



# The Moon: New Views on Formation and Impact History

Pham Nguyen  
Michigan State University  
and  
Albion College

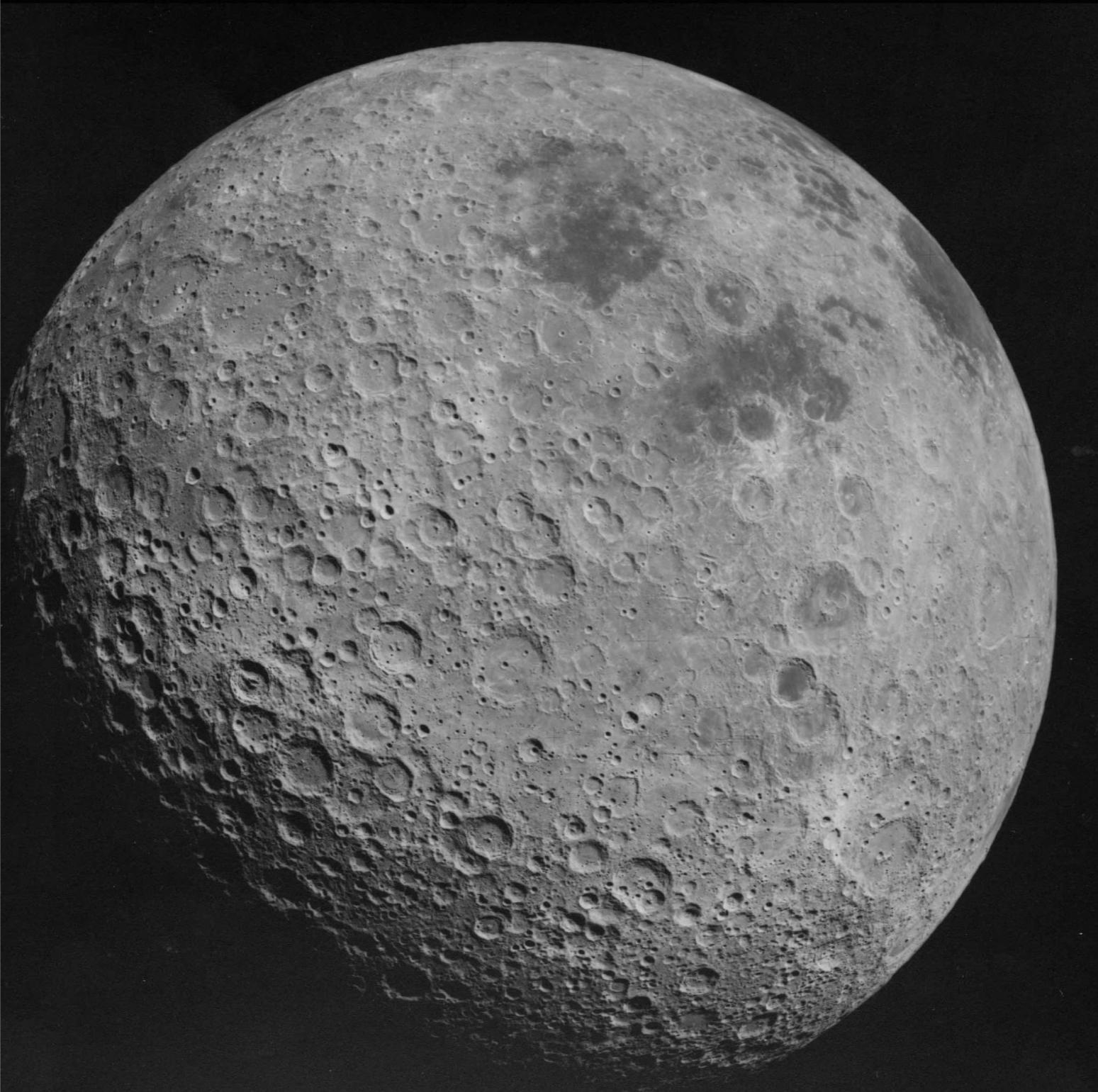
# Outline

Moon  
Formation  
(~ 4.51 Ga)



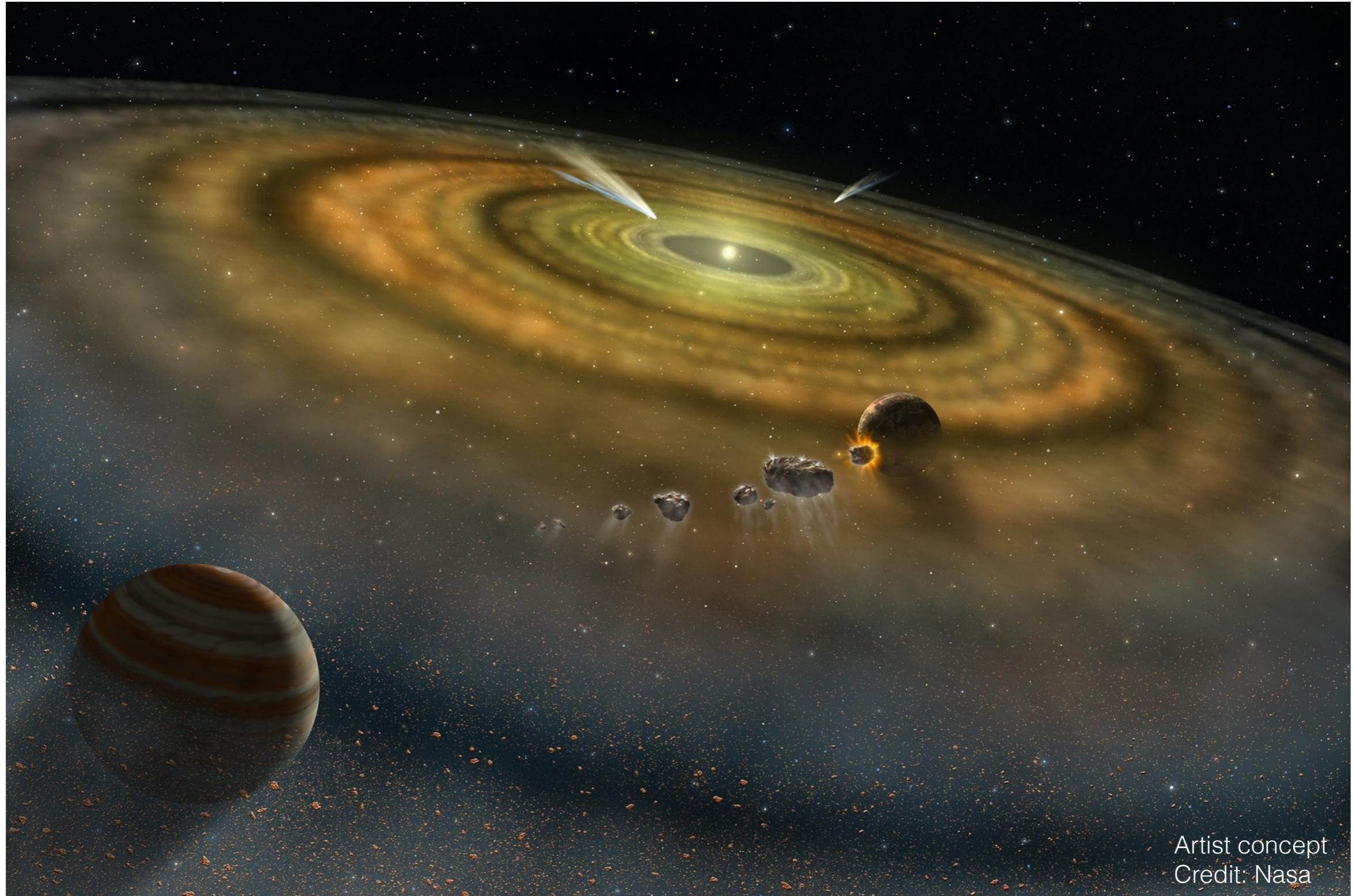
# Why Study the Moon?

# It has a well preserved impact history



Far side of the Moon taken  
during Apollo 16  
Credit: NASA

# Informs dynamical models of the solar system



Artist concept  
Credit: Nasa

# How do we Study the Moon?

# Lunar Samples

Apollo Missions

~382 Kg



Apollo 16 sample 60025

Credit NASA/Johnson Space Center photograph

S72-42187

Luna Program

~0.326 Kg



20 cm portion of Luna 20 core sample

Credit NASA/Johnson Space Center photograph

S73-17207

# Lunar Samples (Cont.)

Meteorites (~150 found)



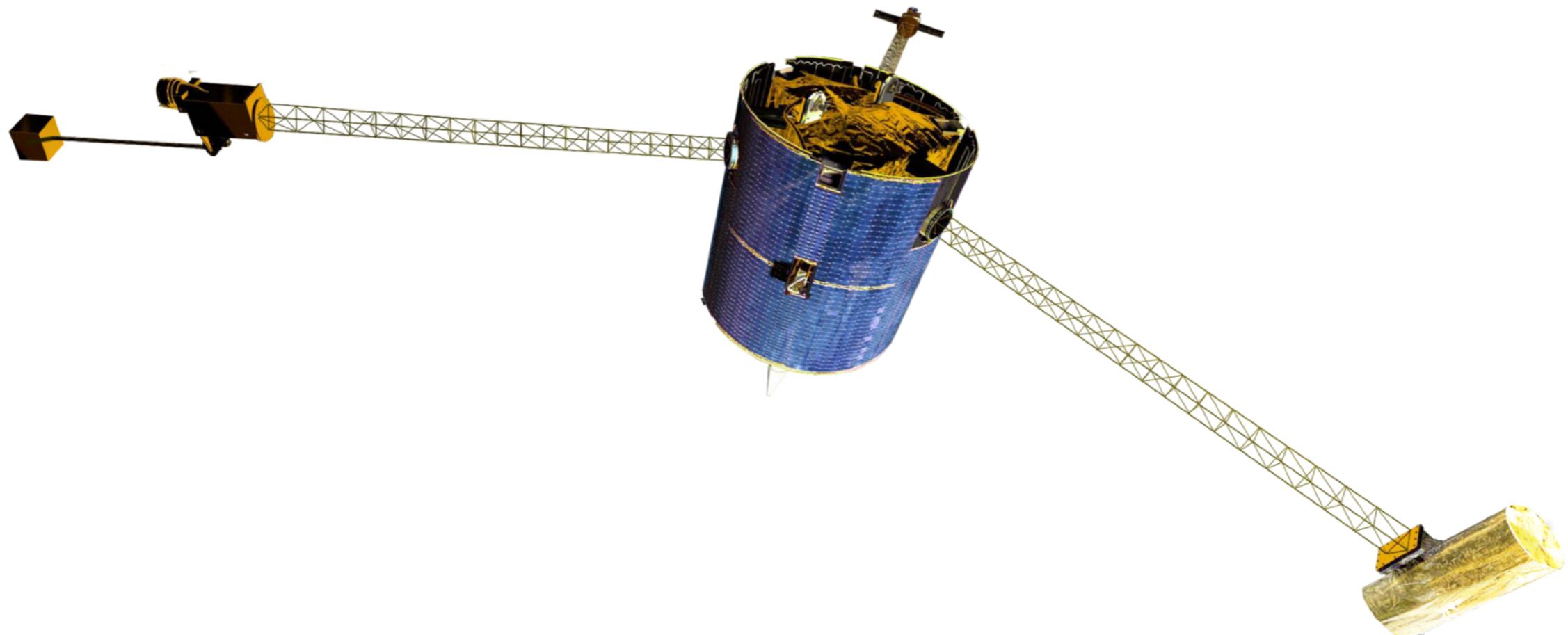
MAC 88105 a 663 gram sample found in Antarctica

