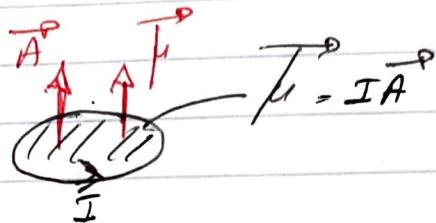


Lect 21

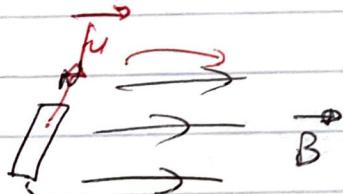
Magnetic Materials, Dia-Para -& Ferromagnetism

* External field = vacuum field.

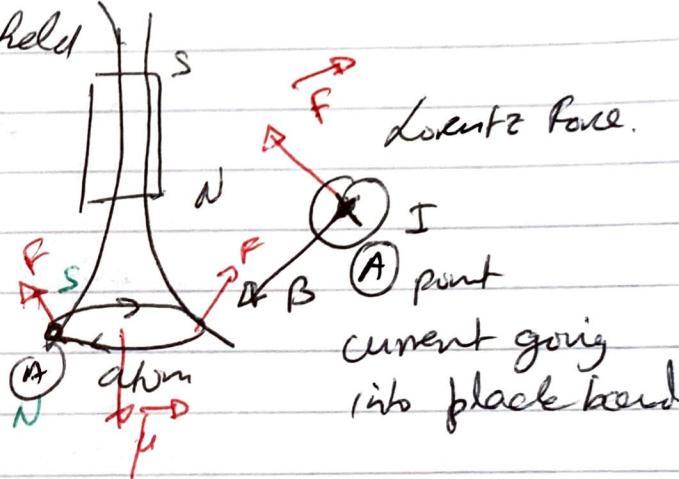
* Magnetic dipole



- dia magnetism (induce dipoles oppose vacum field
(nothing to do with diam's, diam))
(lower rem external field)
- paramagnetism (real)



↳ hair bigger than
vacum field



(∴ Net Force ↑ upwards)

- * Ferromagnetism (permanent dipole moments)
Lo Domains

$$- \overrightarrow{B} = \underline{k_M} \overrightarrow{B_{\text{vac}}}$$

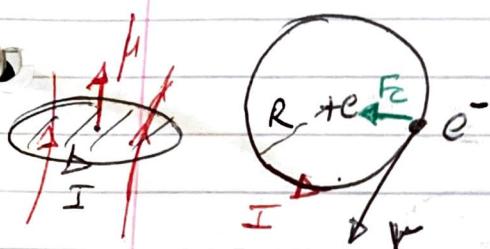
Lorentzian permeability.

$$k_M = 1 + \frac{\chi_M}{\chi_{\text{mag}}} \quad \begin{matrix} \text{susceptibility} \\ \text{magnetic susceptibility} \end{matrix}$$

Lect 22

- Maxwell's Equations, 600 Differenzen
- Field inside magnet - Eg Ferromagnetic, paramagnetic.

$$\vec{B} = \mu_m \frac{\vec{B}_{\text{vac}}}{(\text{ext})}$$



$$m_e = 9.11 \times 10^{-31} \text{ kg}$$

$$q_e = 1.6 \times 10^{-19} \text{ C}$$

$$R = 5 \cdot 10^{-6} \text{ m.}$$

(curr radius)

$$\mu = IA, A = \pi R^2 = 8 \times 10^{-21} \text{ m}^2$$

(Magnetic

dipole moment) $F_c = \frac{e^2}{4\pi\epsilon_0 R^2} = \frac{mv^2}{R}$ } centrifugal force

$$v = \sqrt{\frac{e^2}{m 4\pi\epsilon_0 R}} \approx v = 2.3 \times 10^6 \text{ m/s}$$

$$(\text{time}) T = \frac{2\pi R}{v} = 1.4 \times 10^{-16} \text{ s}$$

$$I = \frac{dQ}{dT} = \frac{e}{T} \approx 1.1 \times 10^{-3} \text{ A.} \therefore \underline{\mu \approx 9.3 \times 10^{-24} \text{ Am}^2}$$

Biot-Savart law

- $\vec{B} = \vec{B}_{\text{vac}} + \vec{B}'$ depends on B_{vac} & Temperature
 align dipoles.

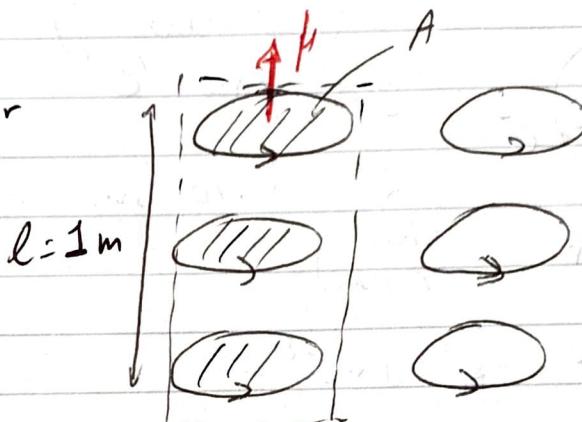
$$\vec{F} \\ \vec{B}' \propto \vec{B}_{\text{vac}}$$

$$\therefore \vec{B}' = \chi_M \vec{B}_{\text{vac}} \quad \vec{B} = (1 + \chi_M) \vec{B}_{\text{vac}}$$

$$B_{\text{vac}} = K_M \vec{B}_{\text{vac}}$$

>

$$\mu = 2\mu_{\text{Bohr}}$$



Solenoid: $\vec{B} = \mu_0 \frac{I N}{l}$ $= \mu_0 I A N$

$$N = 10^{29} \text{ m}^{-3}$$

$$\text{Vol} = A \text{ m}^3$$

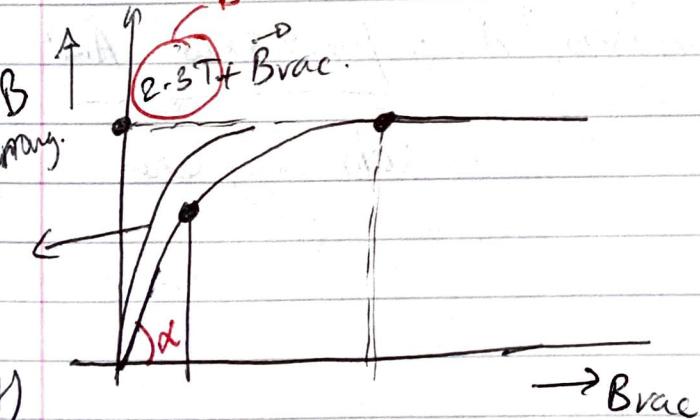
$$= \mu_0 2 \mu_{\text{Bohr}} N \\ \approx 2 \cdot 3 \text{ T}$$

