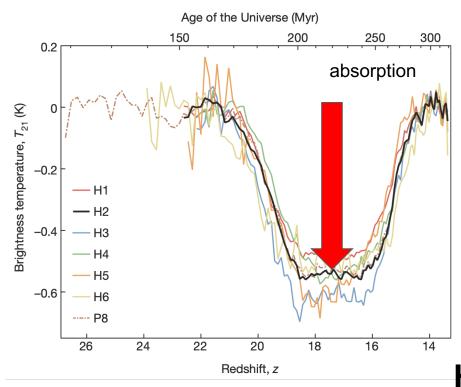
How we can explain the result of EDGES?

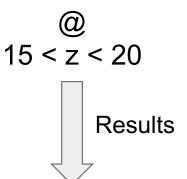
IGM theory group

木村和貴, Nicolas Ledos, 寺口遼, Dongsheng Sun, 伊東拓実, 西垣萌香

We talk about...



EDGES observed HI absorption

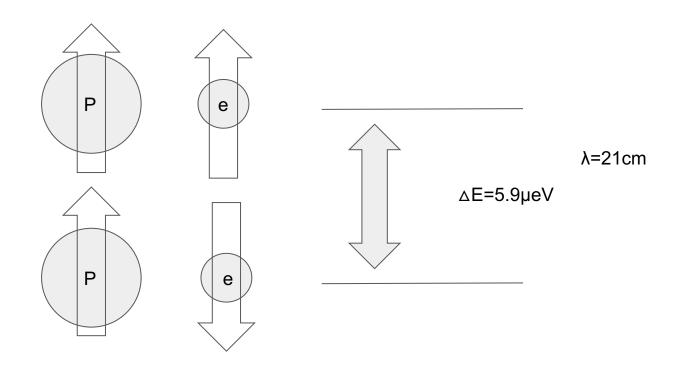


More than twice the maximum value expected in the standard cosmological model.



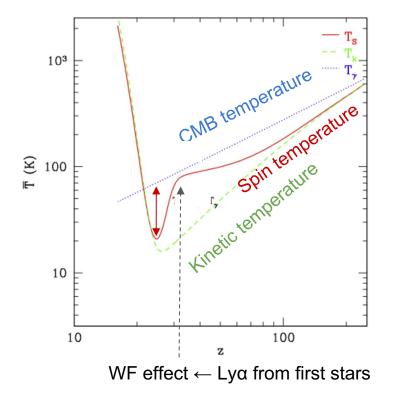
How we can explain this result?

21cm line: hyperfine transition of HI (neutral hydrogen)



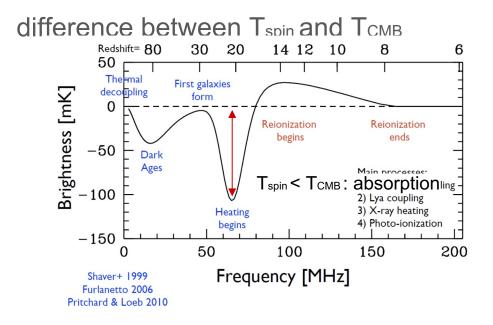
Spin Temperature

... defined through $n_{\uparrow\uparrow}$ / $n_{\uparrow\downarrow}$

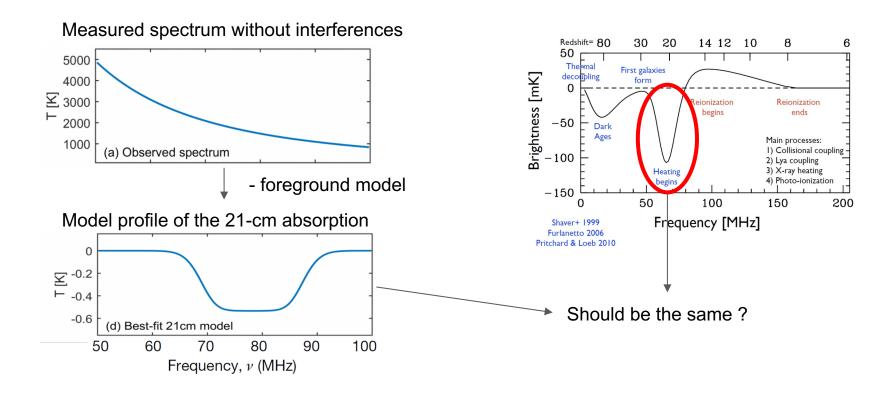


$$T_S^{-1} = \frac{T_{\gamma}^{-1} + x_c T_K^{-1} + x_{\alpha} T_c^{-1}}{1 + x_c + x_{\alpha}}$$

Observable: Brightness temperature



EDGES: Detection of 21cm Signal



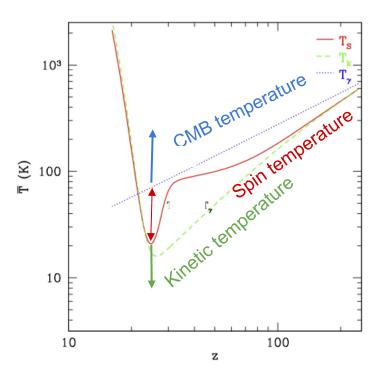
Detection of 21cm Signal?

=> Differential brightness temperature:

$$\delta T_b \propto (T_s - T_\gamma)/T_s$$

=> Only two solutions seems to explain higher absorption :

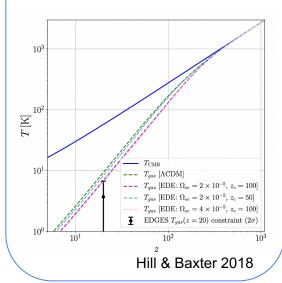
- Higher CMB temperature,
- Lower Gas temperature.



Existing scenarios ~ To decrease gas temperature

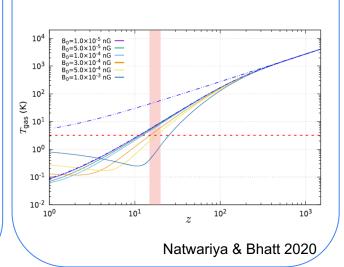
 DM-baryon interaction cooled by b-DM scattering 1+z10 -200 T₂₁ [mK] -400 -600 Dark Cosmic Cosmic reionization dawn 100 ν [MHz] Barkana 2018

Early dark energy scenario
Colder due to the cosmic expansion



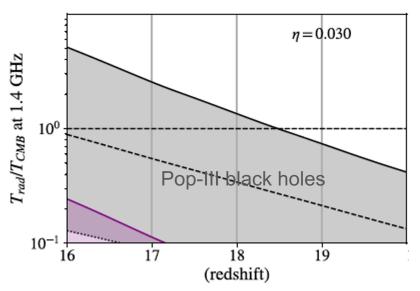
Primordial magnetic field amplification

Effective cooling due to turbulence (alpha-effect).



Existing scenarios ~ To enhance radio background temperature, Τγ

Radio emission: Radio loud AGN, Galaxy, Unresolved source



Ewall-Wice et al. 2018

Group work task: Propose new scenario for EDGES

scenario with supernova → already proposed(Jana et al. 2019).

combine some ideas

→ already proposede.g DM + PBH (Halder et al. 2021)

→ how about radio emitter + Magnetic Field ?

T_{Spin} < T_{Kinetic} ?

→ no idea about such physical process

Summary

• EDGES obtained larger absorption signal of 21cm line than expected.

There are many scenarios which explain EDGES observation.

- We discussed new scenarios.
 - some scenarios are already proposed.
 - how about T_{Spin} < T_{Kinetic}