

# Non-Relativistic Anisotropic Magnetoresistance

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Anisotropic magnetoresistance (AMR) is a manifestation of magnetic-order-induced symmetry lowering of conductivity tensor. While the AMR in simple ferromagnets is usually considered to be a relativistic effect (it relies on spin-orbit interaction), we show that symmetry lowering, similar in spirit, can also happen in a non-relativistic limit. Using tight-binding models and Boltzmann transport theory, we investigate systems with multiple magnetic sublattices, including both collinear and non-collinear antiferromagnets, as well as ferrimagnetic configurations. We show that AMR and related anisotropies can emerge purely from magnetic order, without spin-orbit interaction, and may reach appreciable magnitudes. The findings are supported by case studies on toy-model lattices and real materials such as MnN, Mn<sub>3</sub>Sn, and EuTe<sub>2</sub>, and are further interpreted through symmetry analysis based on Neumann's principle.

## I. INTRODUCTION

Anisotropic magnetoresistance (AMR), first observed by William Thomson in 1857<sup>1</sup>, describes the dependence of resistivity in ferromagnetic materials such as cobalt and nickel on the direction of magnetization. Since then, AMR has remained an active subject of research<sup>2</sup>. In its conventional form, AMR is observed in ferromagnets (FMs), where it manifests as a two-fold (180°-periodic) angular dependence of the resistivity on the angle  $\varphi$  between the current and the magnetization direction<sup>3</sup>:

$$\frac{\Delta\rho}{\rho} \propto \cos 2\varphi \quad (1)$$

This implies that, for example, if the magnetization in an otherwise isotropic medium ( $\sigma_{xx} = \sigma_{yy}$  above the magnetic ordering temperature) is oriented along the  $x$ -direction, the conductivity tensor  $\sigma = \rho^{-1}$  becomes anisotropic with  $\sigma_{xx} \neq \sigma_{yy}$ .

The conventional realization of AMR, as described in Eq. 1, relies on spin-orbit coupling (SOC), which couples the spin of electrons to the crystal lattice. However, this description is not exhaustive. In this work, we demonstrate that similar symmetry-lowering effects in the conductivity tensor can also arise in the absence of SOC, provided the magnetic order involves multiple magnetic sublattices (MSLs), including both collinear and non-collinear magnetically ordered systems.

### A. Definition of AMR

More generally, AMR can be defined as the change in symmetry of the conductivity tensor  $\sigma_{ij}$  due to magnetic ordering. This includes the standard formulation in Eq. 1, but excludes anisotropies arising from purely orbital effects in non-magnetic systems or surface states in topological insulators<sup>2</sup>. Historically, AMR has also appeared under various names in the literature—such as spontaneous magnetoresistance anisotropy

(SMA)—though these are not always consistently defined throughout literature<sup>2</sup>. The above definition offers a general framework encompassing these different terminologies.

We distinguish two commonly realized forms of AMR:

1. *Rotation of magnetic order*: This is the most widely used experimental approach. AMR is measured while the magnetic order is rotated via an applied magnetic field. Such studies often focus on systems with collinear magnetic order, including FMs, antiferromagnets (AFMs) like MnTe<sup>4,5</sup> and CuMnAs<sup>6–8</sup>, or ferrimagnets such as Mn<sub>4</sub>N<sup>9</sup>. These systems are characterized by a single spin axis (SSA), either the magnetization  $\vec{M}$  or the Néel vector  $\vec{L}$ . Let  $\varphi_1 \neq \varphi_2 \pm n\pi$  be two different orientations of the SSA, then AMR is present if  $\sigma_{ii}(\varphi_1) \neq \sigma_{ii}(\varphi_2)$  for some direction  $i$ . This mechanism necessarily requires SOC and is thus relativistic.
2. *Measurement of conductivity along different directions*: AMR can also arise when the conductivity is measured along two perpendicular directions for a fixed magnetic configuration  $\varphi$ . For instance,  $\sigma_{xx}(\varphi) \neq \sigma_{yy}(\varphi)$ . This anisotropy can either be a spontaneous effect<sup>10</sup> or induced by external manipulation of the magnetic configuration. In **simple FMs with a single MSL**, in the absence of SOC, no such anisotropy is induced under global spin rotation, since spin and lattice are decoupled<sup>11</sup>. However, in systems with multiple MSLs, even without SOC, magnetic ordering can break crystal symmetries and lead to anisotropic conductivity. This fact is illustrated in Fig. 1

Both approaches are often considered equivalent but we will discuss situations where they are not. Role of SOI to be discussed in Sec. I.D (simultaneous rotations of all mmoms cannot change  $\sigma_{ij}$  if SOI is absent).

## B. Categories of AMR

Beyond the manner of realization, AMR can be further categorized based on its physical origin:

*a. Intrinsic vs. Extrinsic.* Intrinsic AMR refers to symmetry-breaking effects that are scattering-independent, whereas extrinsic AMR arises from spin-dependent scattering. For a long time, the extrinsic mechanism received more attention, and only a few works acknowledge scattering-independent contributions to the AMR<sup>12–17</sup>. Frequency-dependent studies<sup>12,13</sup> allow these two contributions to be distinguished experimentally: the extrinsic contribution scales roughly with  $1/\omega$ , while the intrinsic one is frequency-independent. Intrinsic effects are well known in the anomalous Hall effect (AHE) and spin Hall effect (SHE), where they are linked to Berry curvature<sup>18,19</sup>. In contrast, the origin of intrinsic AMR is related to the anisotropy of the Fermi surface, which will be explored in Sec. III. **There's a recent claim<sup>48</sup> that Berry curvature can also contribute to intrinsic AMR... think about it.**

*b. Non-Crystalline vs. Crystalline.* The AMR signal in polycrystalline samples follows the  $\cos 2\varphi$  dependence of Eq. 1 and is called non-crystalline AMR. In samples with a high crystalline quality, more complex angular dependencies emerge, such as four-fold symmetries in e.g. Ni<sup>20</sup>, Co<sub>2</sub>MnGa<sup>21,22</sup>, and (Ga,Mn)As<sup>23</sup>, six-fold symmetries e.g. in hexagonal MnTe<sup>4,5</sup>, and sometimes even higher symmetries<sup>5,24</sup>. Although such features have been known for decades, they are sometimes misinterpreted as magnetocrystalline anisotropy<sup>2</sup> or titled to be newly discovered effects<sup>25</sup>. Analysis of higher angular harmonics are usually conducted in systems with an SSA, but can be expanded effortlessly to include MSLs as shown in appendix C.

This shall only serve as a short introduction to the topic of Anisotropic Magnetoresistance. A more comprehensive overview can be found in Ref.<sup>2</sup>.

## C. Beyond Spin-Orbit Coupling

This study builds on earlier findings (see Sec. 4.2.3 of Ref.<sup>2</sup>) where we demonstrated the first realization of non-relativistic AMR. Recently, related effects have been explored, including spin textures<sup>26</sup> and spin-orbit-torque-like phenomena<sup>11</sup> in non-collinear magnets without SOC. In the non-collinear antiferromagnet Mn<sub>3</sub>Sn, the strong spin Hall effect (SHE)<sup>18,27</sup> and local (sublattice-projected) Edelstein effect<sup>11</sup> were found to prevail also in absence of SOC<sup>11,28</sup> indicating that complex magnetic order can mimic relativistic effects. Non-relativistic effects in the novel framework of altermagnetism were explored in Ref.<sup>29</sup>. Related phenomena, specifically the non-relativistic parity-breaking of Fermi surfaces leading to conductivity anisotropies, were investigated in a separate study<sup>30</sup>, which focuses on the symmetry-based characterization of magnetic phases exhibiting such features.

These are referred to as *p-wave magnets* by the authors, with the emphasis placed on the symmetry classification of this class.

