

Neutron stars from skyrmion branes

Paul Leask

School of Mathematics, University of Leeds, Leeds, LS2 9JT, England, UK

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Table of Contents

1 Motivation

Neutron stars from
skyrmion branes

Paul Leask

2 Skyrme model

Motivation

Skyrme model

3 Linking in the Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

4 Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

5 Skyrmion crystals and phases of skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

6 Quantum skyrmion crystals and the symmetry energy

Neutron stars

7 Neutron stars

Final remarks

8 Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Motivation



Motivation

- Main aim: Describe baryonic matter on all scales from **finite atomic nuclei to dense infinite nuclear matter**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Main aim: Describe baryonic matter on all scales from **finite atomic nuclei** to **dense infinite nuclear matter**
- The Skyrme model can be used to model **neutron crystals**, which exist under high pressure inside neutron stars [*Nucl. Phys. B* **262** 133–143 (1985)]

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Within the Skyrme framework for various crystals, the neutron stars so far have been generically **crustless**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Can we obtain a **single EoS** that yields neutron stars with crusts?

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Motivation

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- Can we obtain a **single EoS** that yields neutron stars with crusts?
- Can these neutron stars have sufficient maximal masses?

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrme model



Skyrme model

- It was initially believed that descriptions of the low energy regime of QCD must contain explicit quarks

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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⇒ Degrees of freedom are **hadrons** (mesons and baryons)

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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⇒ Degrees of freedom are **hadrons** (mesons and baryons)
 - In the large N_c -limit, QCD can be reduced to an effective field theory of mesons
 - Skyrme's idea [*Proc. R. Soc. Lond. A* **260** 127-138 (1961)]: effective mesonic Lagrangian involving only **pions**, with **baryons emerging as stable topological solitons**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Skyrme model

- The theory has $N_f = 2$ flavours of quarks (u,d) that make up the pion fields

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- The theory has $N_f = 2$ flavours of quarks (u,d) that make up the pion fields
- These are encoded in the Skyrme field

$$\varphi = \begin{pmatrix} \sigma + i\pi_3 & i\pi_1 + \pi_2 \\ i\pi^1 - \pi_2 & \sigma - i\pi_3 \end{pmatrix} \in \text{SU}(2)$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- This is a map $\varphi : \mathbb{R}^3 \rightarrow \text{SU}(2)$ with the constraint $\sigma^2 + \vec{\pi} \cdot \vec{\pi} = 1$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Skyrme model

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrme model

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- One-point compactification of space $\mathbb{R}^3 \cup \{\infty\} \cong S^3$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrme model

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 - One-point compactification of space $\mathbb{R}^3 \cup \{\infty\} \cong S^3$
- \Rightarrow Skyrme field is now a map $\varphi : S^3 \rightarrow \mathrm{SU}(2)$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrme model

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$$\text{SU}(2) \ni \begin{pmatrix} \sigma + i\pi_3 & i\pi_1 + \pi_2 \\ i\pi^1 - \pi_2 & \sigma - i\pi_3 \end{pmatrix} \leftrightarrow (\sigma, \pi_1, \pi_2, \pi_3) \in S^3$$

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrme model

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- \Rightarrow Skyrme field is now a map $\varphi : S^3 \rightarrow \text{SU}(2) \cong S^3$
- Disjoint homotopy classes labelled by $B \in \pi_3(S^3) = \mathbb{Z}$
- \Rightarrow Fields are **topologically stable** and B is identified with the **baryon number**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Generalized Skyrme Lagrangian

- Skyrme's original model:

$$\mathcal{L}_{24} = \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta])$$
$$L_\mu = \varphi^\dagger \partial_\mu \varphi \in \mathfrak{su}(2)$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- This is $(\text{SU}(2) \times \text{SU}(2))/\mathbb{Z}_2 \cong \text{SO}(4)$ invariant and the **pions are massless**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- This is $(\text{SU}(2) \times \text{SU}(2))/\mathbb{Z}_2 \cong \text{SO}(4)$ invariant and the **pions are massless**
- Boundary condition $\varphi(\vec{x} \rightarrow \infty) = \mathbb{I}_2$ spontaneously breaks chiral $\text{SO}(4)$ symmetry to an isospin $\text{SO}(3)$ symmetry, which acts on $\vec{\pi}$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Generalized Skyrme Lagrangian

- The standard massive Skyrme model includes the **pion mass potential**:

$$\mathcal{L}_{024} = -\frac{F_\pi^2 m_\pi^2}{8\hbar^3} \text{Tr}(\mathbb{I}_2 - \phi) + \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta])$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Generalized Skyrme Lagrangian

- Generalized Skyrme model includes a sextic term [*Phys. Lett. B* **154**, 101–106 (1985)], which is related to the ω -Skyrme model [*Phys. Lett. B* **137**, 251–256 (1984)]:

$$\begin{aligned}\mathcal{L}_{0246} = & -\frac{F_\pi^2 m_\pi^2}{8\hbar^3} \text{Tr}(\mathbb{I}_2 - \varphi) + \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta]) \\ & - \pi^4 \lambda^2 g^{\mu\nu} \mathcal{B}_\mu \mathcal{B}_\nu, \quad \lambda^2 = g_\omega^2 / (2\pi^4 m_\omega^2)\end{aligned}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Generalized Skyrme Lagrangian

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- Exhibits short range ω -meson-like repulsion while still describing scalar meson effects

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Exhibits short range ω -meson-like repulsion while still describing scalar meson effects
- Baryon d.o.f. not explicitly visible → topology: Homotopy invariant \leftrightarrow Baryon number

$$\pi_3(\text{SU}(2)) = \mathbb{Z} \ni B = \int_{\mathbb{R}^3} d^3x \sqrt{-g} \mathcal{B}^0, \quad \mathcal{B}^\mu = \frac{1}{24\pi^2 \sqrt{-g}} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(L_\nu L_\rho L_\sigma)$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Generalized Skyrme Lagrangian

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- Exhibits short range ω -meson-like repulsion while still describing scalar meson effects
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- Baryons realized as non-perturbative excitations of the pions ⇒ solutions of the Euler–Lagrange field equations - topological solitons (**skyrmions**)

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Generalized Skyrme model

- We are interested in **static** solutions and adopt the usual Skyrme units of length $\tilde{L} = 2\hbar/eF_\pi$ and energy $\tilde{E} = F_\pi/4e$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- In Skyrme units the energy-momentum tensor is

$$\begin{aligned}
 T_{\mu\nu} &= -\frac{2}{\sqrt{-g}} \frac{\partial(\sqrt{-g}\mathcal{L}_{0246})}{\partial g^{\mu\nu}} \quad \frac{\pi^4 \lambda^2 e^4 F_\pi^2}{2\hbar^3} = c_6 \boxtimes \\
 &= -\text{Tr}(L_\mu L_\nu) - \frac{1}{4}g^{\alpha\beta} \text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6 \mathcal{B}_\mu \mathcal{B}_\nu + g_{\mu\nu} \mathcal{L}_{0246}
 \end{aligned}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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$$T_{\mu\nu} = - \text{Tr}(L_\mu L_\nu) - \frac{1}{4}g^{\alpha\beta} \text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6 \mathcal{B}_\mu \mathcal{B}_\nu + g_{\mu\nu} \mathcal{L}_{0246}$$

- The adimensional static energy is thus ($T_{00} = \mathcal{E}_{\text{stat}} + \mathcal{E}_{\text{kin}}$)

$$\begin{aligned} M_B(\varphi, g) &= \int_{\mathbb{R}^3} d^3x \sqrt{-g} \mathcal{E}_{\text{stat}} \\ &= \int_M d^3x \sqrt{-g} \left\{ -\frac{1}{2}g^{ij} \text{Tr}(L_i L_j) - \frac{1}{16}g^{ik}g^{jl} \text{Tr}([L_i, L_j][L_k, L_l]) \right. \\ m &= \frac{2m_\pi}{F_\pi e} \rightarrow +m^2 \text{Tr}(\mathbb{I}_2 - \varphi) + c_6 \frac{\epsilon^{ijk}\epsilon^{abc}}{(24\pi^2\sqrt{-g})^2} \text{Tr}(L_i L_j L_k) \text{Tr}(L_a L_b L_c) \Big\} \end{aligned}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Generalized Skyrme model

- We are interested in **static** solutions and adopt the usual Skyrme units of length $\tilde{L} = 2\hbar/eF_\pi$ and energy $\tilde{E} = F_\pi/4e$
- In Skyrme units the energy-momentum tensor is

$$T_{\mu\nu} = - \text{Tr}(L_\mu L_\nu) - \frac{1}{4}g^{\alpha\beta} \text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6 \mathcal{B}_\mu \mathcal{B}_\nu + g_{\mu\nu} \mathcal{L}_{0246}$$

- The adimensional static energy is thus

$$\begin{aligned} M_B(\phi, g) = & \int_{\mathbb{R}^3} d^3x \sqrt{-g} \left\{ -\frac{1}{2}g^{ij} \text{Tr}(L_i L_j) - \frac{1}{16}g^{ik}g^{jl} \text{Tr}([L_i, L_j][L_k, L_l]) \right. \\ & \left. + m^2 \text{Tr}(\mathbb{I}_2 - \phi) + c_6 \frac{\epsilon^{ijk}\epsilon^{abc}}{(24\pi^2\sqrt{-g})^2} \text{Tr}(L_i L_j L_k) \text{Tr}(L_a L_b L_c) \right\} \end{aligned}$$

- Skyrmions are **energy minimizing** static solutions of the Euler–Lagrange equations associated to M_B

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Linking in the Skyrme model



Baryon number as the linking number of vortices

- The Skyrme field can be written in terms of two vortices $\psi_1, \psi_2 \in \mathbb{C}$ as

$$\varphi = \begin{pmatrix} \psi_1 & -\bar{\psi}_2 \\ \psi_2 & \bar{\psi}_1 \end{pmatrix}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Consider the Hopf map $H : S^3 \rightarrow S^2$ due to the Hopf fibration $S^1 \hookrightarrow S^3 \rightarrow S^2$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- The map $\varphi : \mathbb{R}^3 \cup \{\infty\} \cong S^3 \rightarrow S^3$ of degree B has Hopf charge $Q = B$ under the Hopf map $H : S^3 \rightarrow S^2$ [*Phys. Rev. D* **101**, 065011 (2020)]

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Antipodal points on S^2 identified as **vortex lines** (zeros) and are linked $Q = B$ times

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Skyrmion solutions

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



$B = 1$ 🦔 skyrmion [*Proc. R. Soc. Lond. A* **260** 127-138 (1961)]