(# density of
atoms/
molecules)

$A N$ = # of windings/m

> * Not on scale

$$\vec{B}'$$

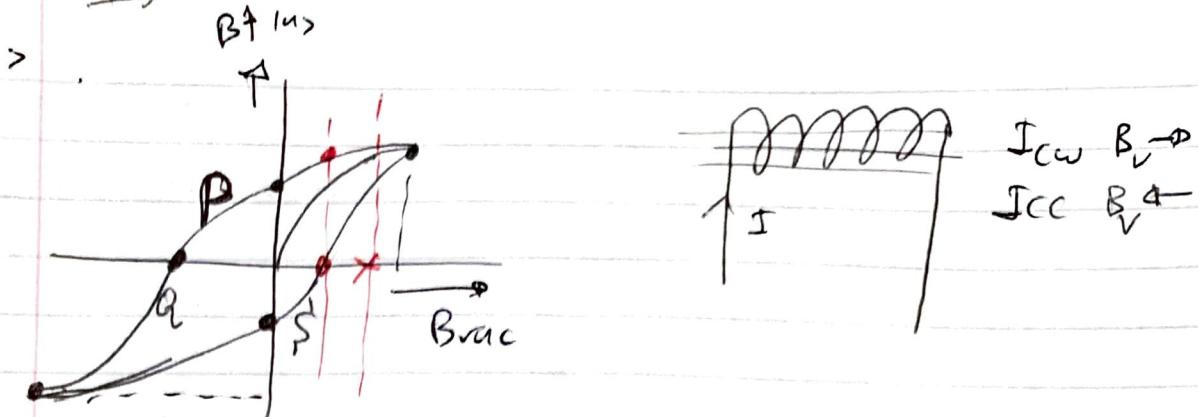


$$B_{\text{vac}} = \mu_0 I / N / l$$

$$K_M = 10^3$$

$$\tan \alpha = K_M = 10^3$$

Hysteresis Curve



(CW) At location P

$$Br_{ac} = 0 \\ B' \rightarrow$$

(CCW) Q $Br_{ac} \leftarrow$

$$B = 0 \\ \text{Total field} \\ \left. \begin{array}{l} \text{Field} \\ \text{Field} \end{array} \right\} \text{Vectorially added} \\ \therefore \phi$$

* Effect of Magnetic Materials on Maxwell Equations ??

Maxwell's equations

$$- \oint \vec{E} \cdot d\vec{A} = \frac{Q_{\text{free}}}{\epsilon_0}$$

Why use an inductor
instead of
resistor?
power

$$- \oint \vec{B} \cdot d\vec{l} = 0 \quad * \text{magnetic monopoles} != 1$$

$$- \oint \vec{E} \cdot d\vec{l} = - \frac{d\phi_B}{dt} \quad \text{no Faraday's law}$$

$$\text{Ampere's Law} - \oint \vec{B} \cdot d\vec{l} = \mu_0 (I_{\text{pen}} + \epsilon_0 \kappa \frac{d\phi_E}{dt})$$

material.

e.g. ferromagnetic ...

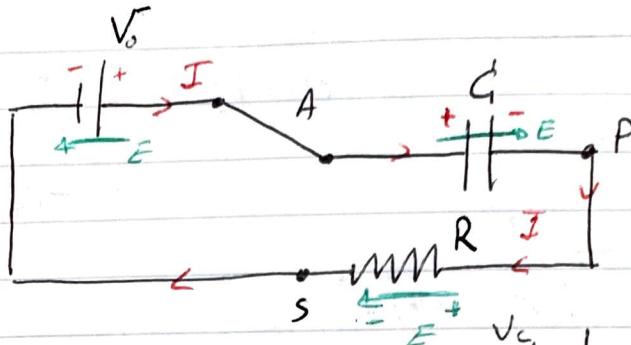
Sheet 24

- Transformers, Car coils, RC Circuits

» RC circuit

$$V_C = V_A - V_P$$

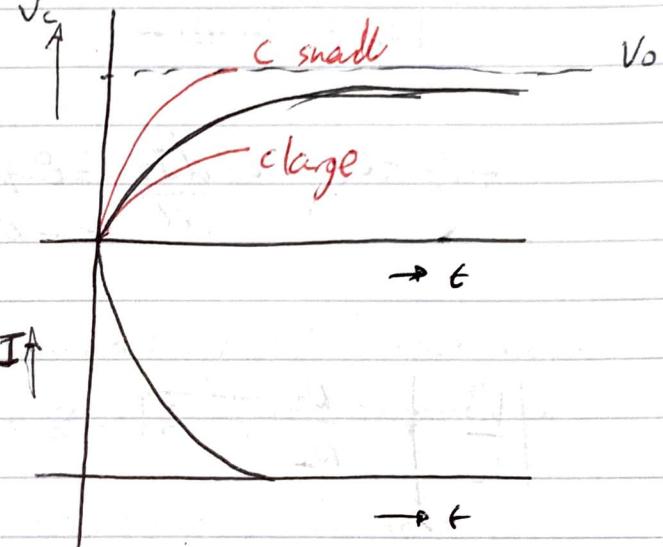
$$V_R = IR = V_P - V_S$$



At $t=0$, $V_C = 0$

$t > 0$, $V_C \uparrow$, $I \downarrow$

$t \rightarrow \infty$, $V_C = V_0$, $I = 0$



$$\oint \vec{E} \cdot d\vec{l} = 0$$

$$I = \frac{dQ}{dt} \quad V_C = \frac{Q}{C}$$

$$\therefore +V_C + IR - V_0 = 0.$$

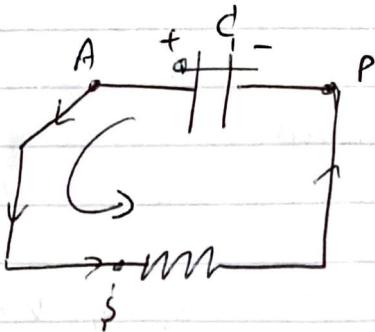
$$\frac{Q}{C} + \frac{dQ}{dt} R - V_0 = 0 \quad \text{~differential equation.}$$

