Abstract

We further develop and extend a recent perturbative framework for neutrino oscillations in uniform matter density so that the resulting oscillation probabilities are accurate for the complete matter potential versus baseline divided by neutrino energy plane. This extension also gives the exact oscillation probabilities in vacuum for all values of baseline divided by neutrino energy. The expansion parameter used is related to the ratio of the solar to the atmospheric Δm^2 scales but with a unique choice of the atmospheric Δm^2 such that certain first-order effects are taken into account in the zeroth-order Hamiltonian. Using a mixing matrix formulation, this framework has the exceptional feature that the neutrino oscillation probability in matter has the same structure as in vacuum, to all orders in the expansion parameter. It also contains all orders in the matter potential and $\sin \theta_{13}$. It facilitates immediate physical interpretation of the analytic results, and makes the expressions for the neutrino oscillation probabilities extremely compact and very accurate even at zeroth order in our perturbative expansion. The first and second order results are also given which improve the precision by approximately two or more orders of magnitude per perturbative order.

Analytic and compact perturbative expressions for neutrino oscillations in matter

Peter B. Denton

Pheno 2016

May 9, 2016

Work done with S. Parke and H. Minakata. 1604.08167 github.com/PeterDenton/Nu-Pert





A perturbative description of neutrino oscillations in matter

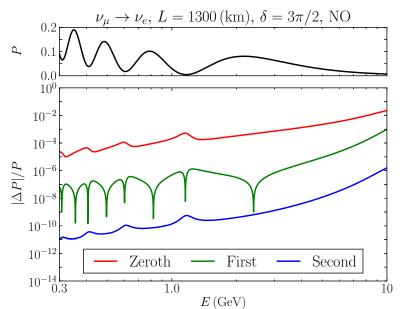
Several goals:

1. Use $\epsilon \simeq \Delta m_{21}^2/\Delta m_{31}^2 \approx 0.03$ as an expansion parameter.

1604 08167

- 2. Minimize the number of additional 'matter' terms (2).
- 3. Extremely accurate for all channels.
- 4. Exact in the vacuum limit.
- Reproduce the known CPV term in matter.

Precision order-by-order



Neutrino oscillations in matter

The electron flavor interacts differently in matter than the others through a G_F type process.

$$H^m=rac{1}{2E}\left[Uegin{pmatrix} m_1^2 & & \ & m_2^2 & \ & & m_3^2 \end{pmatrix}U^\dagger+egin{pmatrix} a & & \ & & \end{pmatrix}
ight]$$

We could write

$$H^{m} = \frac{1}{2E} U^{m} \begin{pmatrix} \lambda_{1} & & \\ & \lambda_{2} & \\ & & \lambda_{3} \end{pmatrix} U^{m\dagger}$$

where the λ_i are the masses squared in matter and $\theta_{12}^m, \theta_{13}^m, \theta_{23}^m$, and δ^m and the angles and phase in matter.

$$U \equiv U_{23}(\theta_{23}, \delta) U_{13}(\theta_{13}) U_{12}(\theta_{12})$$

This has been done.

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$$\lambda_{1} = \frac{A}{3} - \frac{1}{3}\sqrt{A^{2} - 3B} \frac{S}{S} - \frac{\sqrt{3}}{3}\sqrt{A^{2} - 3B}\sqrt{1 - \frac{S^{2}}{3}}$$

$$\lambda_{2} = \frac{A}{3} - \frac{1}{3}\sqrt{A^{2} - 3B} \frac{S}{S} + \frac{\sqrt{3}}{3}\sqrt{A^{2} - 3B}\sqrt{1 - \frac{S^{2}}{3}}$$

$$\lambda_{3} = \frac{A}{3} + \frac{2}{3}\sqrt{A^{2} - 3B} \frac{S}{S}$$

This has been done.

