

Quantum Optics, Homework 1

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Scully 1.1 The radiation field in an empty cubic cavity of side L satisfies the wave equation

$$\nabla^2 \mathbf{A} - \frac{1}{c^2} \frac{\partial^2 \mathbf{A}}{\partial t^2} = 0 \quad (1)$$

together with the Coulomb gauge condition $\nabla \cdot \mathbf{A} = 0$. Show that the solution that satisfies the boundary conditions has components

$$\begin{aligned} A_x(\mathbf{r}, t) &= A_x(t) \cos(k_x x) \sin(k_y y) \sin(k_z z), \\ A_y(\mathbf{r}, t) &= A_y(t) \sin(k_x x) \cos(k_y y) \sin(k_z z), \\ A_z(\mathbf{r}, t) &= A_z(t) \sin(k_x x) \sin(k_y y) \cos(k_z z), \end{aligned} \quad (2)$$

where $\mathbf{A}(t)$ is independent of position and the wave vector \mathbf{k} has components given by Eq. (1.1.21). Hence show that the integers n_x, n_y, n_z in Eq. (1.1.21) are restricted in that only one of them can be zero at a time.

Solution The boundary condition concerning the vector potential is

$$\mathbf{n} \times (\mathbf{A}_1 - \mathbf{A}_2) = 0, \quad \mathbf{n} \cdot (\nabla \times \mathbf{A}_1 - \nabla \times \mathbf{A}_2) = 0.$$

Since under the Coulomb gage condition

$$\mathbf{E} = -\frac{\partial \mathbf{A}}{\partial t}, \quad \mathbf{B} = \nabla \times \mathbf{A}$$

and both of them vanish outside the cavity, the vector potential is a spacial and temporal constant outside the cavity. We are free to add a global constant to \mathbf{A} , and the most convenient choice is to let $\mathbf{A} = 0$ outside the cavity, so the boundary condition is just

$$\mathbf{n} \times \mathbf{A} = 0, \quad \mathbf{n} \cdot (\nabla \times \mathbf{A}) = 0$$

inside the cavity.

The first condition reads $A_x = A_y = 0$ on the $z = 0$ plane and $z = L$ plane, $A_x = A_z = 0$ on the $y = 0$ plane and the $y = L$ plane, and $A_y = A_z = 0$ on the $x = 0$ plane and $x = L$ plane, which *do not* mix the three components of \mathbf{A} together, so solving (1) with the boundary condition $\mathbf{n} \times \mathbf{A} = 0$ is just solving

$$\nabla^2 A_i - \frac{1}{c^2} \frac{\partial^2 A_i}{\partial t^2} = 0, \quad i = x, y, z \quad (3)$$

separately for all the three components. The problem for A_x is

$$\left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) A_x - \frac{1}{c^2} \frac{\partial^2 A_x}{\partial t^2} = 0, \quad A_x|_{y=0} = A_x|_{y=L} = A_x|_{z=0} = A_x|_{z=L} = 0. \quad (4)$$

By separation of variables we seek a solution in the form of

$$A_x(\mathbf{r}, t) = A_{x,xt}(x, t) A_{xy}(y) A_{xz}(z),$$

and the equations for $A_{x,xt}, A_{xy}, A_{xz}$ are

$$\left(\frac{\partial^2}{\partial x^2} - k_y^2 - k_z^2 \right) A_{x,xt} - \frac{1}{c^2} \frac{\partial^2 A_{x,xt}}{\partial t^2} = 0, \quad \frac{\partial^2 A_{xy}}{\partial y^2} + k_y^2 = 0, \quad \frac{\partial^2 A_{xz}}{\partial z^2} + k_z^2 = 0,$$

respectively. Imposing boundary conditions to (4) to A_{xy} and A_{xz} we obtain

$$A_{xy}(y) = \text{const} \times \sin(k_y y), \quad A_{xz}(z) = \text{const} \times \sin(k_z z),$$

where

$$k_y = \frac{2\pi n_y}{L}, \quad k_z = \frac{2\pi n_z}{L}, \quad n_y, n_z \in \mathbb{Z}, \quad (5)$$

and we find A_x has the form

$$A_x(\mathbf{r}, t) = A_{x,xt}(x, t) \sin(k_{xy}y) \sin(k_{xz}z), \quad k_{xy} = \frac{2\pi n_{xy}}{L}, k_{xz} = \frac{2\pi n_{xz}}{L}, \quad n_{xy}, n_{xz} \in \mathbb{Z}. \quad (6)$$

Similarly we solve (3) for A_y and the result is

$$A_y(\mathbf{r}, t) = A_{y,yt}(y, t) \sin(k_{yx}x) \sin(k_{yz}z), \quad k_{yx} = \frac{2\pi n_{yx}}{L}, k_{yz} = \frac{2\pi n_{yz}}{L}, \quad n_{yx}, n_{yz} \in \mathbb{Z}, \quad (7)$$

and for A_z the result is

$$A_z(\mathbf{r}, t) = A_{z,zt}(z, t) \sin(k_{zx}x) \sin(k_{zy}y), \quad k_{zx} = \frac{2\pi n_{zx}}{L}, k_{zy} = \frac{2\pi n_{zy}}{L}, \quad n_{zx}, n_{zy} \in \mathbb{Z}, \quad (8)$$

Now the Coulomb gauge condition is just

$$\frac{\partial A_{x,xt}}{\partial x} \sin(k_{xy}y) \sin(k_{xz}z) + \frac{\partial A_{y,yt}}{\partial y} \sin(k_{yx}x) \sin(k_{yz}z) + \frac{\partial A_{z,zt}}{\partial z} \sin(k_{zx}x) \sin(k_{zy}y) = 0.$$

This equation holds only when

$$k_{xy} = k_{zy}, \quad k_{xz} = k_{yz}, \quad k_{yx} = k_{zx},$$

and

$$\frac{\partial A_{x,xt}}{\partial x} \propto \sin(k_{yx}x), \quad \frac{\partial A_{y,yt}}{\partial y} \propto \sin(k_{xy}y), \quad \frac{\partial A_{z,zt}}{\partial z} \propto \sin(k_{yz}z).$$

We therefore rename k_{yx} as k_x , k_{xz} as k_z , and k_{xy} as k_y , and we have

$$A_{x,xt} \propto \cos(k_x x), \quad A_{y,yt} \propto \cos(k_y y), \quad A_{z,zt} \propto \cos(k_z z),$$

so finally, we have

$$A_x(\mathbf{r}, t) = \text{some function of } t \times \cos(k_x x) \sin(k_y y) \sin(k_z z)$$

and similar equations for A_y and A_z , where k_x, k_y, k_z all satisfy Eq. (1.1.21), and thus we have proved (2).

Now if $n_x = n_y = n_z$, then $\mathbf{k} = 0$. Hence $\sin(k_i r_i)$, $i = x, y, z$ are constantly zero, and (2) is an all-zero trivial solution. If two of the integers n_x, n_y, n_z are zero, say, $n_x = n_y = 0$, that still makes (2) an all-zero solution. That explains why only one of them can be zero at a time.

Scully 1.2 If A and B are two noncommuting operators that satisfy the conditions

$$[[A, B], A] = [[A, B], B] = 0, \quad (9)$$

then show that

$$\begin{aligned} e^{A+B} &= e^{-\frac{1}{2}[A,B]} e^A e^B \\ &= e^{+\frac{1}{2}[A,B]} e^B e^A. \end{aligned} \quad (10)$$

This is a special case of the so-called Baker-Hausdorff theorem of group theory.