- The hedgehog field is $\varphi(\vec{x}) = \exp(if(r)\vec{x} \cdot \vec{\tau})$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Profile function $f(r)$ must satisfy the B.C.s $f(\infty) = 0$ and $f(0) = \pi$
- The hedgehog solution has baryon number $B = 1$ since

$$B = -\frac{1}{2\pi^2} \int_0^\infty \frac{\sin^2 f}{r^2} \frac{df}{dr} 4\pi r^2 dr = \frac{1}{\pi} f(0) = 1$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrme crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- For the hedgehog ansatz, the (massless) static energy is

$$M_1 = 4\pi \int_0^\infty \left[r^2 \left(\frac{df}{dr} \right)^2 + 2 \sin^2 f \left(1 + \left(\frac{df}{dr} \right)^2 \right) + \frac{\sin^4 f}{r^2} \right] dr$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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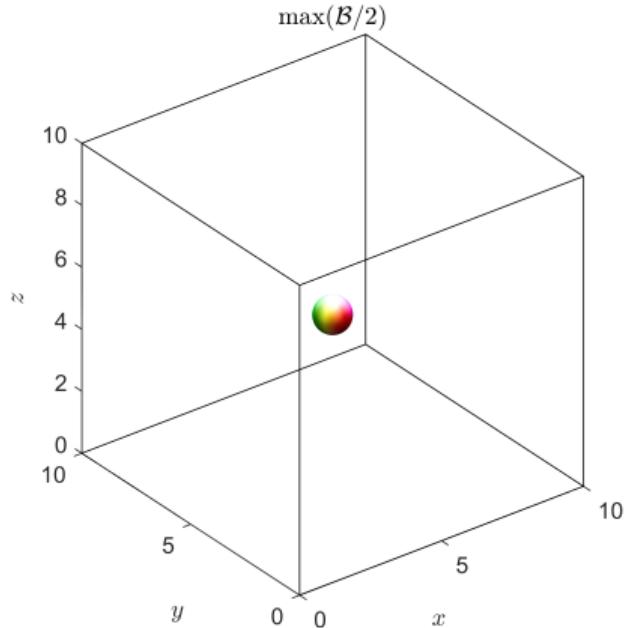
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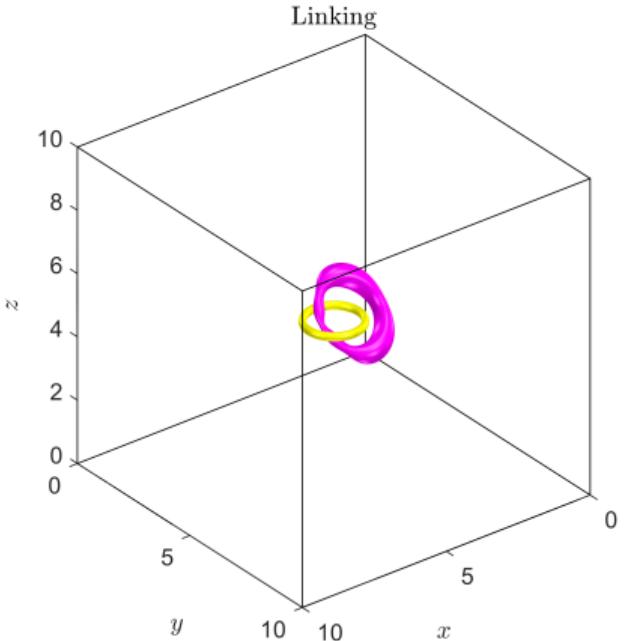
- E-L equations reduce to a 2nd order non-linear ODE that can only be solved numerically,

$$(r^2 + 2 \sin^2 f) \frac{d^2 f}{dr^2} + 2r \frac{df}{dr} + \sin 2f \left[\left(\frac{df}{dr} \right)^2 - 1 - \frac{\sin^2 f}{r^2} \right] = 0$$

$B = 1$ hedgehog skyrmion [*Proc. R. Soc. Lond. A* **260** 127-138 (1961)]



(a) Isosurface plot of the baryon density \mathcal{B}_0



(b) Linking of two preimages



How to construct larger Skyrmions

- Asymptotic interactions of two $B = 1$ skyrmions have preferred orientation (**attractive channel**) [*Commun. Math. Phys.* **245** 123–147 (2004)]

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Quite large skyrmions (up to $B = 108$) have been constructed by gluing α -particles together [*Proc. R. Soc. A* **463** 261–279 (2007)] or using a multi-layer RMA based on the **Skyrme crystal** [*Phys. Rev. D* **87** 0850834 (2013)]

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrme solutions

Skyrme crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Skyrmions from rational maps [*Nucl. Phys. B* **510**, 507–537 (1998)]

- Rational maps are functions from $S^2 \rightarrow S^2$, whereas as $\varphi : \mathbb{R}^3 \rightarrow S^3$.

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmiⁿ solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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$$\varphi(r, z) = \exp \left[\frac{if(r)}{1 + |R|^2} \begin{pmatrix} 1 - |R|^2 & 2\bar{R} \\ 2R & |R|^2 - 1 \end{pmatrix} \right],$$

Neutron stars from
skyrmiⁿ branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmiⁿ solutions

Skyrmiⁿ crystals
and phases of
skyrmiⁿ matter

Quantum skyrmiⁿ
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- $R(z) = p(z)/q(z)$ is a rational map of degree $B = \max(\deg p, \deg q)$.
- Rational maps for $B = 1, \dots, 4$:

B	$R(z)$	Symmetry
1	z	$O(3)$
2	z^2	$O(2) \times \mathbb{Z}$
3	$\frac{z^3 - \sqrt{3}iz}{\sqrt{3}iz^2 - 1}$	T_d
4	$\frac{z^4 + 2\sqrt{3}iz^2 + 1}{z^4 - 2\sqrt{3}iz^2 + 1}$	O_b

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Substituting the RMA into the massless static energy functional yields

$$M_B = 4\pi \int_0^\infty r^2 \left\{ \left(\frac{df}{dr} \right)^2 + 2B \sin^2 f \left(\left(\frac{df}{dr} \right)^2 + 1 \right) + \mathcal{I} \frac{\sin^4 f}{r^2} \right\} dr,$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- \mathcal{I} is the purely angular integral to be minimised for choice of rational map R :

$$\mathcal{I} = \frac{1}{4\pi} \int \left(\frac{1+|z|^2}{1+|R|^2} \left| \frac{dR}{dz} \right| \right)^4 \frac{2idz d\bar{z}}{(1+|z|^2)^2}.$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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$$M_B = 4\pi \int_0^\infty r^2 \left\{ \left(\frac{df}{dr} \right)^2 + 2B \sin^2 f \left(\left(\frac{df}{dr} \right)^2 + 1 \right) + \mathcal{I} \frac{\sin^4 f}{r^2} \right\} dr,$$

- \mathcal{I} is the purely angular integral to be minimised for choice of rational map R :

$$\mathcal{I} = \frac{1}{4\pi} \int \left(\frac{1+|z|^2}{1+|R|^2} \left| \frac{dR}{dz} \right| \right)^4 \frac{2idz d\bar{z}}{(1+|z|^2)^2}.$$

- Optimising \mathcal{I} and the profile function $f(r)$ gives approximate Skyrmions, but further numerical relaxation is required to find true Skyrmions.

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

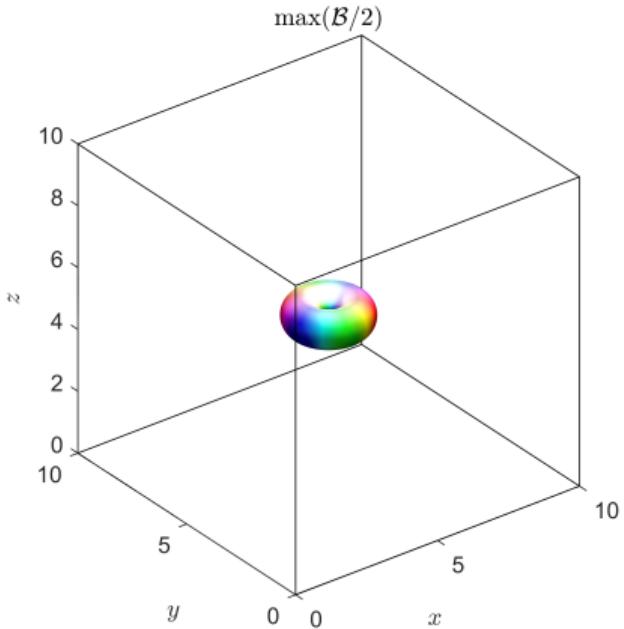
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and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

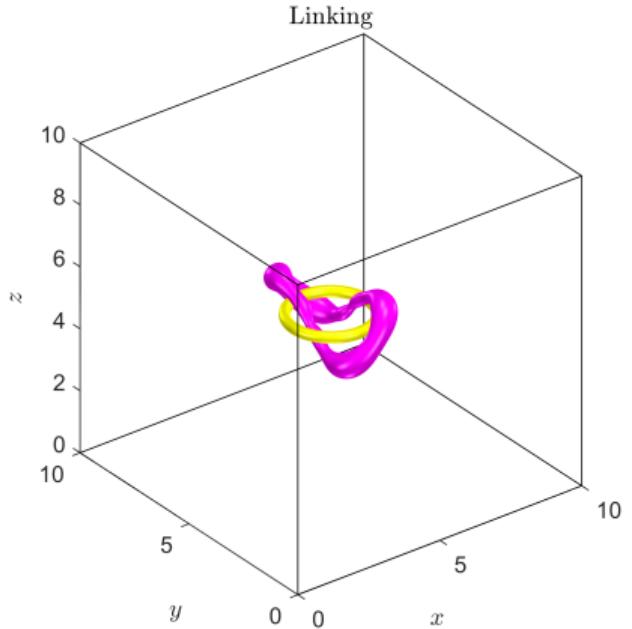
Neutron stars

Final remarks

$B = 2$ skyrmion [*Phys. Lett. B* **195**, 235–239 (1987)]



(a) Isosurface plot of the baryon density \mathcal{B}_0



(b) Linking of two preimages

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmiⁿ solutions

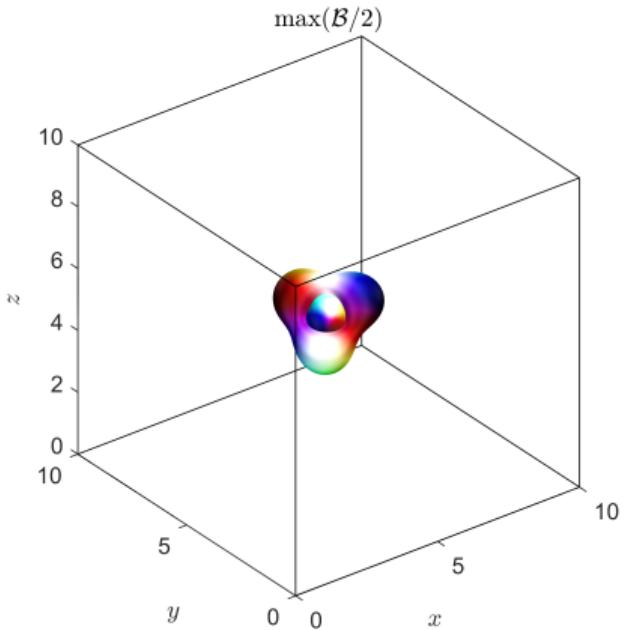
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and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

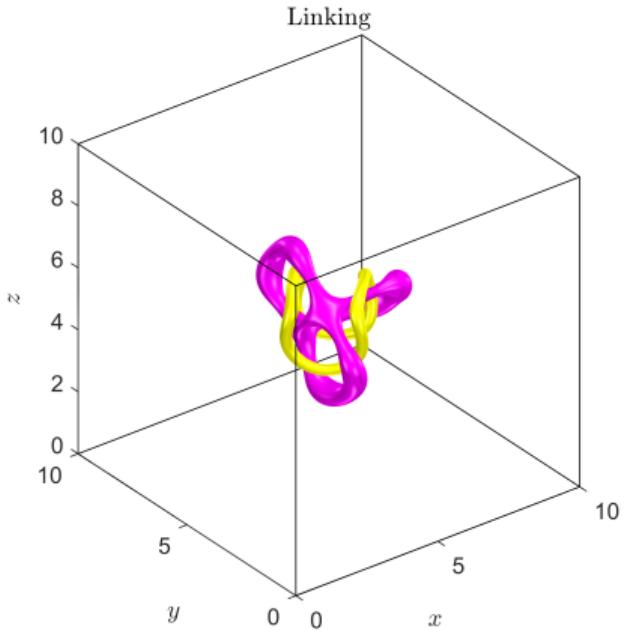
Neutron stars

Final remarks

$B = 3$ skyrmion [*Phys. Lett. B* **235**, 147–152 (1990)]