$$\begin{split} \lambda_1 &= \frac{A}{3} - \frac{1}{3}\sqrt{A^2 - 3B} S - \frac{\sqrt{3}}{3}\sqrt{A^2 - 3B}\sqrt{1 - S^2} \\ \lambda_2 &= \frac{A}{3} - \frac{1}{3}\sqrt{A^2 - 3B} S + \frac{\sqrt{3}}{3}\sqrt{A^2 - 3B}\sqrt{1 - S^2} \\ \lambda_3 &= \frac{A}{3} + \frac{2}{3}\sqrt{A^2 - 3B} S \\ A &= \Delta m_{21}^2 + \Delta m_{31}^2 + a \\ B &= \Delta m_{21}^2 \Delta m_{31}^2 + a \left[c_{13}^2 \Delta m_{31}^2 + (c_{12}^2 c_{13}^2 + s_{13}^2) \Delta m_{21}^2 \right] \end{split}$$

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Alternative solutions

- Numerical methods.
 - Good for experiments.
 - Important understanding can be missed (magic baseline)

P. Huber, W. Winter, hep-ph/0301257 A. Smirnov, hep-ph/0610198

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- Numerical methods.
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 - P. Huber, W. Winter, hep-ph/0301257

A. Smirnov, hep-ph/0610198

- Perturbative expansion
 - ▶ Small matter potential: $a/\Delta m^2$.
 - $ightharpoonup s_{13}, s_{13}^2.$

A. Cervera, et. al., hep-ph/0002108

H. Minakata, 0910.5545

K. Asano, H. Minakata, 1103.4387

 $\Delta m_{21}^2/\Delta m_{21}^2 \sim 0.03.$

J. Arafune, J. Sato, hep-ph/9607437

A. Cervera, et. al., hep-ph/0002108

M. Freund, hep-ph/0103300 E. Akhmedov, et. al., hep-ph/0402175

H. Minakata, S. Parke, 1505.01826

PBD, H. Minakata, S. Parke, 1604.08167

► Start with $\epsilon = \frac{\Delta m_{21}^2}{\Delta m_{ee}^2} = 0.03$.

$$\Delta m_{ee}^2 = \Delta m_{31}^2 - s_{12}^2 \Delta m_{21}^2$$

H. Nunokawa, S. Parke, R. Zukanovich Funchal, hep-ph/0503283

- ► Start with $\epsilon = \frac{\Delta m_{21}^2}{\Delta m_{ee}^2} = 0.03$.
- ▶ Perform one fixed and two variable rotations: (θ_{23}, δ) , ϕ , ψ .

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- ▶ Perform one fixed and two variable rotations: (θ_{23}, δ) , ϕ , ψ .
- Write the probabilities with simple L/E dependence:

$$P(\nu_{\alpha} \to \nu_{\beta}) = \delta_{\alpha\beta} - \sum_{i < j} \Re \left[U_{\alpha i} U_{\beta j}^* U_{\alpha j}^* U_{\beta i} \right] \sin^2 \Delta_{ij}$$
$$+ 8\Im \left[U_{\alpha 1} U_{\beta 2}^* U_{\alpha 2}^* U_{\beta 1} \right] \sin \Delta_{32} \sin \Delta_{31} \sin \Delta_{21}$$

Nonvanishing Wronskian \Rightarrow fewest number of L/E functions.

$$\Delta m^2_{ee} = \Delta m^2_{31} - s^2_{12} \Delta m^2_{21}$$

H. Nunokawa, S. Parke, R. Zukanovich Funchal, hep-ph/0503283

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{4E}$$

C. Jarlskog, PRL 1985

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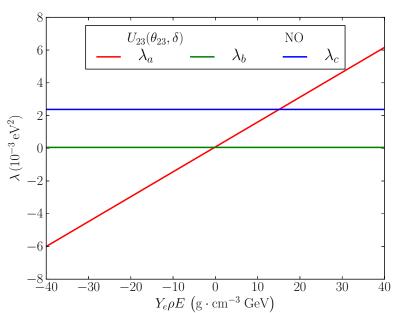
Clear that the CPV term is $\mathcal{O}[(L/E)^3]$ not $\mathcal{O}[(L/E)^1]$.

