

# Emergent Dimensions in Background Independent Quantum Gravity

Kassahun Betre<sup>†</sup>, Patrick Wells<sup>†</sup>

<sup>†</sup>*Pepperdine University, 24255 Pacific Coast Hwy, Malibu, CA 90263, USA*

## ABSTRACT

Abstract goes here.

## I. INTRODUCTION

Background-independent models of spacetime geometry are increasingly common in the search for a quantum theory of gravity. If General Relativity is only an effective theory,

Introduction [1], [2], [3].

## II. SECTION 1

### A. Evolution of the Model

Given an undirected, loopless graph  $G$  with vertices  $V(G) = \{1...N\}$  and edges  $E(G)$ , a Hamiltonian  $H$  can be defined on the graph as follows:

$$H = J \sum_{i,j \in V(G), i \neq j} (d_i - d_j)^2 + K \sum_{v \in V(G)} d_v \quad (1)$$

Where  $J$  and  $K$  are weighting constants, and  $d_i$  simply represents the degree (number of connected edges) of a given node. The first of the sum forces the graph to be regular in the low-temperature limit, while the second sum pushes the graph to be more or less connected depending on the value of  $K$ . As  $K$  goes to infinity, the expectation value of the average node degree drops to zero. Conversely, as  $K$  tends towards negative infinity, the expectation value of the average node degree should go to  $N - 1$ . A set of seed graphs are evolved according to this Hamiltonian using a Monte-Carlo simulation with a Metropolis algorithm.

### III. GEOMETRICAL PROPERTIES OF GRAPHS

Geometrical properties of the graphs considered are defined as in [5]. This definition of dimensionality is computationally problematic, not due to the complexity of the algorithm but rather due to its fundamentally recursive nature. Calculating the dimensionality of a complete graph  $K_N$  on  $N$  nodes requires  $N!$  recursive calls, which quickly becomes computationally intractable as  $N$  increases. Of course,  $K_N$  is just an  $N-1$ -dimensional simplex, but no such simplification exists for highly dense graphs close to  $K_N$ .

### IV. CONCLUSION

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