**Emphasize the fact that w/o SOI, the simultaneous rotation of magnetic moments will not change  $\sigma_{ij}$ .**

## D. Organization

In this work, we explore how AMR can arise in the absence of SOC by focusing on both intrinsic and extrinsic mechanisms:

- **Intrinsic AMR:** We analyze how magnetic-order-induced anisotropy can manifest when scattering is disregarded in three different ways: As a spontaneous effect in the collinear antiferromagnet MnN, due to manipulation of magnetic moments by an external magnetic field in idealized toy models and Mn<sub>3</sub>Sn, and in the extreme case of a magnetic-field-dependent metal-insulator-transition (MIT) in EuTe<sub>2</sub>. The spontaneous effect is investigated by an *ab initio* calculation. We investigate the idealized toy models (e.g., kagome and triangular lattices) and Mn<sub>3</sub>Sn using a tight-binding model and Boltzmann transport calculations. We assume isotropic scattering by applying the relaxation time approximation (RTA). We apply Neumann's principle to determine which symmetries must be broken to realize AMR in the absence of SOC.
- **Extrinsic AMR:** We study AMR resulting from scattering on (aligned) magnetic impurities. We assume that typically unordered impurities are only weakly coupled to the lattice and can be aligned by a weak external magnetic field, which does not overcome the exchange interactions of the host material. **Explain-our-scenario...** Our motivation here is the off-stoichiometry of Mn<sub>3+x</sub>X non-collinear AFMs (typically,  $x \approx 0.2$ , see early work of Kren et al. '75 or other references in Sec. III.C). For  $x = 0$  there would be no such impurities but when  $x > 0$  they must be incorporated into the crystal somehow — and likely not couple to the "underlying matrix" (or at least, couple differently).

This paper is organized as follows: In Sec. II, we will introduce the theoretical background and the methodology used in this work. In Sec. III, we will discuss the intrinsic AMR, followed by the discussion of extrinsic AMR in Sec. IV. The summary and conclusion follow in the final Sec. V. Details of *ab initio* calculations are given in Appendix A, while more details about the tight-binding model are given in Appendix B.

## II. METHODS

### A. Formalism

We employ a simplistic tight-binding model which only consists of a hopping and an exchange term<sup>11</sup>:

$$H = - \sum_{i,j} \sum_{\alpha} t_{ij} \hat{c}_i^{\alpha\dagger} \hat{c}_j^{\alpha} + J \sum_i \sum_{\alpha,\beta} (\vec{\sigma} \cdot \hat{m}_i)_{\alpha\beta} \hat{c}_i^{\alpha\dagger} \hat{c}_i^{\beta} \quad (2)$$

where  $t_{ij}$  is the hopping parameter from site  $i$  to  $j$ ,  $\alpha$  and  $\beta$  are the spin indices,  $\hat{c}_i^{\alpha}(\dagger)$  is an annihilation (creation) operator at site  $i$  with spin  $\alpha$ ,  $J$  is the Heisenberg exchange constant,  $\vec{\sigma}$  the vector of the Pauli spin matrices and  $\hat{m}_i$  the magnetization direction unit vector at site  $i$ .

The conductivity is then calculated using the Boltzmann equation<sup>31</sup>.

$$\sigma_{ij} = e^2 \sum_n \int_{1stBZ} \frac{d^3k}{(2\pi)^3} \delta(E_n(\vec{k}) - E_F) \frac{1}{\hbar \Gamma_{n,\vec{k}}} \times v_{n,i}(\vec{k}) v_{n,j}(\vec{k}) \quad (3)$$

where  $e$  is the elementary charge,  $E_n(\vec{k})$  is the  $k$ -dependent Eigen energy of the  $n$ th-band,  $E_F$  is the Fermi energy,  $\Gamma_{n,\vec{k}}$  is the scattering rate and  $v_{n,i}$  is the  $i$ -th component of the Fermi velocity in the  $n$ -th band. The delta distribution evaluates the integral over the first Brillouin zone (1st BZ) at the Fermi surface. The Fermi velocity is calculated by:

$$v_{n,i} = \frac{1}{\hbar} \frac{\partial E_n(\vec{k})}{\partial k_i} \quad (4)$$

The scattering rate is obtained by using Fermi's Golden Rule:

$$\Gamma_{n,\vec{k}} = \frac{2\pi}{\hbar} N_{scat} \sum_{n'} \int_{1stBZ} \frac{d^3k'}{(2\pi)^3} \delta(E_{n'}(\vec{k}') - E_n(\vec{k})) \times |M_{nn'}^{\vec{k}\vec{k}'}|^2 (1 - \cos \theta_{vv'}) \quad (5)$$

where  $N_{scat}$  is the volume density of the scatterers,  $M_{nn'}^{\vec{k}\vec{k}'}$  is the transition matrix element and  $\cos \theta_{vv'} = \frac{\vec{v}_n(\vec{k}) \cdot \vec{v}_{n'}(\vec{k}')}{|\vec{v}_n(\vec{k})| |\vec{v}_{n'}(\vec{k}')|}$ . The transition matrix element is calculated by:

$$M_{nn'}^{\vec{k}\vec{k}'} = \langle \psi_{n,\vec{k}} | \hat{M} | \psi_{n',\vec{k}'} \rangle \quad (6)$$

where  $\psi_{n,\vec{k}}$  is the wave function for the Eigenenergy value  $E_n(\vec{k})$ <sup>31</sup>.

### B. Symmetry Analysis

We will analyze the real-space symmetries of various magnetic configurations—both for the toy models and

Mn<sub>3</sub>Sn—using the open-source code *Symmetr*<sup>32</sup>. The software identifies the symmetry group of each magnetic configuration and returns the corresponding generator matrices. To understand the origin of anisotropic magnetoresistance (AMR), we compare the symmetries of configurations that exhibit isotropic conductivity with those that show AMR. The key idea is to identify which symmetries must be broken to enable AMR.

At the core of this analysis lies Neumann's principle, which states that a tensor representing a macroscopic physical property of a crystal must be invariant under the symmetry operations of that crystal<sup>21</sup>.

## III. INTRINSIC AMR

In this section, we discuss intrinsic AMR. Subsection III A focuses on MnN as an example of spontaneous AMR in a collinear antiferromagnet. In subsection III B, we investigate various magnetic configurations—both collinear and non-collinear, compensated and non-compensated—on kagome and triangular toy lattices. Using a tight-binding model and the Boltzmann formalism in the relaxation-time approximation (RTA), we analyze Fermi surface anisotropy and conductivity. We also examine the real-space symmetries underlying these models. In subsection III C, we apply the same framework to the real material Mn<sub>3</sub>Sn. Finally, subsection D 1 explores how large intrinsic AMR can arise near a magnetic-field-induced metal-insulator transition in EuTe<sub>2</sub>.

### A. MnN: Example of AMR as a Spontaneous Effect

Since magnetic order is a spontaneous effect, realizing AMR due to the measurement of conductivity along different directions can also make it a spontaneous effect. This is illustrated in Fig. 1: In a cubic crystal without magnetic order, the conductivity remains isotropic  $\sigma_{xx} = \sigma_{yy} = \sigma_{zz}$ . A FM magnetic order (in a cubic crystal still exhibits isotropic conductivity in absence of SOC, since global spin rotations do not alter the symmetry of the conductivity tensor. If more MSLs are considered, the conductivity can become anisotropic if the magnetic order breaks the 90° real-space rotational symmetry (see (Fig. 1c). This is the case of cubic MnN, which has a rock-salt structure (see 1e): Without magnetic order, the conductivity would remain isotropic. However, the cation (manganese) magnetic moments, prefer an A-type AFM order and in choosing the direction of the ferromagnetic planes (e.g.  $xy$ -planes), anisotropy arises (in that case,  $\sigma_{zz}$  is different from  $\sigma_{xx} = \sigma_{yy}$ )<sup>34</sup>. This effect in itself is non-relativistic in nature.