Solution To show (10) it is sufficient to prove

$$e^A e^B = e^{A+B+\frac{1}{2}[A,B]}, \quad (11)$$

as $[A, B]$ commutes with both A and B and therefore commute with e^A and e^B . To prove (11), we define an operator function $G(x)$ by

$$e^{xA} e^{xB} = e^{G(x)}. \quad (12)$$

When $x = 0$, $e^{G(x)} = 1$, so $G(x) = 0$, and we have Taylor series of $G(x)$ at $x = 0$:

$$G(x) = xG_1 + x^2G_2 + \cdots$$

Now consider the trivial equation

$$e^{-xB}e^{-xA}\frac{d}{dx}(e^{xA}e^{xB}) = e^{-G(x)}\frac{d}{dx}e^{G(x)}, \quad (13)$$

where LHS is

$$\begin{aligned} e^{-xB}e^{-xA}\frac{d}{dx}(e^{xA}e^{xB}) &= e^{-xB}e^{-xA}(Ae^{xA}e^{xB} + e^{xA}Be^{xB}) \\ &= e^{-xB}Ae^{xB} + B. \end{aligned} \quad (14)$$

Now we need an important equation:

$$e^GAe^{-G} = A + [G, A] + \frac{1}{2!}[G, [G, A]] + \cdots + \frac{1}{n!}\underbrace{[G, [G, [G, \dots [G, A]]] \dots]}_{n \text{ times}} + \cdots. \quad (15)$$

We use the proof given by [2]. By definition we have

$$\begin{aligned} e^GAe^{-G} &= \sum_{m=0}^{\infty} \frac{1}{m!} G^m A \sum_{n=0}^{\infty} \frac{1}{n!} (-G)^n \\ &= \sum_{m=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{m!n!} (-1)^n G^m A G^n \\ &= \sum_{s=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{(s-n)!n!} (-1)^n G^{s-n} A G^n \\ &= \sum_{s=0}^{\infty} \frac{1}{s!} \sum_{n=0}^{\infty} \binom{s}{n} (-1)^n G^{s-n} A G^n \\ &= A + [G, A] + \cdots + \frac{1}{s!} \sum_{n=0}^s \binom{s}{n} (-1)^n G^{s-n} A G^n + \cdots, \end{aligned}$$

so the problem is whether

$$\sum_{k=0}^n \binom{n}{k} (-1)^k G^{n-k} A G^k = \underbrace{[G, [G, [G, \dots [G, A]]] \dots]}_{n \text{ times}}. \quad (16)$$

(16) can be shown recursively. The cases of $n = 1, 2$ are trivial. Suppose (16) for an arbitrary n , then

$$\begin{aligned} &\underbrace{[G, [G, [G, \dots [G, A]]] \dots]}_{n+1 \text{ times}} \\ &= \sum_{k=0}^n \binom{n}{k} (-1)^k G^{n-k} [G, A] G^k \\ &= \sum_{k=0}^n \binom{n}{k} (-1)^k G^{n-k+1} A G^k - G^{n-k} A G^{k+1} \\ &= \sum_{k=1}^n \binom{n}{k} (-1)^k (G^{n-k+1} A G^k + G^{n+1} A - \sum_{k=1}^n \binom{n}{k-1} (-1)^{k-1} G^{n-k+1} A G^k - (-1)^n A G^{n+1}) \\ &= G^{n+1} A - (-1)^n A G^{n+1} + \sum_{k=1}^n \left(\binom{n}{k} + \binom{n}{k-1} \right) (-1)^k G^{n-k+1} A G^k \\ &= G^{n+1} A - (-1)^n A G^{n+1} + \sum_{k=1}^n \binom{n+1}{k} (-1)^k G^{n-k+1} A G^k \\ &= \sum_{k=0}^{n+1} \binom{n+1}{k} (-1)^k G^{n-k+1} A G^k, \end{aligned}$$

which is just (16) with n replaced by $n + 1$. So (16) has been proven.

Now due to (15), the RHS of (14) is

$$B + e^{-xB} A e^{xB} = B + A + [-xB, A].$$

All higher order terms contain $[B, [B, A]]$ and therefore vanish. The RHS of (13) can be evaluated as

$$e^{-G(x)} \frac{d}{dx} e^{G(x)} = e^{-G(x)} \sum_{m=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{(m+n+1)!} G^m G' G^n,$$

and by the integral formula

$$\int_0^1 dy (1-y)^n y^m = \frac{n!m!}{(n+m+1)!}$$

we have

$$\begin{aligned} e^{-G(x)} \frac{d}{dx} e^{G(x)} &= e^{-G(x)} \sum_{m=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{m!n!} \int_0^1 dy (1-y)^n y^m G^m G' G^n \\ &= e^{-G(x)} \int_0^1 dy e^{(1-y)G} G' e^{yG} \\ &= \int_0^1 dy e^{-yG} G' e^{yG}, \end{aligned}$$

and now it can be evaluated using (15) as

$$\begin{aligned} e^{-G(x)} \frac{d}{dx} e^{G(x)} &= \int_0^1 dy \left(G' + y[G', G] + \frac{y^2}{2} [[G', G], G] + \mathcal{O}(y^3) \right) \\ &= G' + \frac{1}{2} [G', G] + \frac{1}{3!} [[G', G], G] + \dots \end{aligned}$$

As $G(x)$ can be expanded into

$$G'(x) = G_1 + 2xG_2 + \dots,$$

We have

$$e^{-G(x)} \frac{d}{dx} e^{G(x)} = G_1 + 2xG_2 + \frac{1}{2}x^2[G_2, G_1] + \dots$$

So finally, by rewriting its LHS and RHS, from (13) we find that

$$B + A + x[A, B] = G_1 + 2xG_2 + \mathcal{O}(x^2),$$

and thus

$$G_1 = A + B, \quad G_2 = \frac{1}{2}[A, B], \quad G_n = 0 \text{ for } n > 2.$$

So

$$G(x) = A + B + \frac{1}{2}[A, B],$$

and by taking $x = 1$ in (12) we complete the proof of (11), therefore have proved (10).

Scully 1.4 If $f(a, a^\dagger)$ is a function which can be expanded in a power series of a and a^\dagger , then show that (a) $[a, f(a, a^\dagger)] = \frac{\partial f}{\partial a^\dagger}$, (b) $[a^\dagger, f(a, a^\dagger)] = -\frac{\partial f}{\partial a}$, (c) $e^{-\alpha a^\dagger} f(a, a^\dagger) e^{\alpha a^\dagger} = f(ae^\alpha, a^\dagger e^{-\alpha})$ where α is a parameter.

Solution Suppose

$$f(a, a^\dagger) = \sum_{m,n \geq 0} f_{mn} a^m (a^\dagger)^n. \quad (17)$$

We do not include terms that start with a^\dagger , because by $[a, a^\dagger] = 1$ we can move all a operators to the left. We can also move all a operators to the right, and write down a similar expansion

$$f(a, a^\dagger) = \sum_{m,n \geq 0} g_{mn} (a^\dagger)^m a^n. \quad (18)$$

(a) We have

$$\begin{aligned}
[a, f(a, a^\dagger)] &= \sum_{m,n \geq 0} f_{mn} a^m [a, (a^\dagger)^m] \\
&= \sum_{m,n \geq 0} f_{mn} a^m ([a, a^\dagger] (a^\dagger)^{m-1} + a^\dagger [a, (a^\dagger)^{m-1}]) \\
&= \sum_{m,n \geq 0} f_{mn} a^m ((a^\dagger)^{m-1} + a^\dagger [a, (a^\dagger)^{m-1}]) \\
&= \sum_{m,n \geq 0} f_{mn} a^m ((a^\dagger)^{m-1} + a^\dagger (a^\dagger [a, (a^\dagger)^{m-2}] + [a, a^\dagger] (a^\dagger)^{m-2})) \\
&= \sum_{m,n \geq 0} f_{mn} a^m ((a^\dagger)^{m-1} + a^\dagger (a^\dagger [a, (a^\dagger)^{m-2}] + (a^\dagger)^{m-2})) \\
&= \sum_{m,n \geq 0} f_{mn} a^m (2(a^\dagger)^{m-1} + (a^\dagger)^2 [a, (a^\dagger)^{m-2}]) \\
&= \dots
\end{aligned}$$