Credit: NASA photo S89-38379

# Lunar Orbiters

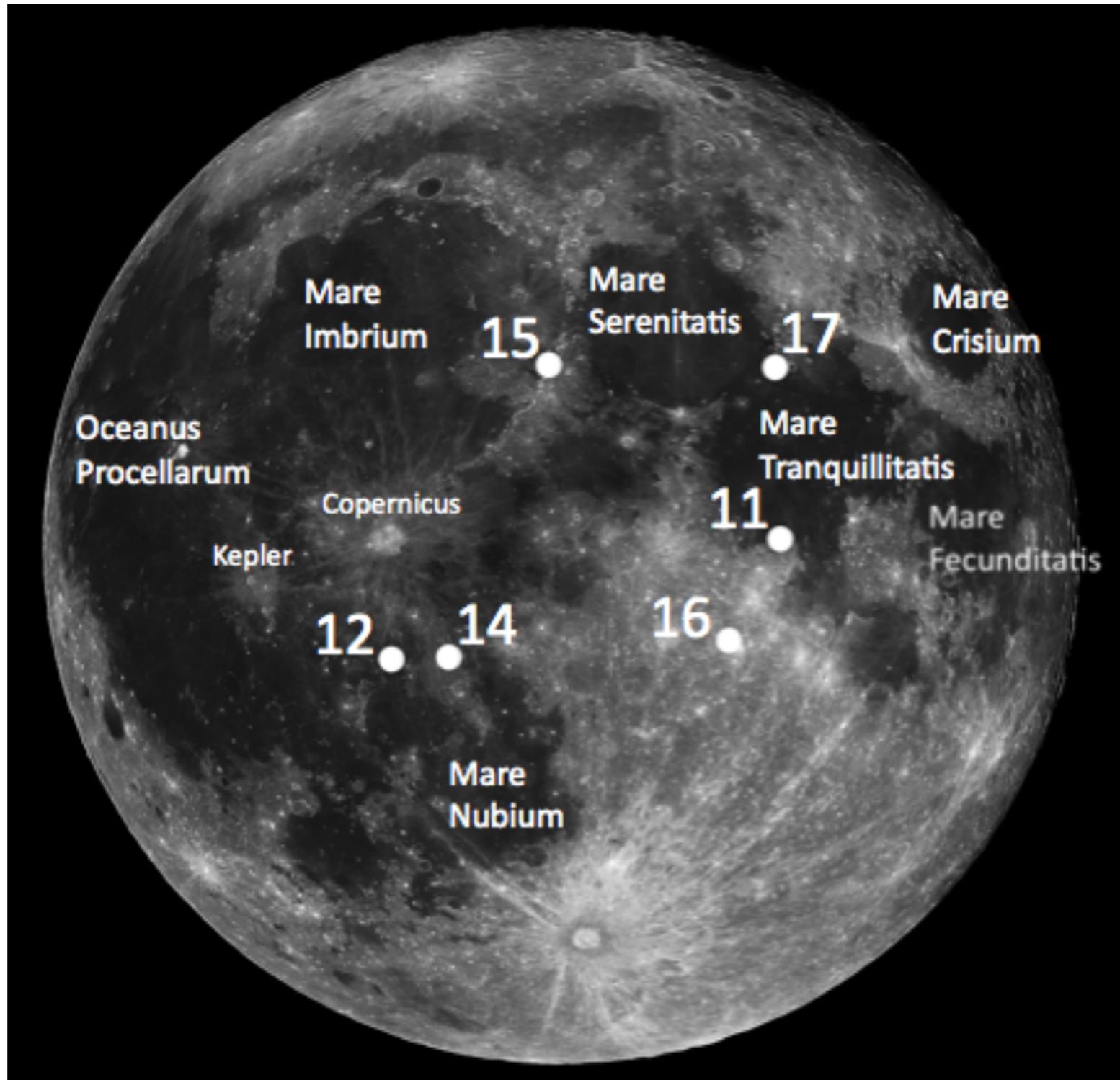
Many lunar orbiter missions including:

- Lunar Prospector (1998-1999)
- Lunar Reconnaissance Orbiter (2009-Present)
- Gravity Recovery and Interior Laboratory (2011-2012)



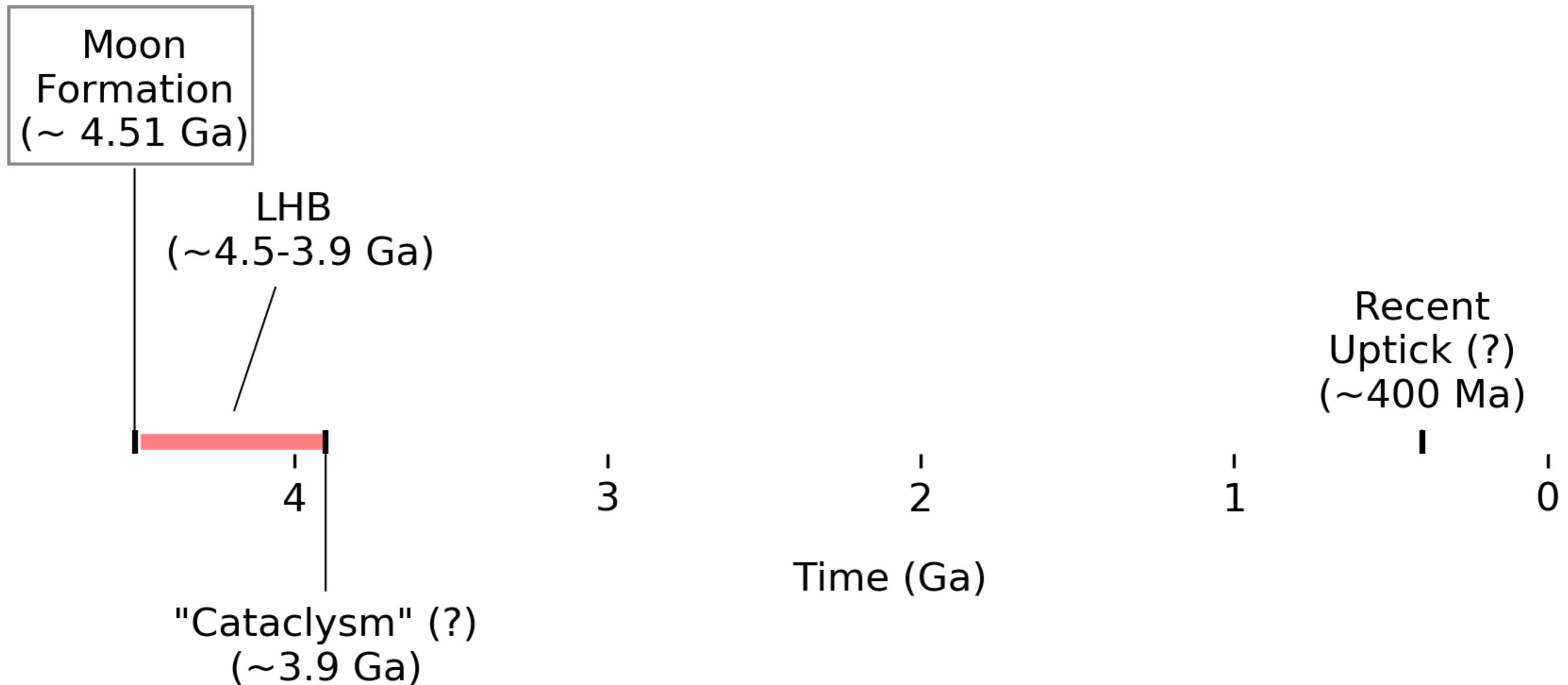
So...what is the Moon like?

# Surface Features of the Moon



Credit: Dustin Scriven

# Outline

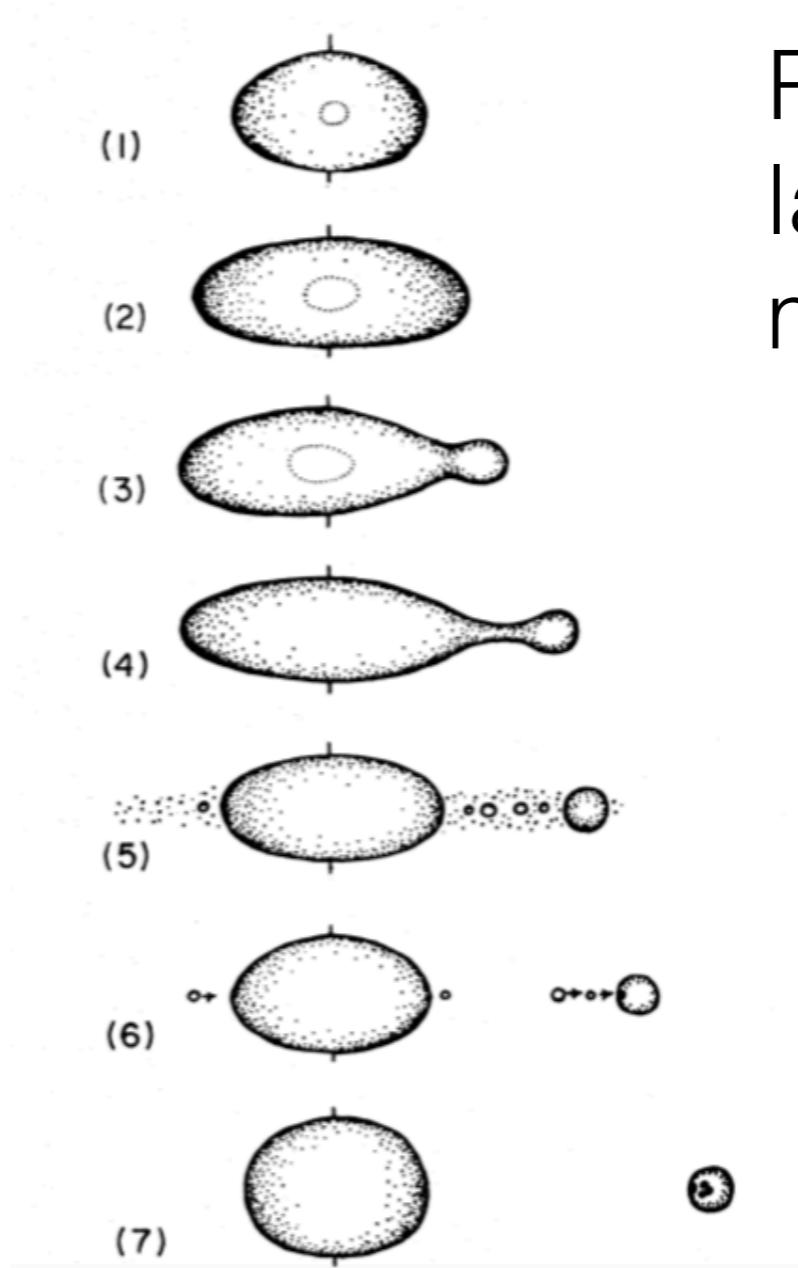


# To Make a Moon

What aspects of the Moon need to be explained?

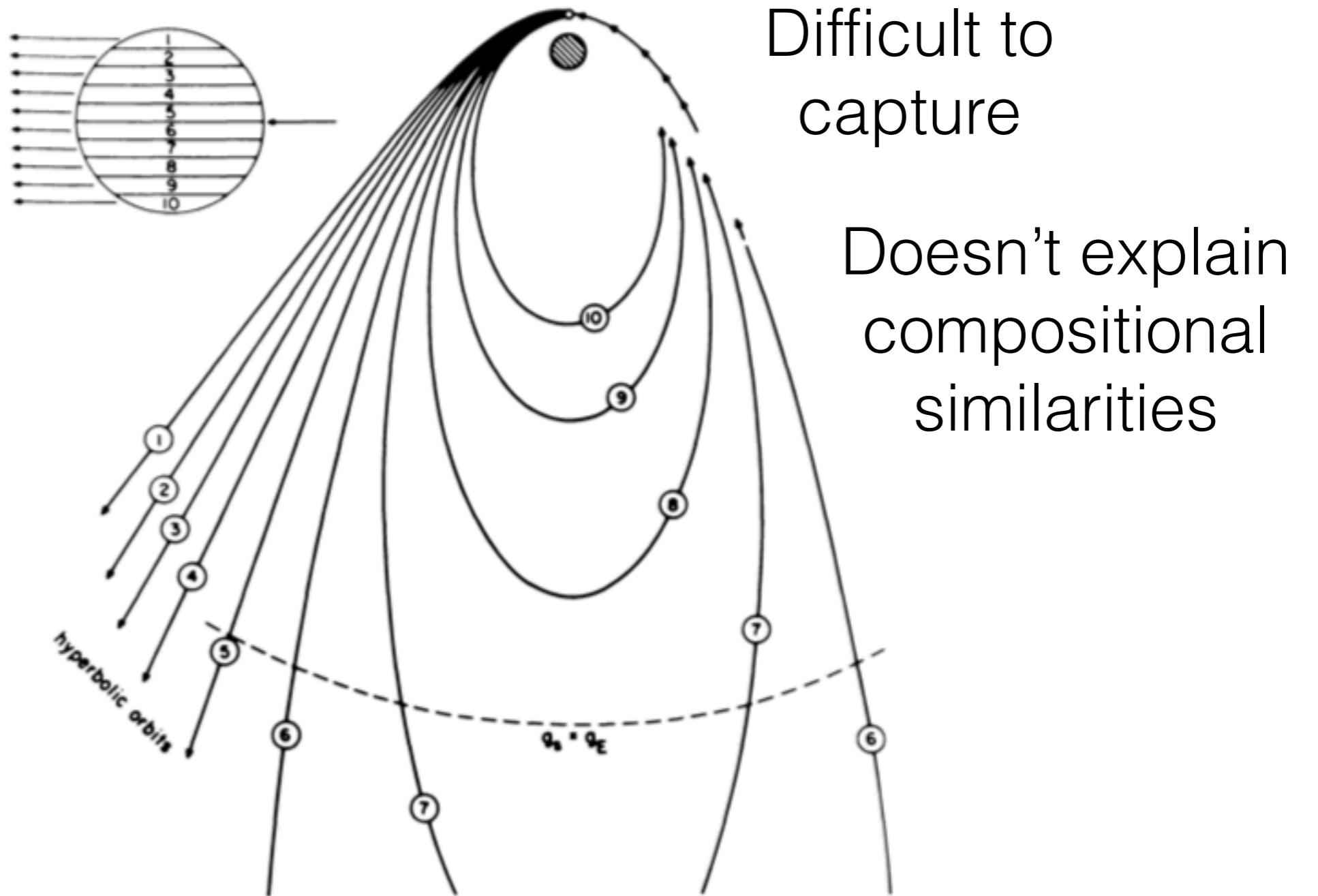
- Angular momentum
- Composition (lack of volatiles)
- Small (iron) core