$$\Delta m_{\rm ee}^2 = \Delta m_{31}^2 - s_{12}^2 \Delta m_{21}^2$$

H. Nunokawa, S. Parke, R. Zukanovich Funchal, hep-ph/0503283

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{\Delta F}$$

Eigenvalues in matter: two rotations are needed

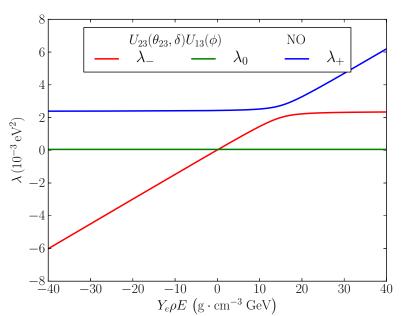


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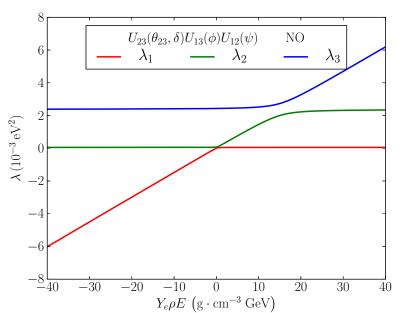
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Eigenvalues in matter: two rotations are needed



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Three rotations

- 1. Perform a constant $U_{23}(\theta_{23}, \delta)$ rotation.
 - U_{23} commutes with the matter potential.
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- 2. Diagonalize the diagonal and $\mathcal{O}(\epsilon^0)$ off-diagonal terms with $U_{13}(\phi)$.
 - $\phi_{\rm vac} = \theta_{13}.$
 - Expansion parameter is $c_{12}s_{12}\frac{\Delta m_{21}^2}{\Delta m^2}=0.015$.

H. Minakata, S. Parke, 1505.01826

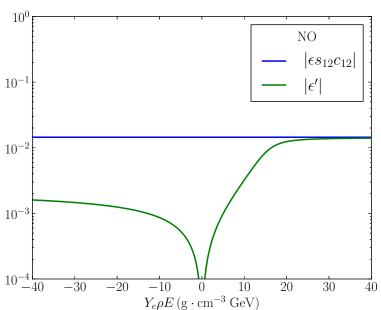
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H. Minakata, S. Parke, 1505.01826

- 3. Diagonalize the terms non-zero in matter with $U_{12}(\psi)$.
 - $\psi_{\rm vac} = \theta_{12}.$
 - Expansion parameter is now $\epsilon' = c_{12} s_{12} s_{(\phi \theta_{13})} \frac{\Delta m_{21}^2}{\Delta m_{22}^2} < 0.015$.

Expansion parameter



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Simplicity at zeroth order

Simple L/E dependence:

$$P = \delta^{\alpha\beta} + 4C_{21}^{\alpha\beta}\sin^2\Delta_{21} + 4C_{31}^{\alpha\beta}\sin^2\Delta_{31} + 4C_{32}^{\alpha\beta}\sin^2\Delta_{32} + 8D^{\alpha\beta}\sin\Delta_{21}\sin\Delta_{31}\sin\Delta_{32}$$

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{\Delta E}$$

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Zeroth order is equivalent to the vacuum probability with

$$heta_{13} o \phi$$
 $heta_{12} o \psi$ and $\Delta m_{ij}^2 o \Delta \lambda_{ij}$

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{\Delta E}$$

CPV term

It is known that the exact CPV term in matter is

$$\pm 8\sin\delta c_{12}s_{12}c_{13}^2s_{13}c_{23}s_{23}\prod_{i>j}\frac{\Delta m_{ij}^2}{\Delta\lambda_{ij}^{(\mathrm{ex})}}\sin\Delta_{32}\sin\Delta_{31}\sin\Delta_{21}$$

V. Naumov, Int. J. Mod. Phys. 1992

P. Harrison, W. Scott, hep-ph/9912435

Our expression reproduces this order by order in ϵ' .