$$- Q = V_0 C (1 - e^{-t/RC})$$

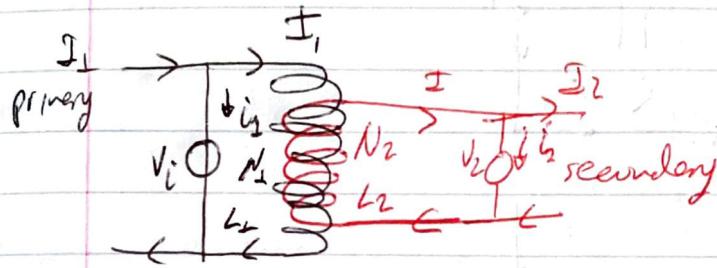
$$I = \frac{dQ}{dt} = V_0 C / (f) C / (-) \frac{1}{RC} e^{-t/RC}$$

$$I = \frac{V_0}{R} e^{-t/RC}$$

$$- V_C = \frac{Q}{C} = V_0 (1 - e^{-t/RC})$$



Transformers AC



$$0 - V_1 = -L \frac{dI_1}{dt} = E_{ind\ 1}$$

$$= -N_1 \frac{d\phi_B}{dt}$$

$$+V_2 + 0 = -L_2 \frac{dI_2}{dt} = E_{ind\ 2}$$

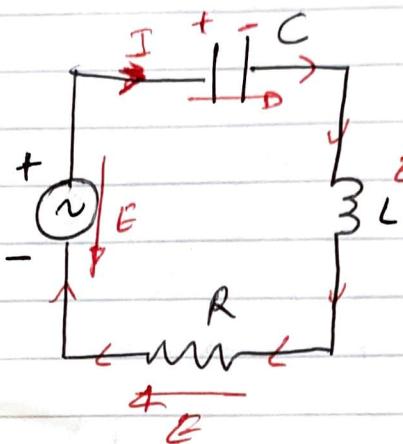
$$= -N_2 \frac{d\phi_B}{dt}$$

$$\left| \frac{V_2}{V_1} \right| = \frac{N_2}{N_1}$$

Lect 25

- Driven LRC Circuits, Metal Detectors

> LRC



$$V_o = V_0 \cos(\omega t)$$

$$\oint E \cdot d\ell \neq \phi$$

$$-V_C + 0 + IR - V_0 \cos(\omega t) = -\frac{L dI}{dt}$$

$$I = \frac{dQ}{dt}$$

$$\Rightarrow \frac{d^2Q}{dt^2} + \frac{RdQ}{dt} + \frac{Q}{C} = V_0 \cos(\omega t) \quad //$$

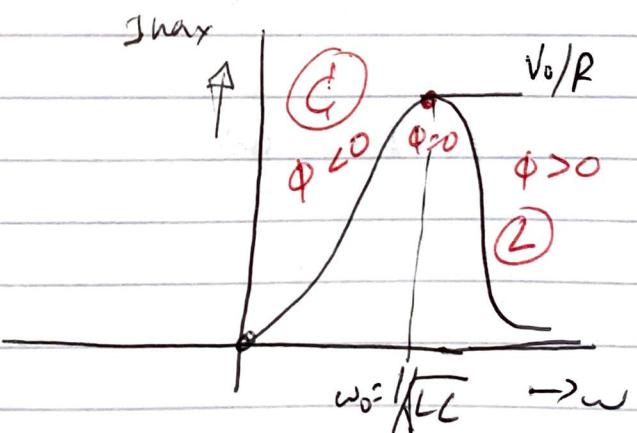
$$I = \frac{V_0}{\sqrt{R^2 + (C\omega L - \frac{1}{\omega C})^2}} \cos(\omega t - \phi) \quad (\text{effective resistance})$$

$$\Rightarrow Z = \sqrt{R^2 + X^2} = \text{impedance}$$

$$\tan \phi = \frac{\omega L - 1/\omega C}{R} \quad \left. \right\} \rightarrow X_{\text{reactance}} = \omega L - 1/\omega C$$

$$\left(\omega_0 = \frac{1}{\sqrt{LC}} = \text{resonance} \right) \quad : \quad Z = R$$

$$I_{\text{max}} = \frac{V_0}{R} \quad \phi = 0$$



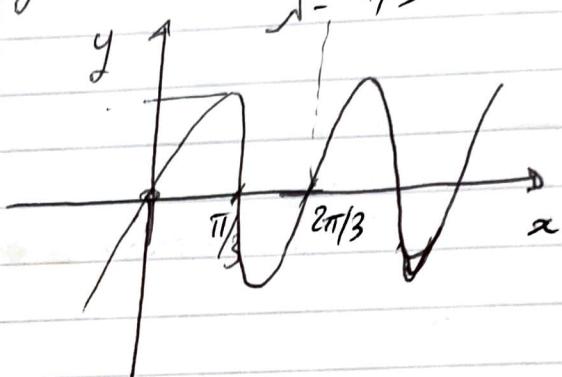
Lesson 26

- Travelling waves, standing waves, musical instruments

$$v + x \rightarrow x - vt$$

$$-x \rightarrow x + vt$$

$$y = 2 \sin(3)x$$



$$k = \frac{2\pi}{\lambda} = 3 \rightarrow \text{wave number}$$

Standing Wave

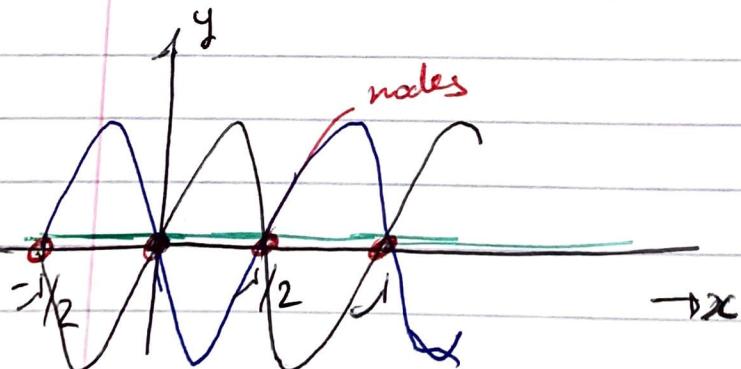
$$y_1 = y_0 \sin(kx - \omega t)$$

$$y_2 = y_0 \sin(kx + \omega t)$$

$$y = y_1 + y_2 = 2y_0 \underbrace{\sin(kx)}_{\text{spatial information}} \underbrace{\cos(\omega t)}_{\text{time information}}$$