As the expansion of $[a, (a^\dagger)^m]$ goes, we get a sequence of equations in the form of

$$\begin{aligned}
[a, f(a, a^\dagger)] &= \sum_{m,n \geq 0} f_{mn} a^m ((a^\dagger)^{m-1} + a^\dagger [a, (a^\dagger)^{m-1}]) \\
&= \sum_{m,n \geq 0} f_{mn} a^m (2(a^\dagger)^{m-1} + a^\dagger [a, (a^\dagger)^{m-2}]) \\
&= \dots,
\end{aligned}$$

and the sequence stops at

$$\sum_{m,n \geq 0} f_{mn} a^m (m(a^\dagger)^{m-1} + a^\dagger [a, (a^\dagger)^{m-m}]),$$

so we have

$$[a, f(a, a^\dagger)] = \sum_{m,n \geq 0} f_{mn} a^m m(a^\dagger)^{m-1},$$

and since

$$\frac{\partial f}{\partial a^\dagger} = \sum_{m,n \geq 0} f_{mn} a^m \frac{\partial (a^\dagger)^m}{\partial a^\dagger},$$

we arrive at the conclusion that

$$[a, f(a, a^\dagger)] = \frac{\partial f}{\partial a^\dagger}. \tag{19}$$

(b) The logic is similar to (a) but this time the expansion is

$$\begin{aligned}
[a^\dagger, f(a, a^\dagger)] &= \sum_{m,n \geq 0} [a^\dagger, a^m] (a^\dagger)^n \\
&= \sum_{m,n \geq 0} ([a^\dagger, a] a^{m-1} + a[a^\dagger, a^{m-1}]) \\
&= \sum_{m,n \geq 0} (-a^{m-1} + a[a^\dagger, a^{m-1}]) \\
&= \sum_{m,n \geq 0} (-a^{m-1} + a(a[a^\dagger, a^{m-2}] + [a^\dagger, a] a^{m-2})) \\
&= \sum_{m,n \geq 0} (-a^{m-1} + a((a[a^\dagger, a^{m-2}]) - a^{m-2})) \\
&= \sum_{m,n \geq 0} (-2a^{m-1} + a^2[a^\dagger, a^{m-2}]) \\
&= \dots,
\end{aligned}$$

or to be concise,

$$\begin{aligned} [a^\dagger, f(a, a^\dagger)] &= \sum_{m,n \geq 0} (-a^{m-1} + a[a^\dagger, a^{m-1}]) \\ &= \sum_{m,n \geq 0} (-2a^{m-1} + a^2[a^\dagger, a^{m-2}]) \\ &= \dots, \end{aligned}$$

and this time the sequence stops at

$$[a^\dagger, f(a, a^\dagger)] = \sum_{m,n \geq 0} (-ma^{m-1} + a^2[a^\dagger, a^{m-m}]),$$

and since

$$-ma^{m-1} = -\frac{\partial a^m}{\partial a}$$

we have

$$[a^\dagger, f(a, a^\dagger)] = -\frac{\partial f}{\partial a}. \quad (20)$$

(c) We need to use results proved in Problem 1.5. By (23), we have

$$ae^{-\alpha a^\dagger a} = e^{-\alpha} e^{-\alpha a^\dagger a} a = (e^{-\alpha} a) e^{-\alpha a^\dagger a}, \quad (21)$$

and by taking its conjugate transpose we have

$$a^\dagger e^{-\alpha^* a^\dagger a} e^{-\alpha^*} = e^{-\alpha^* a^\dagger a} a^\dagger,$$

and since α is an arbitrary parameter we can redefine it and get

$$a^\dagger e^{-\alpha a^\dagger a} e^{-\alpha} = e^{-\alpha a^\dagger a} a^\dagger = (a^\dagger e^{-\alpha}) e^{-\alpha a^\dagger a}. \quad (22)$$

Substituting (18) into $e^{-\alpha a^\dagger a} f(a, a^\dagger) e^{\alpha a^\dagger a}$, and applying (21) and (22) iteratively, we have

$$\begin{aligned} e^{-\alpha a^\dagger a} f(a, a^\dagger) e^{\alpha a^\dagger a} &= \sum_{n,m \geq 0} g_{mn} e^{-\alpha a^\dagger a} (a^\dagger)^m a^n e^{\alpha a^\dagger a} \\ &= \sum_{n,m \geq 0} g_{mn} (e^{-\alpha} a^\dagger) e^{-\alpha a^\dagger a} (a^\dagger)^{m-1} a^{n-1} e^{\alpha a^\dagger a} (ae^\alpha) \\ &= \sum_{n,m \geq 0} g_{mn} (e^{-\alpha} a^\dagger) (e^{-\alpha} a^\dagger) e^{-\alpha a^\dagger a} (a^\dagger)^{m-2} a^{n-2} e^{\alpha a^\dagger a} (ae^\alpha) (ae^\alpha) \\ &= \dots \\ &= \sum_{m,n \geq 0} g_{mn} (e^{-\alpha} a^\dagger)^m e^{-\alpha a^\dagger a} e^{\alpha a^\dagger a} (e^\alpha a)^n \\ &= \sum_{m,n \geq 0} g_{mn} (e^{-\alpha} a^\dagger)^m (e^\alpha a)^n \\ &= f(ae^\alpha, a^\dagger e^{-\alpha}), \end{aligned}$$

which finishes the proof.

Scully 1.5 Show that

$$\begin{aligned} [a, e^{-\alpha a^\dagger a}] &= (e^{-\alpha} - 1) e^{-\alpha a^\dagger a} a, \\ [a^\dagger, e^{-\alpha a^\dagger a}] &= (e^\alpha - 1) e^{-\alpha a^\dagger a} a^\dagger, \end{aligned} \quad (23)$$

where α is a parameter.

Solution The most convenient way is not to invoke (19) and (20), but to check the matrix elements under the $\{|n\rangle\}$ basis. We have

$$\begin{aligned}
[a, e^{-\alpha a^\dagger a}] |n\rangle &= a e^{-\alpha a^\dagger a} |n\rangle - e^{-\alpha a^\dagger a} a |n\rangle \\
&= a e^{-\alpha n} |n\rangle - e^{-\alpha a^\dagger a} \sqrt{n} |n-1\rangle \\
&= e^{-\alpha n} \sqrt{n} |n-1\rangle - e^{-\alpha(n-1)} \sqrt{n} |n-1\rangle \\
&= (e^{-\alpha} - 1) e^{-\alpha(n-1)} \sqrt{n} |n-1\rangle \\
&= (e^{-\alpha} - 1) e^{-\alpha a^\dagger a} \sqrt{n} |n-1\rangle \\
&= (e^{-\alpha} - 1) e^{-\alpha a^\dagger a} a |n\rangle,
\end{aligned}$$

for all $n = 0, 1, \dots$, and thus

$$[a, e^{-\alpha a^\dagger a}] = (e^{-\alpha} - 1) e^{-\alpha a^\dagger a} a. \quad (24)$$

The same procedure applies for the second equation. We have

$$\begin{aligned}
[a^\dagger, e^{-\alpha a^\dagger a}] |n\rangle &= a^\dagger e^{-\alpha a^\dagger a} |n\rangle - e^{-\alpha a^\dagger a} a^\dagger |n\rangle \\
&= a^\dagger e^{-\alpha n} |n\rangle - e^{-\alpha a^\dagger a} \sqrt{n+1} |n+1\rangle \\
&= e^{-\alpha n} \sqrt{n+1} |n+1\rangle - e^{-\alpha(n+1)} \sqrt{n+1} |n+1\rangle \\
&= (e^\alpha - 1) e^{-\alpha(n+1)} \sqrt{n+1} |n+1\rangle \\
&= (e^\alpha - 1) e^{-\alpha a^\dagger a} \sqrt{n+1} |n+1\rangle \\
&= (e^\alpha - 1) e^{-\alpha a^\dagger a} a^\dagger |n\rangle,
\end{aligned}$$

for all $n = 0, 1, \dots$, and thus

$$[a^\dagger, e^{-\alpha a^\dagger a}] = (e^\alpha - 1) e^{-\alpha a^\dagger a} a^\dagger. \quad (25)$$

Scully 1.6 Show that the free-field Hamiltonian

$$\mathcal{H} = \hbar \nu \left(a^\dagger a + \frac{1}{2} \right) \quad (26)$$

can be written in terms of the number states as

$$\mathcal{H} = \sum_n E_n |n\rangle \langle n|,$$

and hence

$$e^{i\mathcal{H}t/\hbar} = \sum_n e^{iE_n t/\hbar} |n\rangle \langle n|. \quad (27)$$

Solution An n -photon state

$$|n\rangle = \frac{1}{\sqrt{n!}} (a^\dagger)^n |0\rangle \quad (28)$$

is an eigenstate of (26), since by applying (26) on (28) we have

$$\begin{aligned}
\mathcal{H} |n\rangle &= \hbar \nu a^\dagger a |n\rangle + \frac{1}{2} \hbar \nu |n\rangle \\
&= \hbar \nu a^\dagger \sqrt{n} |n-1\rangle + \frac{1}{2} \hbar \nu |n\rangle \\
&= \hbar \nu \sqrt{n} \sqrt{n-1+1} |n-1+1\rangle + \frac{1}{2} \hbar \nu |n\rangle \\
&= \hbar \nu \left(n + \frac{1}{2} \right) |n\rangle \\
&= E_n |n\rangle,
\end{aligned}$$

so $|n\rangle$ is an eigenstate of \mathcal{H} with energy E_n . Since all $|n\rangle$ are orthogonal to each other and are uniform, we have

$$\mathcal{H} = \sum_n E_n |n\rangle \langle n|. \quad (29)$$

Since \mathcal{H} is diagonal under the $\{|n\rangle\}$ basis, so is $e^{i\mathcal{H}t/\hbar}$ and hence we have (27).