Calculations based on density functional theory (DFT as detailed in Appendix A,  $a = b = c$ ), show that  $\hbar\omega_{xx}^p =$

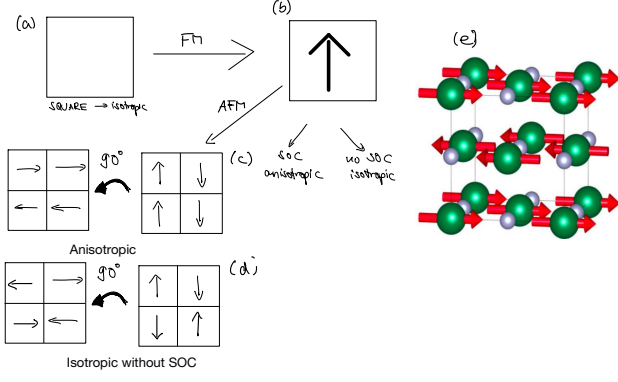


FIG. 1. (a) A square lattice without magnetic order exhibits isotropic conductivity  $\sigma_{xx} = \sigma_{yy}$ . (b) A FM magnetic order allows for AMR  $\sigma_{xx} \neq \sigma_{yy}$  only in presence of SOC. (c) Including multiple MSL allows for AMR in a collinear AFM even in absence of SOC if the  $90^\circ$  real-space rotation symmetry is broken by the magnetic order. (d) If the  $90^\circ$  real-space rotation symmetry is conserved by the magnetic order, the conductivity is isotropic in absence of SOC. (e) Crystalline structure with magnetic order of MnN, reproduced from Ref.<sup>33</sup>. The magnetic order breaks the  $90^\circ$  real-space rotation symmetry and allows for non-relativistic AMR.

5.87 eV and  $\hbar\omega_{zz}^p = 5.23$  eV so that, assuming isotropic scattering,  $\sigma_{xx}/\sigma_{zz} - 1 \approx 26\%$  according to Eq. 8.

Magnetic order leads to a distortion of lattice but this effect has only minor influence on such transport anisotropy — for example,  $a/c = 0.4256/0.4189$  nm changes  $\hbar\omega_{xx}^p$  to 5.88 eV and  $\hbar\omega_{zz}^p$  to 5.10 eV.

### B. Intrinsic AMR due to Manipulation of the Magnetic Order

We are choosing the RTA in the Boltzmann formalism<sup>35</sup> and replace Eq. 5 by:

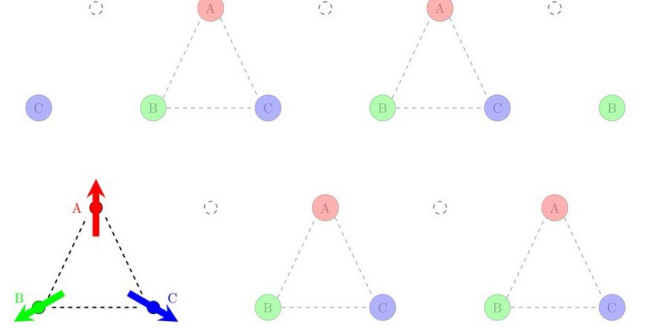
$$\tau = \frac{1}{\hbar\Gamma_{n,\vec{k}}} \quad (7)$$

where the relaxation time  $\tau$  is constant and thus isotropic. This means that the anisotropy can only enter through the Fermi velocity contribution (or anisotropic plasma frequencies  $\omega^p$ ). We can simplify Eq. 3 to:

$$\sigma_{ii} \propto \int_{FS} \sum_n dk_F v_{n,i}^2(k_F) \propto (\omega_{ii}^p)^2 \quad (8)$$

where  $\vec{k}_F$  is the wave vector at the Fermi surface (Fermi vector). We made use of the facts that we only consider the longitudinal conductivity  $\sigma_{ii}$  in two dimensional systems, the delta distribution evaluates the integral over the first Brillouin zone at the Fermi surface (FS), and the scattering rate is obtained by RTA.  $\sigma_{xx} \neq \sigma_{yy}$  if the integral of  $\sum_n v_{n,x}^2$  and  $\sum_n v_{n,y}^2$  over the Fermi

(a) kagome lattice



(b) triangular lattice

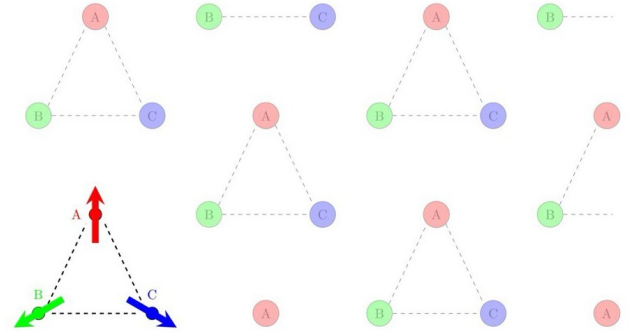


FIG. 2. Schematic image of (a) the kagome lattice and (b) the triangular lattice with non-collinear magnetic order. The three magnetic moments in the magnetic unit cell are shown in red (A), green (B) and blue (C). The net magnetic moment is zero. While the magnetic unit cell of both configurations is the same, the regular vacancy of the kagome lattice (dashed circle) distinguishes the two. In practice, the vacancy can be filled with a non-magnetic atom.

circle are not the same. This is generally achieved by anisotropic Fermi surfaces, whereas the FS must neither be spherical (which is perfect isotropic) nor show the symmetry of the system, since the conductivity must reflect the crystal symmetry due to Neumann's principle<sup>21</sup> (e.g. a hexagonal FS in a hexagonal material is insufficient).

We start by analyzing various magnetic configurations on a kagome and triangular lattice. First, we assume a non-collinear (no SSA), coplanar ( $m_{z,i} = 0, \forall i$ ) and fully compensated ( $\sum_i \vec{m}_i = 0$ ) magnetic order as indicated in Fig. 2. This can be realized by considering a simple next-neighbor (NN) Heisenberg exchange term  $\sum_{\langle ij \rangle} -J_{ij} \vec{S}_i \cdot \vec{S}_j$  with  $J_{ij} = -1$ . On a kagome and a triangular lattice, the non-collinear magnetic order will form to avoid frustration. In a first step, we test configurations that are non-collinear and fully compensated, which leaves us with two possibilities: First, a global spin rotation and second, a permutation of moments as indicated in Figs. 3 (c) and (e). The global spin rotation