spatial information

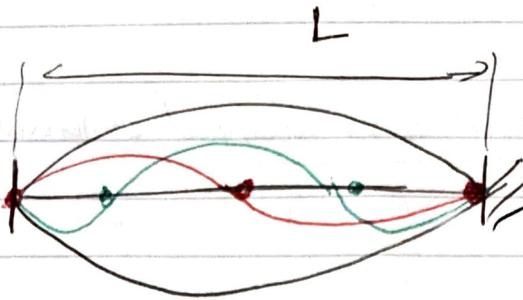
time information



$$t=0$$

$$-t = 1/4T$$

$$t = 1/2T$$



$$\text{Fundamental } f_1 = \frac{2L}{\lambda} f_1 = \frac{v}{2L}$$

$$2^{\text{nd}} \text{ Harmonic } f_2 = L f_2 = \frac{v}{\lambda}$$

$$3^{\text{rd}} \text{ Harmonic } f_3 = 3f_1$$

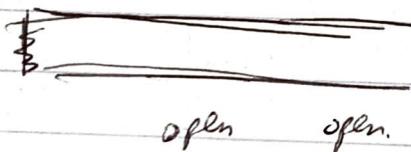
$$f_n = \frac{2L}{\lambda} = \frac{2L}{n} ; f_n = \frac{nV}{2L}$$

$$V = \sqrt{\frac{\text{Tension}}{\text{mass/length}}}$$

(wind)

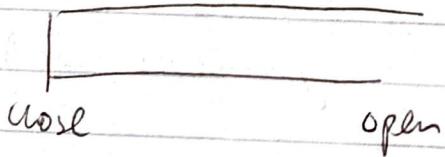


$$V = 340 \text{ ms}^{-1} \text{ (speed of sound)}$$



open open.

$$V = \sqrt{\frac{\text{Tens}}{M}} \sim \text{molecular weight}$$



close open

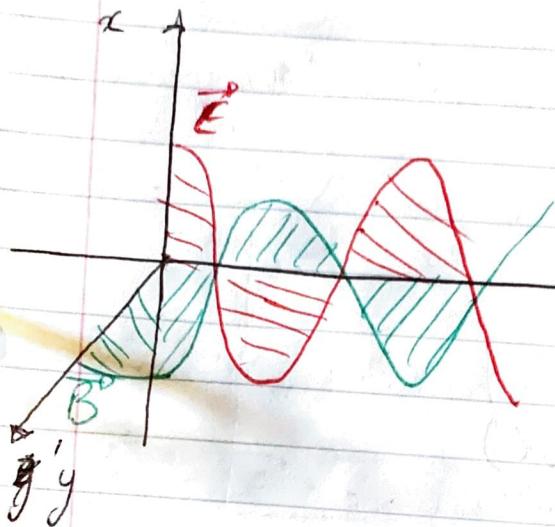
Leet 27

- Destructive Resonance, Electromagnetic Waves, Speed of light

$$- \vec{E} = E_0 \hat{x} \cos(kz - \omega t)$$

$$- \vec{B} = B_0 \hat{y} \cos(kz - \omega t)$$

$$\hat{x} \times \hat{y} = \hat{z}$$



$$\rightarrow v = \frac{\omega}{k} = c$$

$$\rightarrow z \quad \lambda = \frac{2\pi}{k}$$

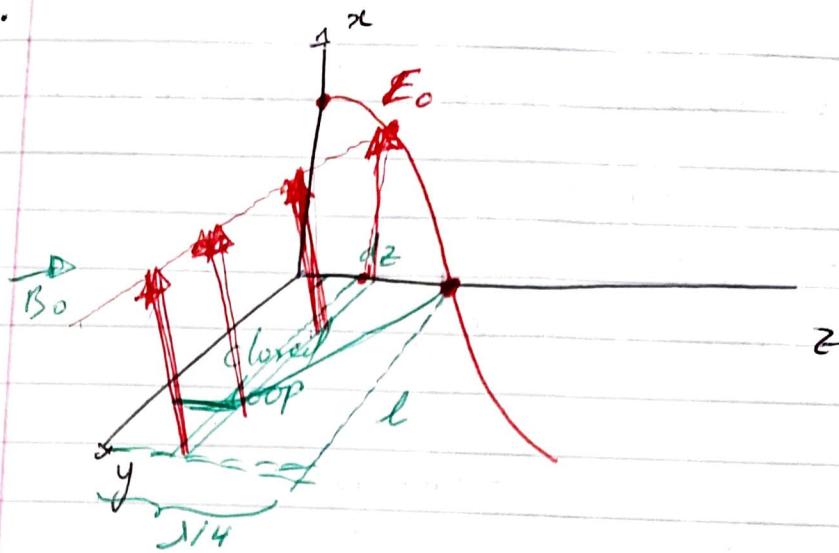
$$B_0 = \frac{E_0}{c}$$

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}}$$

$$\epsilon_0 = 8.85 \times 10^{-12}$$

$$\mu_0 = 1.26 \times 10^{-6}$$

$$c = 2.99 \times 10^8 \text{ ms}^{-1}$$



$$\oint \vec{B} \cdot d\vec{l} = \epsilon_0 \mu_0 \frac{d\phi_E}{dt}$$

$$\phi_E = \int_0^{\lambda/4} l \cdot dz E_0 \cos(kz - \omega t)$$

$$\frac{d\phi_E}{dt} = l E_0 (\omega) \int_0^{\lambda/4} (-\sin(kz - \omega t)) dz$$

$$\frac{d\phi_E}{dt} = \frac{l E_0 \omega}{k} \left[-\cos(kz) \right]_0^{\lambda/4} = l E_0 c$$