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmiⁿ solutions

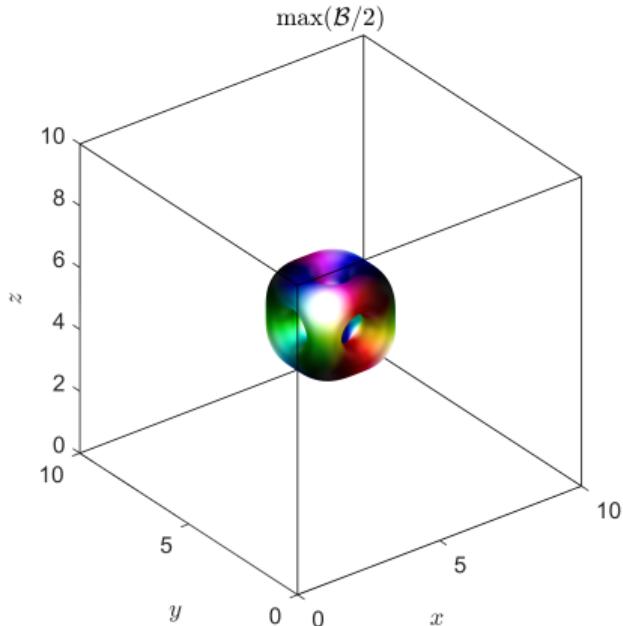
Skyrmiⁿ crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

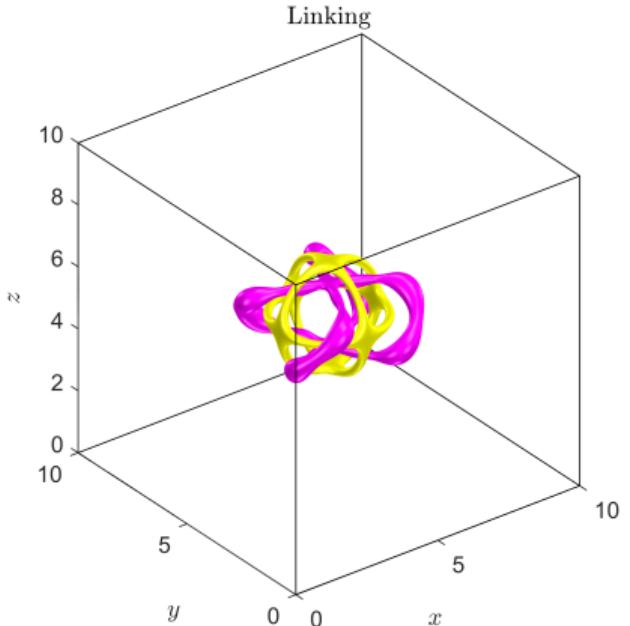
Neutron stars

Final remarks

$B = 4$ skyrmion [*Phys. Lett. B* **235**, 147–152 (1990)]



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrmion crystals and phases of skyrmion matter



Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
⇒ We need to understand **phases** and **phase transitions** of nuclear matter

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
- ⇒ We need to understand **phases** and **phase transitions** of nuclear matter
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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
⇒ We need to understand **phases** and **phase transitions** of nuclear matter
- Ground state of dense nuclear matter has a **crystalline** structure in the classical approximation
- In order to determine skyrmion crystals, we first need to define what a crystal really is!

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Skyrmion crystals

- Skyrme crystals are energy minimizing maps

$$\varphi : \mathbb{R}^3 / \Lambda_\circ \rightarrow \mathrm{SU}(2), \quad \Lambda_\circ = \left\{ n_1 \vec{X}_1 + n_2 \vec{X}_2 + n_3 \vec{X}_3 : n_i \in \mathbb{Z} \right\}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- They are critical and stable w.r.t. variations of the lattice Λ about Λ_0

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Key idea [*Comm. Math. Phys.* **332** 355-377 (2014)]: Identify all 3-tori via diffeomorphism (with $T^3 \cong \mathbb{R}^3 / \mathbb{Z}^3$)

$$F : (T^3, g) \rightarrow (\mathbb{R}^3 / \Lambda, d), \quad (x^1, x^2, x^3) \mapsto x^1 \vec{X}_1 + x^2 \vec{X}_2 + x^3 \vec{X}_3$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- The metric on T^3 is the pullback $g = F^* d = g_{ij} dx^i dx^j$, $g_{ij} = \vec{X}_i \cdot \vec{X}_j$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrmion crystals

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrmion crystals

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- Vary metric g with $g_0 = F^* d \iff$ vary lattice Λ_s with $\Lambda_0 = \Lambda$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Skyrmion crystals

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- Vary metric g with $g_0 = F^* d \iff$ vary lattice Λ with $\Lambda_0 = \Lambda$
- Energy minimized over variations of $g \iff$ energy minimizing period lattice Λ_0

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Summary of [J. Math. Phys. 64 103503 (2023)]

- For fixed \mathcal{L}_{024} -field ϕ , there always **exists** a critical point of $M_B(\phi, g)$ w.r.t. variations of g and it is in fact a **unique** c.p. (generalizes to \mathcal{L}_{0246} -model)

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Four crystal solutions were found for unit cells with charge $B_{\text{cell}} = 4$
- These are the φ_{FCC} , φ_α , φ_{string} and φ_{brane} crystals
- The φ_{FCC} -crystal [*Phys. Lett. B* **208** 491–494 (1988)] can be obtained from a Fourier series-like expansion as an initial configuration [*Nucl. Phys. A* **501** 801–812 (1989)],

$$\sigma = -c_1 c_2 c_3, \quad \pi_1 = s_1 \sqrt{1 - \frac{s_2^2}{2} - \frac{s_3^2}{2} + \frac{s_2^2 s_3^2}{3}}, \quad \text{and cyclic,}$$

where $s_i = \sin(2\pi x^i/L)$ and $c_i = \cos(2\pi x^i/L)$, with initial metric $g = L^2 \mathbb{I}_3$.

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

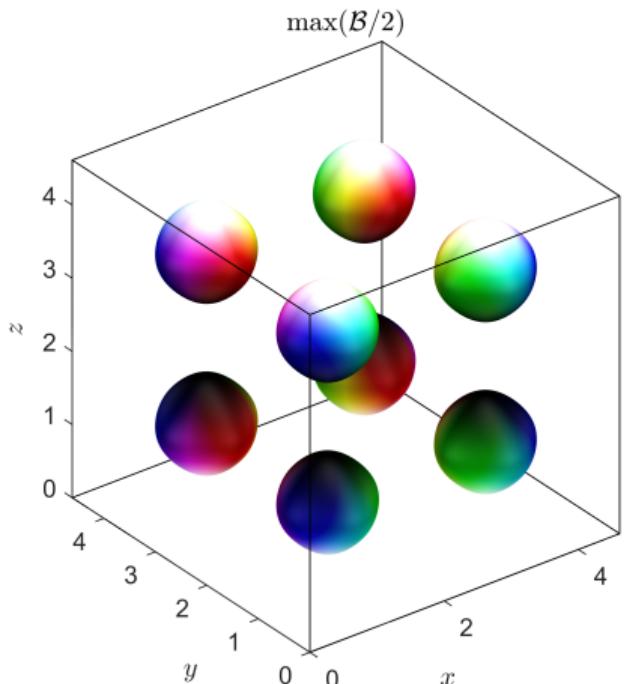
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

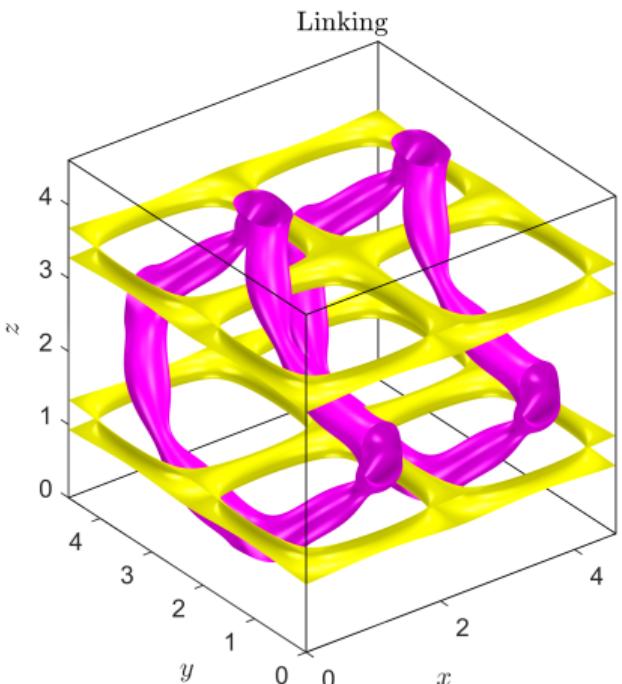
Neutron stars

Final remarks

Half-skyrmion (FCC) crystal [*Phys. Lett. B* **208** 491–494 (1988)]



(a) Isosurface plot of the baryon density \mathcal{B}_0



(b) Linking of two preimages

Neutron stars from
skyrmi branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmi solutions

Skyrmi crystals
and phases of
skyrmi matter

Quantum skyrmi
crystals and the
symmetry energy

Neutron stars

Final remarks



Summary of [J. Math. Phys. 64 103503 (2023)]

- From φ_{FCC} , the other three crystals can be constructed by applying a chiral $\text{SO}(4)$ transformation $Q \in \text{SO}(4)$, such that $\varphi = Q\varphi_{\text{FCC}}$, and minimizing M_B w.r.t. variations of φ and g

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the Skyrme model

Skyrmion solutions

Skyrmion crystals and phases of skyrmion matter

Quantum skyrmion crystals and the symmetry energy

Neutron stars

Final remarks

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- The φ_{brane} -crystal is the lowest energy solution at all baryon densities $n_B = B_{\text{cell}}/V_{\text{cell}}$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- ⇒ Should yield a **lower compression modulus** than previous studies

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- ⇒ Should yield a **lower compression modulus** than previous studies
- ⇒ **Brane crystal** is an ideal candidate for **dense nuclear matter**

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

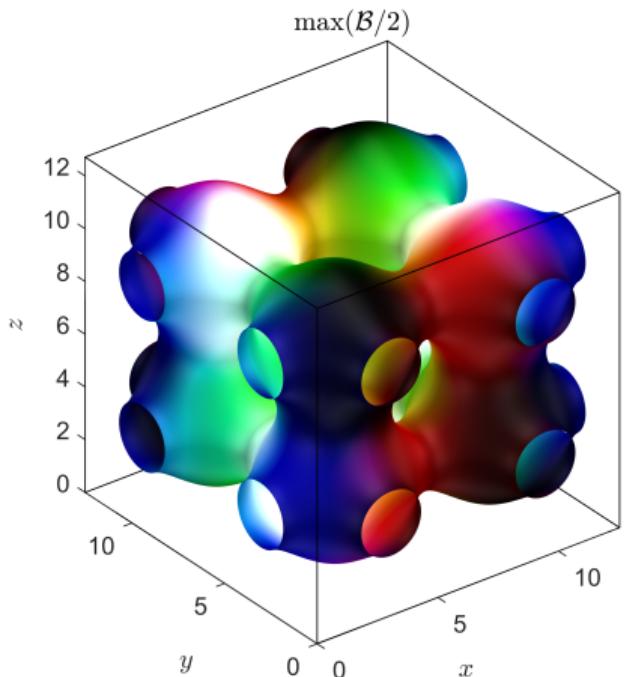
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