Precision

At the first oscillation minimum and maximum:

DUNE: NO, $\delta = 3\pi/2$		First min	First max
$P(u_{\mu} ightarrow u_{e})$		0.0047	0.081
E (GeV)		1.2	2.2
$\frac{ \Delta P }{P}$	Zeroth	5×10^{-4}	4×10^{-4}
	First	3×10^{-7}	2×10^{-7}
	Second	6×10^{-10}	5×10^{-10}

Conclusions

- ▶ Analytic expression for neutrino oscillations with precision and clarity.
- Clear that exactly two matter rotations is the correct number.
- Our perturbative parameter is < 0.015 and is 0 in vacuum.
- ▶ Form emphasizes the L/E dependence of oscillations at all orders.
- ▶ Zeroth order expression matches the form of the vacuum expression.
- ▶ Zeroth order is sufficient for current and planned experiments.

Backups

Eigenvalues

Tilde basis after $U_{23}(\theta_{23}, \delta)$:

$$\lambda_a = a + (s_{13}^2 + \epsilon s_{12}^2) \Delta m_{ee}^2$$

$$\lambda_b = \epsilon c_{12}^2 \Delta m_{ee}^2$$

$$\lambda_c = (c_{13}^2 + \epsilon s_{12}^2) \Delta m_{ee}^2$$

Hat basis after $U_{13}(\phi)$:

$$\lambda_{\mp} = \frac{1}{2} \left[(\lambda_a + \lambda_c) \mp \text{sign}(\Delta m_{\text{ee}}^2) \sqrt{(\lambda_c - \lambda_a)^2 + 4(s_{13}c_{13}\Delta m_{\text{ee}}^2)^2} \right]$$
 $\lambda_0 = \lambda_b$

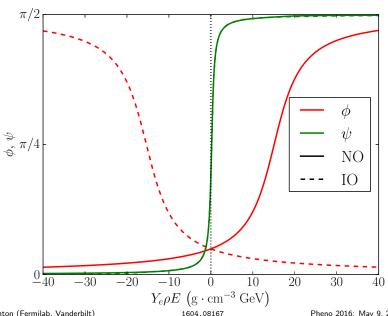
Check basis after $U_{12}(\psi)$:

$$\lambda_{1,2} = rac{1}{2} \left[(\lambda_0 + \lambda_-) \mp \sqrt{(\lambda_0 - \lambda_-)^2 + 4(\epsilon c_{(\phi - \theta_{13})} c_{12} s_{12} \Delta m_{\mathsf{ee}}^2)^2} \right]$$
 $\lambda_3 = \lambda_+$

Phases

$$egin{aligned} c_{2\phi} &= rac{\lambda_c - \lambda_{\mathsf{a}}}{\lambda_+ - \lambda_-} \qquad s_{2\phi} = s_{2 heta_{13}} rac{\Delta m_{\mathsf{ee}}^2}{\lambda_+ - \lambda_-} \ c_{2\psi} &= rac{\lambda_0 - \lambda_-}{\lambda_2 - \lambda_1} \qquad s_{2\psi} = s_{2 heta_{12}} \epsilon c_{(\phi - heta_{13})} rac{\Delta m_{\mathsf{ee}}^2}{\lambda_2 - \lambda_1} \end{aligned}$$

The two phases in matter



$\lambda_{1,2} - \psi$ interchange

From the shape of $U_{12}(\psi)$, it is clear that the probabilities are invariant under a simultaneous interchange of

$$\lambda_1 \leftrightarrow \lambda_2$$
 and $\psi \to \psi \pm \frac{\pi}{2}$

Since only even powers of ψ trig functions $(c_{\psi}^2, s_{\psi}^2, c_{\psi}s_{\psi}, c_{2\psi}, s_{2\psi})$ appear in the probabilities, the sign degeneracy is irrelevant. More usefully, we can write that the probabilities are invariant under the simultaneous interchange of

$$\lambda_1 \leftrightarrow \lambda_2 \qquad c_\psi^2 \leftrightarrow s_\psi^2 \qquad ext{and} \qquad c_\psi s_\psi o -c_\psi s_\psi$$

This interchange constrains C_{21} , and C_{32} is then easily calculated from C_{31} .