$$B_0 \cdot l = \epsilon_0 \mu_0 = l E_0 c$$

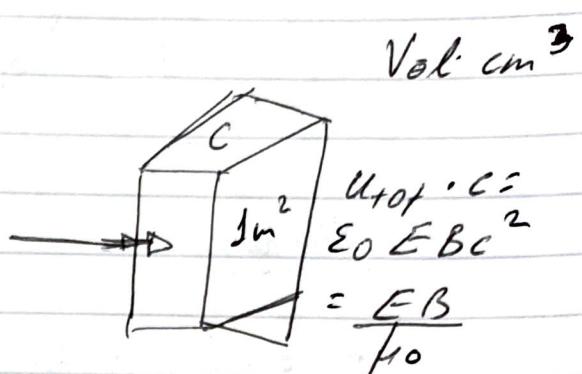
Lect 28

- Poynting Vector, Oscillating charges, Polarization, Radiation pressure

$$- U_E = \frac{1}{2} \epsilon_0 E^2 \quad [\text{J/m}^3]$$

$$- U_B = \frac{1}{2\mu_0} B^2 \quad [\text{J/m}^3]$$

$$= \frac{1}{2\mu_0} \frac{E^2}{c^2} = \frac{1}{2} \epsilon_0 E^2$$



$$\text{Total } U_{tot} = \epsilon_0 E^2 = \epsilon_0 E B c$$

[J/m² sec]

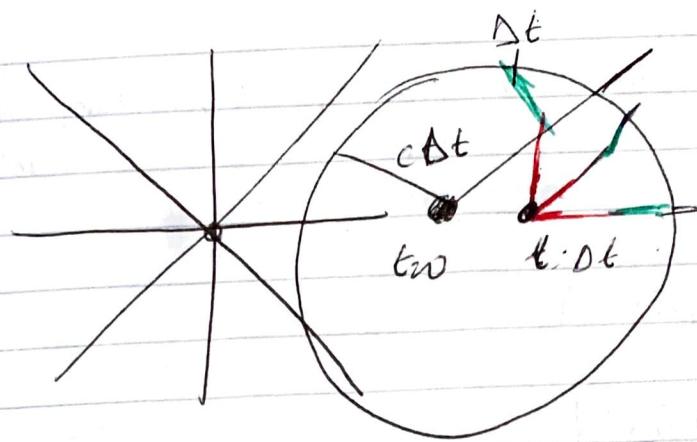
$$\vec{S} = \frac{\vec{E} \times \vec{B}}{\mu_0} \quad [\text{W/m}^2]$$

(Poynting
vector)

$$\langle \vec{S} \rangle = \frac{1}{2} \frac{\epsilon_0 B_0}{\mu_0} = \frac{1}{2} \frac{\epsilon_0 E^2}{\mu_0 c}$$

$$\bar{\epsilon}_0 = 100 \text{ V/m}$$

$$\langle \vec{S} \rangle = \frac{100^2}{2\mu_0 c} = 13 \text{ W/m}^2$$

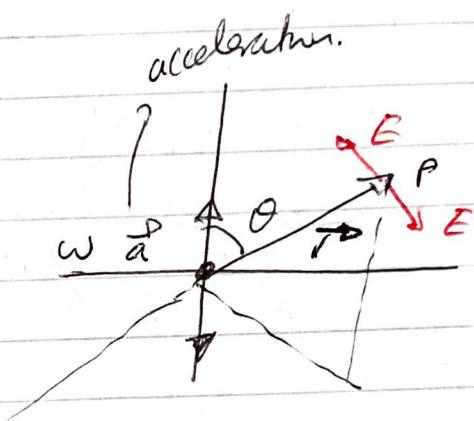


- Photons:

$$p = \frac{\text{energy of photon}}{c}$$

$$E \propto q \frac{a \sin \theta}{r}$$

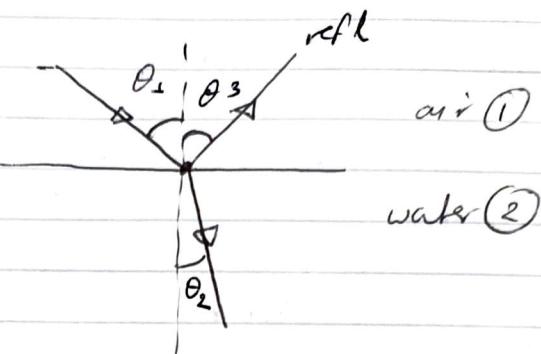
$$S \propto \frac{q^2 a^2 \sin^2 \theta}{r^2}$$



Lect 29

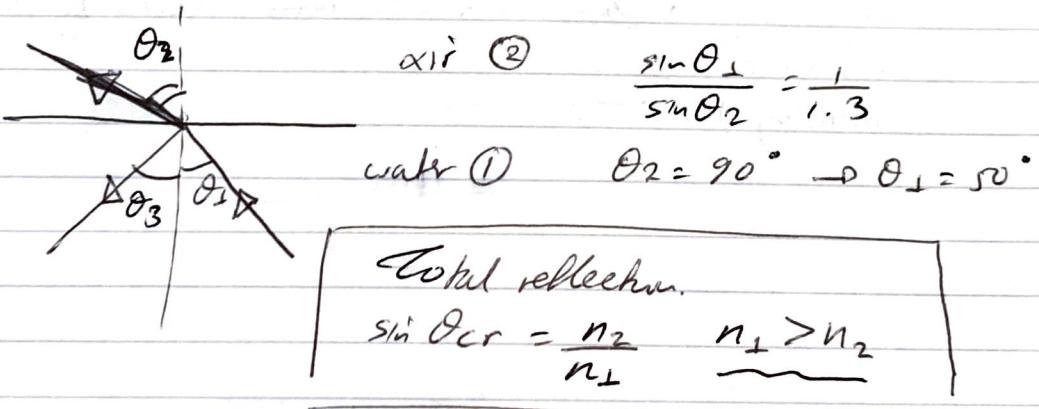
- Snell's Law, Index of Refraction, Huygen's Principle, Illusion of Color

>



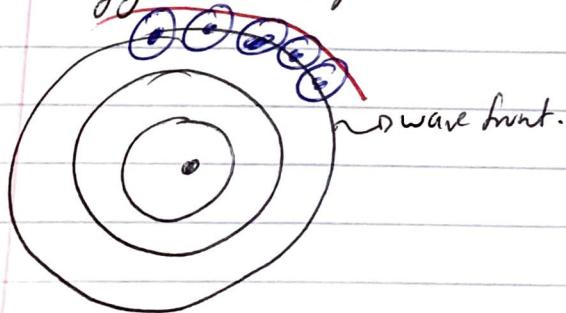
$$\text{Snell's Law: } \frac{\sin \theta_1}{\sin \theta_2} = 1.3 \quad (\text{air} \rightarrow \text{water}) = \frac{n_2}{n_1} \approx \text{refraction.}$$