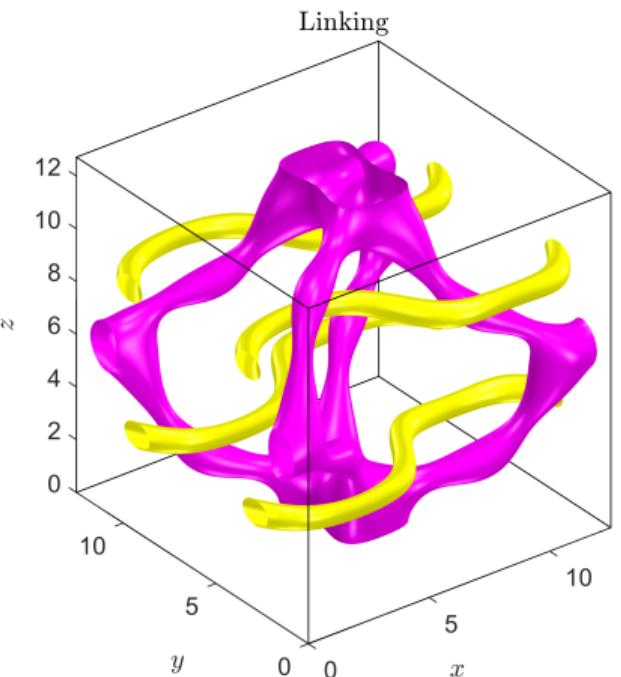
Neutron stars

Final remarks

Brane or domain wall crystal [J. Math. Phys. **64** 103503 (2023)]



(a) Isosurface plot of the baryon density \mathcal{B}_0



(b) Linking of two preimages



Varying the metric on \mathbb{T}^3

- Let g_ϵ be a smooth one-parameter family of metrics on \mathbb{T}^3 with $g_0 = F^*d$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Varying the metric on \mathbb{T}^3

- Let g_s be a smooth one-parameter family of metrics on \mathbb{T}^3 with $g_0 = F^*d$
- Set $\delta g = \partial_s g_s|_{s=0} \in \Gamma(\odot^2 T^*\mathbb{T}^3)$ (symmetric 2-covariant tensor field on \mathbb{T}^3)

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- Inner product on the space of 2-covariant tensor fields $\langle A, B \rangle_g = A_{ij} g^{jk} B_{kl} g^{li}$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Varying the metric on \mathbb{T}^3

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- First variation of M_B w.r.t. g_s is

$$\frac{dM_B(\varphi, g_s)}{ds} \Big|_{s=0} = \int_{\mathbb{T}^3} d^3x \sqrt{g} \langle S(\varphi, g), \delta g \rangle_g, \quad S(\varphi, g) \in \Gamma(\odot^2 T^* \mathbb{T}^3)$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Varying the metric on \mathbb{T}^3

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- $S(\varphi, g)$ is the **stress-energy tensor**:

$$\begin{aligned} S_{ij} &= \frac{1}{2} \left[m^2 \text{Tr}(\text{Id} - \varphi) - \frac{1}{2} g^{kl} \text{Tr}(L_k L_l) - \frac{1}{16} g^{km} g^{ln} \text{Tr}([L_k, L_l][L_m, L_n]) - c_6 (B_0)^2 \right] g_{ij} \\ &\quad + \frac{1}{2} \text{Tr}(L_i L_j) + \frac{1}{8} g^{kl} \text{Tr}([L_i, L_k][L_j, L_l]). \end{aligned}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Numerical minimization of the field and lattice

- Fix $\varphi : \mathbb{T}^3 \rightarrow \mathrm{SU}(2)$ and think of the energy as a map $E_\varphi : \mathrm{SPD}_3 \rightarrow \mathbb{R}$ such that
$$E_\varphi := M_B(\varphi|_{\text{fixed}}, g)$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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$$E_\varphi := M_B(\varphi|_{\text{fixed}}, g)$$
- We use arrested Newton flow on SPD_3 to minimize E_φ w.r.t. g

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- We use arrested Newton flow on SPD_3 to minimize E_φ w.r.t. g
- Explicitly, we are solving the system of 2nd order ODEs

$$\frac{d^2}{ds^2} \Big|_{s=0} (g_{ij})_s = -\frac{\partial E_\varphi}{\partial g_{ij}} = - \int_{\mathbb{T}^3} d^3x \sqrt{g} S_\varphi^{ij}, \quad (g_{ij})_0 = \vec{X}_i \cdot \vec{X}_j$$

where $S_\varphi \equiv S(\varphi|_{\text{fixed}}, g)$ is the fixed field stress tensor

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- In conjunction, we minimize $M_B(\varphi, g|_{\text{fixed}})$ w.r.t. φ for some initial field φ_0

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Numerical minimization of the field and lattice

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 ⇒ Laddering of minimizations

Neutron stars from
skyrmion branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

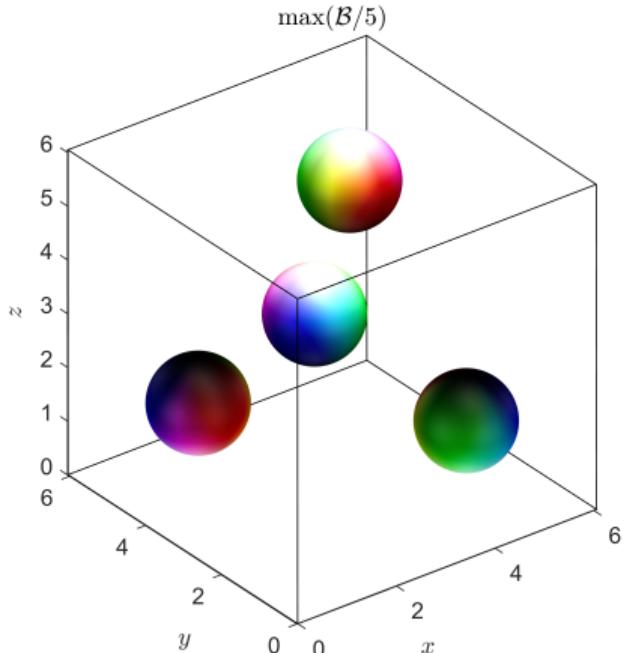
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

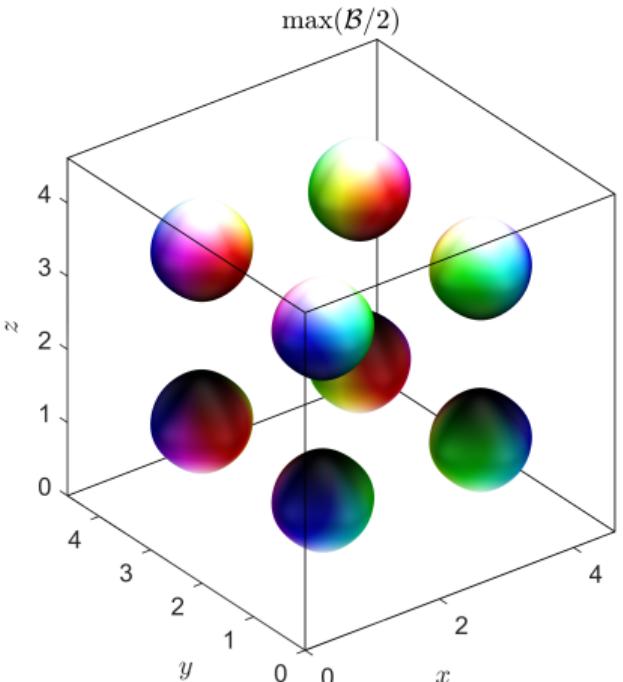
Neutron stars

Final remarks

An example: the FCC to half-skymion crystal



(a) Initial configuration of four $B = 1$ hedgehogs arranged in the attractive channel on a cubic grid



(b) Relaxed final solution of the cubic crystal of half-skymions

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Paul Leask

Motivation

Skyrme model

Linking in the Skyrme model

Skyrmion solutions

Skyrmion crystals and phases of skyrme matter

Quantum skyrme crystals and the symmetry energy

Neutron stars

Final remarks



Phases of skyrmion matter

- Consider fixed baryon density n_B variations of $M_B(\varphi, g)$ w.r.t. g

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter

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- This process enables us to determine an **energy-density** curve
- This is key to obtaining an **equation of state** within our framework

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skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

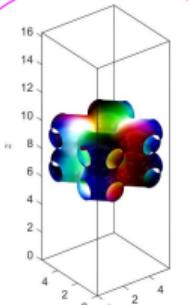
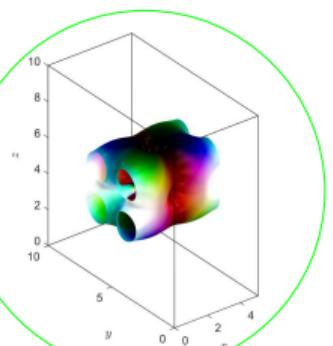
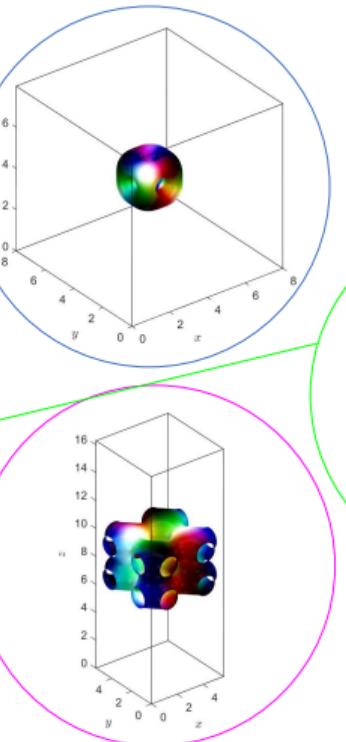
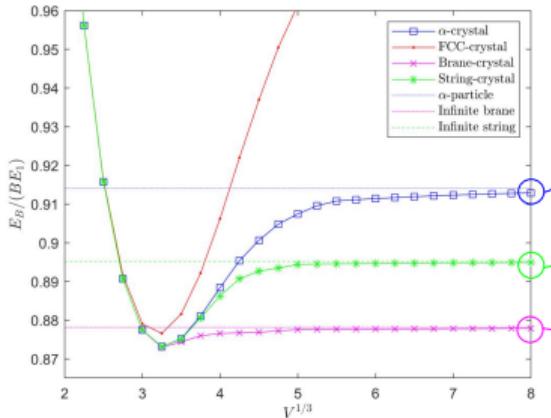
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Phases of skyrmion matter



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Quantum skyrmion crystals and the symmetry energy

Isospin quantization

- Non-renormalizable theory \Rightarrow isospin asymmetry is included by semi-classically quantizing isospin collective coordinates: $\phi(x) \mapsto \hat{\phi}(x, t) = A(t)\phi(x)A^\dagger(t)$ [*Nucl. Phys. B* **262** 133–143 (1985)]

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Can use a mean-field approximation of a large chunk ($B = N_{\text{cell}}B_{\text{cell}}$) in a generic quantum state with fixed eigenvalue [*Phys. Rev. D* **106** 114031 (2022)]

$$I_3 = \frac{(Z - N)}{2} = -\frac{(1 - 2\gamma_p)}{2}N_{\text{cell}}B_{\text{cell}}, \quad \gamma_p \text{ is the proton fraction}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Isospin quantization

