

# Finger formation at the base of ash clouds: A linear stability analysis

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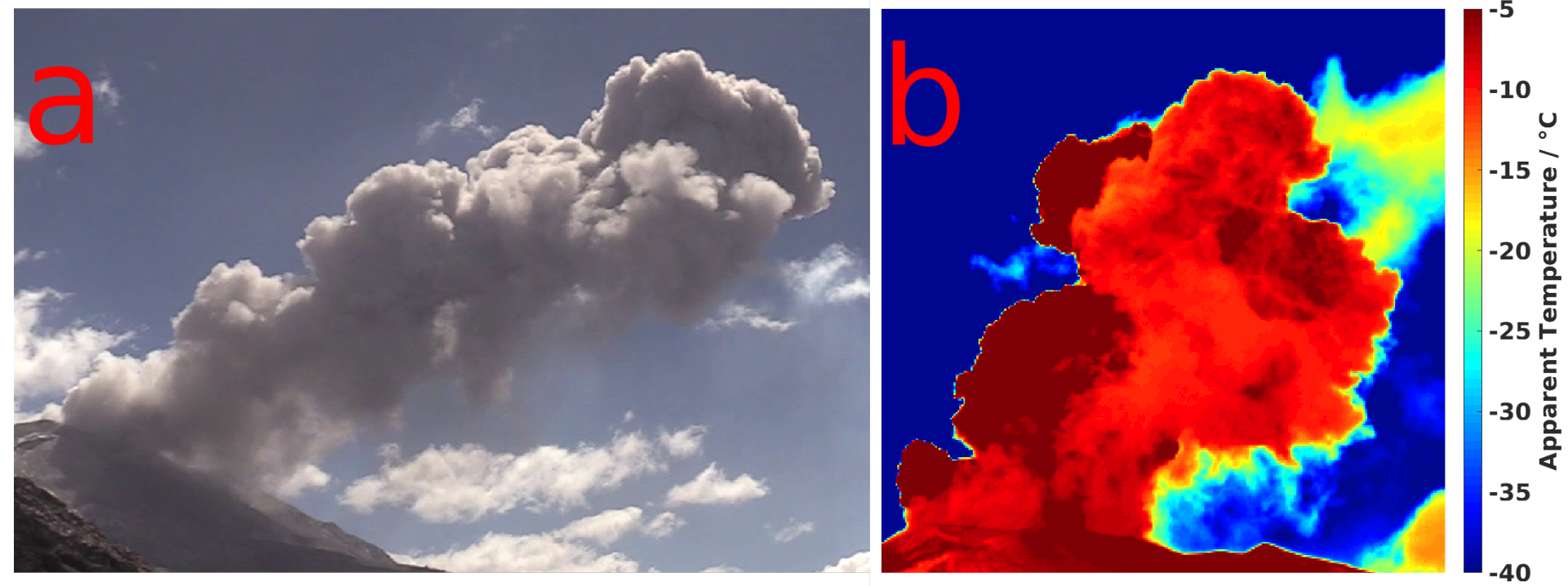


Figure 1. a) Visible and b) infrared images of ash clouds from Vulcanian explosions of Sabancaya volcano, Peru. Fingers can be seen in both images.

## Equations of motion

Conservation of mass  $\nabla \cdot \mathbf{u}(\mathbf{x}, t)$

Conservation of momentum  $\frac{\partial \mathbf{u}(\mathbf{x}, t)}{\partial t} + [\mathbf{u}(\mathbf{x}, t)] \cdot \nabla \mathbf{u}(\mathbf{x}, t) = \frac{\nabla P(\mathbf{x}, t)}{\rho_0} + \nu \nabla^2 \mathbf{u}(\mathbf{x}, t) - g' \hat{\mathbf{z}}$

Conservation of particles  $\frac{\partial \phi(\mathbf{x}, t)}{\partial t} + [\mathbf{u}(\mathbf{x}, t)] \cdot \nabla \phi(\mathbf{x}, t) - \mathbf{U}_p \cdot \nabla \phi(\mathbf{x}, t) = D_p \nabla^2 \phi(\mathbf{x}, t)$

