 44 (Mata Sánchez et al. 2022) 45 (Del Santo et al. 2023) 46 (Fabian et al. 2021) 47 (Muñoz-Darias et al. 2019) 48 (Yoshitake et al. 2024) 49 (Della Valle et al. 1998) 50 (Jiménez-Ibarra et al. 2019) 51 (Charles et al. 2019) 52 (Sánchez et al. 2015) 53 (Bogensberger et al. 2020) 54 (Draghis et al. 2024) 55 (Xu et al. 2018b) 56 (Mata Sánchez et al. 2024) 57 (Muñoz-Darias & Ponti 2022) 58 (King et al. 2015) 59 (Muñoz-Darias et al. 2016) 60 (Shaw et al. 2022) 61 (Muñoz-Darias et al. 2018) 62 (Dubus et al. 2001) 63 (Connors et al. 2020) 64 (Chiang et al. 2012) 65 (Welsh et al. 2002) 66 (Hynes et al. 2002)

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