# Fission

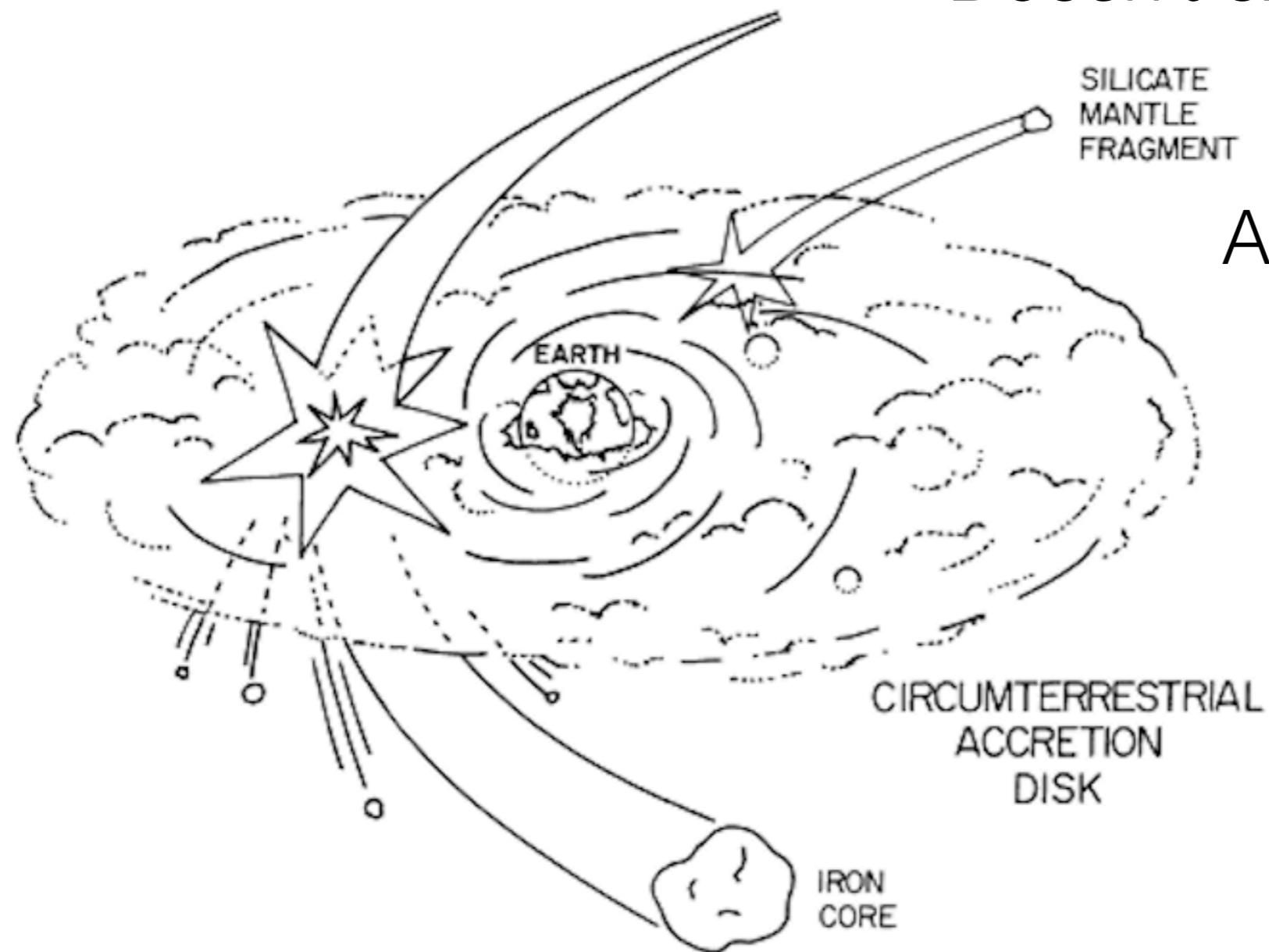


Requires very  
large angular  
momentum

# Capture



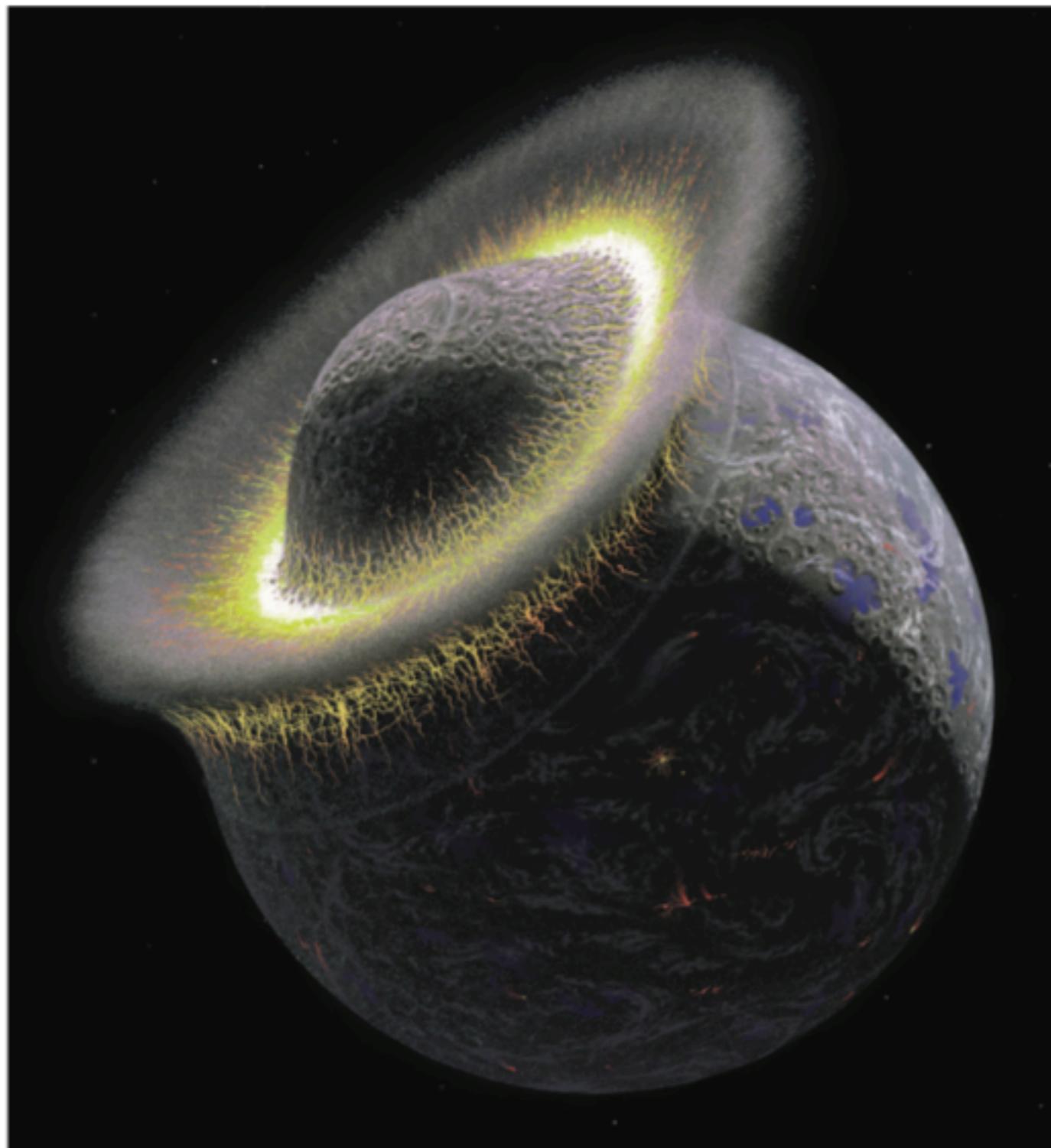
# Binary Accretion



Doesn't explain small core

Angular momentum  
issues

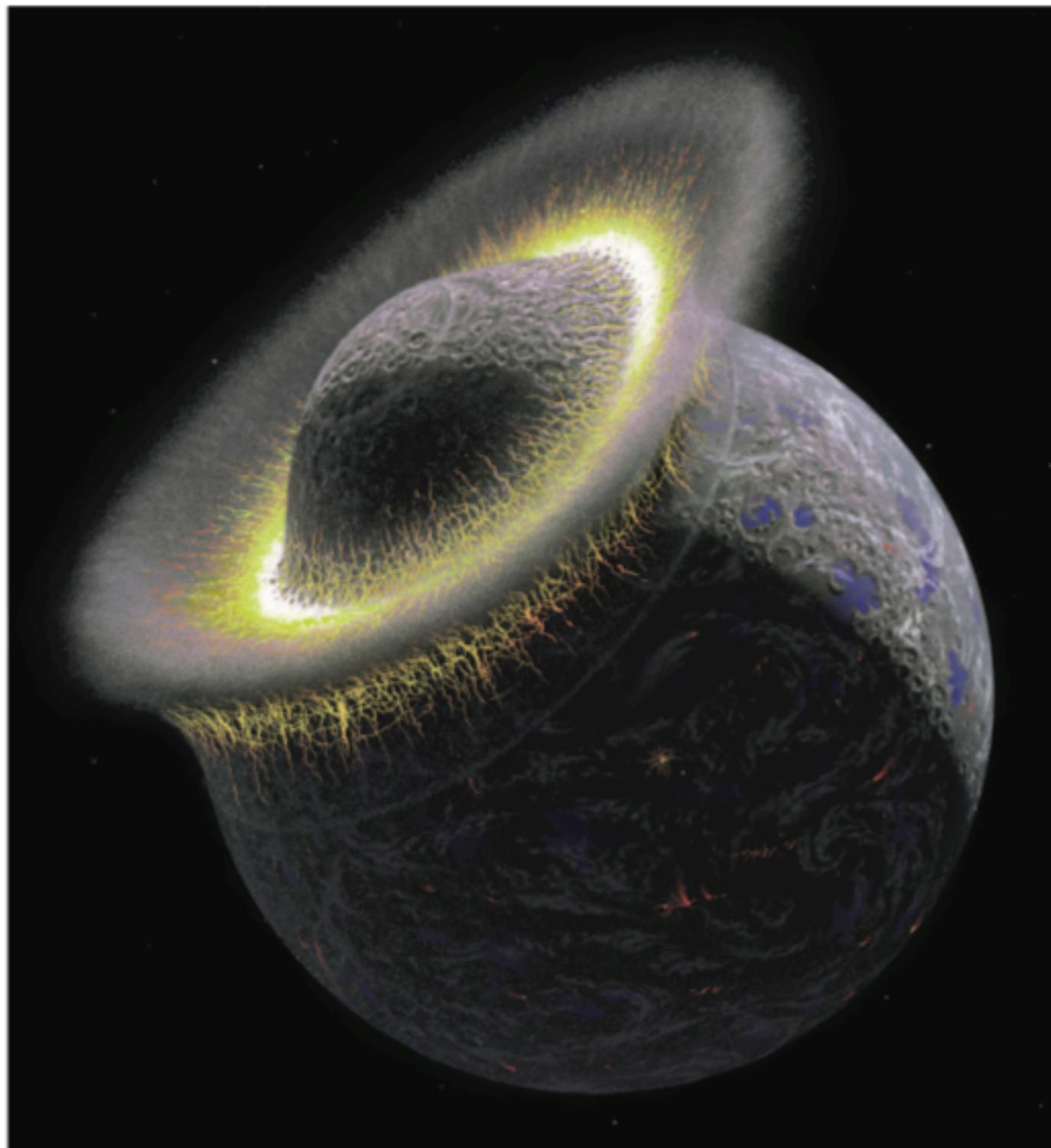
# Giant Impact



General Model:

- Mars sized impactor ('Theia')
- Angled impact
- Moon accretes from debris disk

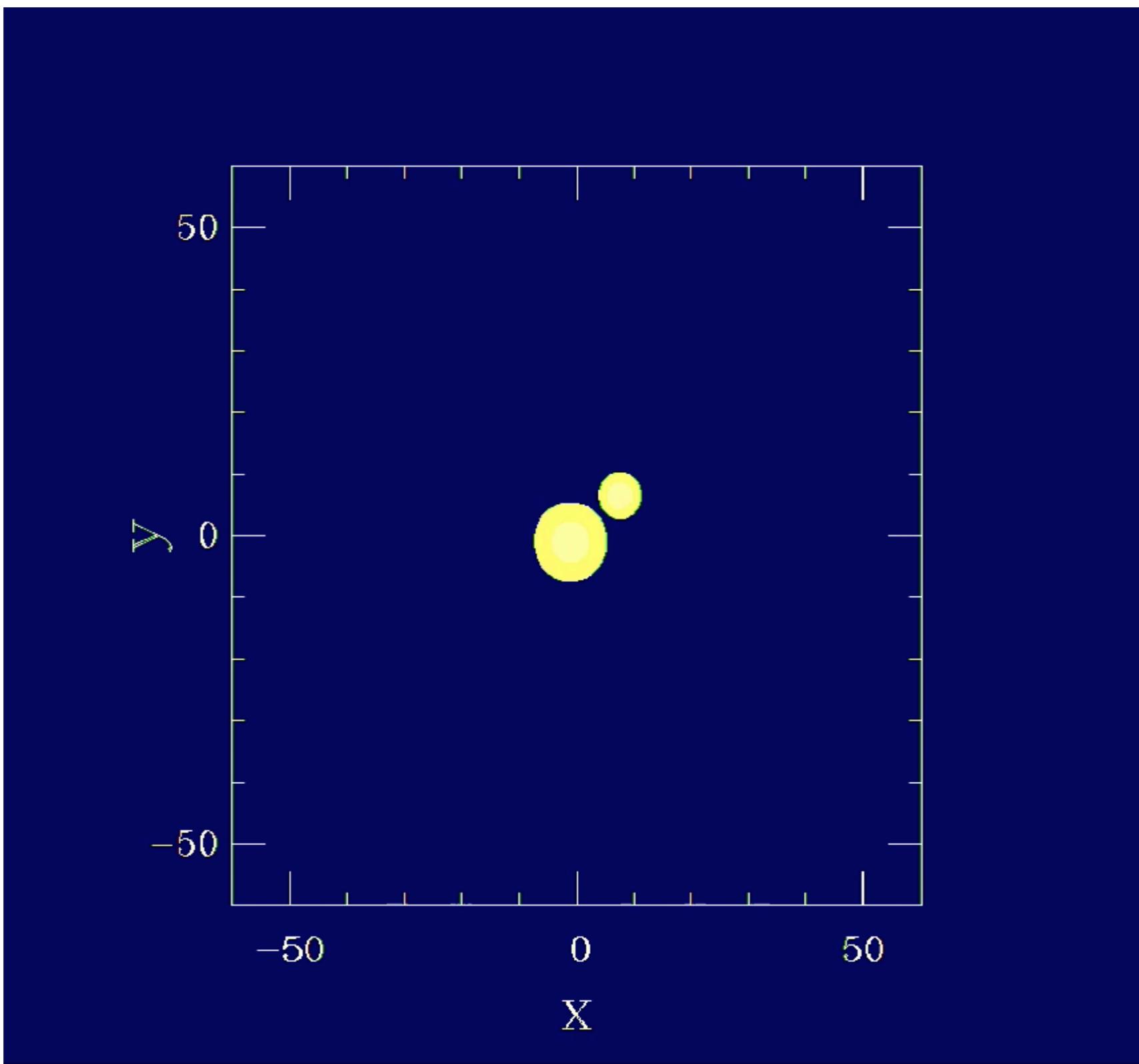
# Giant Impact



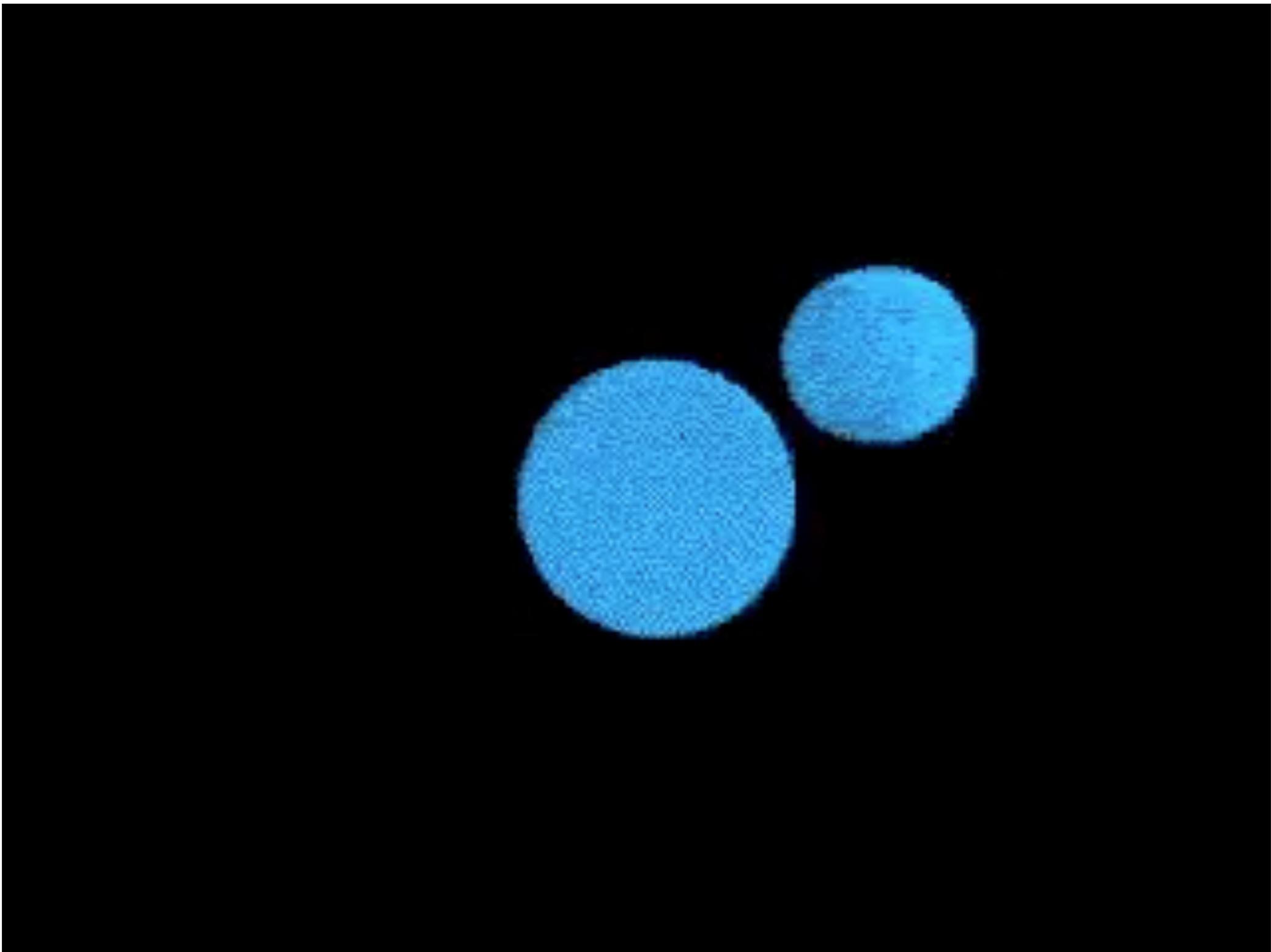
## Features:

- Angular moment
- Depleted iron core
- Common in early solar system

# Impact Simulation

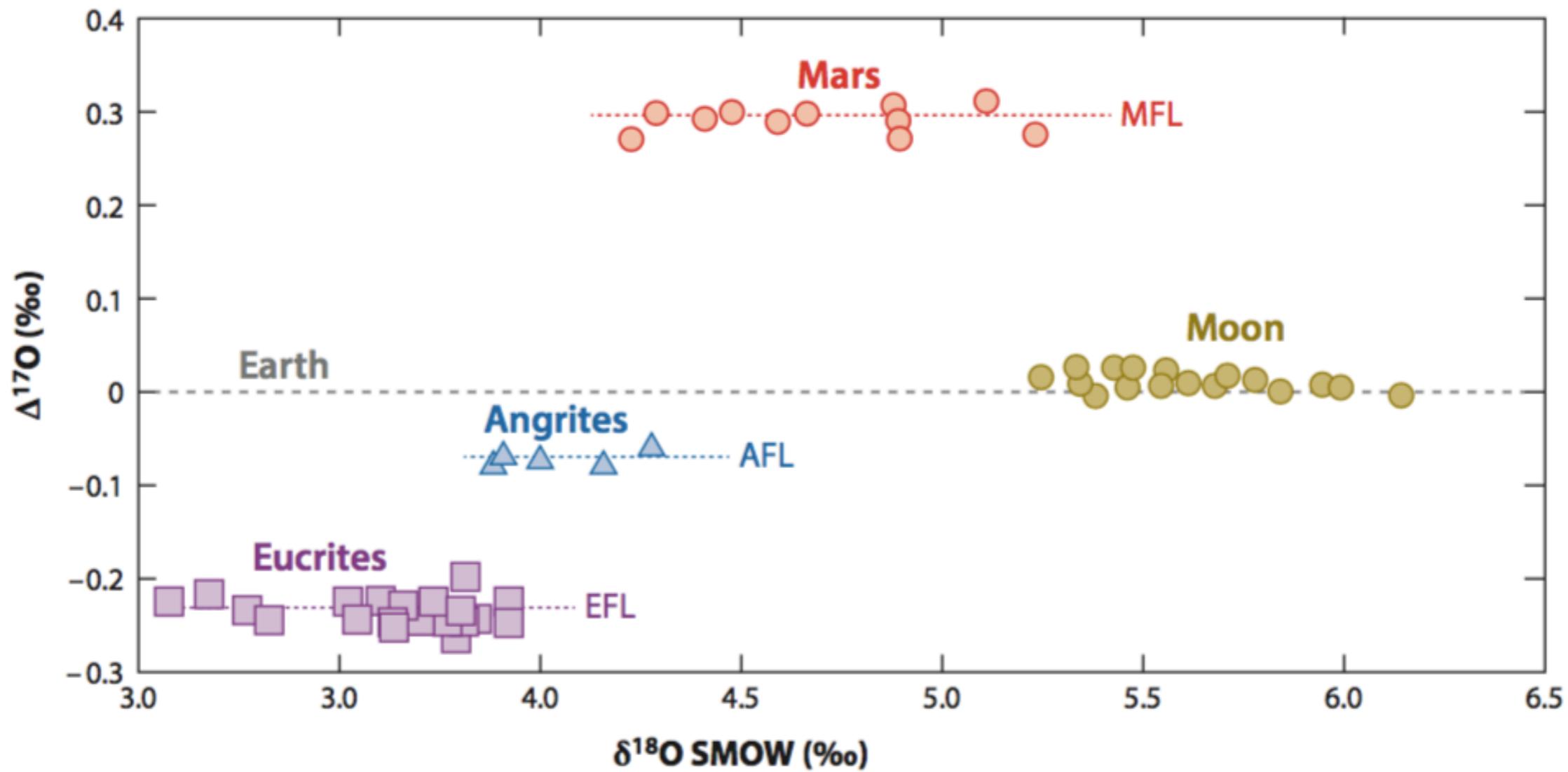


# Impact Simulation (Cont.)



# Geochemistry Challenge

The Moon is TOO similar to Earth!

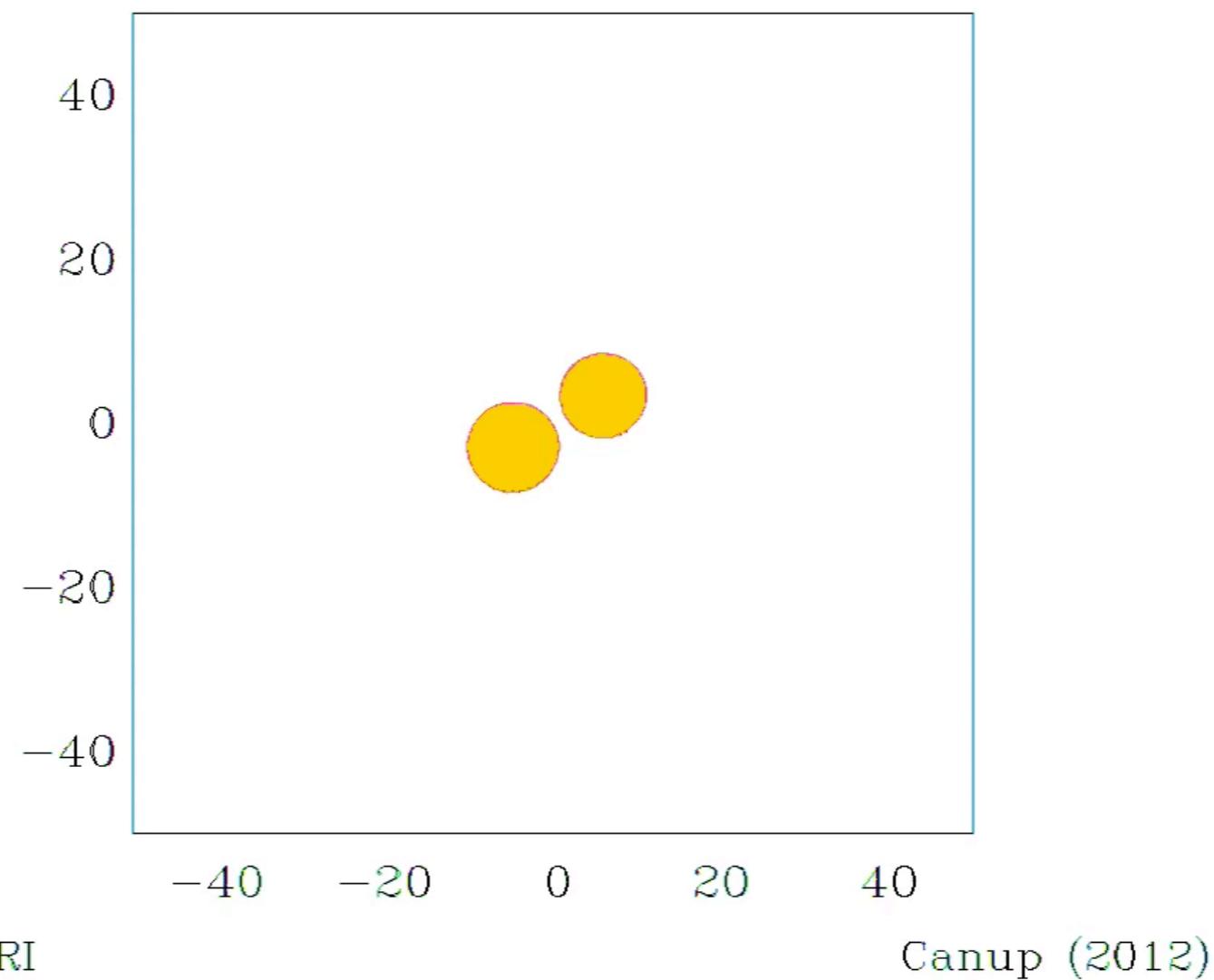


Asphaug 2014

Also similar in Ti, W, K, O...