Zeroth order coefficients

$\nu_{\alpha} \rightarrow \nu_{\beta}$	$(\mathcal{C}_{21}^{lphaeta})^{(0)}$
$ u_e ightarrow u_e$	$-c_\phi^4 s_\psi^2 c_\psi^2$
$ u_{\mu} ightarrow u_{e}$	$c_{\phi}^2 s_{\psi}^2 c_{\psi}^2 (c_{23}^2 - s_{\phi}^2 s_{23}^2) + c_{2\psi} J_r^m \cos \delta$
$ u_{\mu} \rightarrow \nu_{\mu} $	$-(c_{23}^2c_{\psi}^2 + s_{23}^2s_{\phi}^2s_{\psi}^2)(c_{23}^2s_{\psi}^2 + s_{23}^2s_{\phi}^2c_{\psi}^2) -2(c_{23}^2 - s_{\phi}^2s_{23}^2)c_{2\psi}J_{rr}^m\cos\delta + (2J_{rr}^m\cos\delta)^2$

$\nu_{\alpha} \rightarrow \nu_{\beta}$	$(C_{31}^{lphaeta})^{(0)}$	$(D^{lphaeta})^{(0)}$
$ u_{e} ightarrow u_{e}$	$-c_\phi^2 s_\phi^2 c_\psi^2$	0
$ u_{\mu} ightarrow u_{e}$	$s_{\phi}^2 c_{\phi}^2 c_{\psi}^2 s_{23}^2 + J_r^m \cos \delta$	$-J_r^m \sin \delta$
$ u_{\mu} \rightarrow \nu_{\mu} $	$-c_{\phi}^{2}s_{23}^{2}(c_{23}^{2}s_{\psi}^{2}+s_{23}^{2}s_{\phi}^{2}c_{\psi}^{2}) \\ -2s_{23}^{2}J_{r}^{m}\cos\delta$	0

$$J_r^m \equiv s_{\psi} c_{\psi} s_{\phi} c_{\phi}^2 s_{23} c_{23}$$
$$J_{rr}^m \equiv J_r^m / c_{\phi}^2$$

General form of the first order coefficients

$$(C_{21}^{\alpha\beta})^{(1)} = \epsilon' \Delta m_{ee}^2 \left(\frac{F_1^{\alpha\beta}}{\Delta \lambda_{31}} + \frac{F_2^{\alpha\beta}}{\Delta \lambda_{32}} \right)$$

$$(C_{31}^{\alpha\beta})^{(1)} = \epsilon' \Delta m_{ee}^2 \left(\frac{F_1^{\alpha\beta} + G_1^{\alpha\beta}}{\Delta \lambda_{31}} - \frac{F_2^{\alpha\beta}}{\Delta \lambda_{32}} \right)$$

$$(C_{32}^{\alpha\beta})^{(1)} = \epsilon' \Delta m_{ee}^2 \left(-\frac{F_1^{\alpha\beta}}{\Delta \lambda_{31}} + \frac{F_2^{\alpha\beta} + G_2^{\alpha\beta}}{\Delta \lambda_{32}} \right)$$

$$(D^{\alpha\beta})^{(1)} = \epsilon' \Delta m_{ee}^2 \left(\frac{K_1^{\alpha\beta}}{\Delta \lambda_{31}} - \frac{K_2^{\alpha\beta}}{\Delta \lambda_{32}} \right)$$

$$K_1^{\alpha\beta} = \begin{cases} 0 & \alpha = \beta \\ \mp s_{23}c_{23}c_{\phi}s_{\psi}^2(c_{\phi}^2c_{\psi}^2 - s_{\phi}^2)\sin\delta & \alpha \neq \beta \end{cases}$$

where the minus sign is for $\nu_{\mu} \rightarrow \nu_{e}$. Peter B. Denton (Fermilab, Vanderbilt)