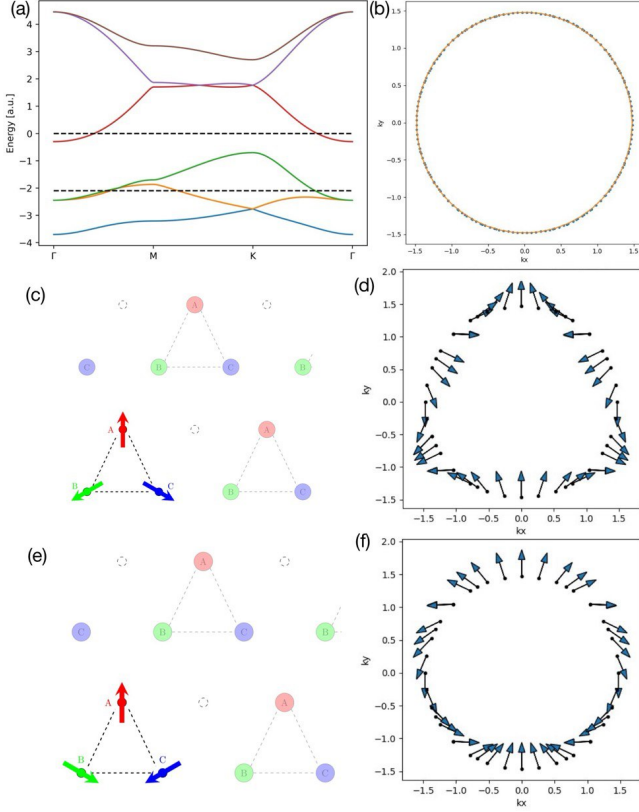


FIG. 3. Results for two compensated magnetic configuration on the kagome lattice. (a) The band structure is spin-split. Permutation of magnetic moments does not change the band structure. (b) The Fermi surface at  $E_F = 0$  (as indicated by the dashed line in (a)) is represented by the blue dots. The orange circle is shown for illustration. Since the FS and the circle overlap, this system has an isotropic FS and does not show intrinsic AMR. (c), (e) The magnetic unit cells of two compensated magnetic configurations and (d), (f) their respective  $k$ -space spin texture at the same Fermi level. While the band structure and Fermi surface is unaffected by permutation within the magnetic unit cell, the spin texture changes. The spin texture is entirely in the  $xy$ -plane. The  $\hat{z}$ -component (not shown) of the spin texture is zero.

does not show any effect, because the lack of SOC. The permutation, while changing the spin texture (see 3 (d) and (f)), keeps the band structure (BS) and Fermi surface (FS) unchanged. Intrinsic AMR cannot be found. **Even the 'permutation' is in fact a global spin rotation (around  $y$ , right?)**

In the next step, we are individually rotating the magnetic moments within the  $xy$ -plane as shown in Fig. 4: While moment  $A$  remains fix, moment  $B$  is rotated counterclockwise and moment  $C$  is rotated clockwise, illustrating the effect of a magnetic field in the negative  $-\hat{y}$ -direction. Moments  $B$  and  $C$  are rotated by the same angle  $\alpha$ . Using this notion, we would like to draw the attention to a few special cases:  $\alpha = 0^\circ(120^\circ)$  corresponds to the compensated states shown in Fig. 3 (c) (Fig. 3 (e)),

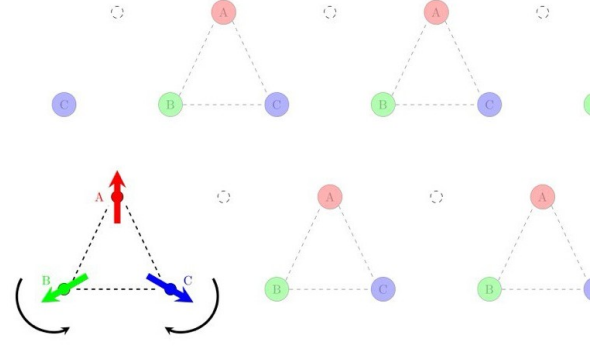


FIG. 4. Illustration of the rotation of the moments  $B$  (counterclockwise) and  $C$  (clockwise) by the same angle  $\alpha$

$\alpha = 240^\circ$  corresponds to the ferromagnetic states with magnetization along  $+\hat{y}$  and  $\alpha = 60^\circ$  corresponds to a collinear ferrimagnetic state. For all  $\alpha \neq 0^\circ, 120^\circ, 240^\circ$  a partially compensated in-plane magnetization, or weak ferromagnetism (WF), exists.

We found an anisotropic FS (see Fig. 5 for the example of  $\alpha = 36^\circ$ ) and thus, intrinsic AMR for all WF cases. The isotropic cases are the fully compensated and the FM case. Applying the *Symmetr* code<sup>32</sup> to get the generators of the real-space symmetries, we find that the WF cases break both a  $60^\circ$  and  $120^\circ$  rotational symmetries in the  $xy$ -plane, while the fully compensated and the FM case, conserve at least one of them. Different to our previous anticipation<sup>2</sup>, the anisotropy in the collinear ferrimagnetic case ( $\alpha = 60^\circ$ ) indicates that the non-collinear magnetic order is not needed to generate the effect, if the real-space symmetry is low enough. The magnetization in the WF cases is likely not sufficient to be a SSA, as mentioned before, but might play a role as a "symmetry breaking parameter". Similar results were observed in MnTe, where a besides the collinear antiferromagnetic ordering, a *weak ferromagneticlike signal* (WFL) was observed. The AHE in MnTe was not attributed due to the WFL, however the AHE is due to an imbalance of magnetic domains with Hall contributions that cancel out. This imbalance causes the WFL and thus the WFL serves as measure of this imbalance<sup>36</sup>, similar to our magnetization serving as measure of the symmetry breaking.

**Discuss Fig. 6 in an appropriate context.** Proposing an experiment 'in real material' may be tricky because details of the magnetic anisotropy can play important role.

Now we are turning towards the triangular lattice (see Fig. 2 (b)). Here, the same operations were applied as for the kagome lattice. In either the fully compensated states, the FM state and the non-compensated WF cases, no anisotropic FS and thus, no intrinsic AMR was found. Analyzing the real-space symmetries using *Symmetr*<sup>32</sup>, we find that all investigated cases are conserving the  $60^\circ$  or  $120^\circ$  rotational symmetry.

In summary, we have investigated various magnetic



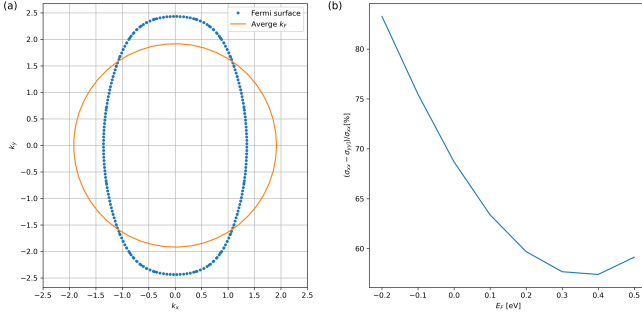


FIG. 5. Anisotropy induced by magnetic order. The magnetic configuration corresponds to a rotation of the green and blue moments by  $\alpha = 36^\circ$  as shown in Fig. 4 (or 6 (a)). The small magnetization that ensues can serve as a measure of symmetry breaking. (a) Fermi surface (blue) at  $E_F = 0.1$  showing a pronounced anisotropy between  $\hat{x}$  and  $\hat{y}$ -directions. The orange circle is shown as reference. (b) AMR vs. Fermi energy in the same system. Despite being a qualitative model, limited quantitative statements are possible. Energy scale directly comparable to Fig. 3 (a)

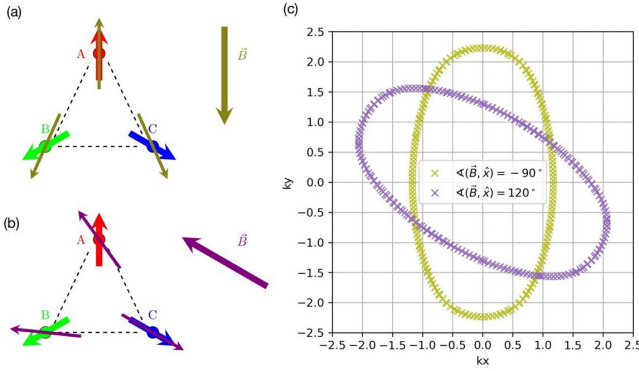


FIG. 6. Rotation of the warped Fermi surface: (a) The magnetic field (ochre) was applied in  $-\hat{y}$ -direction (or  $\langle \vec{B}, \hat{x} \rangle = -90^\circ$ ) rotating the moments B and C towards the magnetic field, while leaving A in its original positions. The rotated moments are shown in ochre. (b) Simultaneously, the magnetic field (violet) was applied along  $\langle \vec{B}, \hat{x} \rangle = 120^\circ$ . The rotated moments are shown in violet. (c) The resulting Fermi surfaces (ochre and violet, respectively) are hence rotated by  $60^\circ$ , and are not overlapping, resulting in a differing AMR.

configurations on a kagome and a triangular lattice. The configurations that break the  $60^\circ$  and  $120^\circ$  real-space rotational symmetry in the  $xy$ -plane exhibit an anisotropic FS and thus, AMR. These are the non-compensated configurations in the kagome lattice (both non-collinear and collinear ferrimagnetic), yet none of the investigated triangular configuration. This underlines that, while spin and lattice are largely decoupled<sup>11</sup> in absence of SOC, the magnetic order in general does can change the symmetry of the conductivity tensor.