Scully 2.1 Show that

$$a^\dagger |\alpha\rangle\langle\alpha| = \left(\alpha^* + \frac{\partial}{\partial\alpha} \right) |\alpha\rangle\langle\alpha|, \quad (30)$$

and

$$|\alpha\rangle\langle\alpha| a = \left(\alpha + \frac{\partial}{\partial\alpha^*} \right) |\alpha\rangle\langle\alpha|. \quad (31)$$

Solution By definition

$$|\alpha\rangle\langle\alpha| = e^{-|\alpha|^2} \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha^n (\alpha^*)^m}{\sqrt{m!n!}} |n\rangle\langle m|,$$

so

$$\begin{aligned} a^\dagger |\alpha\rangle\langle\alpha| &= e^{-|\alpha|^2} \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha^n (\alpha^*)^m}{\sqrt{m!n!}} a^\dagger |n\rangle\langle m| \\ &= e^{-|\alpha|^2} \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha^n (\alpha^*)^m}{\sqrt{m!n!}} \sqrt{n+1} |n+1\rangle\langle m| \\ &= e^{-|\alpha|^2} \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{\alpha^n (\alpha^*)^m}{\sqrt{m!(n+1)!}} (n+1) |n+1\rangle\langle m| \\ &= e^{-|\alpha|^2} \sum_{n=0}^{\infty} \frac{1}{\sqrt{(n+1)!}} \frac{\partial}{\partial\alpha} \alpha^{n+1} |n+1\rangle \sum_{m=0}^{\infty} \frac{(\alpha^*)^m}{\sqrt{m!}} \langle m| \\ &= e^{-|\alpha|^2} \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \left(\frac{\partial}{\partial\alpha} \alpha^n \right) |n\rangle \sum_{m=0}^{\infty} \frac{(\alpha^*)^m}{\sqrt{m!}} \langle m|. \end{aligned}$$

Note that the factors after \sum_m do not include α^* , so they can be regarded as constants. We have

$$\begin{aligned} e^{-|\alpha|^2} \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \frac{\partial}{\partial\alpha} \alpha^n |n\rangle &= e^{-|\alpha|^2} \frac{\partial}{\partial\alpha} \sum_{n=0}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \\ &= \frac{\partial}{\partial\alpha} \left(e^{-|\alpha|^2} \sum_{n=0}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \right) - \left(\frac{\partial e^{-|\alpha|^2}}{\partial\alpha} \right) \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \\ &= \frac{\partial}{\partial\alpha} \left(e^{-|\alpha|^2} \sum_{n=0}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \right) - (-\alpha^* e^{-|\alpha|^2}) \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle. \end{aligned}$$

Therefore, we have

$$\begin{aligned} a^\dagger |\alpha\rangle\langle\alpha| &= e^{-|\alpha|^2} \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \left(\frac{\partial}{\partial\alpha} \alpha^n \right) |n\rangle \sum_{m=0}^{\infty} \frac{(\alpha^*)^m}{\sqrt{m!}} \langle m| \\ &= \frac{\partial}{\partial\alpha} \left(e^{-|\alpha|^2} \sum_{n=0}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \right) \sum_{m=0}^{\infty} \frac{(\alpha^*)^m}{\sqrt{m!}} \langle m| - (-\alpha^* e^{-|\alpha|^2}) \sum_{n=1}^{\infty} \frac{1}{\sqrt{n!}} \alpha^n |n\rangle \sum_{m=0}^{\infty} \frac{(\alpha^*)^m}{\sqrt{m!}} \langle m| \\ &= \frac{\partial}{\partial\alpha} |\alpha\rangle\langle\alpha| + \alpha^* |\alpha\rangle\langle\alpha|, \end{aligned}$$

and thus we have proved that

$$a^\dagger |\alpha\rangle\langle\alpha| = \left(\alpha^* + \frac{\partial}{\partial\alpha} \right) |\alpha\rangle\langle\alpha|.$$

The scope of $\frac{\partial}{\partial\alpha}$ covers $|\alpha\rangle\langle\alpha|$. By taking the conjugate transpose we can immediately show (31), since α and α^* have exactly the same status in $|\alpha\rangle\langle\alpha|$, and taking the conjugate transpose of $f(\alpha, \alpha^*)$ is just to exchange α and α^* , $|m\rangle$ and $|n\rangle$, so after the conjugate transpose $|\alpha\rangle\langle\alpha|$ is still $|\alpha\rangle\langle\alpha|$ while $\frac{\partial}{\partial\alpha}$ turns into $\frac{\partial}{\partial\alpha^*}$, i.e.

$$\left(\frac{\partial}{\partial\alpha} |\alpha\rangle\langle\alpha| \right)^\dagger = \left(\frac{\partial}{\partial\alpha} |\alpha\rangle\langle\alpha| \right) \Big|_{\alpha \leftrightarrow \alpha^*, \text{bra} \leftrightarrow \text{ket}} = \frac{\partial}{\partial\alpha^*} |\alpha\rangle\langle\alpha|.$$

Scully 3.3 Show that the Wigner-Weyl distribution $W(\alpha, \alpha^*)$ can be expressed in terms of the P -representation $P(\alpha, \alpha^*)$ via the relation

$$W(\alpha, \alpha^*) = \frac{2}{\pi} \int d^2\beta P(\beta, \beta^*) \exp(-2|\alpha - \beta|^2). \quad (32)$$

Solution Both $W(\alpha, \alpha^*)$ and $P(\alpha, \alpha^*)$ can be defined in terms of integral transforms. We have

$$P(\alpha, \alpha^*) = \frac{1}{\pi^2} \int d^2\beta e^{-i\beta\alpha^* - i\beta^*\alpha} \text{Tr} \left(e^{i\beta^*a} e^{i\beta a^\dagger} \rho \right)$$

and

$$W(\alpha, \alpha^*) = \frac{1}{\pi^2} \int d^2\beta e^{-i\beta\alpha^* - i\beta^*\alpha} \text{Tr} \left(e^{i\beta a^\dagger + i\beta^*a} \rho \right).$$

By the BCH formula and the completeness of coherent states we have

$$\begin{aligned} \text{Tr} \left(e^{i\beta a^\dagger + i\beta^*a} \rho \right) &= \text{Tr} \left(e^{i\beta^*a} e^{i\beta a^\dagger} e^{|\beta|^2/2} \rho \right) \\ &= e^{|\beta|^2/2} \text{Tr} \left(e^{i\beta a^\dagger} \rho e^{i\beta^*a} \right) \\ &= \frac{1}{\pi} e^{|\beta|^2/2} \int d^2\gamma \langle \gamma | e^{i\beta a^\dagger} \rho e^{i\beta^*a} | \gamma \rangle \\ &= \frac{1}{\pi} e^{|\beta|^2/2} \int d^2\gamma e^{i\beta\gamma^*} \langle \gamma | \rho | \gamma \rangle e^{i\beta^*\gamma} \\ &= e^{|\beta|^2/2} \int d^2\gamma e^{i\beta\gamma^* + i\beta^*\gamma} Q(\gamma, \gamma^*), \end{aligned}$$

so

$$\begin{aligned} W(\alpha, \alpha^*) &= \frac{1}{\pi^2} \int d^2\gamma \int d^2\beta e^{-i\beta\alpha^* - i\beta^*\alpha} e^{|\beta|^2/2} e^{i\beta\gamma^* + i\beta^*\gamma} Q(\gamma, \gamma^*) \\ &= \frac{1}{\pi^2} \int d^2\gamma \int d^2\beta e^{|\beta|^2/2 + i\beta(\gamma^* - \alpha^*) + i\beta^*(\gamma - \alpha)} Q(\gamma, \gamma^*). \end{aligned}$$