First order coefficients

$\nu_{\alpha} \rightarrow \nu_{\beta}$	$\mathcal{F}_1^{lphaeta}$	
$ u_{e} ightarrow u_{e}$	$-2c_\phi^3 s_\phi s_\psi^3 c_\psi$	
$ u_{\mu} ightarrow u_{e} $	$c_{\phi}s_{\psi}^{2}[s_{\phi}s_{\psi}c_{\psi}(c_{23}^{2}+c_{2\phi}s_{23}^{2}) \ -s_{23}c_{23}(s_{\phi}^{2}s_{\psi}^{2}+c_{2\phi}c_{\psi}^{2})\cos\delta]$	
$ u_{\mu} \rightarrow \nu_{\mu} $	$ 2c_{\phi}s_{\psi}(s_{23}^{2}s_{\phi}c_{\psi} + s_{23}c_{23}s_{\psi}\cos\delta) \times $ $ (c_{23}^{2}c_{\psi}^{2} - 2s_{23}c_{23}s_{\phi}s_{\psi}c_{\psi}\cos\delta + s_{23}^{2}s_{\phi}^{2}s_{\psi}^{2}) $	

$\nu_{\alpha} \rightarrow \nu_{\beta}$	$\mathcal{G}_1^{lphaeta}$
$ u_{e} ightarrow u_{e}$	$2s_{\phi}c_{\phi}s_{\psi}c_{\psi}c_{2\phi}$
$ u_{\mu} ightarrow u_{e}$	$-2s_{\phi}c_{\phi}s_{\psi}(s_{23}^{2}c_{2\phi}c_{\psi}-s_{23}c_{23}s_{\phi}s_{\psi}\cos\delta)$
$ u_{\mu} \rightarrow \nu_{\mu} $	$-2c_{\phi}s_{\psi}(s_{23}^2s_{\phi}c_{\psi}+s_{23}c_{23}s_{\psi}\cos\delta) \ imes (1-2c_{\phi}^2s_{23}^2)$

Note on the PMNS matrix

The PMNS matrix typically has the CP phase associated with θ_{13} , while our unitary matrix has $\delta_{\rm CP}$ associated with θ_{23} . The standard form is related to our form by multiplying the third row by $e^{i\delta}$ and the third column by $e^{-i\delta}$.

Our formulation is useful because after the U_{23} rotation the resulting Hamiltonian of the "tilde" basis is real.

B. Pontecorvo, Sov. Phys. JETP 7 1958

Z. Maki, M. Nakagawa, S. Sakata, Prog. Theor. Phys. 28 1962

Neutrino oscillations in vacuum

Neutrino oscillations in matter are described by this Hamiltonian written in the flavor basis:

$$H = \frac{1}{2E} U \begin{pmatrix} m_1^2 & & \\ & m_2^2 & \\ & & m_3^2 \end{pmatrix} U^\dagger$$

where the unitary mixing matrix is parameterized by

$$U = \begin{pmatrix} 1 & c_{23} & s_{23}e^{i\delta} \\ -s_{23}e^{-i\delta} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & s_{13} \\ 1 & 1 \\ -s_{13} & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} \\ -s_{12} & c_{12} \\ & & 1 \end{pmatrix}$$

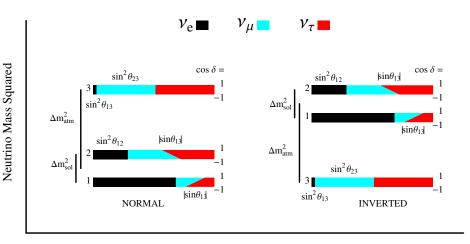
$$= \begin{pmatrix} c_{12}c_{13} & c_{13}s_{12} & s_{13} \\ -c_{23}s_{12} - e^{i\delta}c_{12}s_{13}s_{23} & c_{12}c_{23} - e^{i\delta}s_{12}s_{13}s_{23} & e^{i\delta}c_{13}s_{23} \\ e^{-i\delta}s_{12}s_{23} - c_{12}c_{23}s_{13} & -c_{23}s_{12}s_{13} - e^{-i\delta}c_{12}s_{23} & c_{13}c_{23} \end{pmatrix}$$