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Isospin quantization

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- $I = I_3$ minimizes the isospin energy since by definition $I^2 \geq I_3^2$
- The isospin correction (per unit cell) to the energy of the crystal is found to be

$$E_{\text{iso}} = \frac{b^2}{8U_{33}}B_{\text{cell}}^2\delta^2$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Symmetry energy

- The asymmetry of matter is determined by the isospin asymmetry parameter

$$\delta = (N - Z)/(N + Z) = 1 - 2\gamma_p$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Symmetry energy

- The asymmetry of matter is determined by the isospin asymmetry parameter
 $\delta = (N - Z)/(N + Z) = 1 - 2\gamma_p$
- Binding energy per baryon number of asymmetric nuclear matter is given by

$$\frac{E}{B}(n_B, \delta) = E_N(n_B) + S_N(n_B)\delta^2 + \mathcal{O}(\delta^3), \quad n_0 = 0.160 \text{ fm}^{-3}$$
$$E_N(n_0) = 923 \text{ MeV}$$
$$S_N(n_0) \approx 30 \text{ MeV}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Symmetry energy

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Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- The symmetry energy S_N dictates how the binding energy changes going from symmetric ($\delta = 0$) to asymmetric ($\delta \neq 0$) nuclear matter

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- The symmetry energy S_N dictates how the binding energy changes going from symmetric ($\delta = 0$) to asymmetric ($\delta \neq 0$) nuclear matter
- It is obtained from the quantum isospin energy $S_N(n_B) = \frac{E_{\text{iso}}}{B_{\text{cell}}\delta^2} = \frac{b^2}{8U_{33}}V_{\text{cell}}n_B$

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmi
on crystals and the
symmetry energy

Neutron stars

Final remarks

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$$\frac{E}{B}(n_B, \delta) = E_N(n_B) + S_N(n_B)\delta^2 + \mathcal{O}(\delta^3),$$

$$\begin{aligned} n_0 &= 0.160 \text{ fm}^{-3} \\ E_N(n_0) &= 923 \text{ MeV} \\ S_N(n_0) &\approx 30 \text{ MeV} \end{aligned}$$

- The isospin symmetric binding energy is defined by $E_N = M_B/B$
- The symmetry energy S_N dictates how the binding energy changes going from symmetric ($\delta = 0$) to asymmetric ($\delta \neq 0$) nuclear matter
- It is obtained from the quantum isospin energy $S_N(n_B) = \frac{E_{\text{iso}}}{B_{\text{cell}}\delta^2} = \frac{\hbar^2}{8U_{33}}V_{\text{cell}}n_B$
- In [arXiv:2306.04533 (2023)], at saturation we find

$$n_0 = 0.160 \text{ fm}^{-3}, E_N(n_0) = 912 \text{ MeV} \text{ and } S_N(n_0) = 22.7 \text{ MeV}$$

Neutron stars from
skyrmion branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

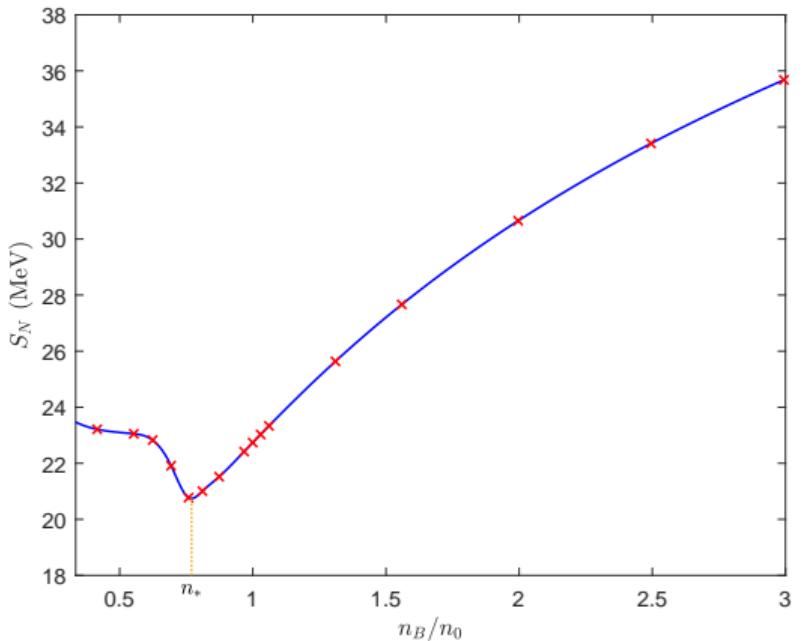
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Symmetry energy and the cusp structure



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

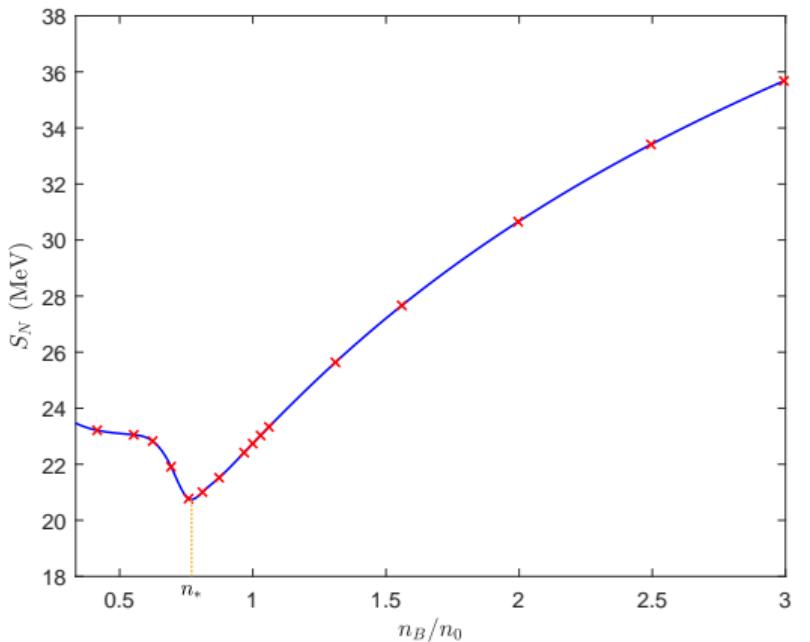
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Symmetry energy and the cusp structure



- **Cusp below saturation at $n_* \sim 3n_0/4$**

Neutron stars from skyrme branes

Paul Leask

Motivation

Skyrme model

Linking in the Skyrme model

Skyrmion solutions

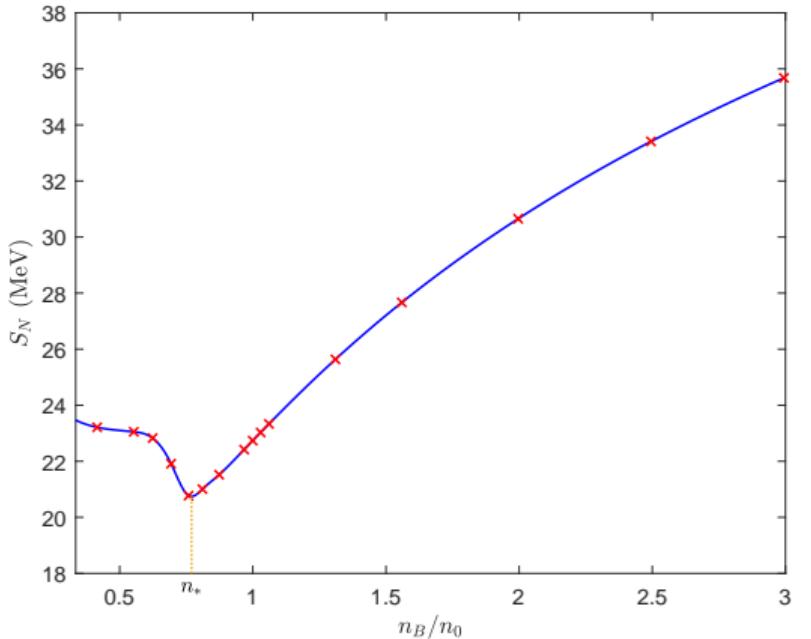
Skyrmion crystals and phases of skyrme matter

Quantum skyrmion crystals and the symmetry energy

Neutron stars

Final remarks

Symmetry energy and the cusp structure



- **Cusp below saturation at**
 $n_* \sim 3n_0/4$
- Symmetry energy at zero density
 $S_N(0) = 23.77 \text{ MeV}$ (finite symmetric nucl. mat.)

Neutron stars from skyrmi
on branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

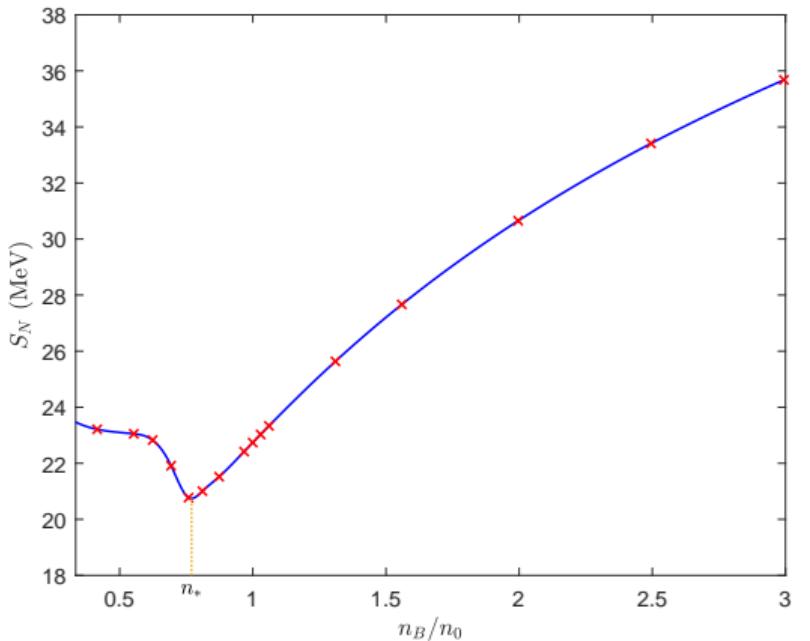
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Symmetry energy and the cusp structure



- **Cusp below saturation at**
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- Symmetry energy at zero density
 $S_N(0) = 23.77 \text{ MeV}$ (finite symmetric nucl. mat.)
- Bethe–Weizsäcker SEMF
asymmetry energy $E_A = a_A \delta^2 B$

Neutron stars from skyrmi
on branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