$n_1 \approx \text{incident}$



* Fiber-optic cables

* Huygen's Principle



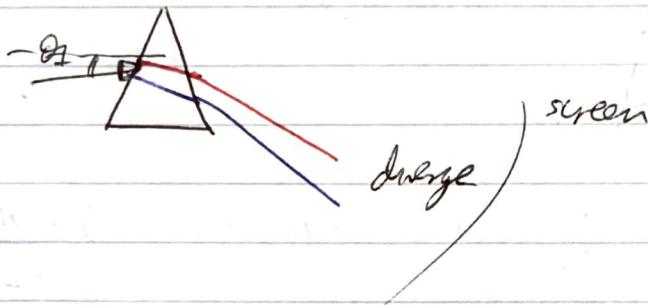
$$V = \frac{1}{\sqrt{\epsilon_0 \mu_0 K M}} = \frac{C}{\sqrt{K M}} = \frac{C}{n}$$

↓
dielectric constant → negative permeability.

Water

$$n_{red} = 1.331$$

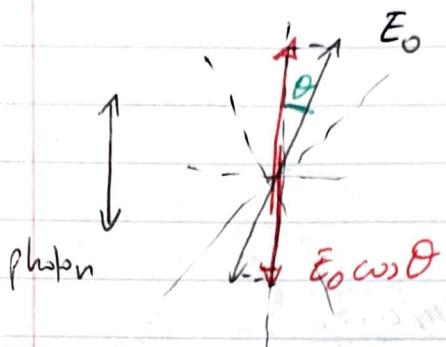
$$n_{blue} = 1.343$$



- Babinet Top. draw interference ~~diffraction~~ colors by pair different phases (X)
(interesting).

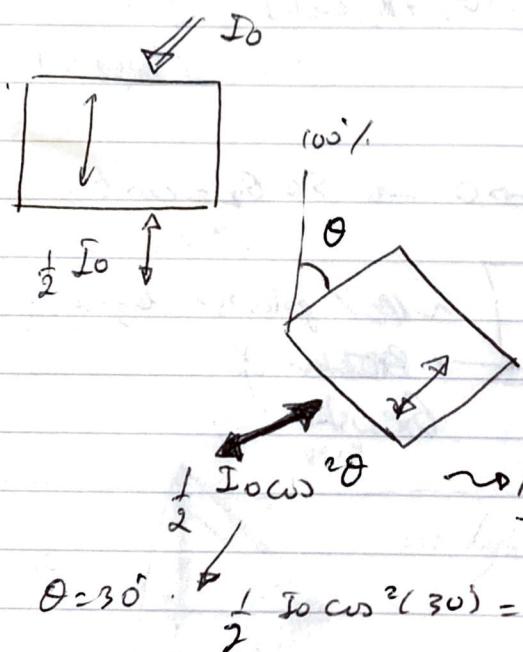
sheet 30

- Polarizers, Malus's Law, light scattering, Blue skies, Red sunsets



$$S \propto E_0^2$$

$$I_0 \text{ unpol} \rightarrow \frac{1}{2} I_0 100\% \text{ pol}$$

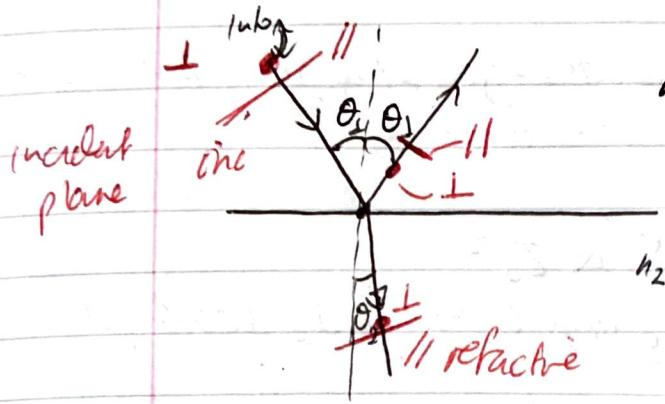


$$\theta = 30^\circ \quad \frac{1}{2} I_0 \cos^2(30^\circ) =$$

$$\theta = 0^\circ \quad \frac{1}{2} I_0$$

$$\theta = 90^\circ \quad \phi \text{ at } (90) = 0 \quad \text{across polarizers}$$

* All arguments perpendicular



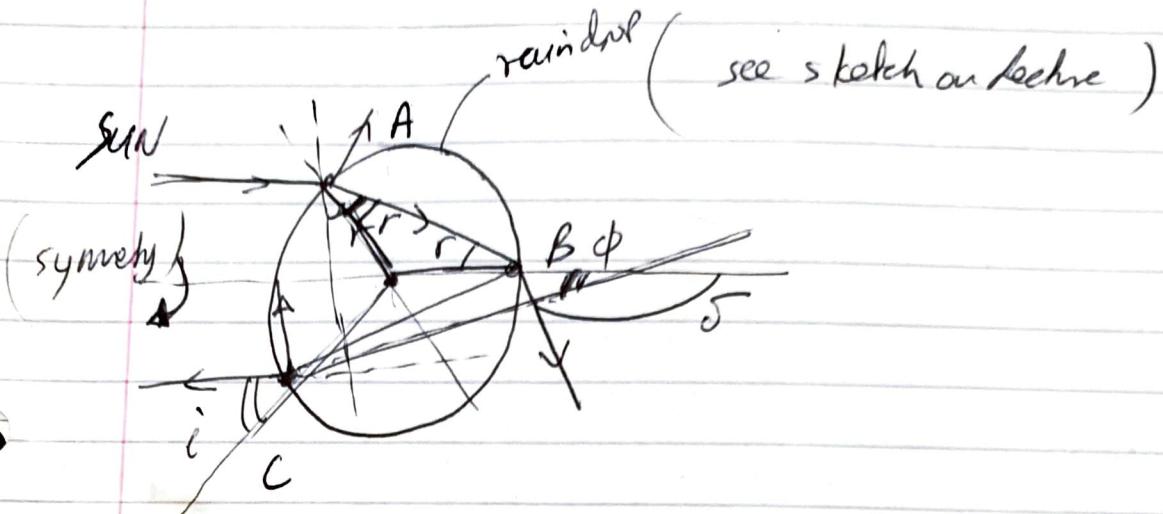
$$\frac{\sin \theta_s}{\sin \theta_2} \left[\frac{n_2}{n_s} = \tan \theta_{\perp} \right] \sim 100\% \text{ polarised light}$$