Neutrino oscillations in vacuum: example

For example, it is easy to calculate the exact $\nu_e \rightarrow \nu_e$ disappearance channel in vacuum.

$$\begin{split} P(\nu_{\text{e}} \rightarrow \nu_{\text{e}}) &= 1 \\ &- 4c_{12}^2 s_{12}^2 c_{13}^4 \sin^2 \Delta_{21} \\ &- 4c_{12}^2 c_{13}^2 s_{13}^2 \sin^2 \Delta_{31} \\ &- 4s_{12}^2 c_{13}^2 s_{13}^2 \sin^2 \Delta_{32} \end{split}$$

Known neutrino properties



Fractional Flavor Content varying $\cos \delta$

O. Mena, S. Parke, hep-ph/0312131

A constant rotation in the 2-3 plane

Change basis by rotating $H^m o \tilde{H}^m$ by $U_{23}(\theta_{23}, \delta)$

- $ightharpoonup U_{23}$ commutes with the matter potential,
- $ightharpoonup \tilde{H}^m$ is now entirely real.

$$\lambda_{a} = a + (s_{13}^{2} + \epsilon s_{12}^{2}) \Delta m_{ee}^{2}$$
 $\lambda_{b} = \epsilon c_{12}^{2} \Delta m_{ee}^{2}$
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 $\lambda_{c} = (c_{13}^{2} + \epsilon s_{12}^{2}) \Delta m_{ee}^{2}$

$$\Delta m_{\mathrm{ee}}^2 = \Delta m_{31}^2 - s_{12}^2 \Delta m_{21}^2$$
 $\epsilon = \frac{\Delta m_{21}^2}{\Delta m_{22}^2} pprox 0.03$

H. Nunokawa, S. Parke, R. Zukanovich Funchal, hep-ph/0503283

A rotation in the 1-3 plane

Change basis again by rotating $\tilde{H}^m o \hat{H}^m$ by $U_{13}(\phi)$

•
$$\phi(a=0)=\theta_{13}$$
.

$$\lambda_{\mp} = \frac{1}{2} \left[(\lambda_a + \lambda_c) \mp \operatorname{sign}(\Delta m_{ee}^2) \sqrt{(\lambda_c - \lambda_a)^2 + 4(s_{13}c_{13}\Delta m_{ee}^2)^2} \right]$$

$$\lambda_0 = \lambda_b$$

A rotation in the 1-2 plane

Change basis again by rotating $\hat{H}^m \to \check{H}^m$ by $U_{12}(\psi)$

- $\psi(a=0)=\theta_{12}$
- We have now created our own $U^m(\theta_{23}, \delta, \phi, \psi)$.

$$\lambda_{1,2} = \frac{1}{2} \left[(\lambda_0 + \lambda_-) \mp \sqrt{(\lambda_0 - \lambda_-)^2 + 4(\epsilon c_{(\phi - \theta_{13})} c_{12} s_{12} \Delta m_{\text{ee}}^2)^2} \right]$$
 $\lambda_3 = \lambda_+$

Exact neutrino oscillations in matter: mixing angles

$$\begin{split} s_{12}^{m2} &= \frac{-\left(\lambda_2^2 - \alpha \lambda_2 + \beta\right) \Delta \lambda_{31}}{\left(\lambda_1^2 - \alpha \lambda_1 + \beta\right) \Delta \lambda_{32} - \left(\lambda_2^2 - \alpha \lambda_2 + \beta\right) \Delta \lambda_{31}} \\ s_{13}^{m2} &= \frac{\lambda_3^2 - \alpha \lambda_3 + \beta}{\Delta \lambda_{31} \Delta \lambda_{32}} \\ s_{23}^{m2} &= \frac{s_{23}^2 E^2 + c_{23}^2 F^2 + 2c_{23}s_{23}c_{\delta}EF}{E^2 + F^2} \\ e^{-i\delta^m} &= \frac{c_{23}^2 s_{23}^2 \left(e^{-i\delta} E^2 - e^{i\delta} F^2\right) + \left(c_{23}^2 - s_{23}^2\right) EF}{\sqrt{\left(s_{23}^2 E^2 + c_{23}^2 F^2 + 2EFc_{23}s_{23}c_{\delta}\right) \left(c_{23}^2 E^2 + s_{23}^2 F^2 - 2EFc_{23}s_{23}c_{\delta}\right)}} \end{split}$$