Conservation of solute  $\frac{\partial s(\mathbf{x}, t)}{\partial t} + \mathbf{u}(\mathbf{x}, t) \cdot \nabla s(\mathbf{x}, t) = D_s \nabla^2 s(\mathbf{x}, t)$

Reduced gravity  $g' = g \left( 1 + \alpha s(\mathbf{x}, t) + \frac{\rho_p \phi(\mathbf{x}, t)}{\rho_0} \right)$

## Introduction

Volcanic ash presents a hazard for infrastructure and human health. Understanding ash settling is therefore crucial for assessing the associated risk. Downward-propagating fingers (Figure 1), hypothesised to form from a gravitational instability at the base of the ash cloud (Figure 2) have been seen at many volcanoes, and in analogue experiments (**Allan Fries**) and numerical models (**Jonathon Lemus**). Here we present a linear stability analysis, that predicts the initial growth rate of the fingers.

## Expansion

$$\mathbf{u}(\mathbf{x}, t) = \hat{\mathbf{u}}(z) e^{i(k_x x + k_y y - \omega t)}$$

$$P(\mathbf{x}, t) = P^{(0)}(\mathbf{x}, t) + \hat{P}(z) e^{i(kx - \omega t)}$$

$$\phi(\mathbf{x}, t) = \phi^{(0)}(\mathbf{x}, t) + \hat{\phi}(z) e^{i(kx - \omega t)}$$