### C. A Material Example: $\text{Mn}_3\text{Sn}$

$\text{Mn}_3\text{Sn}$  is a well-studied non-collinear antiferromagnet where stoichiometry is very important: see Kasia Gas et al. '25 [Elaborate on this...](#) It features a double-layer kagome lattice and a net moment of (close to<sup>37</sup>) zero. It rose to attention due to its considerable potential for future spintronics applications: Strong, and intrinsic AHE and SHE were found<sup>18,27,28,38,39</sup> as well as a strong anomalous Nernst effect<sup>27,28,38</sup> spin-polarized charge currents<sup>40</sup>, and ultrafast spin dynamics<sup>38,39</sup>.

We employ a simultaneous analysis utilizing a tight-binding model and the Boltzmann formalism in RTA as in the previous section. The structure for  $\text{Mn}_3\text{Sn}$  is taken from the MAGNDATA database<sup>41</sup>. We will start by considering the fully compensated case and proceed by tilting the magnetic moments in a fashion that simulates an external magnetic field in the  $x$ -direction (see Fig. 11 (a)), similar to Fig. 4. As a result, similar to the previous case of the fully compensated kagome lattice, the fully compensated  $\text{Mn}_3\text{Sn}$  shows an isotropic FS and does not exhibit AMR. The tilted case shows a small ellipsoid warping of the otherwise isotropic FS (as can be seen by the two-fold symmetry in Fig. 11 (c)) and exhibits thus a small AMR. Employing the symmetry analysis using the *Symmetr* code<sup>32</sup>, we find that the fully compensated  $\text{Mn}_3\text{Sn}$  configuration exhibits  $60^\circ$  and  $120^\circ$  real-space rotation symmetry, while the tilted configuration breaks both. This further confirms our previously made assumptions.

While we have treated  $\text{Mn}_3\text{Sn}$  as a fully compensated coplanar AFM, there are studies about a weak ferromagnetic moment occur in the material<sup>37</sup>. This also opens a debate about how the magnetic field couples to the magnetic moments: While we suggest a tilting of the magnetic moments towards the magnetic field, it is also possible that the WF moment couples to the field and rotates the entire structure simultaneously, as suggested by some studies<sup>42</sup>. However, as emphasized before, a global spin rotation does not exhibit AMR due to the absence of SOC.

In summary, we have analyzed the promising non-collinear AFM  $\text{Mn}_3\text{Sn}$  and showed that if the  $60^\circ$  and  $120^\circ$  real-space rotation symmetries are broken, an anisotropic FS and AMR is induced. This is consistent with our previous findings on a kagome toy model.

## IV. EXTRINSIC AMR

In this section, we investigate the extrinsic (scattering-dependent) AMR. We will revert to the full treatment of scattering via Eq. 5, assuming magnetic impurities pointing in  $\hat{z}$ -direction described by the transition matrix matrix  $\hat{M} = \hat{S}_i \otimes \hat{1}_{N \times N}$ , where  $\hat{S}_i$  is the  $i$ -th Pauli spin matrix and  $\hat{1}_{N \times N}$  is the  $N$ -dimensional identity matrix, where  $N$  is the number of atoms in the unit cell. For

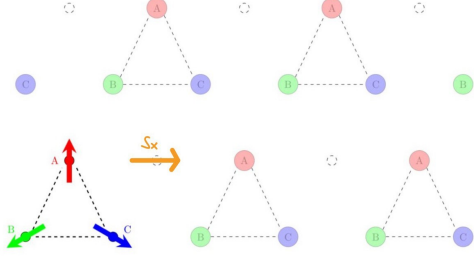


FIG. 7. Illustration of a  $x$ -impurity in a kagome lattice

Lattice	Impurity $\hat{M}$	AMR?	Spin texture
<b>Kagome (comp.)</b>	$x$	<b>Yes</b>	$xy$
<b>Kagome (comp.)</b>	$y$	<b>Yes</b>	$xy$
<b>Kagome (non-comp.)</b>	$x$	<b>Yes</b>	$xy$
<b>Kagome (non-comp.)</b>	$y$	<b>Yes</b>	$xy$
Triangular (comp.)	$y$	divergent	$z$
Triangular (non-comp.)	$y$	No	$zy$

TABLE I. Results of the extrinsic AMR calculations for kagome and triangular lattice in both compensated and non-compensated magnetic configurations. The impurities as introduced in the text. "Yes" means that AMR was found, hence  $\sigma_{xx} \neq \sigma_{yy}$  while for "No" an isotropic behavior was identified, where  $\sigma_{xx} = \sigma_{yy}$ . The value of the AMR ratio is not stated here, as our model is qualitative. "Divergent" means that both  $\sigma_{xx}$  and  $\sigma_{yy}$  are divergent or infinite and AMR cannot be defined, originating from a suppressed scattering rate  $\Gamma \rightarrow 0$ . "Spin texture" indicates the active components of the  $k$ -space spin texture at the FS.  $xy$  indicates that all the spins are within the  $xy$ -plane and thus  $\hat{s}_z = 0$  for all spins. The spin texture serves as a good intuition for whether a certain impurity would lead to suppressed scattering.

simplicity reasons, we will restrict ourselves to impurities either pointing in  $x$ -direction, as illustrated in Fig. 7, which shall be abbreviated as  $x$ -impurities, and impurities in  $y$ -direction, denoted as  $y$ -impurities. It shall be noted that these impurities, which are ferromagnetically aligned in a single direction, introduce a symmetry breaking to the system. This might seem counterintuitive at first, since impurities in a real material would likely be nearly statistically equal distributed among all directions of space. We justify this ferromagnetic alignment by assuming that the magnetic impurities are only weakly coupled to the crystal and can be aligned by an external magnetic field, which is not sufficient to overcome the exchange interaction and thus, to manipulate the other magnetic moments.

As previously in Sec. IIIB, we investigate both a kagome and a triangular lattice. We then calculate the conductivities  $\sigma_{xx}$  and  $\sigma_{yy}$  using the Eqs. 3-6. The results are summarized in table I.

In the kagome case, for both the  $x$ - and the  $y$ -impurity in both the compensated and non-compensated (WF

FIG. 8. New fig. 8 (originally, it was number 9) should contain scatt. rates on FS. Philipp will send it shortly... :-)

non-collinear) cases, non-zero extrinsic AMR was found. In the triangular case, no extrinsic AMR has been identified at all. In some cases the results in Tab. I are denoted by *divergent*, which means that both  $\sigma_{xx}$  and  $\sigma_{yy}$  are divergent or infinite. This originates from a suppressed scattering due to that impurity  $\Gamma \rightarrow 0$  and the fact that the inverse scattering rate is part of the Boltzmann equation Eq. 3. Since our model is very simplistic, it is, however, not expected that in a real system the conductivity would diverge.