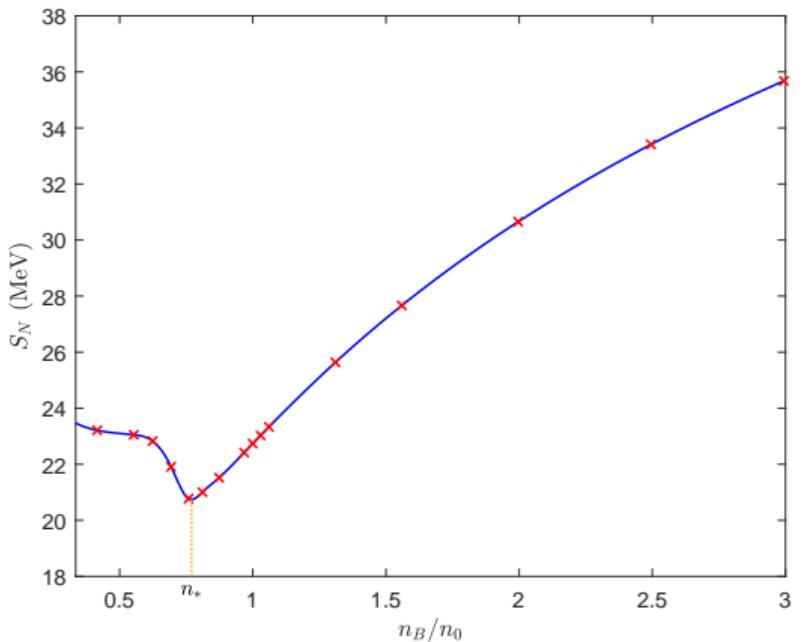
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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$$S_N(0) = 23.77 \text{ MeV}$$
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- Bethe–Weizsäcker SEMF asymmetry energy $E_A = a_A \delta^2 B$
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$$S_N(0) \sim a_A = 23.7 \text{ MeV}$$

Neutron stars from skyrmi
on branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

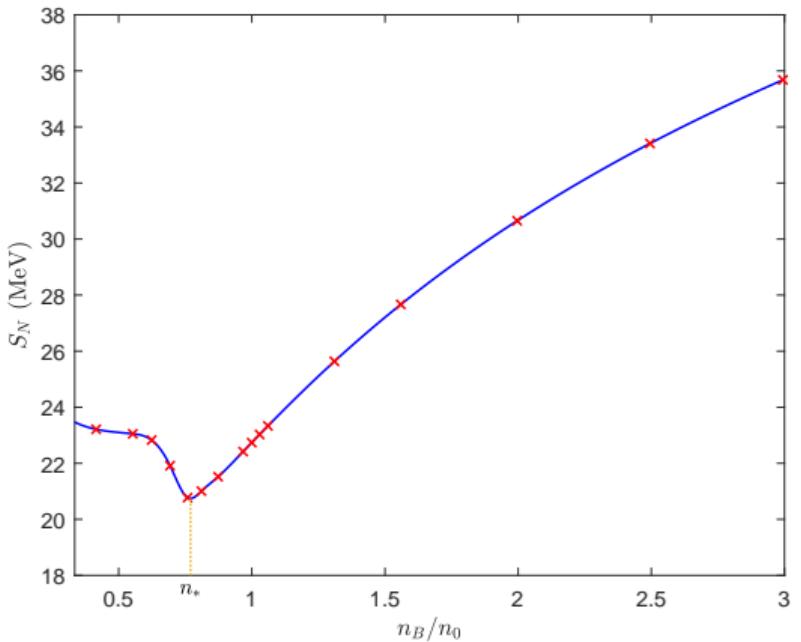
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Symmetry energy at zero density $S_N(0) = 23.77 \text{ MeV}$ (finite symmetric nucl. mat.)
- Bethe–Weizsäcker SEMF asymmetry energy $E_A = a_A \delta^2 B$
- Can identify $S_N(0) \sim a_A = 23.7 \text{ MeV}$
- Cusp origin: **phase transition** between **infinite isospin asymmetric nuclear matter** and somewhat **isolated finite nuclear matter**

[arXiv:2306.04533 (2023)]

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Particle fractions of $npe\mu$ matter in β -equilibrium

- Global **charge neutrality** by including background of charged leptons $n_p = n_e + n_\mu$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium

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- The simultaneous β -decay and electron capture processes allow the calculation of the proton fraction γ_p at a prescribed density n_B [*Phys. Rev. D* **106** 114031 (2022)]

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

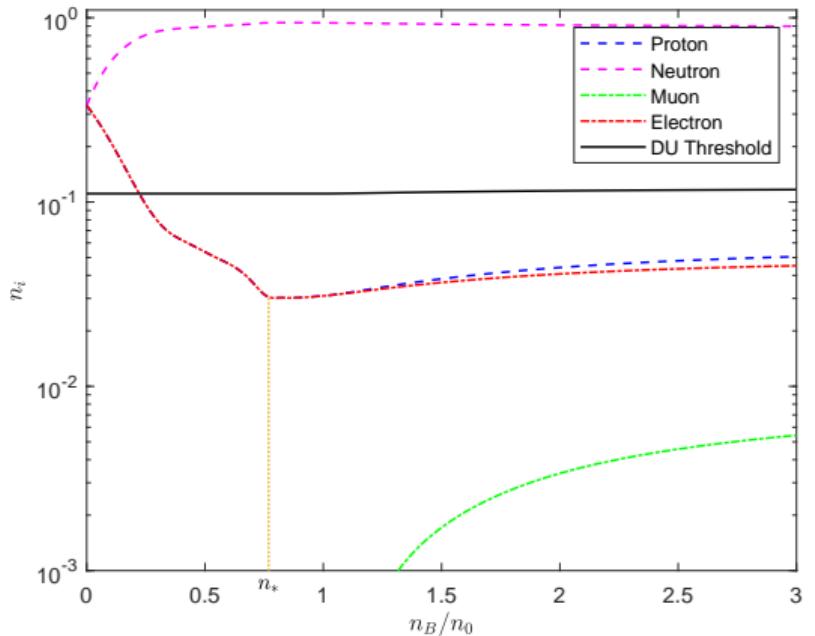
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

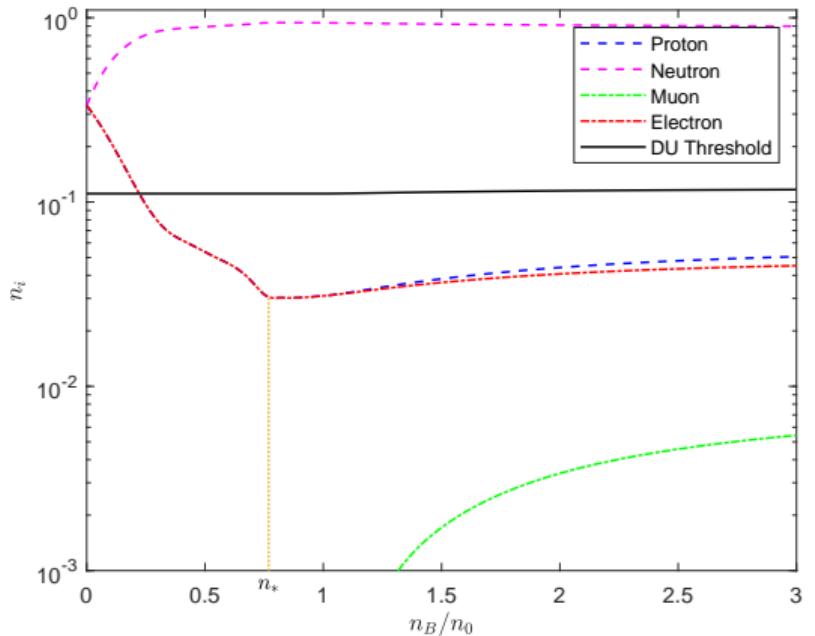
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium



- Cusp also present at n_*

Neutron stars from
skyrmion branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

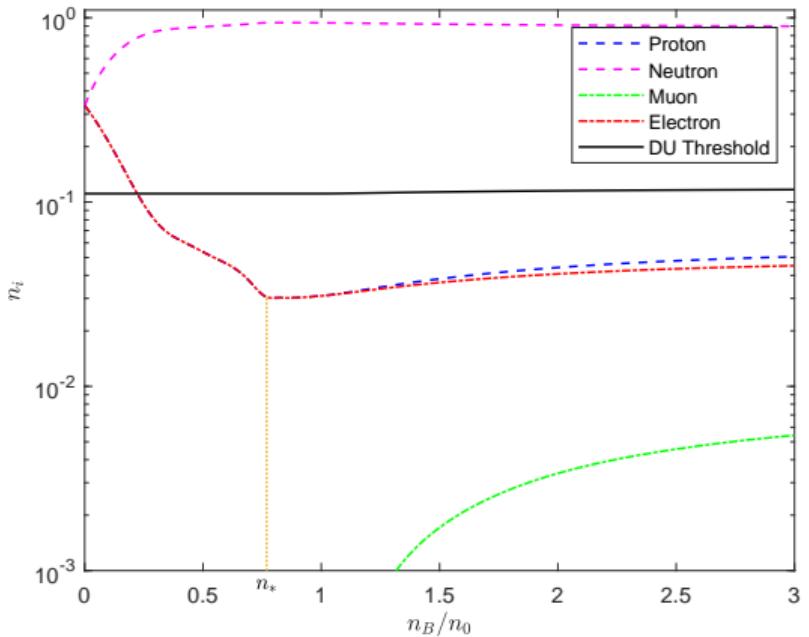
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Reinforces the proposition that the **cusp** indicates the start of a **phase transition** between **infinite asym matter** and **finite sym matter**

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branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

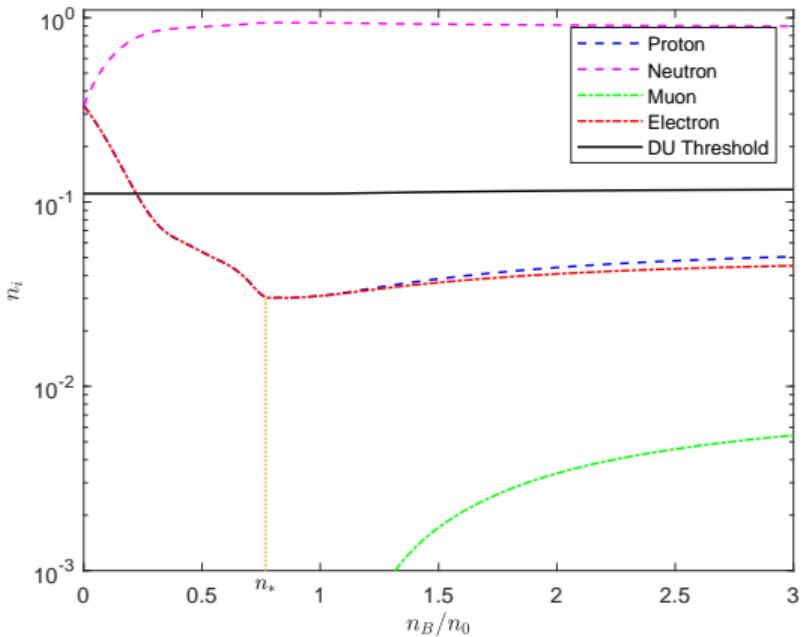
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

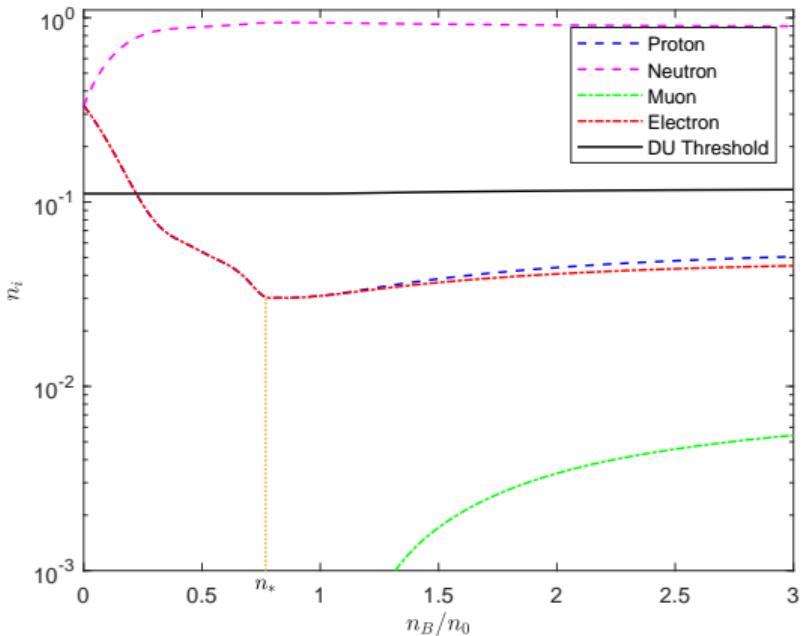
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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skyrmion branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