$$s(\mathbf{x}, t) = s^{(0)}(\mathbf{x}, t) + \hat{s}(z) e^{i(kx - \omega t)}$$

$$\hat{P} \ll P^{(0)}, \hat{\phi} \ll \phi^{(0)}, \hat{s} \ll s^{(0)}$$

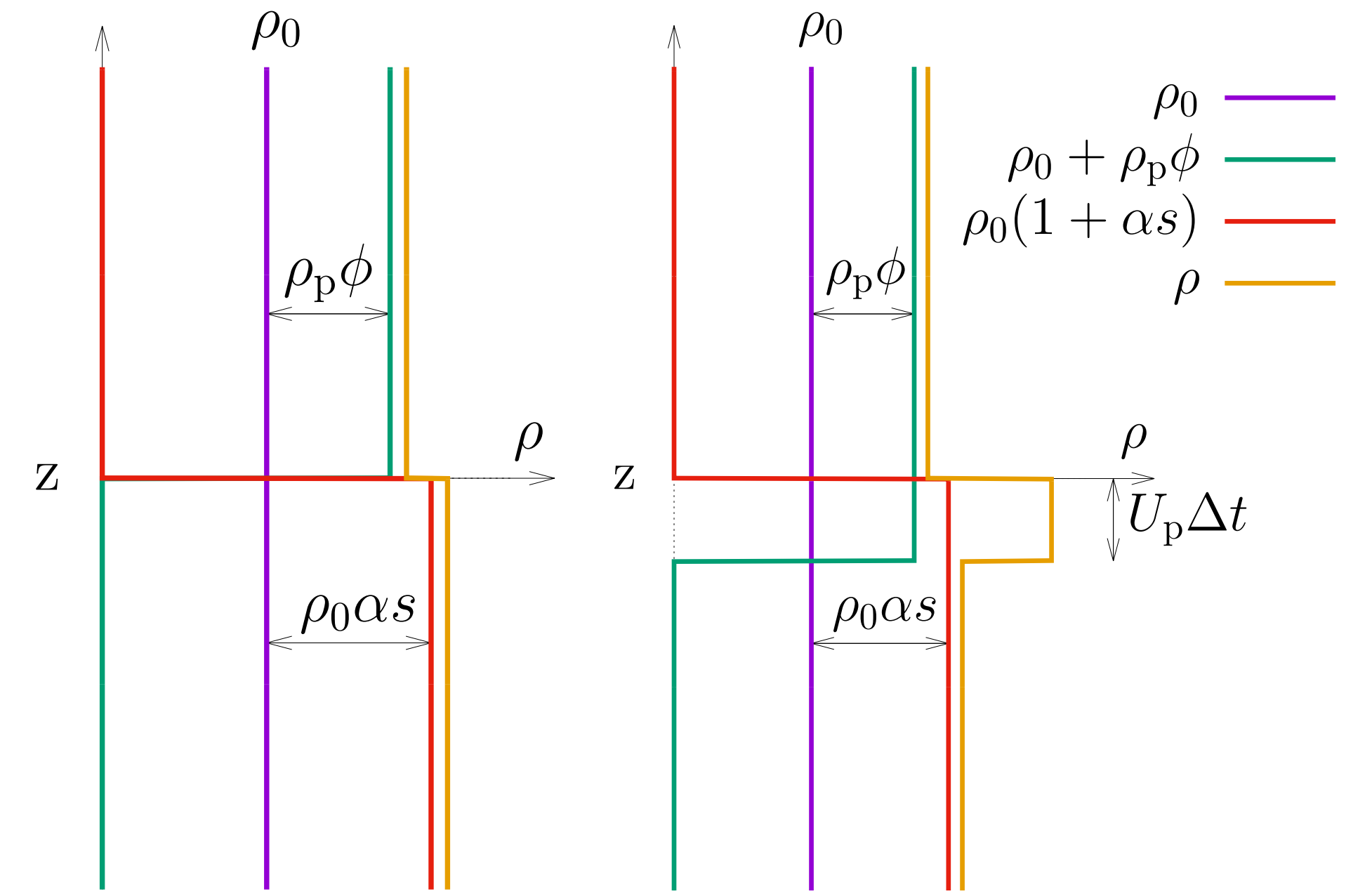


Figure 2. Despite an initial stable configuration, ash settling leads to the formation of a gravitationally unstable **particle boundary layer**.

Unknowns		Parameters	
$\mathbf{u}(\mathbf{x}, t)$	Fluid velocity	$\rho_0$	Reference fluid density
$P(\mathbf{x}, t)$	Pressure	$\nu$	Kinematic viscosity
$\phi(\mathbf{x}, t)$	Particle volume	$\mathbf{U}_p$	Particle settling velocity
$s(\mathbf{x}, t)$	Solute concentration	$\rho_p$	Particle density
Variables		$D_p$	Particle diffusivity
$\mathbf{x}$	Position vector	$D_s$	Solute diffusivity
$t$	time	$g$	Gravity
$\hat{\mathbf{z}}$	Vertical unit vector	$\alpha$	Expansivity