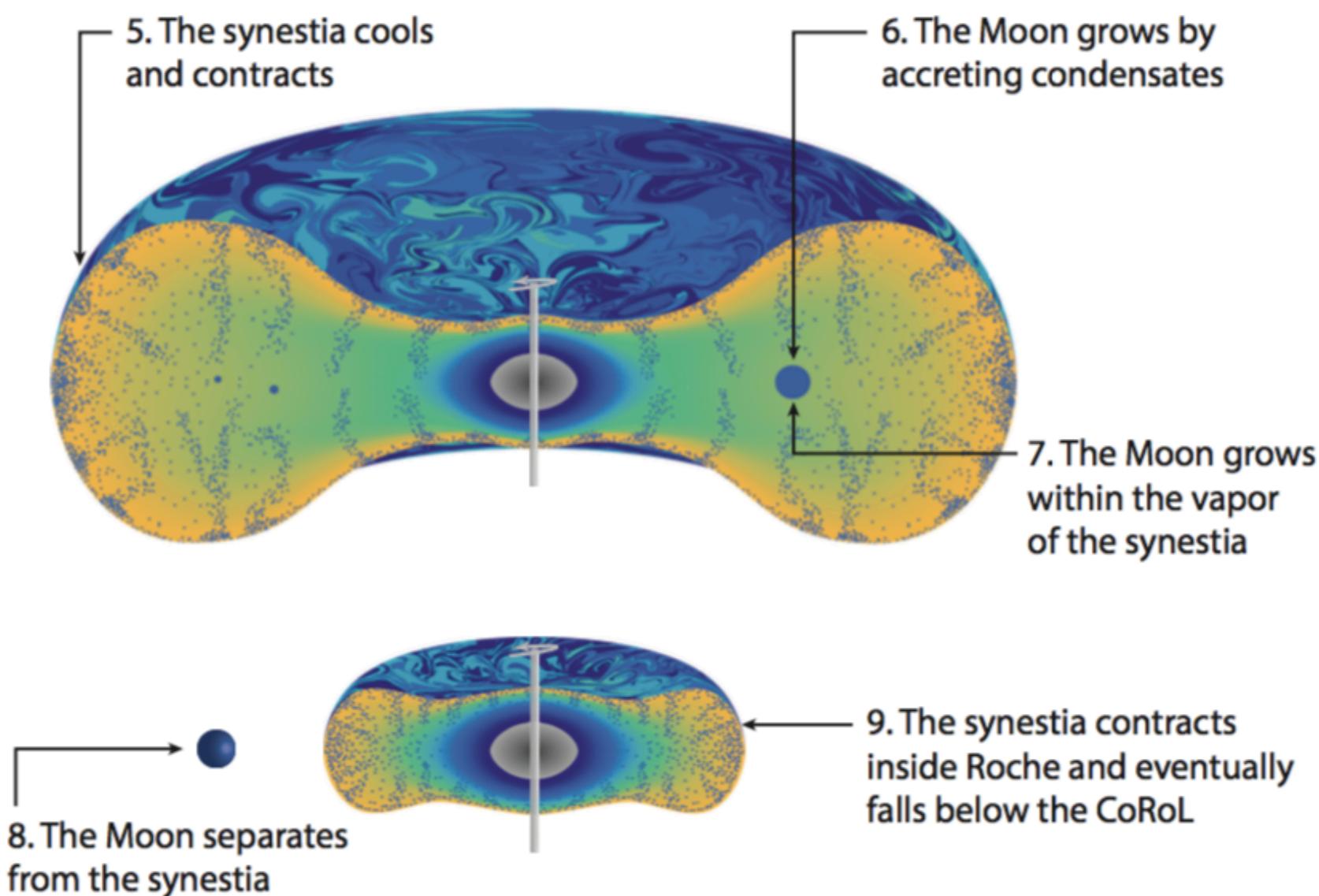
# Alternative Impact Models

Two objects with equal masses



# Alternative Impact Models (Cont.)

Donut shaped cloud of debris ('synestia')



# Outline

Moon  
Formation  
(~ 4.51 Ga)



# The Late Heavy Bombardment

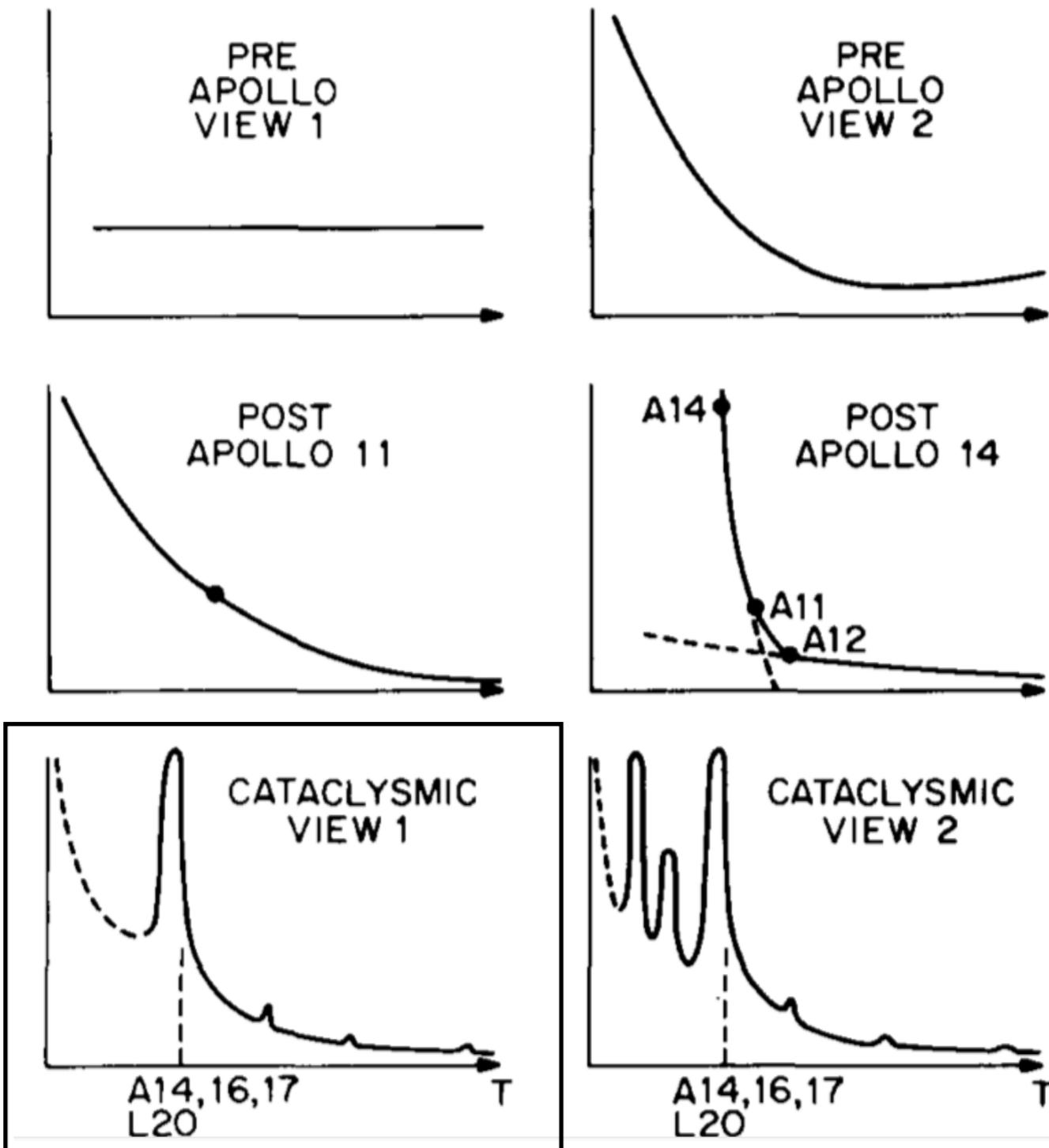
First ~600 million years of solar system



Credit: NASA's Goddard Space Flight Center Conceptual Image Lab

# Models of Impact Flux and the ‘Terminal Cataclysm’

CRATERING RATES

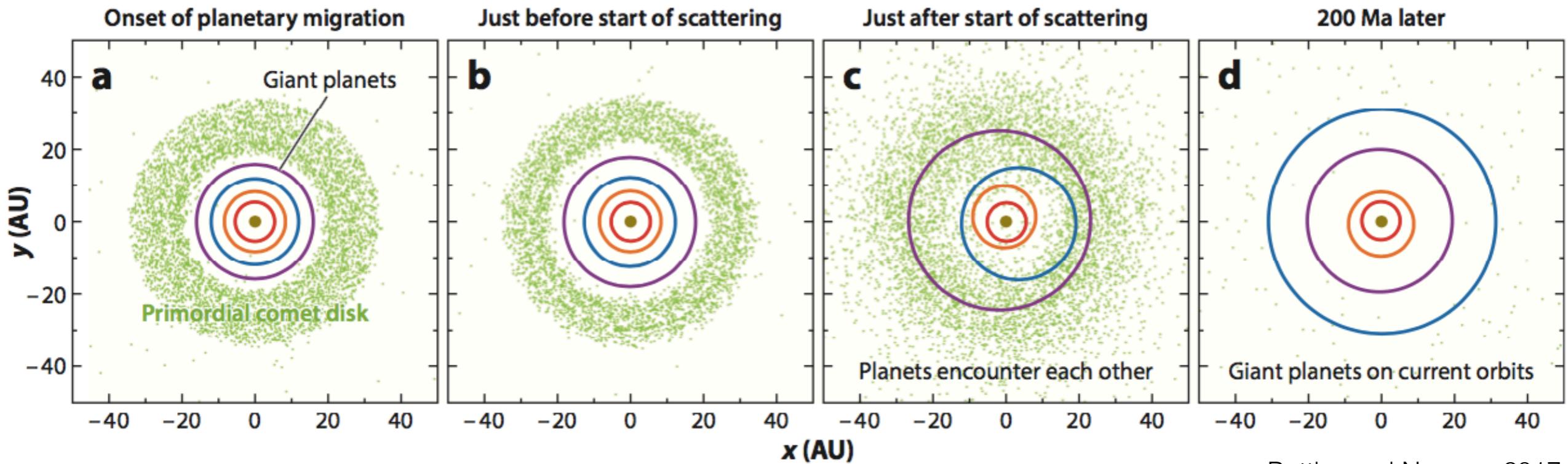


Prevailing view  
from 1990s

Tera et al. 1974

# The Nice Model

How do we explain a late impact flux increase?

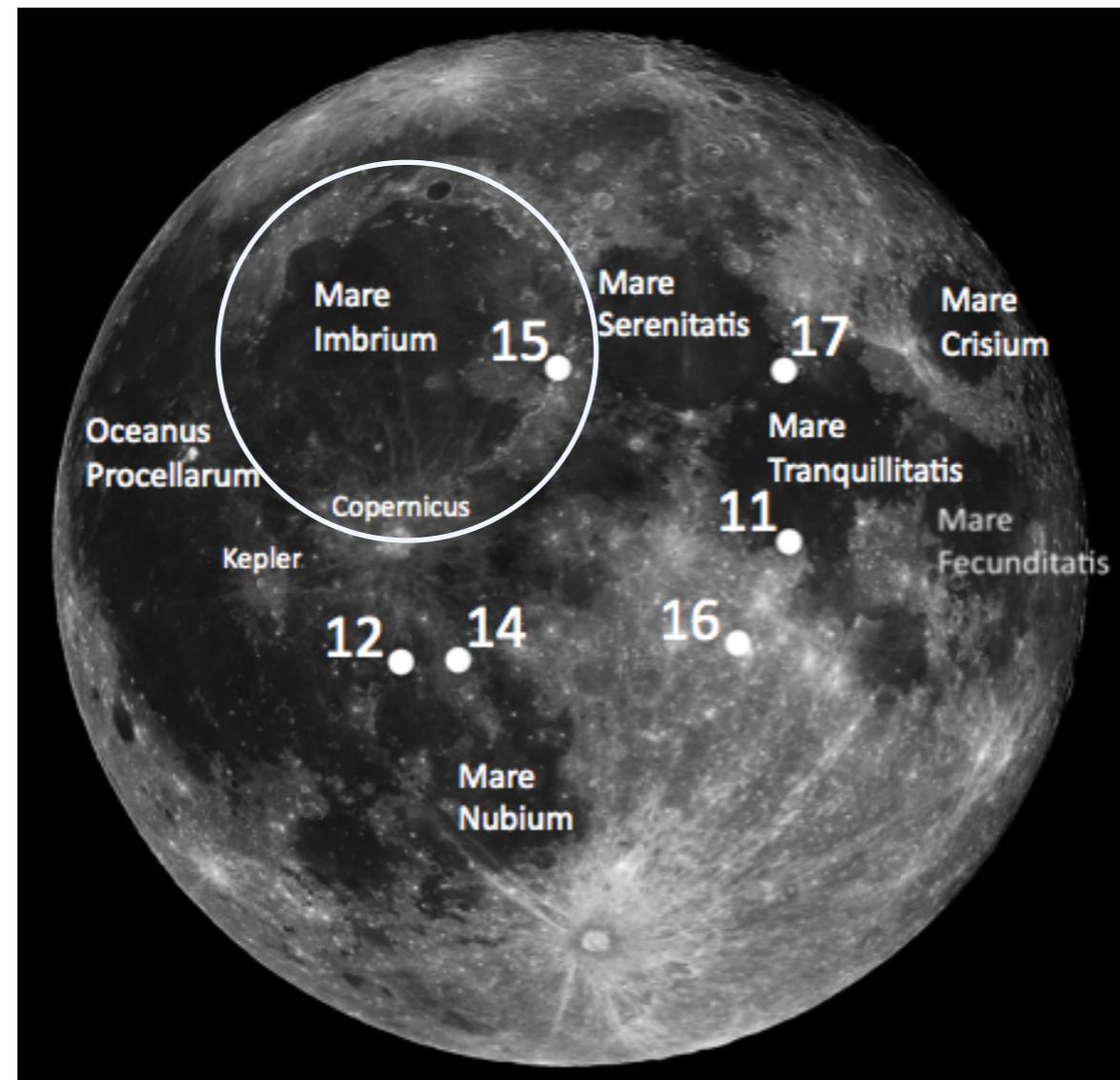


Bottke and Norman 2017

Many modifications e.g. ‘Jumping Jupiter model’