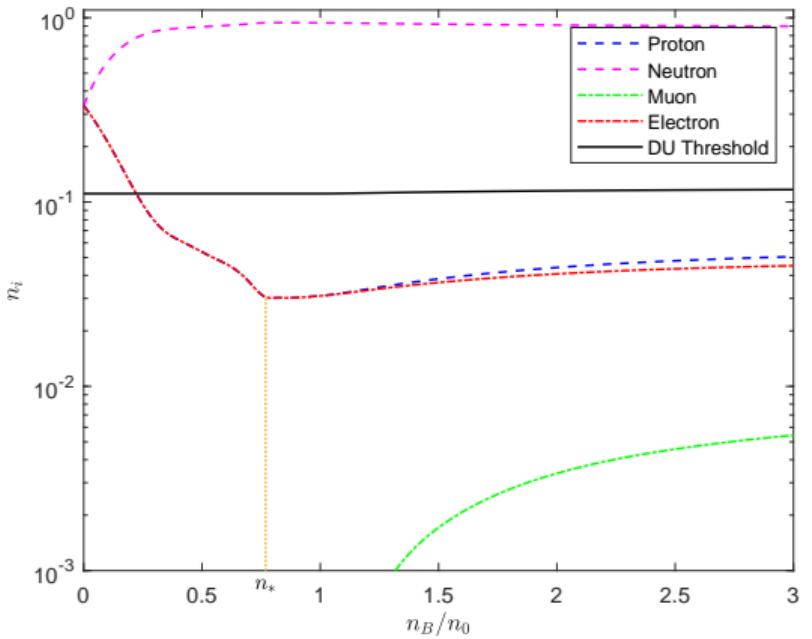
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- **Cusp** also present at n_*
 - Reinforces the proposition that the **cusp** indicates the start of a **phase transition** between **infinite asym matter** and **finite sym matter**
 - The crust of NS is iron rich with $\gamma_p = 0.46$ for ^{56}Fe
 - We find as $n_B \rightarrow 0$ then $\gamma_p = 0.5$
- ⇒ These correspond quite well

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on branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- \Rightarrow Energetically favourable for muons to appear
- The simultaneous β -decay and electron capture processes allow the calculation of the proton fraction γ_p at a prescribed density n_B [*Phys. Rev. D* **106** 114031 (2022)]
 - Energy of a relativistic Fermi gas at zero temperature (lepton energy)

$$E_l(n_B) = \frac{B_{\text{cell}}}{n_B \hbar^3 \pi^2} \int_0^{bk_F} k^2 \sqrt{k^2 + m_l^2} \, dk, \quad k_F = (3\pi^2 n_l)^{1/3}, \quad n_l = \gamma_l n_B$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Particle fractions of $npe\mu$ matter in β -equilibrium

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- Energy per unit cell of β -equilibrated matter

$$E_{\text{cell}}(n_B) = M_B(n_B) + E_{\text{iso}}(n_B) + E_e(n_B) + E_\mu(n_B)$$

Neutron stars from
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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

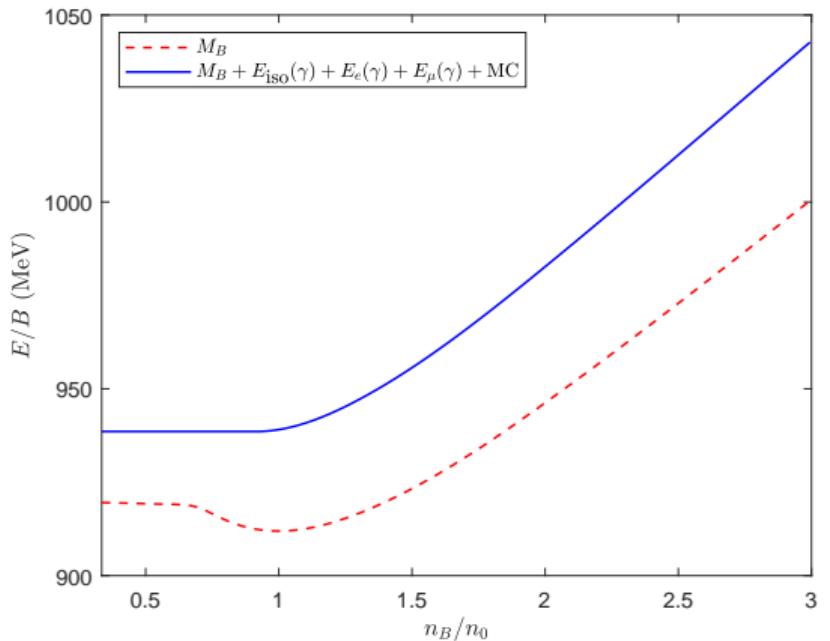
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Isospin asymmetric equation of state



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Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

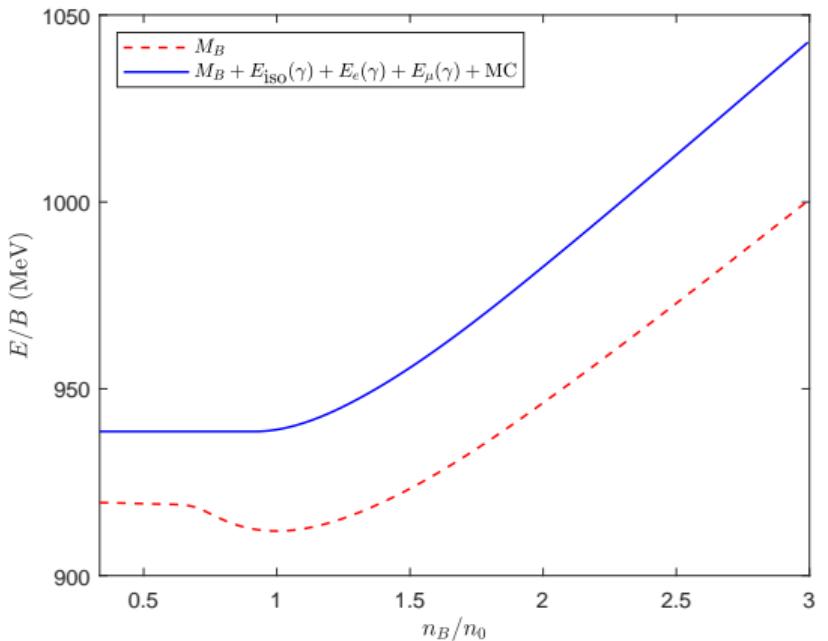
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Isospin asymmetric equation of state



- Can obtain the pressure p and energy density ρ from the $E(n_B)$ curve, with

$$\rho = \frac{E}{V} = \frac{n_B}{B} E_{\text{cell}}$$

$$p = -\frac{\partial E}{\partial V} = \frac{n_B^2}{B} \frac{\partial E_{\text{cell}}}{\partial n_B}$$

Neutron stars from skyrmi
on branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

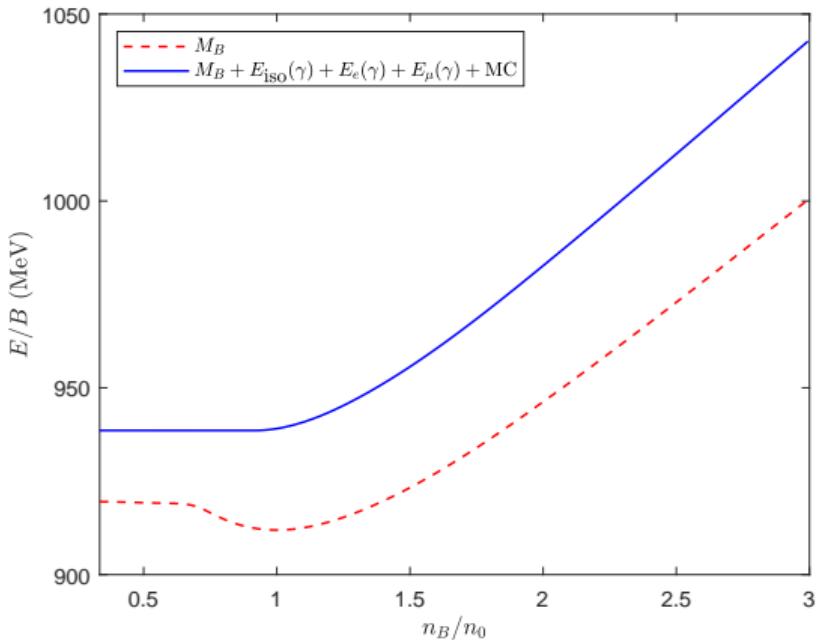
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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⇒ Isospin asymmetric nuclear matter
 EoS $\rho_{\text{brane}} = \rho_{\text{brane}}(p)$

Neutron stars from
 skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
 Skyrme model

Skyrmion solutions

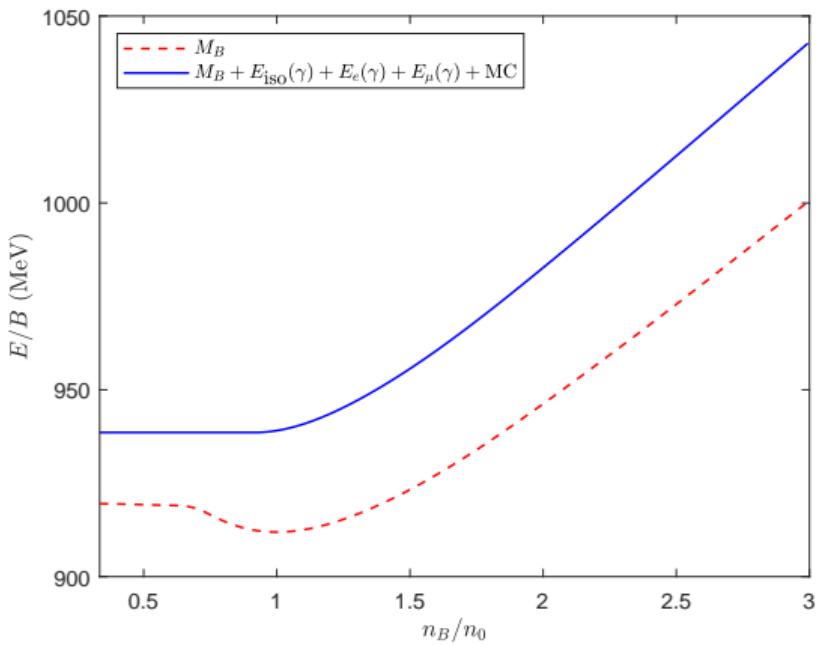
Skyrmion crystals
 and phases of
 skyrmion matter

Quantum skyrmion
 crystals and the
 symmetry energy

Neutron stars

Final remarks

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- \Rightarrow Isospin asymmetric nuclear matter
 EoS $\rho_{\text{brane}} = \rho_{\text{brane}}(p)$
- We will use this EoS to obtain NS within the Skyrme model