\rightarrow Brewster & Brewster

Lect 31

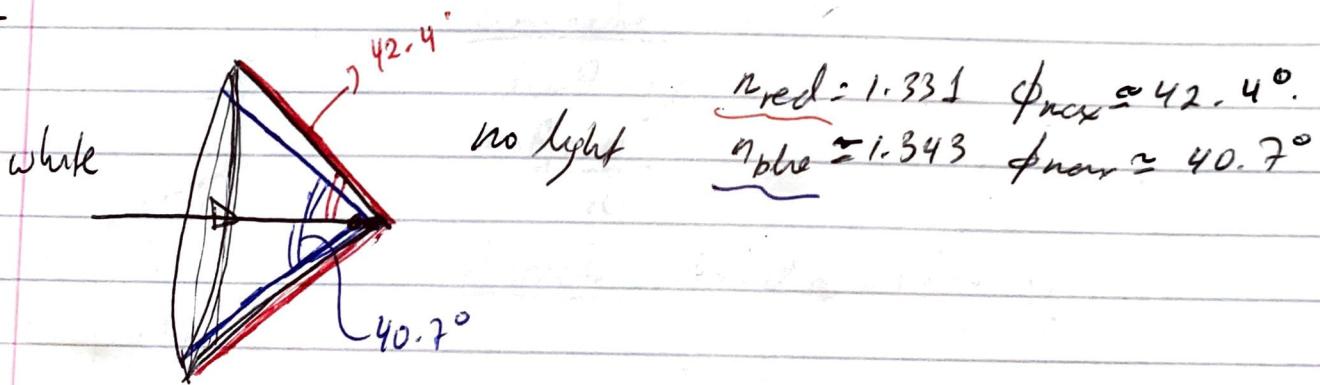
- Rainbows, Fog Bows, Haloes, Glories, sun Dogs

$$\delta = 180^\circ + 2i - 4r$$



$$8 \text{ min} \approx 138^\circ$$

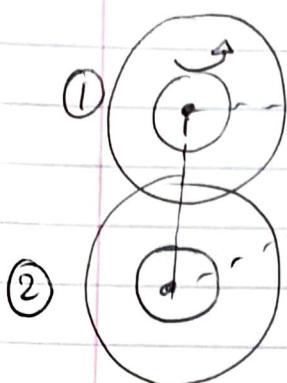
$$\phi_{\max} \approx 42^\circ$$



* Nice explanation in lecture video.

Lect 33

- Double-slit interference, Interferometers



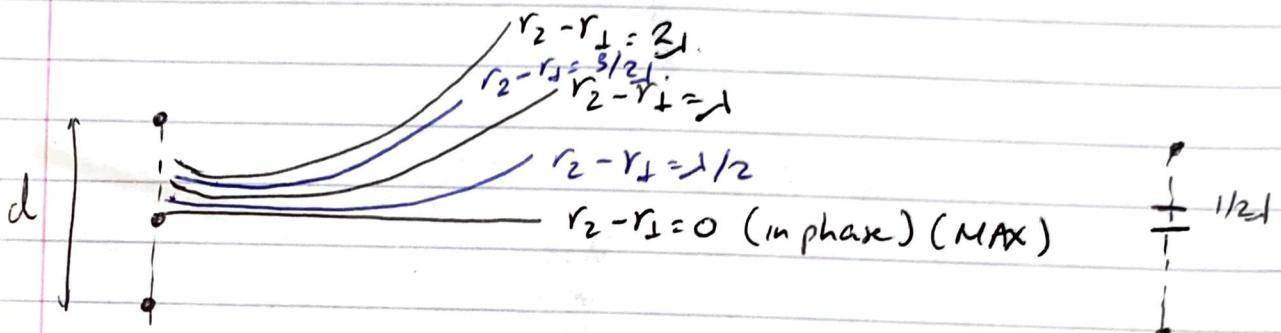
• destructive

$$r_2 - r_1 = \pm \frac{1}{2} \lambda, \pm \frac{3}{2} \lambda, \dots$$

$$r_2 - r_1 = \frac{(2n+1)}{2} \lambda \quad n=0, \pm 1, \pm 2, \dots$$

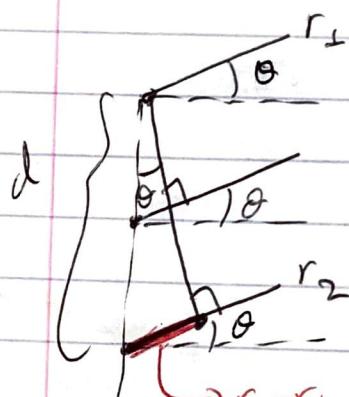
• constructive

$$r_2 - r_1 = n\lambda. \quad n=0, n \neq 1, n \neq 2, \dots$$



$$\#_{\max} = \frac{2d}{\lambda} \uparrow$$

$$\#_{\min} = \frac{1}{2}$$



constructive

$$ds \sin \theta = n\lambda$$

$$\sin \theta = \frac{n\lambda}{d}$$

destructive: $ds \sin \theta = (2n+1) \frac{\lambda}{2}$

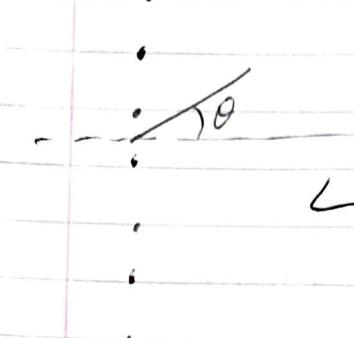
$$\sin \theta_n = (2n+1) \frac{\lambda}{2d}$$

(* Slit-experiment)