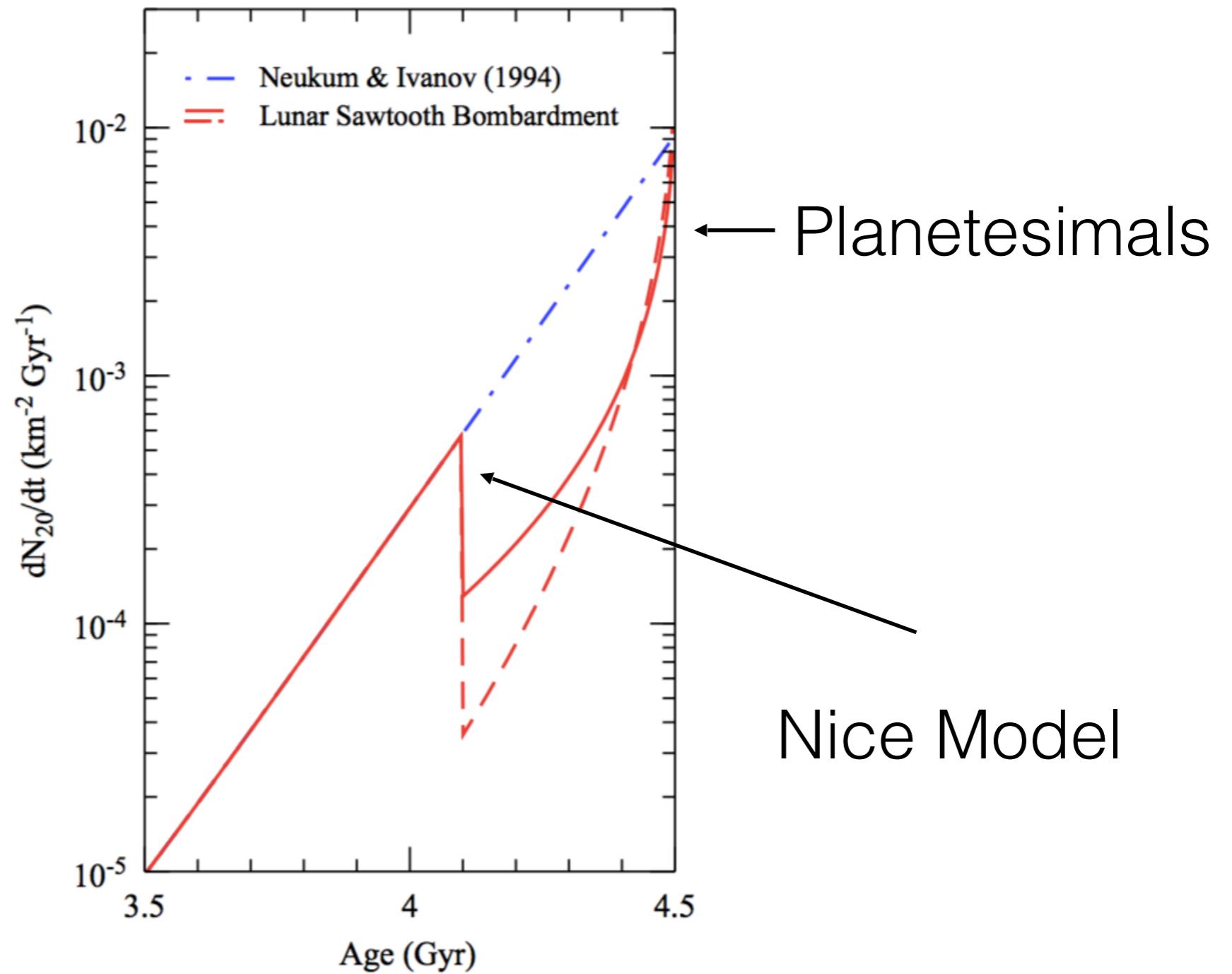
# Did a Terminal Cataclysm Happen?

- Material > 3.9 billion years old
- Contamination from Imbrium
- Orbital data reveals more impact craters



# Sawtooth Model

Modest increase starting ~4.2 - 4.1 Ga

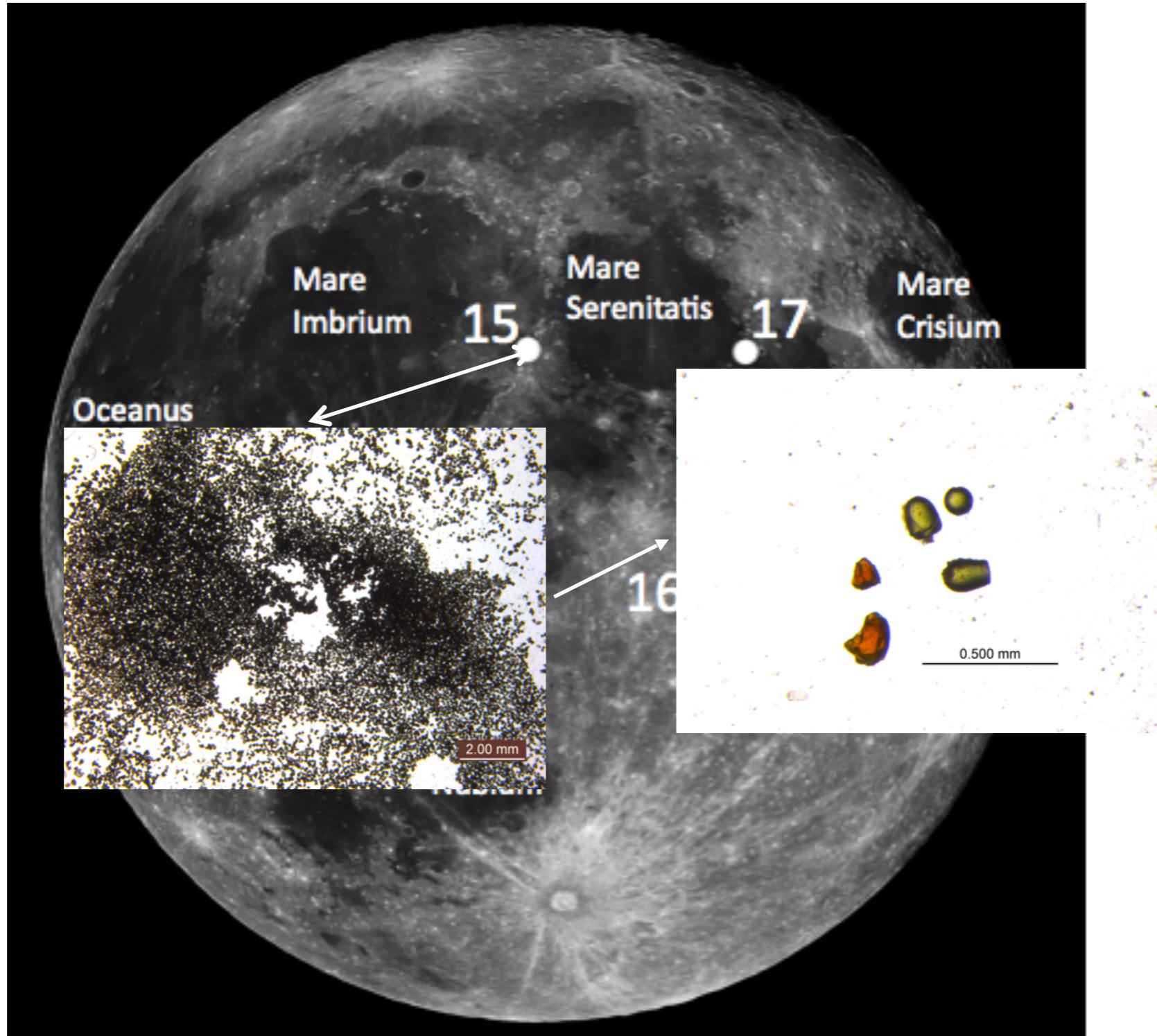


# Outline

Moon  
Formation  
(~ 4.51 Ga)