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Coupling to gravity

- In order to describe neutrons stars within the Skyrme framework, we need to couple the generalized Skyrme model to gravity

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Coupling to gravity

- In order to describe neutrons stars within the Skyrme framework, we need to couple the generalized Skyrme model to gravity
- Introduce the Einstein–Hilbert–Skyrme action

$$S = \frac{1}{16\pi G} \int_{\Sigma} d^4x \sqrt{-g} R + S_{\text{matter}}$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- S_{matter} describes matter inside NS
- NS Interior well described by **perfect fluid** of nearly free neutrons & degenerate gas of electrons:

$$T_{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\delta S_{\text{matter}}}{\delta g^{\mu\nu}} = (\rho(p) + p) u_{\mu} u_{\nu} + p g_{\mu\nu}$$

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Coupling to gravity

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$$T_{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\delta S_{\text{matter}}}{\delta g^{\mu\nu}} = (\rho(p) + p) u_{\mu} u_{\nu} + p g_{\mu\nu}$$

- The energy density ρ and the pressure p are related by the (Brane) crystal EoS
 $\rho(p) = \rho_{\text{brane}}(p)$ [*Phys. Lett. B* **811** 135928 (2020)]

Neutron stars from skyrmi
on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



The Tolman–Oppenheimer–Volkoff system

- Our aim is to calculate M_{\max} and R_{\max} for a NS described by our system

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



The Tolman–Oppenheimer–Volkoff system

- Our aim is to calculate M_{\max} and R_{\max} for a NS described by our system
- Need to solve the Einstein equations $G_{\mu\nu} = 8\pi G T_{\mu\nu}$ for some particular choice of $g_{\mu\nu}$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



The Tolman–Oppenheimer–Volkoff system

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- Need to solve the Einstein equations $G_{\mu\nu} = 8\pi G T_{\mu\nu}$ for some particular choice of $g_{\mu\nu}$
- Simplest case: **static & non-rotating** neutron star

Neutron stars from
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Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Spherically symmetric ansatz of the spacetime metric

$$ds^2 = -A(r)dt^2 + B(r)dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2) = g_{\mu\nu}dx^\mu dx^\nu$$

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Substituting this into the Einstein equations $G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi GT_{\mu\nu}$ yields the TOV system

$$\frac{dA}{dr} = A(r)r \left(8\pi GB(r)p(r) - \frac{1 - B(r)}{r^2} \right)$$

$$\frac{dB}{dr} = B(r)r \left(8\pi GB(r)\rho(p(r)) + \frac{1 - B(r)}{r^2} \right)$$

$$\frac{dp}{dr} = -\frac{p(r) + \rho(p(r))}{2A(r)} \frac{dA}{dr}$$

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

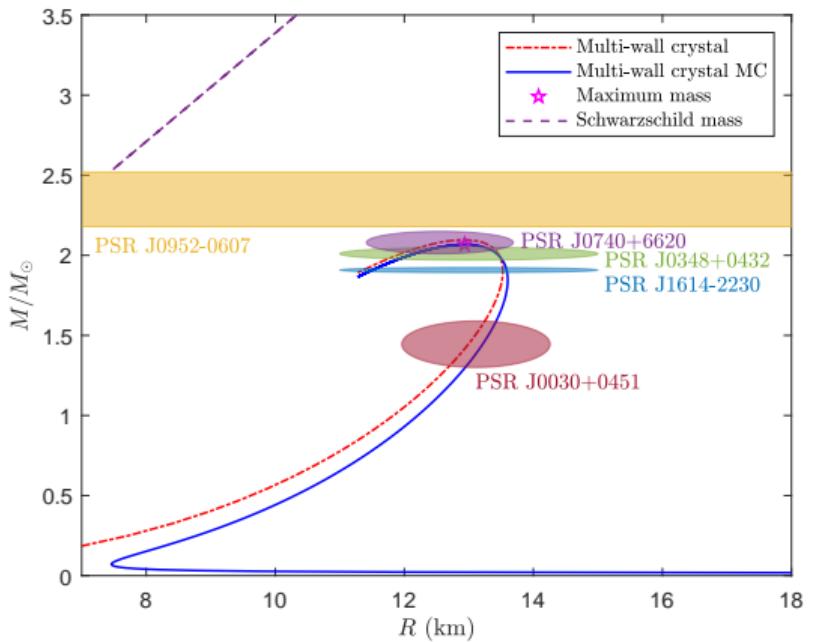
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Neutron star properties and the mass-radius curve



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

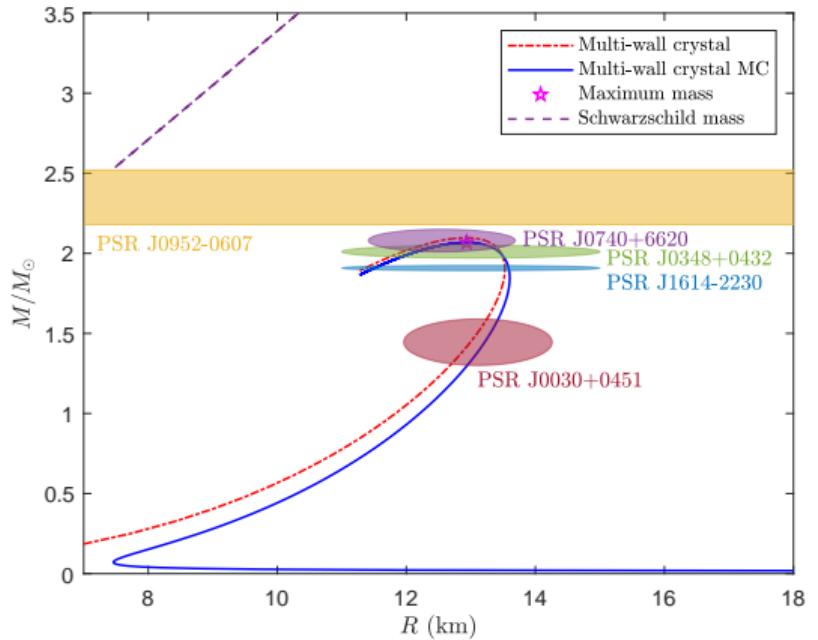
Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

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- Mass M obtained from Schwarzschild metric definition outside the star

$$B(R_{\text{NS}}) = \frac{1}{1 - \frac{2MG}{R_{\text{NS}}}}$$

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Paul Leask

Motivation

Skyrme model

Linking in the Skyrme model

Skyrmion solutions

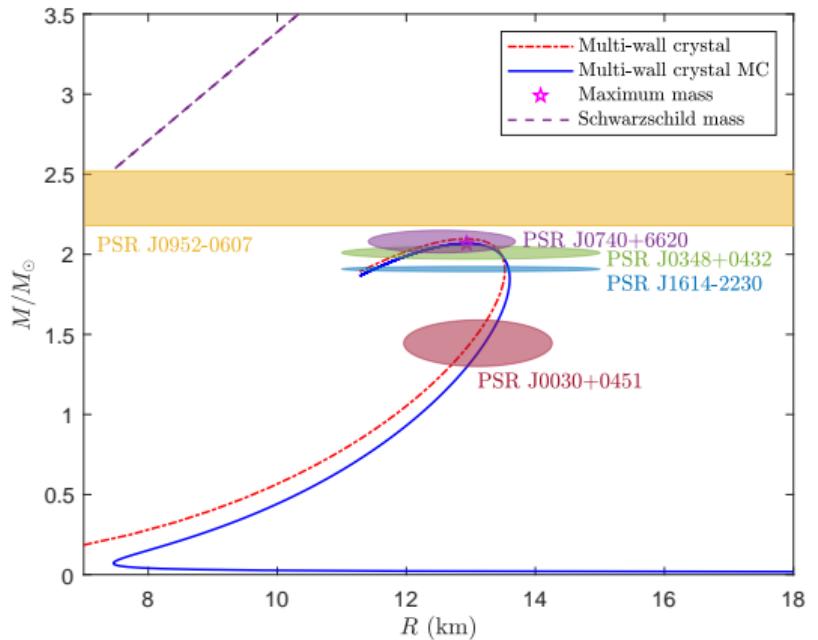
Skyrmion crystals and phases of skyrmi
on matter

Quantum skyrme crystals and the symmetry energy

Neutron stars

Final remarks

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$$B(R_{\text{NS}}) = \frac{1}{1 - \frac{2MG}{R_{\text{NS}}}}$$

- $M_{\text{max}} = 2.0971 M_{\odot}$, occurring for a neutron star of radius $R_{\text{NS}} = 13.12$ km.

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on branes

Paul Leask

Motivation

Skyrme model

Linking in the Skyrme model

Skyrmion solutions

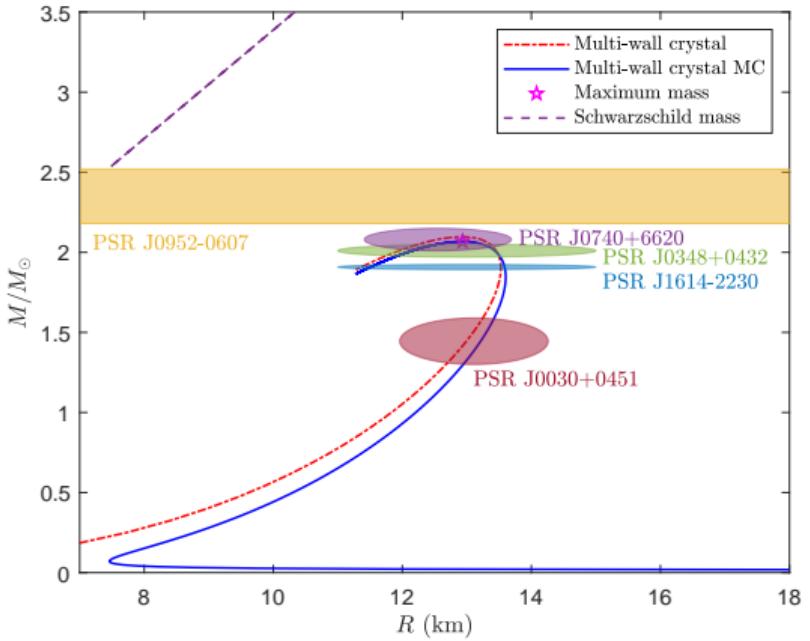
Skyrmion crystals and phases of skyrmi
on matter

Quantum skyrme
crystals and the symmetry energy

Neutron stars

Final remarks

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- $M_{\text{max}} = 2.0971M_{\odot}$, occurring for a neutron star of radius $R_{\text{NS}} = 13.12$ km.
- ⇒ Resulting neutron stars agree well with recent NICER/LIGO observational data

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on branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Final remarks



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- Cusp structure in the symmetry energy observed in the hidden-local-symmetric (HLS) Skyrme model [*Phys. Rev. C* **83** 025206 (2011)]

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skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Final remarks

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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- There is a topological phase transition where the FCC lattice of hedgehog skyrmions fractionalize into half-skyrmions (FCC crystal)
- Analogous to “pseudo-gap” phenomenon in condensed matter physics

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Open problems

- Brane solution improves on compressibility at saturation

Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Open problems

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks



Open problems

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Open problems

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Open problems

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Neutron stars from
skyrmion branes

Paul Leask

Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks

Open problems

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Neutron stars from
skyrmion branes

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Motivation

Skyrme model

Linking in the
Skyrme model

Skyrmion solutions

Skyrmion crystals
and phases of
skyrmion matter

Quantum skyrmion
crystals and the
symmetry energy

Neutron stars

Final remarks