The momentum-space spin textures of the kagome lattice are shown in Fig. 3. Spin textures can be an important tool as already shown in the context of the non-relativistic Edelstein effect<sup>11</sup>. Well, yes... but PRB 80, 134405<sup>47</sup> is much more relevant here. In our case, it can give us a hint about the possibility of AMR due to its simiarity (explain please if you really see it :) The  $i$ -th component of the spin texture is given by:

$$S_i(\vec{k}) = \langle \psi_k | \hat{S}_i | \psi_k \rangle \quad (9)$$

where  $\psi_k$  is the wave function at  $\vec{k}$ . The scattering rate  $\Gamma$  at the same  $\vec{k}$  for a magnetic impurity in direction  $i$  described by  $\hat{S}_i$ :

$$\Gamma_{\vec{k}} \propto \int_{FS} dk' |\langle \psi_k | \hat{S}_i | \psi_k \rangle|^2 \quad (10)$$

where in Eq. 10 we ignored  $\cos \theta_{vv'}$  and the prefactors, and the delta distribution was evaluated. The integration over  $k_z$  does only contribute in a prefactor the system we are looking at is two dimensional. The resulting integral in Eq. 10 is a one dimensional integral over the Fermi circle. As can be seen the scattering rate for a magnetic impurity in direction  $i$  (Eq. 10) and the  $i$ -th component of the spin texture are looking very similar, except that the integral over the Fermi surface adds some level of complexity to it. Despite these small differences, it allows us to explain a part of the results we find in Tab. I. The spin texture of the triangular compensated pointing in the positive  $z$ -direction leading to a suppressed scattering for both  $x$ - and  $y$ -impurities. In the triangular non-compensated (WF) case, the spin texture acquires some  $y$ -component as well. Introducing a  $y$ -impurity will now not lead to a suppressed scattering anymore, yet still not exhibit extrinsic AMR.

In the last part of this section, we investigate at the crystalline AMR, which refers to the anisotropy of conductivity created by the crystalline symmetry. Here, it can be defined as:

$$AMR_{vw}^{cry} = \frac{\sigma_{ww}(\hat{M} = \hat{S}_w)}{\sigma_{vv}(\hat{M} = \hat{S}_v)} \quad (11)$$

where  $v$  and  $w$  are two arbitrary directions (e.g.  $x$  and  $y$ ). The crystalline AMR is thus defined as the quotient as the longitudinal conductivity in  $v$ -direction for a magnetic impurity in the same direction with the longitudinal conductivity in  $w$ -direction for a magnetic impurity in the same direction. In both numerator and denominator are thus longitudinal conductivities with parallel aligned impurity. The resulting anisotropy  $AMR_{vw}^{cry} \neq 1$  would thus arise only from the influence of the crystal directions. Calculating the respective quantities for the kagome lattice, we indeed find non-zero crystalline AMR.

## V. SUMMARY AND CONCLUSIONS

In this paper, we have investigated several pathways to realize an anisotropy of conductivity due to magnetic order in the absence of SOC, thus non-relativistic AMR: The magnetic order must break the real-space rotational symmetry of the crystal. In the case of cubic MnN, the A-type AFM order breaks the  $90^\circ$  rotational symmetry, exhibiting AMR as a spontaneous effect. In a kagome lattice, the  $60^\circ$  and  $120^\circ$  rotational symmetries need to be broken, which can be achieved by manipulating the magnetic moments due to an external field. We have suggested an experimental setup to confirm these findings. We have applied those findings to the well-studied non-collinear AFM and spintronic application candidate  $Mn_3Sn$ , confirming that the  $60^\circ$  and  $120^\circ$  rotational symmetries need to be broken to exhibit AMR. We have discussed that a field-dependent metal-insulator-transition is an extreme version of that effect, as shown in antiferromagnetic  $EuTe_2$ . In this paper's second part, we investigated extrinsic AMR depending on ordered magnetic impurities. We found that a suitable impurity can induce AMR in the previously isotropic fully compensated magnetic configuration on a kagome lattice.

Our study calls for further investigations: The model considerations can be extended to include more MSLs as the lattice per se does not predetermine the number of MSLs<sup>44,45</sup>.

A database of materials exhibiting non-relativistic AMR determined by the principles laid out in this paper is desirable. Good candidates are materials with a more complex magnetic order, such as  $RbFe(MoO_4)_2$  or, e.g., skyrmion host materials. Furthermore, a word of caution: Our model is highly simplistic as it only contains a hopping term and a Heisenberg exchange term. More realistic and complex contributions, such as different atomic orbitals, the crystal field, or the contributions of phonons and magnons, are missing to date, as well as any many-body interactions and correlations. In the exchange, more complicated (often mid-range) pairings of moments are ignored, as well as higher-order contributions such as the Dzyaloshinsky-Moriya interaction, which can play a role in such non-collinear systems. It has to be reminded that the spin moments are not just dipole arrows pointing in a direction, but they are spin

densities, which also contain higher multipoles.

## ACKNOWLEDGMENTS

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## Appendix A: Ab initio calculations

$\hbar\omega_{xx}^p = 5.87$  eV and  $\hbar\omega_{zz}^p = 5.23$  for MnN were obtained using GGA with SOC under assumption of 'perfectly cubic lattice constants'. When SOC is switched off, the anisotropy of the conductivity remains similar, indicating that the AMR is mainly of non-relativistic nature.

## Appendix B: Commentary about the Tight-Binding Model

From Fig. 5 (b), we can see that within our model, limited quantitative statements are possible, at least as long we are keeping within the same lattice and magnetic configurations and keep to a sufficiently small regime. Comparing AMR values quantitatively at vastly different Fermi levels (and thus bands), or the same Fermi levels for different magnetic configuration, or even between different lattices, leads to jumps in the numerical value and does thus not provide ground for a useful analysis. On the contrary, a qualitative statement is always possible: Given a certain lattice and magnetic configuration, in an isotropic case  $\sigma_{xx}/\sigma_{yy} = 1$  for all Fermi energies in all bands, and in an anisotropic case,  $\sigma_{xx}/\sigma_{yy} \neq 1$  for all Fermi energies in all bands (although the precise value of  $\sigma_{xx}/\sigma_{yy}$  in the anisotropic case is very sensitive to small parameter changes). With changing Fermi level  $E_F$ , the form of the FS might change as illustrated in Fig. 9.

While in both toy models (see Sec. III B), we only took next neighbor hopping into account, in the material systems ( $Mn_3Sn$ , see Sec. III C), this did not lead to any realistic results. For  $Mn_3Sn$  we took 20 next neighbors into account. We chose this number by calculating the Fermi surface and AMR for the material systems using fully ferromagnetic moments as input. Since there is no SOC, the conductivity needs to be isotropic, and any anisotropy can be considered an artifact. We started by next neighbor hopping and increased the number of neighbors until isotropy in the FM state was reached.

## Appendix C: An Expanded Phenomenological Model

The angle-dependent form of AMR can be expressed phenomenologically in terms of power expansion of the



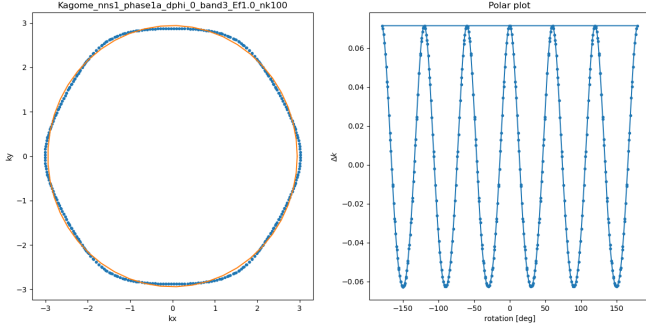


FIG. 9. FS for the fully compensated configuration of the kagome lattice (see Fig. 3 (b)) for a different Fermi level ( $E_F = 1$ ). The FS becomes more hexagonal, which is still isotropic with regards to the conductivity tensor.