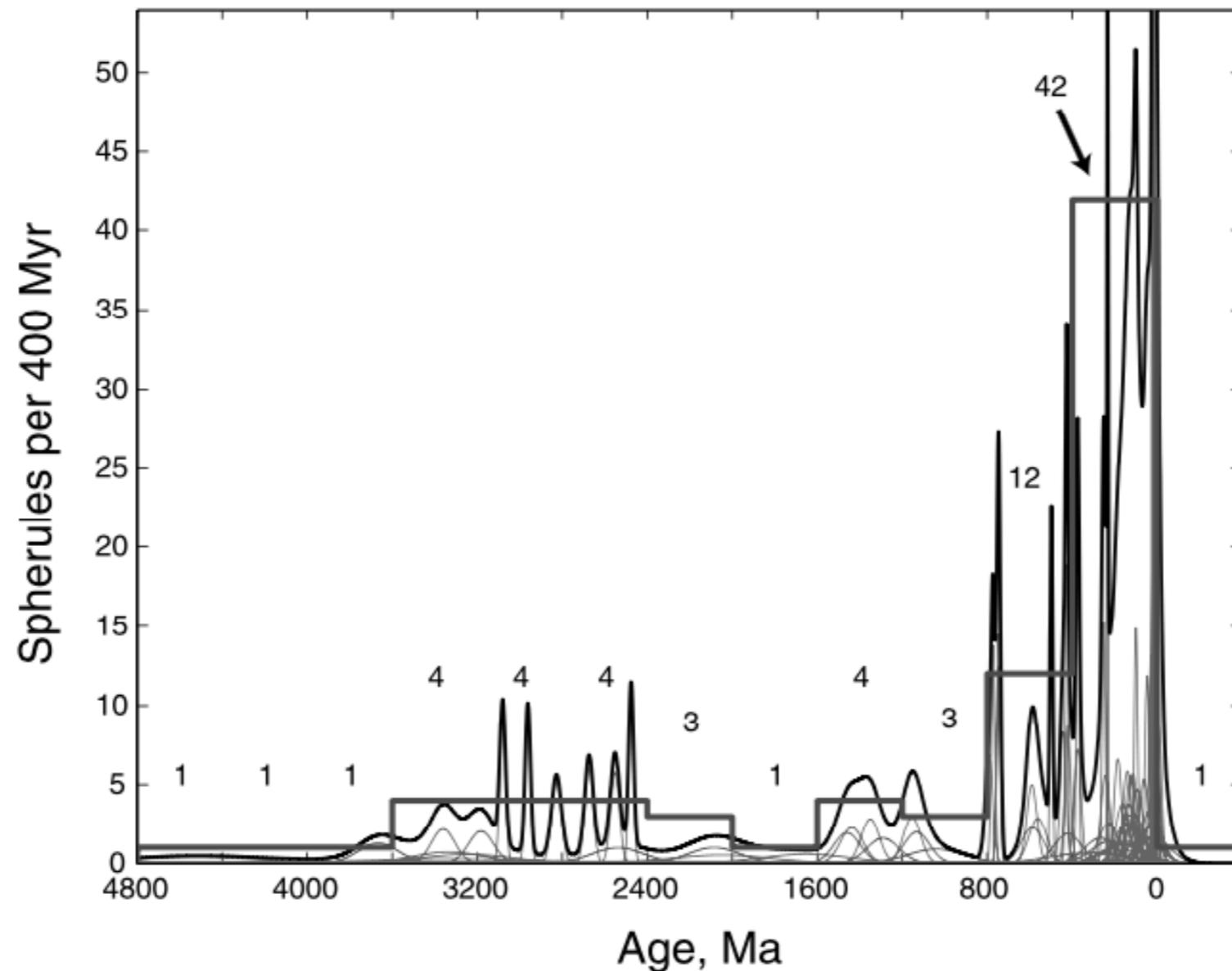
# Lunar Impact Glasses



Credit: Dustin Scriven

# Research Project

Investigate claimed increase in flux over last 0.4 Gy



# Impact Glasses

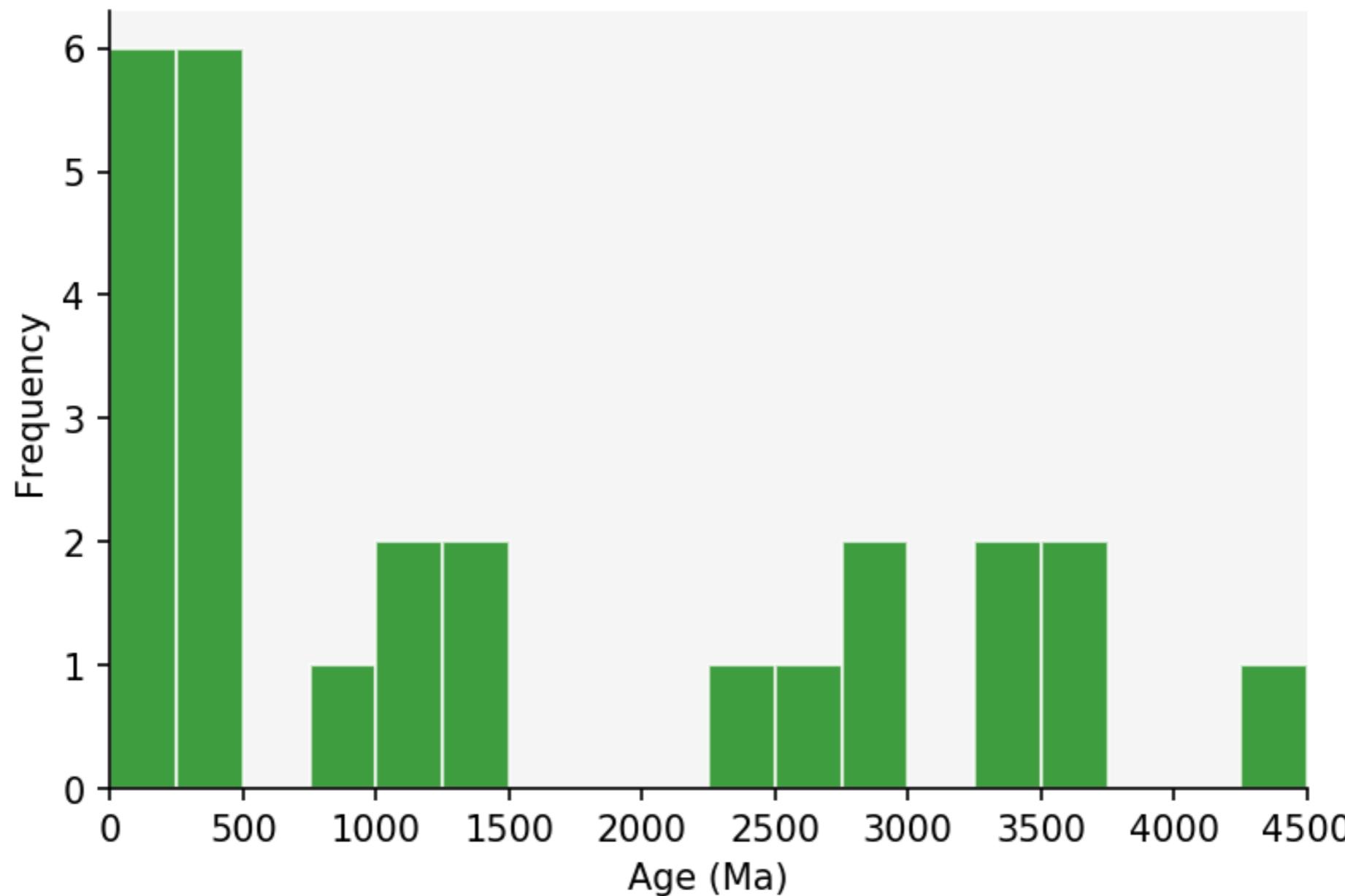
Characteristics:

- Age ( $^{40}\text{Ar}/^{39}\text{Ar}$  dating)
- Composition (MgO, TiO<sub>2</sub>, etc.)
- Shape (shards vs spheres)

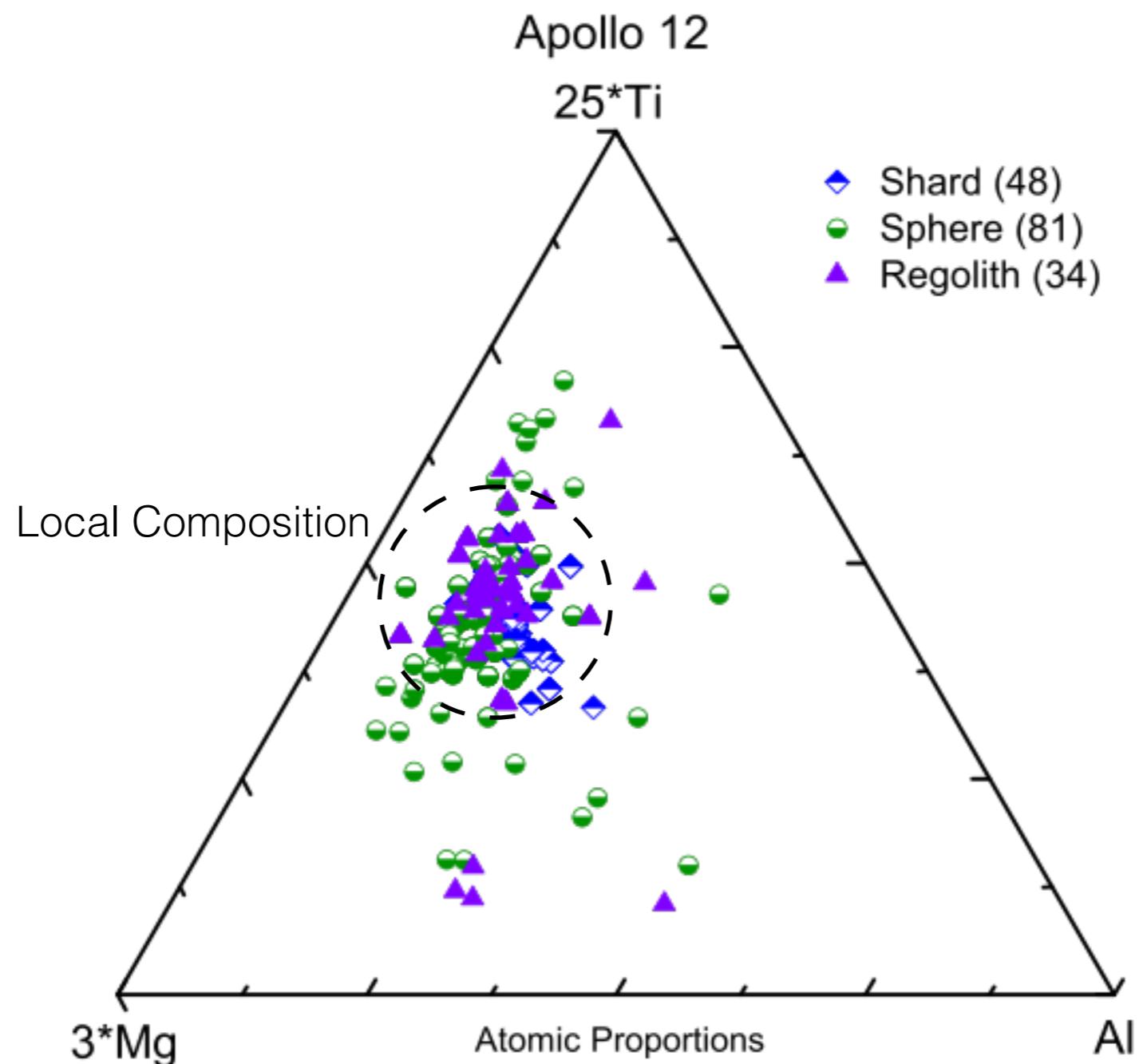
Key question: Are spheres bias to young ages?

# Age Trends

Glasses with Fair or Good Age Quality (Spheres)



# Compositional Trends



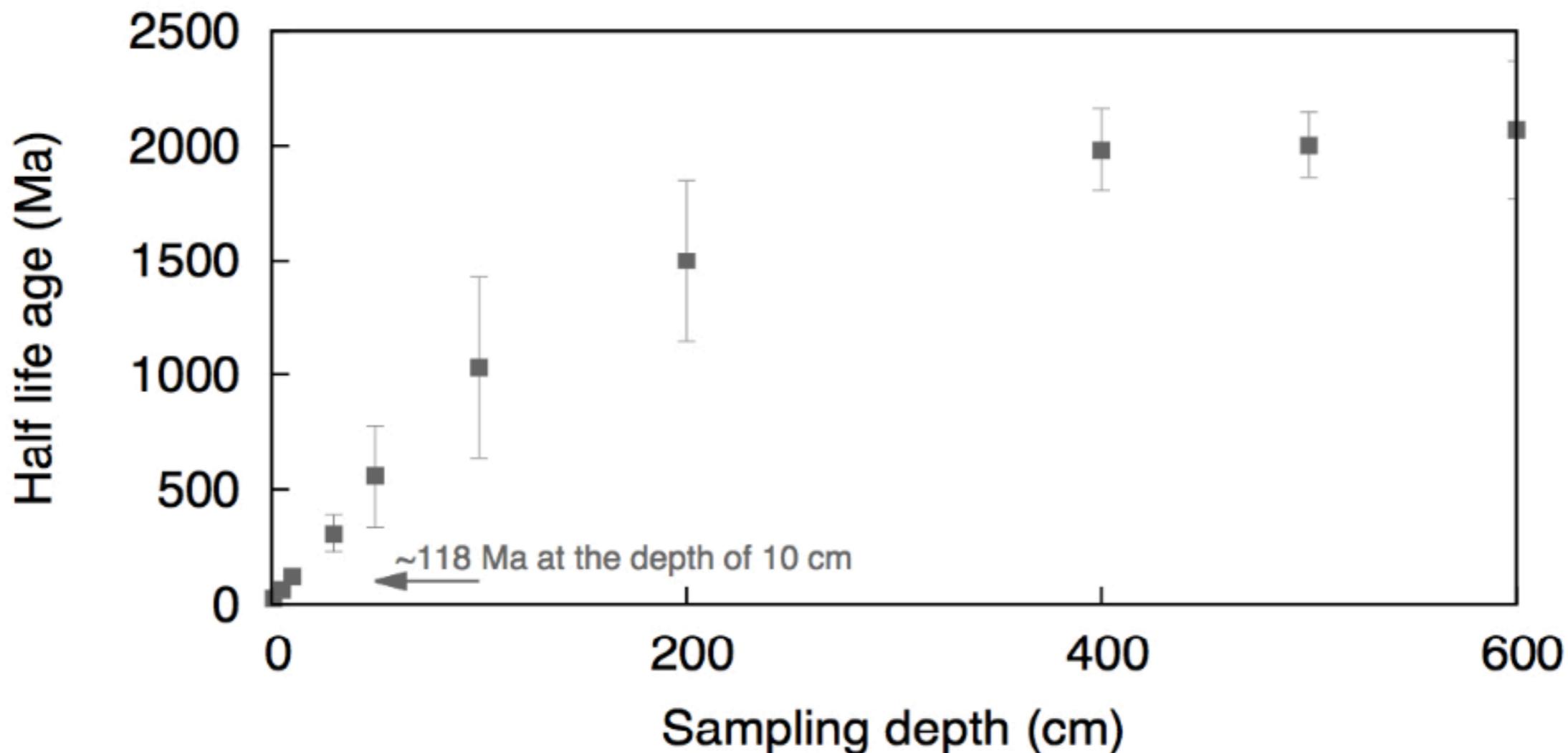
Korotev et al. (2011), Levine et al. (2005), Meyer et al. (1971), and Wentworth et al. (1994)

# Hypothesis

Over time, Impact events destroy older spheres

# Regolith Dynamics Models

Simulations can model impact events and populations of impact glasses



# Conclusion

Apparent increase in impact flux is the result of sampling bias

# Summary

- Moon formed from an impact event
- ‘Terminal Cataclysm’ falling out of favor
- No recent increase in impact flux

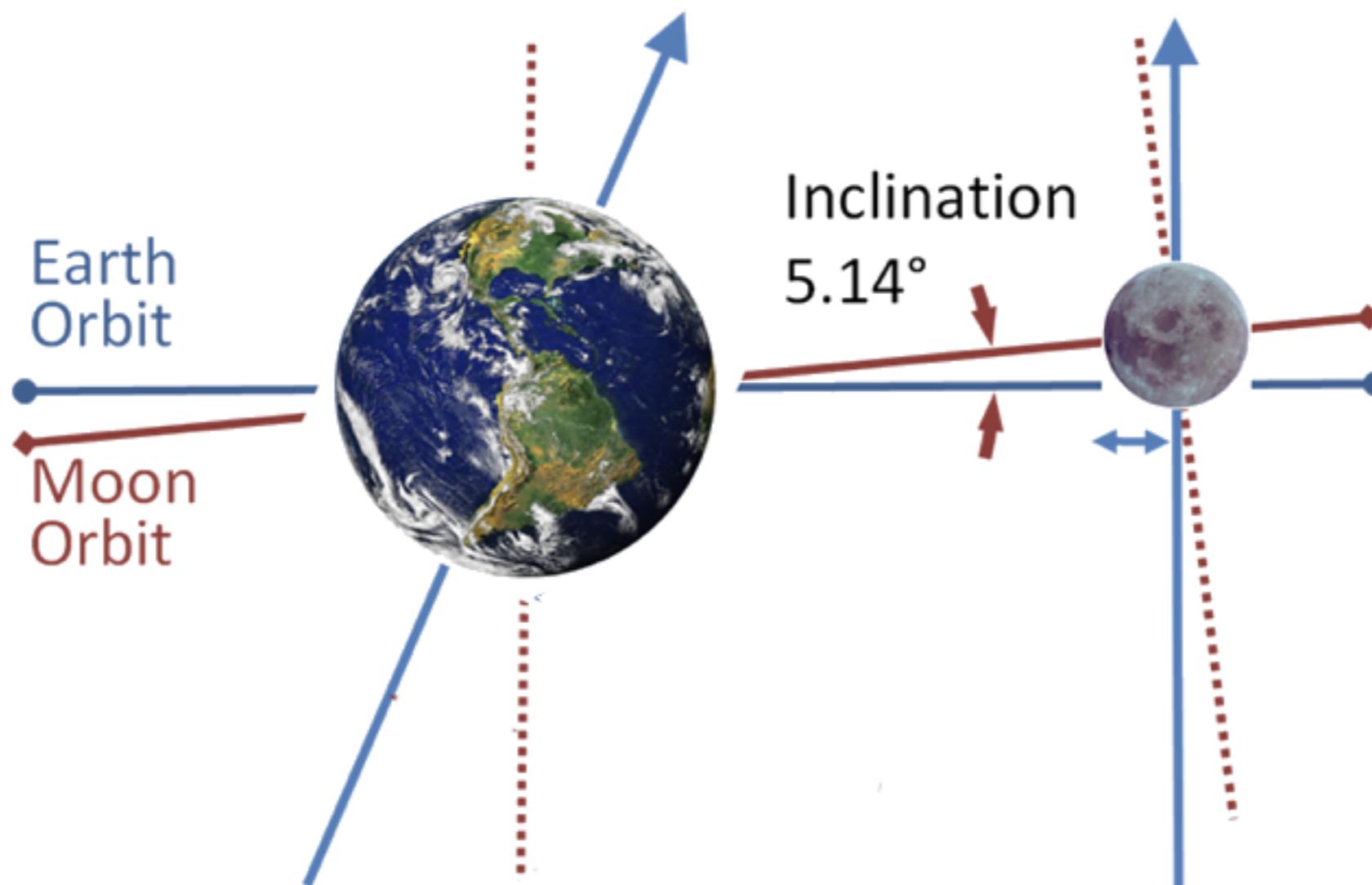
# Acknowledgements

- Dr. Zellner
- Albion College
- NSF Astronomy and Astrophysics grant

Thank you!