Exact neutrino oscillations in matter: mixing angles

$$\begin{split} s_{12}^{m2} &= \frac{-\left(\lambda_{2}^{2} - \alpha\lambda_{2} + \beta\right) \Delta\lambda_{31}}{\left(\lambda_{1}^{2} - \alpha\lambda_{1} + \beta\right) \Delta\lambda_{32} - \left(\lambda_{2}^{2} - \alpha\lambda_{2} + \beta\right) \Delta\lambda_{31}} \\ s_{13}^{m2} &= \frac{\lambda_{3}^{2} - \alpha\lambda_{3} + \beta}{\Delta\lambda_{31}\Delta\lambda_{32}} \\ s_{23}^{m2} &= \frac{s_{23}^{2}E^{2} + c_{23}^{2}F^{2} + 2c_{23}s_{23}c_{\delta}EF}{E^{2} + F^{2}} \\ e^{-i\delta^{m}} &= \frac{c_{23}^{2}s_{23}^{2}\left(e^{-i\delta}E^{2} - e^{i\delta}F^{2}\right) + \left(c_{23}^{2} - s_{23}^{2}\right)EF}{\sqrt{\left(s_{23}^{2}E^{2} + c_{23}^{2}F^{2} + 2EFc_{23}s_{23}c_{\delta}\right)\left(c_{23}^{2}E^{2} + s_{23}^{2}F^{2} - 2EFc_{23}s_{23}c_{\delta}\right)}} \\ \alpha &= c_{13}^{2}\Delta m_{31}^{2} + \left(c_{12}^{2}c_{13}^{2} + s_{13}^{2}\right)\Delta m_{21}^{2} \\ \beta &= c_{12}^{2}c_{13}^{2}\Delta m_{21}^{2}\Delta m_{31}^{2} \\ E &= c_{13}s_{13}\left[\left(\lambda_{3} - \Delta m_{21}^{2}\right)\Delta m_{31}^{2} - s_{12}^{2}\left(\lambda_{3} - \Delta m_{31}^{2}\right)\Delta m_{21}^{2}\right] \\ F &= c_{12}s_{12}c_{13}\left(\lambda_{3} - \Delta m_{31}^{2}\right)\Delta m_{21}^{2} \end{split}$$

Neutrino oscillation probabilities

$$\mathcal{A}(
u_{lpha}
ightarrow
u_{eta})=\sum_{i=1}^{3}U_{lpha i}^{st}U_{eta i}e^{-irac{m_{i}^{2}L}{2E}}$$

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$$\begin{split} P(\nu_{\alpha} \to \nu_{\beta}) &= |\mathcal{A}(\nu_{\alpha} \to \nu_{\beta})|^{2} \\ &= \delta_{\alpha\beta} - \sum_{i < j} \Re \left[U_{\alpha i} U_{\beta j}^{*} U_{\alpha j}^{*} U_{\beta i} \right] \sin^{2} \Delta_{ij} \\ &+ 8 \Im \left[U_{\alpha 1} U_{\beta 2}^{*} U_{\alpha 2}^{*} U_{\beta 1} \right] \sin \Delta_{32} \sin \Delta_{31} \sin \Delta_{21} \end{split}$$

C. Jarlskog, PRL 1985

Nonvanishing Wronskian \Rightarrow fewest number of L/E functions.

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{\Delta E}$$

Neutrino oscillation probabilities

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C. Jarlskog, PRL 1985

Nonvanishing Wronskian \Rightarrow fewest number of L/E functions.

Clear that the CPV term is $\mathcal{O}[(L/E)^3]$ not $\mathcal{O}[(L/E)^1]$.

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{4E}$$

Eigenvalues in matter