magnetization direction<sup>20,23,46</sup> and allows to describe even more complex crystalline AMR signals<sup>5,21,24</sup>. However, these models rely on the existence of a SSA as the magnetization or Néel vector. In non-collinear systems such SSA does not exist - even in case of weak ferromagnetism induced by an applied magnetic field, it would be likely an oversimplification to ignore the effects of the sublattices. "Local" treatment of effects is not new: For instance, basic AMR models in FMs rely on separate contributions for spin up and spin down electrons (two-current models)<sup>2</sup>, or the Edelstein effect in non-collinear  $\text{Mn}_3\text{Sn}$  can be calculated for each sublattice<sup>11</sup>. Such local approach can be applied to AMR by considering the contributions of each magnetic sublattice (MSL) individually in the phenomenological model, which yields:

$$\rho_{yy} = \rho_0 + \sum_{m=1,2,3} \sum_{n=2,4,6,\dots} c_{m,n} \cos(n\alpha_m) \quad (\text{C1})$$

where  $m$  is the index of the MSL,  $n$  is the order of the spherical harmonic,  $c_{m,n}$  is the index of the  $n$ -th harmonic of the  $m$ -th MSL, and  $\alpha_m$  is the angle of the magnetization direction of the  $m$ -th magnetic moment (assuming an in-plane rotation). The coefficients  $c_{m,n}$  can be obtained by fitting where for multiple MSLs measurements for different values of magnetic field  $B$  are necessary. The position of the magnetic moments  $\alpha_m$  can be obtained from SW models. While this is not going to be a main focus of this work, Eq. C1 together with an adequate SW model could help to distinguish the MCA from the AMR.

## Appendix D: Dump

Parts of text and figures moved to here... There was some inconsistency about Fig. 10: is panel (b) from a triangular or square FM lattice? Fig. 11 supports the conclusions of Sec. III.C but it would have to be much improved.

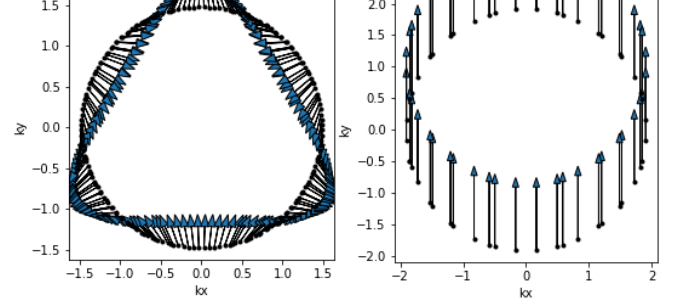


FIG. 10. The  $k$ -space spin textures of both kagome and triangular lattice for the fully compensated magnetic configurations. (a) In the kagome case, the  $z$ -component of the spin (not shown) is zero. (b) The spin texture of the triangular case is shown in the  $(k_x, k_z)$ -plan. All spins point in the  $z$ -direction. [Apply changes to the figure]

### 1. AMR due to Metal-Insulator Transition: $\text{EuTe}_2$

The antiferromagnetic  $\text{EuTe}_2$  undergoes a MIT, where the critical field and temperature are different in the  $ab$ -plane than along the  $c$ -axis. Measuring  $\sigma_{xx}$  while rotating the magnetic field can then cause AMR of up to 40000% due to the insulating MIT<sup>43</sup>. This marks an extreme case of the before discussed case of intrinsic AMR: While in sections IIIB and IIIC, the warping of the FS was responsible for the AMR, here the FS vanishes in the insulating phase.

We are now proposing an experiment to show this effect: Applying the magnetic field along a symmetry axis will tilt the magnetic moments, generate a weak ferromagnetic moment (magnetization) and result in AMR due to an anisotropic Fermi surface (see Fig. 6 (a) and (c), ochre). Now, we are rotating the magnetic field by  $120^\circ$ , so that it points along another symmetry axis. The magnetic moments will tilt in another direction (see Fig. 6 (b)), and result in a rotated FS, which does not overlap with the previous FS (see Fig. 6 (c)), thus leading to another value of AMR. We can now rotate the conductivity measurement from  $\sigma_{xx}$  and  $\sigma_{yy}$  to  $\sigma_{vv}$  and  $\sigma_{ww}$ , where here  $\hat{v} = \hat{x} + 120^\circ$  ( $\hat{w} = \hat{y} + 120^\circ$ ) and recover the original value of the AMR.

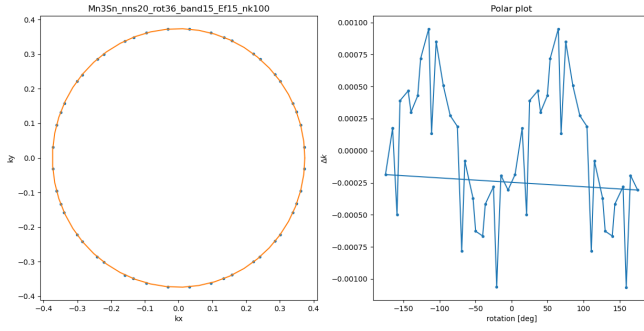


FIG. 11. Non-compensated (WF) case of  $\text{Mn}_3\text{Sn}$ . (a) [to be included] Schematics of the tilted magnetic moments. (b) Simplified FS (blue dots) and, for illustrative reasons, a circle with perfect spherical symmetry (orange) at  $E_F = 15$ . The shown FS is an intersection through the  $k_z = 0$ -plane. (c) Deviation of the FS (blue dots in (b)) from the spherical symmetry (orange in (b)): A weak two-fold symmetry is recognizable, which is sufficient to render the FS sufficiently anisotropic and exhibit AMR