Now by completing the squares we can integrate out the variable β , and obtain

$$\begin{aligned} W(\alpha, \alpha^*) &= \frac{1}{\pi^2} \int d^2\gamma \left(\int d^2\beta e^{|\beta|^2/2 + i\beta(\gamma^* - \alpha^*) + i\beta^*(\gamma - \alpha)} \right) Q(\gamma, \gamma^*) \\ &= \frac{1}{\pi^2} \int d^2\gamma 2\pi e^{2(\gamma^* - \alpha^*)(\gamma - \alpha)} Q(\gamma, \gamma^*), \end{aligned}$$

so we have proved the relation between $W(\alpha, \alpha^*)$ and $Q(\alpha, \alpha^*)$ that

$$W(\alpha, \alpha^*) = \frac{2}{\pi} \int d^2\gamma e^{2|\gamma - \alpha|^2} Q(\gamma, \gamma^*). \quad (33)$$

We also have a similar relation between $P(\alpha, \alpha^*)$ and $Q(\alpha, \alpha^*)$ that

$$Q(\alpha, \alpha^*) = \frac{1}{\pi} \int d^2\beta P(\beta, \beta^*) e^{-|\alpha - \beta|^2}, \quad (34)$$

which is Eq. (3.2.9) in [1]. Now by putting (34) and (33) together we can finally show (32). We have

$$\begin{aligned} W(\alpha, \alpha^*) &= \frac{2}{\pi} \int d^2\gamma e^{2|\gamma - \alpha|^2} \frac{1}{\pi} \int d^2\beta P(\beta, \beta^*) e^{-|\gamma - \beta|^2} \\ &= \frac{2}{\pi^2} \int d^2\beta P(\beta, \beta^*) \int d^2\gamma e^{|\gamma|^2 + \gamma(\beta^* - 2\alpha^*) + \gamma^*(\beta - 2\alpha)} e^{2|\alpha|^2 - |\beta|^2}, \end{aligned}$$

and again, by completing the squares we have

$$\begin{aligned} W(\alpha, \alpha^*) &= \frac{2}{\pi^2} \int d^2\beta P(\beta, \beta^*) \times \pi e^{-(\beta^* - 2\alpha^*)(\beta - 2\alpha)} \times e^{2|\alpha|^2 - |\beta|^2} \\ &= \frac{2}{\pi} \int d^2\beta P(\beta, \beta^*) e^{-2|\alpha|^2 - 2|\beta|^2 + 2\alpha^*\beta + 2\alpha\beta^*}, \end{aligned}$$

so we get (32).

Q-functions for several states Derive the Q -function for Fock state $|n\rangle$, coherent state $|\alpha\rangle$ and the famous “cat” state $C(|\alpha\rangle + |-\alpha\rangle)$. Discuss the corresponding P -functions.

Solution

(a) The Q -function of $|n\rangle$ is

$$\begin{aligned} Q(\alpha, \alpha^*) &= \frac{1}{\pi} \langle \alpha | n \rangle \langle n | \alpha \rangle \\ &= \frac{1}{\pi} \left| e^{-|\alpha|^2/2} \sum_{m=0}^{\infty} \frac{\alpha^m}{\sqrt{m!}} \langle n | m \rangle \right|^2 \\ &= \frac{1}{\pi} \left| e^{-|\alpha|^2/2} \frac{\alpha^n}{\sqrt{n!}} \right|^2 \\ &= \frac{e^{-|\alpha|^2} |\alpha|^{2n}}{\pi n!}. \end{aligned}$$

The corresponding P -function is

$$\begin{aligned} P(\alpha, \alpha^*) &= \frac{e^{|\alpha|^2}}{\pi^2} \int d^2\beta \langle -\beta | n \rangle \langle n | \beta \rangle e^{|\beta|^2} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2} \int d^2\beta \left(\sum_{m=0}^{\infty} e^{-|\beta|^2/2} \frac{(-\beta^*)^m}{\sqrt{m!}} \langle m | n \rangle \right) \left(\sum_{m=0}^{\infty} e^{-|\beta|^2/2} \frac{\beta^m}{\sqrt{m!}} \langle n | m \rangle \right) e^{|\beta|^2} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2} \int d^2\beta \left(e^{-|\beta|^2/2} \frac{(-\beta^*)^n}{\sqrt{n!}} \right) \left(e^{-|\beta|^2/2} \frac{\beta^n}{\sqrt{n!}} \right) e^{|\beta|^2} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2} \int d^2\beta e^{-|\beta|^2} \frac{(-1)^n |\beta|^{2n}}{n!} e^{|\beta|^2} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2 n!} \int d^2\beta (-1)^n |\beta|^{2n} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2 n!} \int d^2\beta \frac{\partial^n}{\partial \alpha^n} \frac{\partial^n}{\partial \alpha^{*n}} e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{\pi^2 n!} \frac{\partial^n}{\partial \alpha^n} \frac{\partial^n}{\partial \alpha^{*n}} \int d^2\beta e^{-\beta\alpha^* + \beta^*\alpha} \\ &= \frac{e^{|\alpha|^2}}{n!} \frac{\partial^n}{\partial \alpha^n} \frac{\partial^n}{\partial \alpha^{*n}} \delta^{(2)}(\alpha), \end{aligned}$$

where we have used the formula

$$\int d^2\beta e^{-\beta\alpha^* + \beta^*\alpha} = \pi^2 \delta^{(2)}(\alpha), \quad (35)$$

which can be proved by substitution of variable $\beta = i\gamma$, and integrate $\text{Re } \gamma$ and $\text{Im } \gamma$ separately. It can be seen that the P -function behaves badly, proportion to the second derivative of δ -function and can be negative near $\alpha = 0$ in certain directions, which is expected since a Fock state $|n\rangle$ with completely determined photon numbers is far from a classical optical field.

(b) The Q -function of $|\alpha\rangle$ is

$$\begin{aligned} Q(\beta, \beta^*) &= \frac{1}{\pi} \langle \beta | \alpha \rangle \langle \alpha | \beta \rangle \\ &= \frac{1}{\pi} |\langle \alpha | \beta \rangle|^2 \\ &= \frac{1}{\pi} e^{-|\alpha - \beta|^2}. \end{aligned}$$

The corresponding P -function is

$$P(\beta, \beta^*) = \delta^{(2)}(\beta - \alpha),$$

as can be directly verified with (34). This is a δ function, and again is expected because a coherent state is a quite “classical” thing, and P -function works best as a probabilistic distribution function for states resembling classical optical fields, so a coherent state must be very “clear” or “sharp” in the P -function representation.

(c) First we normalize the cat state. We have

$$\begin{aligned}
(\langle \alpha| + \langle -\alpha|)(|\alpha\rangle + |-\alpha\rangle) &= 1 + 1 + \langle \alpha|-\alpha\rangle + \langle -\alpha|\alpha\rangle \\
&= 2 + e^{-|\alpha|^2/2 - |\alpha|^2/2 + \alpha^*(-\alpha)} + e^{-|\alpha|^2/2 - |\alpha|^2/2 + (-\alpha^*)\alpha} \\
&= 2 + 2e^{-2|\alpha|^2},
\end{aligned}$$

and hence

$$C = \frac{1}{\sqrt{2 + 2e^{-2|\alpha|^2}}}. \quad (36)$$

Now we calculate the Q -function and P -function of the cat state. The Q -function is

$$\begin{aligned}
Q(\beta, \beta^*) &= \frac{C^2}{\pi} (\langle \beta|\alpha\rangle + \langle \beta|-\alpha\rangle)(\langle \alpha|\beta\rangle + \langle -\alpha|\beta\rangle) \\
&= \frac{C^2}{\pi} (e^{-(|\alpha|^2+|\beta|^2)/2+\beta^*\alpha} + e^{-(|\alpha|^2+|\beta|^2)/2-\beta^*\alpha})(e^{-(|\alpha|^2+|\beta|^2)/2+\beta\alpha^*} + e^{-(|\alpha|^2+|\beta|^2)/2-\beta\alpha^*}) \\
&= \frac{C^2}{\pi} e^{-|\alpha|^2-|\beta|^2} (e^{\beta^*\alpha} + e^{-\beta^*\alpha})(e^{\beta\alpha^*} + e^{-\beta\alpha^*}).
\end{aligned}$$