## Base states

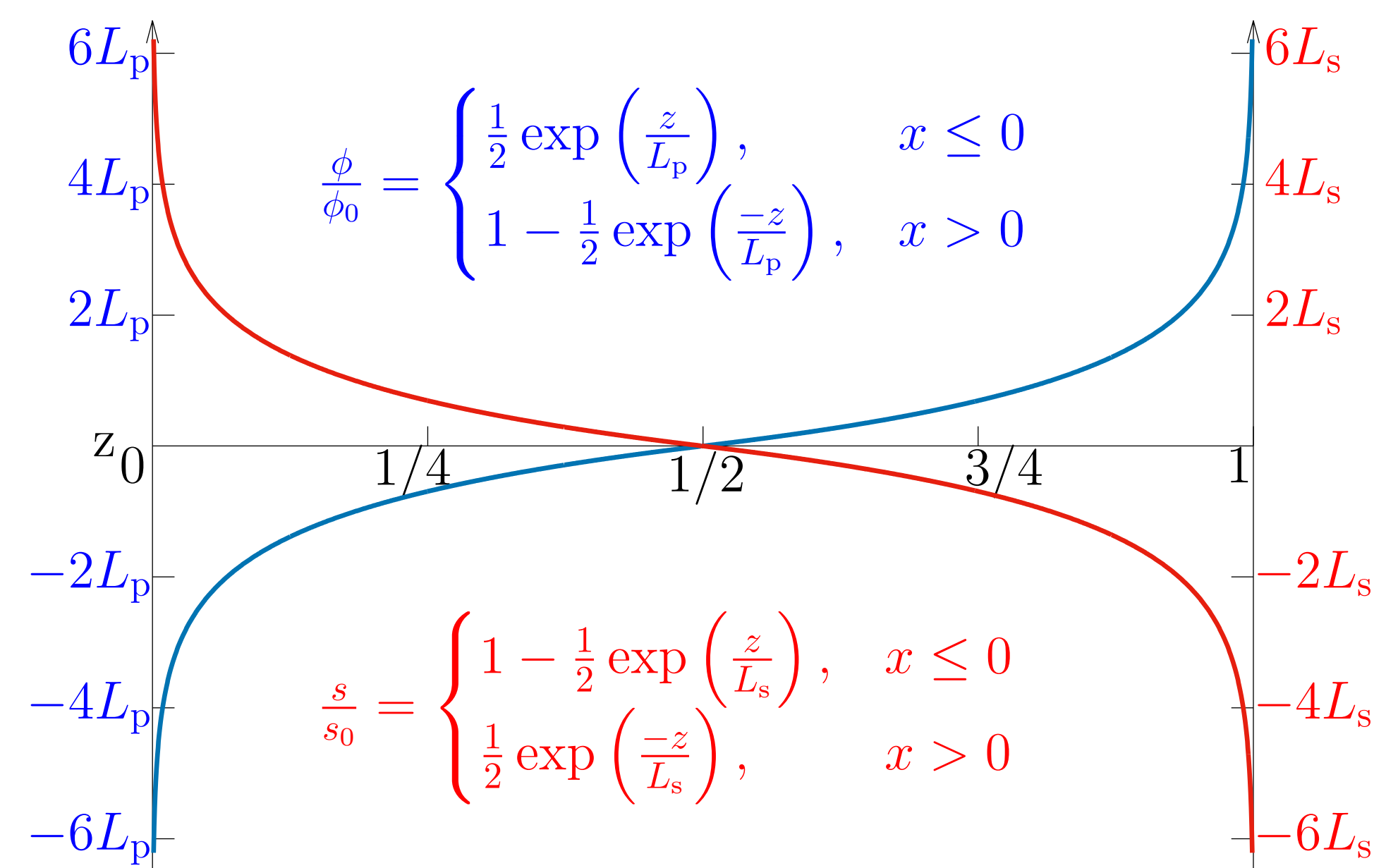


Figure 3. Choice of base states for the particle and solute concentrations. The transitions have length scales  $L_p$  and  $L_s$  respectively.

## Non-dimensionalisation and generalised eigenvalue problem

We nondimensionalise the system using the characteristic length and velocity scales

$$L_c = L_p$$

$$U_c = \frac{gL_p^2}{\nu}$$

This gives seven dimensionless parameters

$$Re = \frac{gL_p^3}{\nu^2}, \quad U'_p = \frac{\nu U_p}{gL_p^2}, \quad L'_s = \frac{L_s}{L_p}$$

$$A_\beta = \frac{\alpha s_0 \nu^2}{gL_p^3}, \quad A_\gamma = \frac{\rho_p \nu^2}{gL_p^3 \rho_0}, \quad Sc_s = \frac{\nu}{D_s}, \quad Sc_p = \frac{\nu}{D_p}$$

Can express problem as a generalised eigenvalue equation

$$\begin{pmatrix} \frac{1}{Re} \left( \frac{d^4}{dz'^4} - 2k'^2 \frac{d^2}{dz'^2} + k'^4 \right) & k'^2 A_\beta & k'^2 A_\gamma \\ \frac{\partial s^{(0)}}{\partial z'} & Re Sc_s \left( k'^2 - \frac{d^2}{dz'^2} \right) & 0 \\ \frac{\partial \phi^{(0)}}{\partial z'} & 0 & Re Sc_p \left( k'^2 - \frac{d^2}{dz'^2} \right) - U'_p \frac{d}{dz'} \end{pmatrix} \begin{pmatrix} \hat{u}_z \\ \hat{s} \\ \hat{\phi} \end{pmatrix} = i\omega \begin{pmatrix} \left( \frac{d^2}{k'^2 - \frac{d^2}{dz'^2}} \right) & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \hat{u}_z \\ \hat{s} \\ \hat{\phi} \end{pmatrix}$$

This is a generalised eigenvalue problem and is solved for the eigenvalue  $\omega = \Omega + i\sigma$ . Solutions where  $\sigma < 0$  are unstable, and the associated growth rate is  $|\sigma|$ . We can therefore predict the fastest growing perturbation wavelength  $\lambda_{\max} = 2\pi/k_{\max}$  which maximises  $|\sigma|$  for different values of the dimensionless parameters.