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- <sup>1</sup> W. Thomson, Proc. R. Soc. Lond. 8, 546.(1857)
  - <sup>2</sup> Philipp Ritzinger, Karel Výborný, R. Soc. Open Sci., 10, 230564 (2023).
  - <sup>3</sup> H. S. Alagoz, J. Desomberg, M. Taheri, F. S. Razavi, K. H. Chow, and J. Jung, Appl. Phys. Lett. 106, 082407 (2015)
  - <sup>4</sup> D. Kriegner, H. Reichlova, J. Grenzer, W. Schmidt, E. Ressouche, J. Godinho, T. Wagner, S. Y. Martin, A. B. Shick, V. V. Volobuev, G. Springholz, V. Holý, J. Wunderlich, T. Jungwirth, and K. Výborný, Phys. Rev. B 96, 214418 (2017).
  - <sup>5</sup> R. D. Gonzalez Betancourt, J. Zubáč, K. Geishendorf, P. Ritzinger, B. Růžicková, T. Kotte, J. Železný, K. Olejník, G. Springholz, B. Büchner, A. Thomas, K. Výborný, T. Jungwirth, H. Reichlová, and D. Kriegner, npj Spintronics, 2, 45 (2024).
  - <sup>6</sup> J. Volný, D. Wagenknecht, J. Železný, P. Hrcuba, E. Duverger-Nedellec, R. H. Colman, J. Kudrnovský, I. Turek, K. Uhlířová, K. Výborný, Phys. Rev. Mat. 4, 064403 (2020).
  - <sup>7</sup> J. Zubáč, Z. Kašpar, F. Krizek, Förster, R. P. Campion, V. Novák, T. Jungwirth, K. Olejník, Phys. Rev. B 104, 184424 (2021).
  - <sup>8</sup> P. Wadley, B. Howells, J. Železný, C. Andrews, V. Hills, R. P. Campion, V. Novák, K. Olejník, F. Maccheronzi, S. S. Dhesi, S. Y. Martin, T. Wagner, J. Wunderlich, F. Freimuth, Y. Mokrousov, J. Kuneš, J. S. Chauhan, M. J. Grzybowski, A. W. Rushforth, K. W. Edmonds, B. L. Gallagher, T. Jungwirth, Science 351, 6273 (2016).
  - <sup>9</sup> K. Kabara, M. Tsunoda, S. Kokado, AIP Adv. 7, 056416 (2017).
  - <sup>10</sup> I. Bakonyi, F. D. Czeschka, L. F. Kiss, V. A. Isnaini, A. T. Krupp, K. Palotás, S. Zsurzsa, L. Péter, arXiv:2203.11568 [cond-mat.mtrl-sci] (2022).
  - <sup>11</sup> R. González-Hernández, P. Ritzinger, K. Výborný, J. Železný, and A. Manchon, Nat. Commun., 15, 7663 (2024).
  - <sup>12</sup> L. Nádvorník, M. Borchert, L. Brandt, R. Schlitz, K. A. de Mare, K. Výborný, I. Mertig, G. Jakob, M. Kläui, S. T.B. Goennenwein, M. Wolf, G. Woltersdorf, T. Kampfrath, Phys. Rev. X 11, 021030 (2021).
  - <sup>13</sup> J.-H. Park, H.-W. Ko, J.-M. Kim, J. Park, S.-Y. Park, Y. Jo, B.-G. Park, S. K. Kim, K.-J. Lee, K.-J. Kim, Sci. Rep. 11, 20884 (2021).
  - <sup>14</sup> T. Kato, Y. Ishikawa, H. Itoh, J.-i. Inoue, Phys. Rev. B 77, 233404 (2008).
  - <sup>15</sup> J. Velev, R. F. Sabirianov, S. S. Jaswal, E. Y. Tsymlal, Phys. Rev. Lett. 94, 127203 (2005).
  - <sup>16</sup> F. L. Zeng, Z. Y. Ren, Y. Li, J. Y. Zeng, M. W. Jia, J. Miao, A. Hoffmann, W. Zhang, Y. Z. Wu, Z. Yuan, Phys. Rev. Lett. 125, 097201 (2020).
  - <sup>17</sup> T. Kato, Y. Ishikawa, H. Itoh, J. Inoue, phys. stat. sol. (b) vol. 244, 12, 4403 - 4406 (2007).
  - <sup>18</sup> Y. Zhang, Y. Sun, H. Yang, J. Železný, S. P. P. Parkin, C. Felser, B. Yan, Phys. Rev. B 95, 075128 (2017).
  - <sup>19</sup> N. Nagaosa, J. Sinova, S. Onoda, A. H. MacDonald, N. P. Ong, Rev. Mod. Phys. 82, 1539 (2010).
  - <sup>20</sup> W. Döring, Ann. Phys. 424, 259-276 (1938).
  - <sup>21</sup> P. Ritzinger, H. Reichlová, D. Kriegner, A. Markou, R. Schlitz, M. Lammel, D. Scheffler, G. H. Park, A. Thomas, P. Středa, C. Felser, S. T. B. Goennenwein, and Karel Výborný, Phys. Rev. B, 104, 094406 (2021)
  - <sup>22</sup> T. Sato, S. Kokado, M. Tsujikawa, T. Ogawa, S. Kosaka, M. Shirai, and M. Tsunoda, Appl. Phys. Express 12, 103005 (2019)
  - <sup>23</sup> E. De Ranieri, A. W. Rushforth, K. Výborný, U. Rana, E. Ahmad, R. P. Campion, C. T. Foxon, B. L. Gallagher, A. C. Irvine, J. Wunderlich, New J. Phys. 10, 065003 (2008).

- <sup>24</sup> P. Nam Hai, D. Sasaki, L. Duc Anh, M. Tanaka, Appl. Phys. Lett. 100, 262409 (2012). DOI:10.1063/1.4730955
- <sup>25</sup> M. Q. Dong, Z.-X. Guo, X. R. Wang, Phys. Rev. B 108, L020401 (2023).
- <sup>26</sup> Bonbien *et al.* J. Phys. D: Appl. Phys. 55, 103002 (2022)
- <sup>27</sup> X. F. Zhou, X. Z. Chen, Y. F. You, L. Y. Liao, H. Bai, R. Q. Zhang, Y. J. Zhou, H. Q. Wu, C. Song, F. Pan, *Phys. Rev. Appl.* 14, 054037 (2020).
- <sup>28</sup> K. Manna, Y. Sun, L. Muechler, J. K  bler, C. Felser, "Heusler, Weyl and Berry", *Nat. Rev. Mater.* 3, 244-256 (2018).
- <sup>29</sup> T. Jungwirth, R. M. Fernandes, E. Fradkin, A. H. MacDonald, J. Sinova, L. Šmejkal, arXiv:2411.00717v2 [cond-mat.mtrl-sci] (2025)
- <sup>30</sup> A. Birk Hellenes, T. Jungwirth, R. Jaeschke-Ubiergo, A. Chakraborty, J. Sinova, L. Šmejkal, arXiv:2309.01607v3 [cond-mat.mes-hall]
- <sup>31</sup> K. V  born  , J. Ku  era, J. Sinova, A. W. Rushforth, B. L. Gallagher, T. Jungwirth, Phys. Rev. B 80, 165204 (2009)
- <sup>32</sup> J.   lezn  , Linear response symmetry. Bitbucket (2024)
- <sup>33</sup> Dunz et al. MnN, DOI: 10.1103/PhysRevResearch.2.013347
- <sup>34</sup> S. Granville et al., Phys. Rev. B 72, 205127 (2005).
- <sup>35</sup> K. V  born  , A. A. Kovalev, J. Sinova, T. Jungwirth, Phys. Rev. B 79, 045427 (2009).
- <sup>36</sup> K. P. Kluczyk, K. Gas, M. J. Grzybowski, P. Skupi  ski, M. A. Borysiewicz, T. Fas, J. Suffczy  ski, J. Z. Domagala, K. Graszka, A. Mycielski, M. Baj, K. H. Ahn, K. V  born  , M. Sawicki, M. Gryglas-Borysiewicz, Phys. Rev. B 110, 155201 (2024)
- <sup>37</sup> S. Tomiyoshi, Y. Yamaguchi, J. Phys. Soc. Jap. 51, 2478 (1982)
- <sup>38</sup> T. Chen, T. Tomita, S. Minami, M. Fu, T. Koretsune, M. Kitatani, I. Muhammad, D. Nishio-Hamane, R. Ishii, F. Ishii, R. Arita, S. Nakatsuji, *Nat. Commun.* 12, 572 (2021).
- <sup>39</sup> S. Nakatsuji, N. Kiyohara, T. Higo, *Nature* 527, 212-215 (2015).
- <sup>40</sup> J.   lezn  , Y. Zhang, C. Felser, B. Yan, *Phys. Rev. Lett.* 119, 187204 (2017).
- <sup>41</sup> <https://www.cryst.ehu.es/magndata/index.php?index=0.199>, 2024-08-23, 12:01
- <sup>42</sup> M. Wu, K. Kondou, T. Chen, S. Nakatsuji, Y. Otani, *AIP Adv.* 13, 045102 (2023).
- <sup>43</sup> H. Yang *et al.*, Phys. Rev. B 104, 214419 (2021).
- <sup>44</sup> J. Rusnacko, D. Gotfryd, J. Chaloupka, Phys. Rev. B 99, 064425 (2019)
- <sup>45</sup> S. Hayami, Y. Yanagi, H. Kusunose, "Spontaneous antisymmetric spin splitting in noncollinear antiferromagnets without spin-orbit coupling". *Phys. Rev. B* vol. 101, p. 220403(R) (2020). DOI: 10.1103/PhysRevB.101.220403
- <sup>46</sup> W. Limmer, J. Daeubler, L. Dreher, M. Glunk, W. Schoch, S. Schwaiger, R. Sauer, Phys. Rev. B 77, 205210 (2008).
- <sup>47</sup> M. Trushin et al., PRB 80, 134405
- <sup>48</sup> M. Q. Dong, Z. X. Song, and Zhi-Xin Guo, Phys. Rev. B 111, 174447 (2025).