Sheet 34

- Diffraction, Gratings, Resolving Power, Angular Resolution

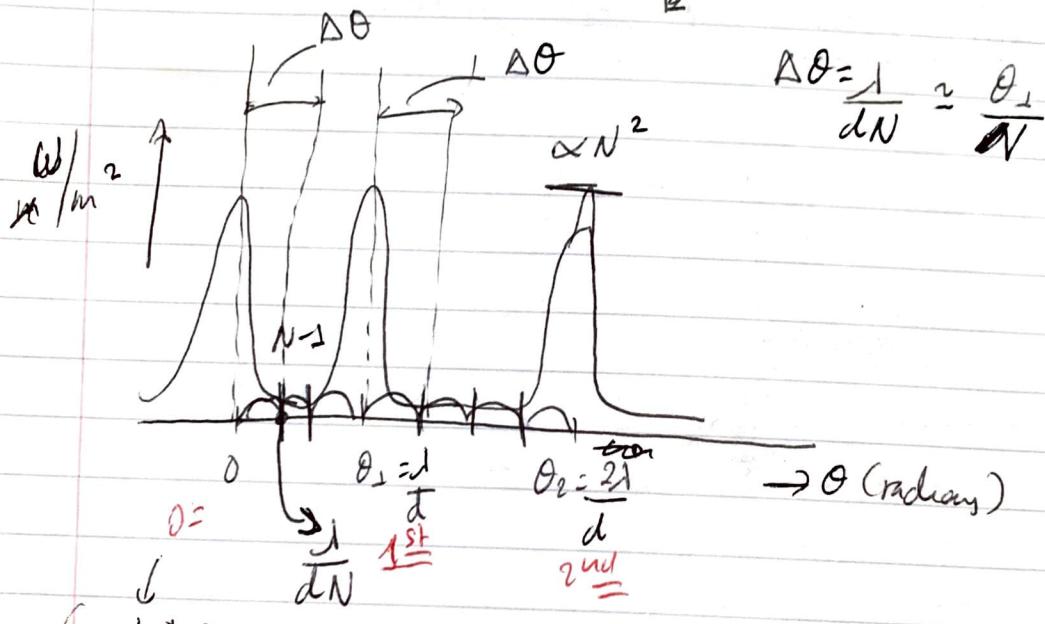
d] :



$$\text{const} = \sin \theta_n = \frac{n\lambda}{d} \approx \theta_n \text{ (radian)}$$

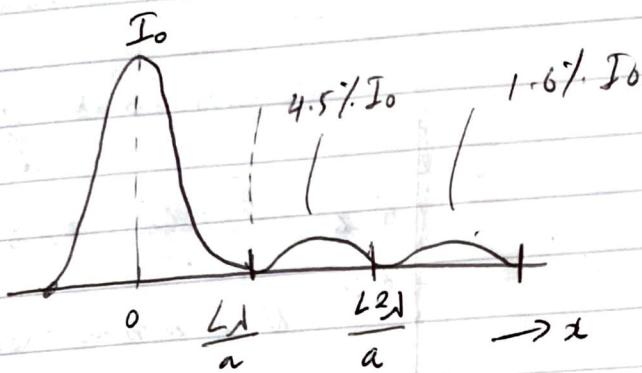
$$\theta_n = \frac{n\lambda}{d} \quad \theta_n \approx \frac{L n \lambda}{d}$$

$N-1$ minima \Rightarrow diff. interference



(white)
colour

* Single Sources



$$x \approx \frac{L\lambda}{a}$$

* smaller (a) ; the wider the bandwidth of the wave

* Angular resolution

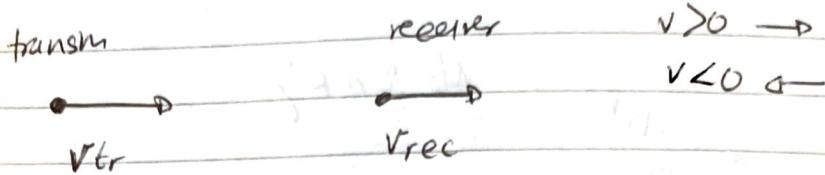
$$\theta = \frac{1.21}{a} \quad \lambda = 5000 \text{ Å}$$

and diffraction limit

Lect 35-

Doppler Effect, Big Bang, Cosmology

* Doppler Effect



$$f' = f \frac{r_s - v_{rec}}{r_s + v_{rec}}$$

$v_s = 340 \text{ m/s}$

$$\rightarrow f' = f \left(\frac{1 - \beta}{1 + \beta} \right)^{1/2} \quad \beta = \frac{v_{\text{relative}}}{c}$$

$\beta > 0$ receding
 $\beta < 0$ approaching.

$$\lambda' = \frac{c}{f'} \Rightarrow \lambda' = \frac{c}{f' \cdot \frac{1 - \beta}{1 + \beta}}$$

$\rightarrow v \cos \theta$ (use in eq.)



$$\lambda' = \lambda \left(\frac{1 + \beta}{1 - \beta} \right)^{1/2}$$

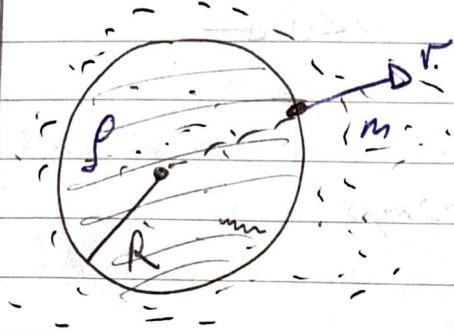
$$\left. \begin{array}{l} \lambda' > \lambda \rightarrow \text{redshift} \\ \lambda' < \lambda \rightarrow \text{blueshift} \end{array} \right\}$$

Note: $1 \text{ Mpc} = 3.26 \times 10^6 \text{ light years}$
 $= 3.1 \times 10^{19} \text{ km.}$

$$+ d = v t_{age} \quad \times \quad \left. \begin{array}{l} t_{age} = \frac{1}{H} \\ v = H \cdot d \end{array} \right\} \text{universe} \approx 14 \text{ billion years}$$

\hookrightarrow Hubble constant

$$H_0 = \frac{72 \text{ km/sec}}{\text{Mpc}}$$



$$M = \frac{4}{3} \pi R^3 \rho$$

ins
sphere

$$\frac{1}{2} m v^2 = \frac{m M G}{R} \quad \left. \begin{array}{l} \\ \end{array} \right\} \text{Total } E = \phi$$

$$\frac{1}{2} v^2 = \frac{4/3 \pi R^3 \rho}{R}$$

$$\frac{v}{R} = H$$

$$\therefore \boxed{\rho_0 = \frac{3}{8\pi G} H_0^2}$$