The P -function is

$$\begin{aligned}
P(\beta, \beta^*) &= C^2 \frac{e^{|\beta|^2}}{\pi^2} \int d^2\gamma (\langle -\gamma|\alpha\rangle + \langle -\gamma|-\alpha\rangle)(\langle \alpha|\gamma\rangle + \langle -\alpha|\gamma\rangle) e^{|\gamma|^2} e^{-\gamma\beta^*+\gamma^*\beta} \\
&= C^2 \frac{e^{|\beta|^2}}{\pi^2} \int d^2\gamma e^{|\gamma|^2} e^{-\gamma\beta^*+\gamma^*\beta} (e^{-(|\gamma|^2+|\alpha|^2)/2-\gamma^*\alpha} + e^{-(|\gamma|^2+|\alpha|^2)/2+\gamma^*\alpha}) \\
&\quad \times (e^{-(|\alpha|^2+|\gamma|^2)/2+\alpha^*\gamma} + e^{-(|\alpha|^2+|\gamma|^2)/2-\alpha^*\gamma}) \\
&= \frac{C^2}{\pi^2} e^{|\beta|^2-|\alpha|^2} \int d^2\gamma e^{-\gamma\beta^*+\gamma^*\beta} (e^{-\gamma^*\alpha} + e^{\gamma^*\alpha})(e^{\alpha^*\gamma} + e^{-\alpha^*\gamma}) \\
&= \frac{C^2}{\pi^2} e^{|\beta|^2-|\alpha|^2} \int d^2\gamma (e^{-\gamma^*(\alpha-\beta)+\gamma(\alpha^*-\beta^*)} + e^{\gamma^*(\alpha+\beta)-\gamma(\alpha^*+\beta^*)} \\
&\quad + (e^{\gamma^*\alpha+\alpha^*\gamma} + e^{-\gamma^*\alpha-\alpha^*\gamma}) e^{-\gamma\beta^*+\gamma^*\beta}).
\end{aligned}$$

By (35) we can find the first two integrals immediately:

$$\int d^2\gamma (e^{-\gamma^*(\alpha-\beta)+\gamma(\alpha^*-\beta^*)} + e^{\gamma^*(\alpha+\beta)-\gamma(\alpha^*+\beta^*)}) = \pi^2 \delta^{(2)}(\alpha - \beta) + \pi^2 \delta^{(2)}(\alpha + \beta).$$

For the third and the fourth term, note that

$$\gamma^* e^{-\gamma\beta^*+\gamma^*\beta} = \frac{d}{d\beta} e^{-\gamma\beta^*+\gamma^*\beta}, \quad \gamma e^{-\gamma\beta^*+\gamma^*\beta} = -\frac{d}{d\beta^*} e^{-\gamma\beta^*+\gamma^*\beta}$$

and therefore

$$\begin{aligned}
\int d^2\gamma (e^{\gamma^*\alpha+\alpha^*\gamma} + e^{-\gamma^*\alpha-\alpha^*\gamma}) e^{-\gamma\beta^*+\gamma^*\beta} &= \int d^2\gamma (e^{\alpha \frac{d}{d\beta} - \alpha^* \frac{d}{d\beta^*}} + e^{-\alpha \frac{d}{d\beta} + \alpha^* \frac{d}{d\beta^*}}) e^{-\gamma\beta^*+\gamma^*\beta} \\
&= \pi^2 (e^{\alpha \frac{d}{d\beta} - \alpha^* \frac{d}{d\beta^*}} + e^{-\alpha \frac{d}{d\beta} + \alpha^* \frac{d}{d\beta^*}}) \delta^{(2)}(\beta),
\end{aligned}$$

so we reach the final result that for a cat state

$$P(\beta, \beta^*) = C^2 e^{|\beta|^2-|\alpha|^2} (\delta^{(2)}(\alpha + \beta) + \delta^{(2)}(\alpha - \beta) + (e^{\alpha \frac{d}{d\beta} - \alpha^* \frac{d}{d\beta^*}} + e^{-\alpha \frac{d}{d\beta} + \alpha^* \frac{d}{d\beta^*}}) \delta^{(2)}(\beta)). \quad (37)$$

Coherent states and dipole radiation Derive in detail the expectation of \mathbf{E} in a system with a electric dipole placed at $\mathbf{r} = 0$ and starting to oscillate at $t = 0$, where the dipole radiation approximation works, i.e. $H_{\text{int}} = -\mathbf{d} \cdot \mathbf{E}$. Consider the $t \rightarrow \infty$ limit.

Solution We already know that the wave function of the optical field, under the interaction picture, is

$$|\psi(t)\rangle = \exp\left(\sum_n (\alpha_n a_n^\dagger - \alpha_n^* a_n)\right) |0\rangle, \quad (38)$$

where

$$\alpha_n = \frac{i}{\hbar} \int_0^t dt' \mathcal{E}_n \mathbf{f}_n^* e^{i\omega_n t'} \cdot (d e^{-i\omega t} + \text{h.c.}). \quad (39)$$

We are going to evaluate $\langle \mathbf{E} \rangle$ under $|\psi\rangle$.

For a free space, n is just the combination of the wave vector \mathbf{k} and the polarization, and the energy spectrum is

$$\omega_{\mathbf{k}\sigma} = c|\mathbf{k}|, \quad (40)$$

and

$$\mathbf{f}_{\mathbf{k}\sigma}(\mathbf{r}) = e^{i\mathbf{k}\cdot\mathbf{r}} \mathbf{e}_\sigma, \quad (41)$$

where for a given \mathbf{k} , we have

$$\mathbf{k} \cdot \mathbf{e}_\sigma = 0 \quad (42)$$

to impose the transverse wave condition. Therefore we have

$$\begin{aligned} \alpha_{\mathbf{k}\sigma} &= \frac{i}{\hbar} \int_0^t dt' \mathcal{E}_{\mathbf{k}\sigma} e^{-i\mathbf{k}\cdot\mathbf{r} + i\omega_{\mathbf{k}} t'} \mathbf{e}_\sigma \cdot (d e^{-i\omega t'} + \text{c.c.})|_{\mathbf{r}=0} \\ &= \frac{1}{\hbar} \mathcal{E}_{\mathbf{k}\sigma} \mathbf{e}_\sigma \cdot \left(d \frac{e^{i(\omega_{\mathbf{k}} - \omega)t} - 1}{\omega_{\mathbf{k}} - \omega} + d^* \frac{e^{i(\omega_{\mathbf{k}} + \omega)t} - 1}{\omega_{\mathbf{k}} + \omega} \right). \end{aligned} \quad (43)$$

The expectation of \mathbf{E} under a coherent state can be written in terms of a sum of different modes in a standard way:

$$\begin{aligned} \langle \mathbf{E} \rangle &= \sum_{\mathbf{k}, \sigma} \mathcal{E}_{\mathbf{k}\sigma} (\mathbf{f}_{\mathbf{k}\sigma}(\mathbf{r}) e^{-i\omega_{\mathbf{k}} t} \langle a_{\mathbf{k}\sigma} \rangle + \mathbf{f}_{\mathbf{k}\sigma}^*(\mathbf{r}) e^{i\omega_{\mathbf{k}} t} \langle a_{\mathbf{k}\sigma}^\dagger \rangle) \\ &= \sum_{\mathbf{k}, \sigma} \mathcal{E}_{\mathbf{k}\sigma} \mathbf{e}_\sigma (e^{i\mathbf{k}\cdot\mathbf{r} - i\omega_{\mathbf{k}} t} \langle a_{\mathbf{k}\sigma} \rangle + e^{-i\mathbf{k}\cdot\mathbf{r} + i\omega_{\mathbf{k}} t} \langle a_{\mathbf{k}\sigma}^\dagger \rangle) \\ &= \sum_{\mathbf{k}, \sigma} \mathcal{E}_{\mathbf{k}\sigma} \mathbf{e}_\sigma (e^{i\mathbf{k}\cdot\mathbf{r} - i\omega_{\mathbf{k}} t} \alpha_{\mathbf{k}\sigma} + e^{-i\mathbf{k}\cdot\mathbf{r} + i\omega_{\mathbf{k}} t} \alpha_{\mathbf{k}\sigma}^*). \end{aligned}$$