# Problems

## Moon's orbital inclination

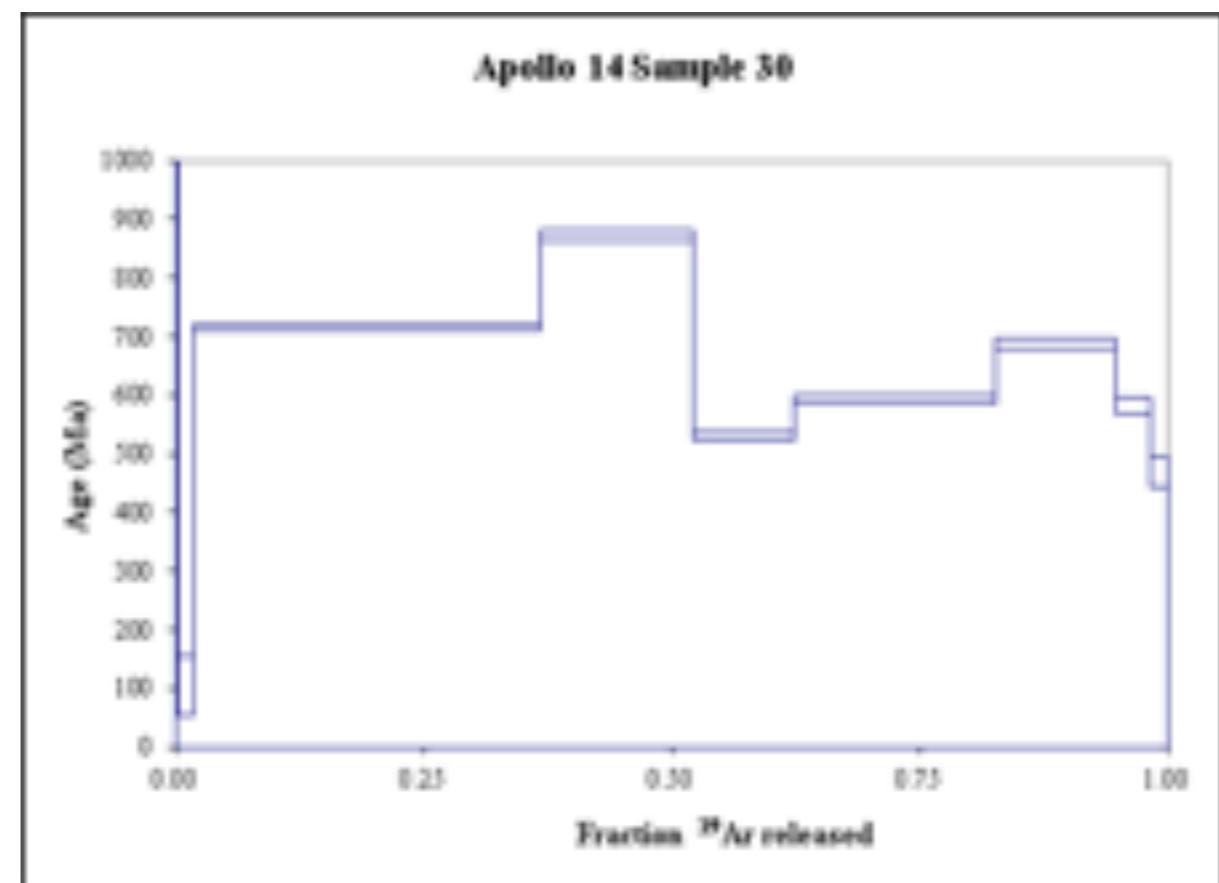
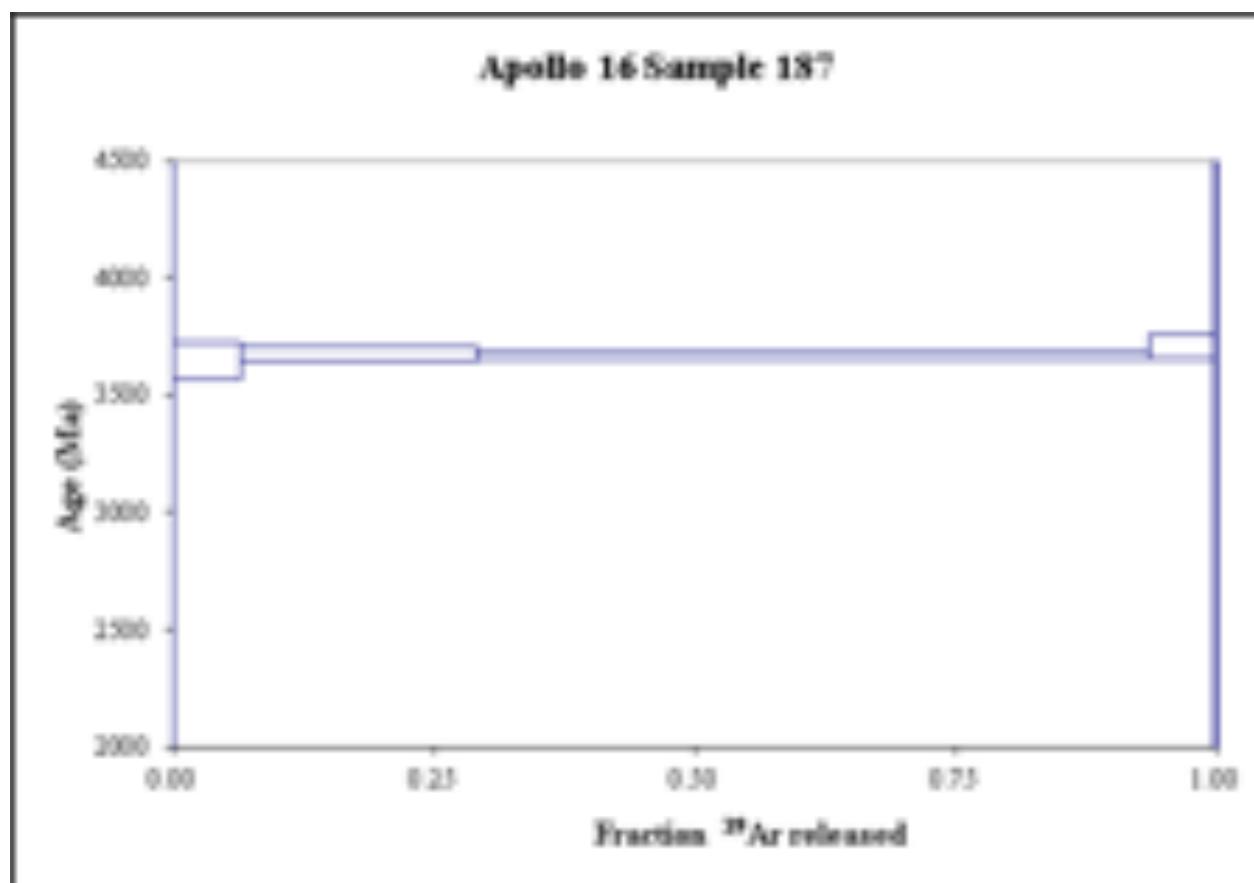


# Solutions

Gravitational interaction with planetesimals

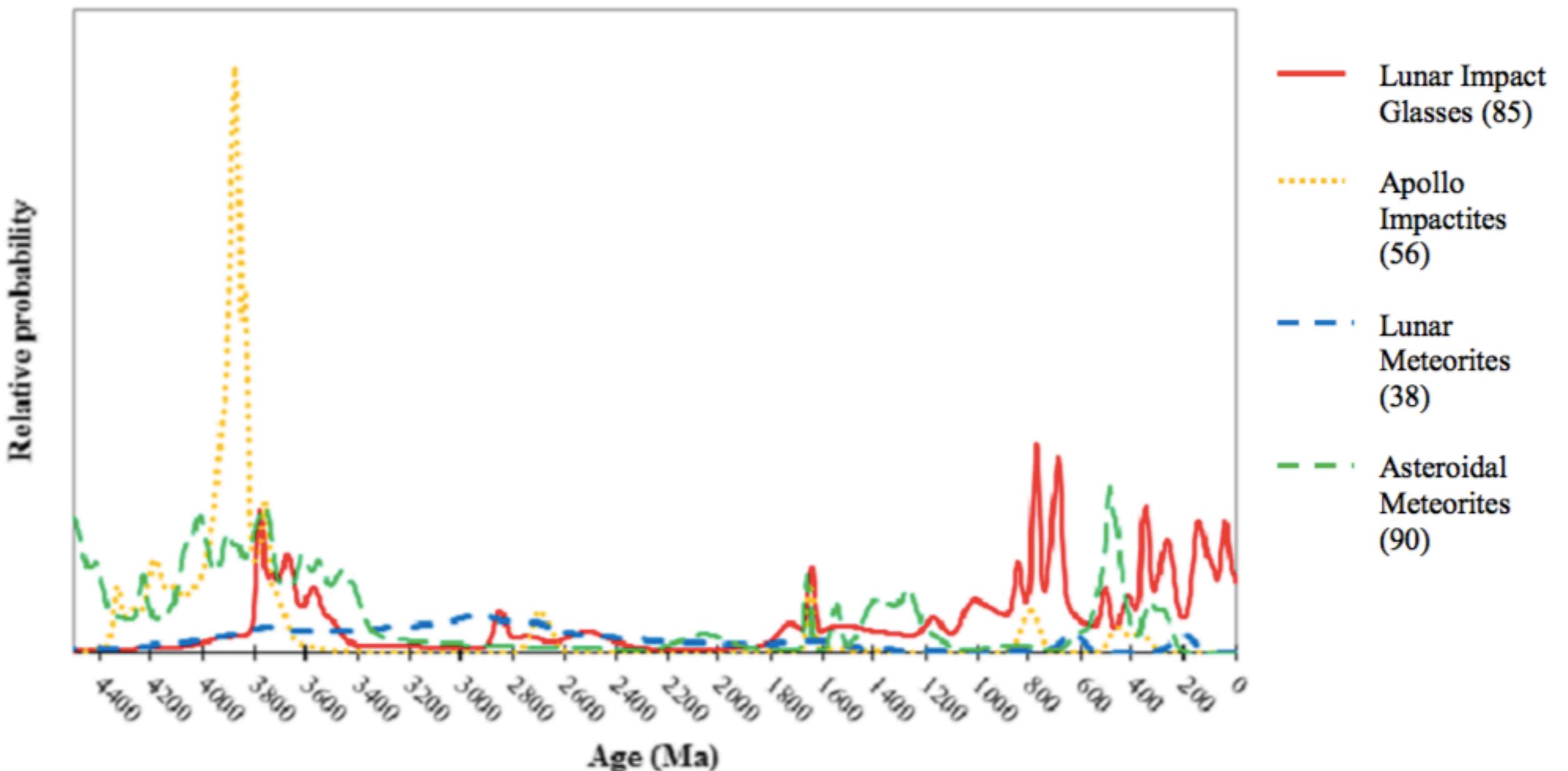
Tidal evolution from the Sun

# Argon Dating



# Timeline

All Impact Data



# Basin Ages

Crater	Age (Ga) (1974–2006)	Age (Ga) (2009-present)
South Pole - Aitken	4.05 – ~4.3	4.0–4.4 (?)
Serenitatis	$3.893 \pm 0.009$	3.83–4.1+
Nectaris	3.89–3.92	3.92–4.2 (?)
Crisium	3.85–3.93	~3.9 (?)
Imbrium	$3.85 \pm 0.02$	3.72–3.93
Orientale	3.77–3.83	3.72–3.93

Zellner 2017