## Special case: Rayleigh-Taylor instability

To validate the model, we consider the special case of a single phase ( $s_0 = 0$ ) in the absense of a settling velocity ( $U'_p = 0$ ). This reduces the problem to that of the well studied Rayleigh-Taylor problem.

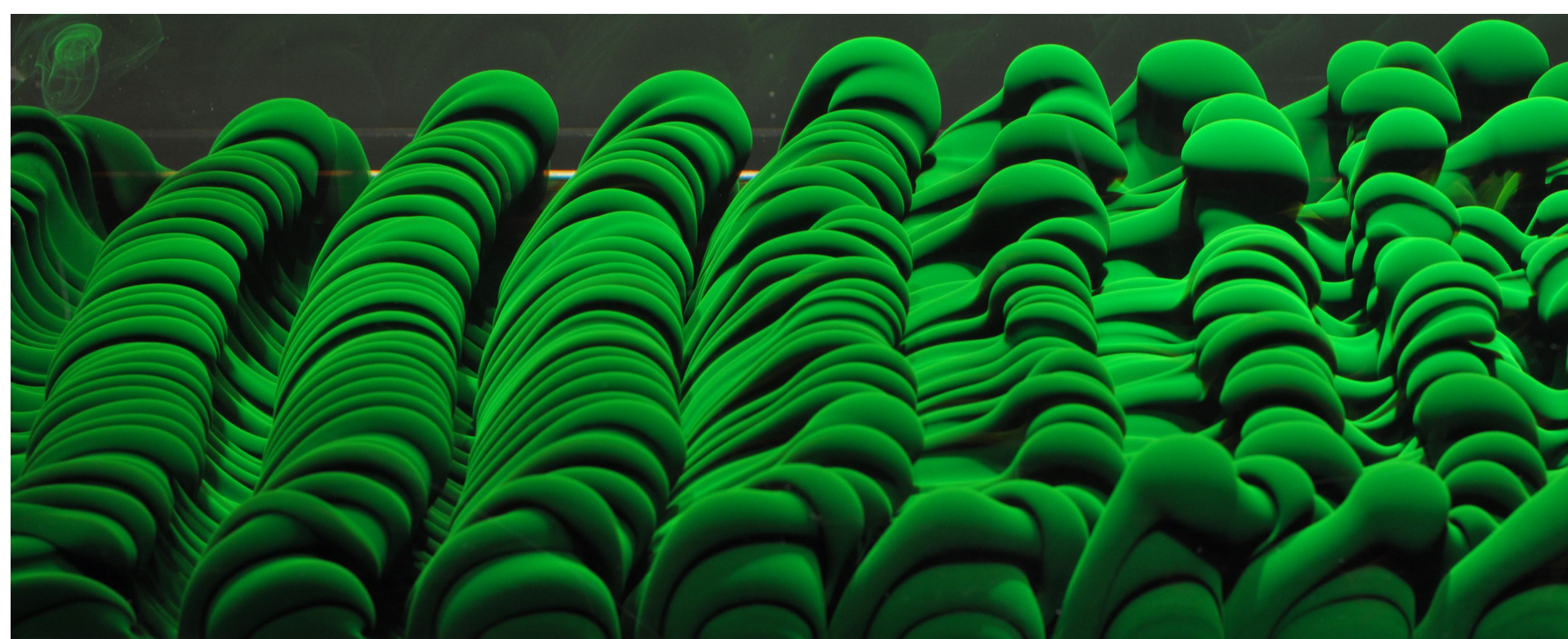


Figure 4. Experimentally produced Rayleigh-Taylor instability. The green fluid is lighter than the overlying clear fluid. Credit: Megan Davies Wykes

Choose sufficiently large  $Re(10^4)$  and  $Sc_p(10^3)$  that system becomes independent of them