Inserting (43) into this equation we have

$$\langle \mathbf{E} \rangle = \sum_{\mathbf{k}, \sigma} \frac{c|\mathbf{k}|}{2\epsilon_0 V} e^{i\mathbf{k}\cdot\mathbf{r} - i\omega_{\mathbf{k}} t} \mathbf{e}_\sigma \mathbf{e}_\sigma \cdot \left(d \frac{e^{i(\omega_{\mathbf{k}} - \omega)t} - 1}{\omega_{\mathbf{k}} - \omega} + d^* \frac{e^{i(\omega_{\mathbf{k}} + \omega)t} - 1}{\omega_{\mathbf{k}} + \omega} \right) + \text{c.c.}$$

Let φ be the phase of the dipole, so

$$\mathbf{d} = d e^{i\varphi} \hat{\mathbf{d}},$$

and since \mathbf{e}_σ is always orthogonal to $\hat{\mathbf{k}}$, $\sum_\sigma \mathbf{e}_\sigma \mathbf{e}_\sigma \cdot \mathbf{d}$ is just \mathbf{d} projected to the plane orthogonal to \mathbf{k} , hence we have

$$\begin{aligned} \langle \mathbf{E} \rangle(\mathbf{r}, t) &= \frac{dc}{2\epsilon_0 V} \sum_{\mathbf{k}} |\mathbf{k}| e^{i\mathbf{k}\cdot\mathbf{r} - i\omega_{\mathbf{k}} t} (\hat{\mathbf{d}} - (\hat{\mathbf{k}} \cdot \hat{\mathbf{d}}) \hat{\mathbf{k}}) \left(e^{i\varphi} \frac{e^{i(\omega_{\mathbf{k}} - \omega)t} - 1}{\omega_{\mathbf{k}} - \omega} + e^{-i\varphi} \frac{e^{i(\omega_{\mathbf{k}} + \omega)t} - 1}{\omega_{\mathbf{k}} + \omega} \right) + \text{c.c.} \\ &= \frac{dc}{2\epsilon_0} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} |\mathbf{k}| e^{i\mathbf{k}\cdot\mathbf{r} - i\omega_{\mathbf{k}} t} (\hat{\mathbf{d}} - (\hat{\mathbf{k}} \cdot \hat{\mathbf{d}}) \hat{\mathbf{k}}) \left(e^{i\varphi} \frac{e^{i(\omega_{\mathbf{k}} - \omega)t} - 1}{\omega_{\mathbf{k}} - \omega} + e^{-i\varphi} \frac{e^{i(\omega_{\mathbf{k}} + \omega)t} - 1}{\omega_{\mathbf{k}} + \omega} \right) + \text{c.c.} \end{aligned} \quad (44)$$

It is hard to calculate the integral explicitly, but it can be observed that the non-zero frequency components are not restricted to ω , which is expected since the dipole starts to oscillate at $t = 0$ so the time translation symmetry is broken, so we do not have the strict $\omega_{\mathbf{k}} = \pm\omega$ relation, but as $\omega_{\mathbf{k}}$ approaches $\pm\omega$ the amplitude increases. In the $t \rightarrow \infty$ limit according to the standard procedure in QFT, we have

$$\frac{e^{i(\omega_{\mathbf{k}} - \omega)t} - 1}{\omega_{\mathbf{k}} - \omega} \rightarrow \frac{1}{\omega - \omega_{\mathbf{k}} + i0^+}, \quad \frac{e^{i(\omega_{\mathbf{k}} + \omega)t} - 1}{\omega_{\mathbf{k}} + \omega} \rightarrow -\frac{1}{\omega + \omega_{\mathbf{k}} + i0^+},$$

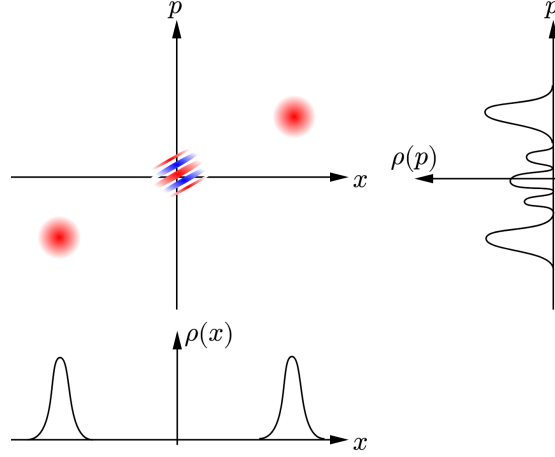


Figure 1: The Wigner function of a cat state, and the corresponding $\rho(x)$ and $\rho(p)$.

so

$$\langle \mathbf{E} \rangle(\mathbf{r}, t) = \frac{dc}{2\epsilon_0} \int \frac{d^3\mathbf{k}}{(2\pi)^3} |\mathbf{k}| e^{i\mathbf{k} \cdot \mathbf{r} - i\omega_{\mathbf{k}} t} (\hat{\mathbf{d}} - (\hat{\mathbf{k}} \cdot \hat{\mathbf{d}}) \hat{\mathbf{k}}) \left(e^{i\varphi} \frac{1}{\omega - \omega_{\mathbf{k}} + i0^+} - e^{-i\varphi} \frac{1}{\omega + \omega_{\mathbf{k}} + i0^+} \right) + \text{c.c.} \quad (45)$$

The time evolution of a cat state In the Schrödinger picture, describe the time evolution of a cat state's Wigner distribution, $\rho(x)$ and $\rho(p)$ in a single mode optical field.

Solution The Wigner function of a cat state have two peaks, corresponding to α and $-\alpha$, and around $\alpha = 0$ there are a series of “squeezed” peaks and valleys, corresponding to the interference between the two components. The scheme is roughly shown in Figure 1 on page 13. $\rho(x)$ can be obtained by integrating out the p variable, so when the two peaks are close in the x coordinate, $\rho(x)$ has strong interference effect, where the two peaks cannot be distinguished clearly due to the peaks created by interference between them; on the other hand, when the two peaks are far from each other in the x coordinate, $\rho(x)$ is simply a two-peak function. Replacing x with p in the last sentence and everything holds.

In a single mode optical field the time evolution is simply to rotate (α, α^*) with angular speed ω . Therefore, the time evolution of the Wigner function of a cat state is just to rotate Figure 1 on page 13 with angular speed ω . In this process, the distance between the two peaks on the x axis goes up and down periodically, so the $\rho(x)$ function transforms between a two-peak function and a multiple-peak function where the main peak occurs at $x = 0$, the peak value decreasing as $|x|$ increases, and so does $\rho(p)$. When $\rho(x)$ is a two-peak function $\rho(p)$ is a multiple-peak function, and vice versa. This is visualized in [3].

Three-photon state Consider a three-photon state

$$|\psi\rangle = b_1^\dagger b_2^\dagger b_3^\dagger |0\rangle. \quad (46)$$

Calculate the three-photon joint probability $P^{(3)}(\mathbf{r}_1, t_2; \mathbf{r}_2, t_2; \mathbf{r}_3, t_3)$. Analyze the form of the three-photon coherence $g^{(3)}(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3)$ when all the three modes are traveling waves with exactly the same frequency.

Solution Let η be the normalizing constant introduced by constants like \mathbf{d} or \hbar in the Fermi's golden rule. The joint probability is

$$\begin{aligned} & P^{(3)}(\mathbf{r}_1, t_2; \mathbf{r}_2, t_2; \mathbf{r}_3, t_3) \\ &= \eta^3 \langle E^-(\mathbf{r}_1, t_1) E^-(\mathbf{r}_2, t_2) E^-(\mathbf{r}_3, t_3) E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) \rangle \\ &= \eta^3 \langle 0 | b_3 b_2 b_1 E^-(\mathbf{r}_1, t_1) E^-(\mathbf{r}_2, t_2) E^-(\mathbf{r}_3, t_3) E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) b_1^\dagger b_2^\dagger b_3^\dagger | 0 \rangle. \end{aligned}$$

By Wick's theorem it can be reduced to the sum of products of two-point correlation functions. Note that since $E^+ \sim b$, it must contract with a creation operator on its right to create a non-zero two-point correlation function, or in other words all the three E^+ operators must contract

with $b_1^\dagger, b_2^\dagger, b_3^\dagger$. So we have

$$P^{(3)}(\mathbf{r}_1, t_2; \mathbf{r}_2, t_2; \mathbf{r}_3, t_3) = \eta^3 \times \sum \text{contractions between } b_1, b_2, b_3 \text{ and the three } E^- \text{'s} \\ \times \sum \text{contraction between the three } E^+ \text{'s and } b_1^\dagger, b_2^\dagger, b_3^\dagger,$$

and we immediately notice that again by Wick's theorem

$$\begin{aligned} & \sum \text{contraction between the three } E^+ \text{'s and } b_1^\dagger, b_2^\dagger, b_3^\dagger \\ &= \langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) b_1^\dagger b_2^\dagger b_3^\dagger | 0 \rangle \\ &= \langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) | \psi \rangle, \end{aligned}$$

and

$$\begin{aligned} & \sum \text{contractions between } b_1, b_2, b_3 \text{ and the three } E^- \text{'s} \\ &= \langle 0 | b_3 b_2 b_1 E^-(\mathbf{r}_1, t_1) E^-(\mathbf{r}_2, t_2) E^-(\mathbf{r}_3, t_3) | 0 \rangle \\ &= \langle \psi | E^-(\mathbf{r}_1, t_1) E^-(\mathbf{r}_2, t_2) E^-(\mathbf{r}_3, t_3) | 0 \rangle, \end{aligned}$$

so

$$P^{(3)}(\mathbf{r}_1, t_2; \mathbf{r}_2, t_2; \mathbf{r}_3, t_3) = \eta^3 \left| \langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) | \psi \rangle \right|^2. \quad (47)$$

When $t_1 = t_2 = t_3 = t$, it is just

$$P^{(3)}(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, t) = \eta^3 \left| \Psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, t) \right|^2, \quad (48)$$

where

$$\Psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, t) = \langle 0 | E^+(\mathbf{r}_3, t) E^+(\mathbf{r}_2, t) E^+(\mathbf{r}_1, t) | \psi \rangle \quad (49)$$

is the “three-photon wavefunction”.

Now we evaluate $\langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) | \psi \rangle$. Again by Wick's theorem

$$\begin{aligned} & \langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) | \psi \rangle \\ &= \sum_{(i,j,k) \in \text{Perm}(\{1,2,3\})} \langle 0 | E^+(\mathbf{r}_i, t_i) b_1^\dagger | 0 \rangle \langle 0 | E^+(\mathbf{r}_j, t_j) b_2^\dagger | 0 \rangle \langle 0 | E^+(\mathbf{r}_k, t_k) b_3^\dagger | 0 \rangle, \end{aligned}$$

where

$$\begin{aligned} \langle 0 | E^+(\mathbf{r}_m, t_m) b_n^\dagger | 0 \rangle &= \sum_i \mathcal{E}_i f_i(\mathbf{r}_m) \langle 0 | b_i e^{-i\omega_i t_m} b_n^\dagger | 0 \rangle \\ &= \mathcal{E}_n f_n(\mathbf{r}_m) e^{-i\omega_n t_m}. \end{aligned}$$

Therefore

$$\begin{aligned} & \langle 0 | E^+(\mathbf{r}_3, t_3) E^+(\mathbf{r}_2, t_2) E^+(\mathbf{r}_1, t_2) | \psi \rangle \\ &= \sum_{(i,j,k) \in \text{Perm}(\{1,2,3\})} \mathcal{E}_1 \mathcal{E}_2 \mathcal{E}_3 f_1(\mathbf{r}_i) f_2(\mathbf{r}_j) f_3(\mathbf{r}_k) e^{-i(\omega_1 t_i + \omega_2 t_j + \omega_3 t_k)}. \end{aligned} \quad (50)$$

Inserting this equation into (47), we have obtained an explicit expression of the three-photon joint probability. Particularly we have

$$\Psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, t) = \mathcal{E}_1 \mathcal{E}_2 \mathcal{E}_3 e^{-i(\omega_1 + \omega_2 + \omega_3)t} \sum_{(i,j,k) \in \text{Perm}(\{1,2,3\})} f_1(\mathbf{r}_i) f_2(\mathbf{r}_j) f_3(\mathbf{r}_k), \quad P^{(3)} = \eta^3 |\Psi|^2. \quad (51)$$

It can be observed that Ψ is just symmetrized direct projects of the three “single photon wave functions” f_1, f_2, f_3 .

The definition of three-photon coherence is

$$g(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3) = \frac{\langle E^-(\mathbf{r}_1, t) E^-(\mathbf{r}_2, t) E^-(\mathbf{r}_3, t) E^+(\mathbf{r}_3, t) E^+(\mathbf{r}_2, t) E^+(\mathbf{r}_1, t) \rangle}{\langle E^-(\mathbf{r}_1, t) E^+(\mathbf{r}_1, t) \rangle \langle E^-(\mathbf{r}_2, t) E^+(\mathbf{r}_2, t) \rangle \langle E^-(\mathbf{r}_3, t) E^+(\mathbf{r}_3, t) \rangle}. \quad (52)$$

We have already known the numerator. As for the denominator, by Wick's theorem we have

$$\begin{aligned} \langle E^-(\mathbf{r}_1, t) E^+(\mathbf{r}_1, t) \rangle &= \langle 0 | b_3 b_2 b_1 E^-(\mathbf{r}_1, t) E^+(\mathbf{r}_1, t) b_1^\dagger b_2^\dagger b_3^\dagger | 0 \rangle \\ &= \mathcal{E}_1^2 |\mathbf{f}_1(\mathbf{r}_1)|^2 + \mathcal{E}_2^2 |\mathbf{f}_2(\mathbf{r}_1)|^2 + \mathcal{E}_3^2 |\mathbf{f}_3(\mathbf{r}_1)|^2. \end{aligned} \quad (53)$$

Note that if E^+ contracts with b_i^\dagger then E^- must contract with b_i , or otherwise we will have $\langle b_j b_k^\dagger \rangle$ factors where $j \neq k$, which evaluates to zero.

When all the three modes are propagating waves with exactly the same frequency ω , we have

$$\mathbf{f}_i(\mathbf{r}) = e^{i\mathbf{k}_i \cdot \mathbf{r}} \mathbf{e}_i, \quad i = 1, 2, 3,$$

and

$$\mathcal{E}_1 = \mathcal{E}_2 = \mathcal{E}_3 = \mathcal{E} = \sqrt{\frac{\hbar\omega}{2\epsilon_0 V}},$$

and therefore

$$\langle E^-(\mathbf{r}_1, t) E^+(\mathbf{r}_1, t) \rangle = 3\mathcal{E},$$

and

$$\begin{aligned} & \langle E^-(\mathbf{r}_1, t) E^-(\mathbf{r}_2, t) E^-(\mathbf{r}_3, t) E^+(\mathbf{r}_3, t) E^+(\mathbf{r}_2, t) E^+(\mathbf{r}_1, t) \rangle \\ &= \left| \mathcal{E}^3 \sum_{(i,j,k) \in \text{Perm}(\{1,2,3\})} e^{i(\mathbf{k}_1 \cdot \mathbf{r}_i + \mathbf{k}_2 \cdot \mathbf{r}_j + \mathbf{k}_3 \cdot \mathbf{r}_k)} \mathbf{e}_1 \mathbf{e}_2 \mathbf{e}_3 \right|^2 \\ &= \mathcal{E}^6 \sum_{\sigma, \sigma' \in \text{Perm}(\{1,2,3\})} \prod_{i=1}^3 e^{i(\mathbf{k}_{\sigma(i)} - \mathbf{k}_{\sigma'(i)}) \cdot \mathbf{r}_i} \mathbf{e}_{\sigma(i)} \cdot \mathbf{e}_{\sigma'(i)}^*. \end{aligned}$$

Note that permuting b_1, b_2, b_3 is equivalent to permuting $\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3$, which justifies the third line. Putting everything together we have

$$g(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3) = \sum_{\sigma, \sigma' \in \text{Perm}(\{1,2,3\})} \prod_{i=1}^3 e^{i(\mathbf{k}_{\sigma(i)} - \mathbf{k}_{\sigma'(i)}) \cdot \mathbf{r}_i} \mathbf{e}_{\sigma(i)} \cdot \mathbf{e}_{\sigma'(i)}^*. \quad (54)$$

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