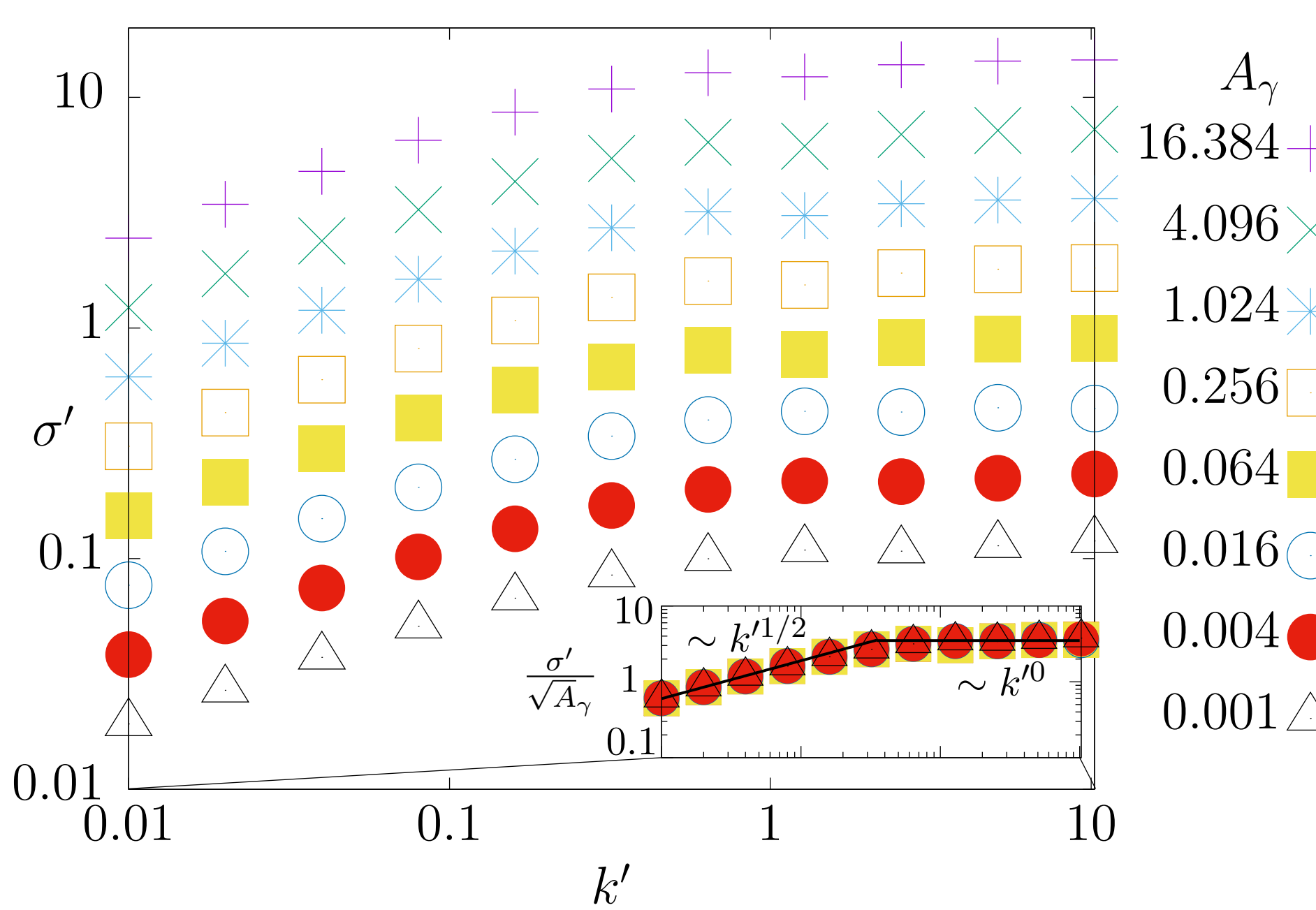


Figure 5. Dispersion relation showing that  $\sigma \sim A_\gamma$ . There are two regimes of  $k'$  dependence, when  $k' < 1$ ,  $\sigma \sim k'^{1/2}$  and when  $k' > 1$ ,  $\sigma \sim k'^0$ .

$$k' < 1 \quad (\lambda > 2\pi L_p)$$

$$k' > 1 \quad (\lambda < 2\pi L_p)$$

Wavelength is larger than the transition region. Scaling behaviour the same as for an infinitesimally thin transition zone.

Wavelength is shorter than the transition region. Growth rate independent of wavenumber.

$$\omega \sim \left( \frac{gk\rho_p}{\rho_0} \right)^{1/2}$$

$$\omega \sim \left( \frac{g\rho_p}{L_p\rho_0} \right)^{1/2}$$