

Analytical Properties of Electrodynamics Response Functions by Prof. Kun Ding

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1 Review of last lectures

The frequency domain Maxwell's equations are

$$\begin{cases} \nabla \cdot \mathbf{D}(\mathbf{r}, \omega) = \rho(\mathbf{r}, \omega), \\ \nabla \times \mathbf{E}(\mathbf{r}, \omega) = i\omega \mathbf{B}(\mathbf{r}, \omega), \\ \nabla \cdot \mathbf{B}(\mathbf{r}, \omega) = 0, \\ \nabla \times \mathbf{H}(\mathbf{r}, \omega) = \mathbf{j}(\mathbf{r}, \omega) - i\omega \mathbf{D}(\mathbf{r}, \omega), \end{cases} \quad (1)$$

the time reversal of which is

$$\begin{cases} \nabla \cdot \mathbf{D}^*(\mathbf{r}, \omega) = \rho^*(\mathbf{r}, \omega), \\ \nabla \times \mathbf{E}^*(\mathbf{r}, \omega) = -i\omega \mathbf{B}^*(\mathbf{r}, \omega), \\ \nabla \cdot \mathbf{B}^*(\mathbf{r}, \omega) = 0, \\ -\nabla \times \mathbf{H}^*(\mathbf{r}, \omega) = \mathbf{j}^*(\mathbf{r}, \omega) - i\omega \mathbf{D}^*(\mathbf{r}, \omega), \end{cases} \quad (2)$$

2 The Kramers–Kronig relations

By the causality condition we have

$$\chi_e(\tau) = \frac{1}{2}(1 + \text{sgn}(\tau))\chi_e(\tau), \quad (3)$$

because $\chi_e(\tau) = 0$ for $\tau < 0$. So we have

$$\begin{aligned} \chi_e(\omega) &= \int d\tau \chi_e(\tau) e^{i\omega\tau} \\ &= \frac{1}{2} \int d\tau e^{i\omega\tau} (1 + \text{sgn}(\tau)) \chi_e(\tau) \\ &= \frac{1}{2} \int d\tau e^{i\omega\tau} \chi_e(\tau) + \frac{1}{2} \int d\tau \text{sgn}(\tau) e^{i\omega\tau} \times \int \frac{d\nu}{2\pi} \chi_e(\nu) e^{-i\nu\tau} \\ &= \frac{1}{2} \int d\tau e^{i\omega\tau} \chi_e(\tau) + \frac{1}{2} \int \frac{d\nu}{2\pi} \chi_e(\nu) \int d\tau \text{sgn}(\tau) e^{i(\omega-\nu)\tau}, \end{aligned}$$

or in other words

$$\int d\tau e^{i\omega\tau} \chi_e(\tau) = \chi_e(\omega) = \int \frac{d\nu}{2\pi} \chi_e(\nu) \int d\tau \text{sgn}(\tau) e^{i(\omega-\nu)\tau}. \quad (4)$$

so we are going to evaluate the Fourier transformation of the sign function.

We present a useful result ahead. Note that

$$\lim_{\epsilon \rightarrow 0^+} \int_a^b \frac{f(x)}{x \pm i\epsilon} dx = \mp i\pi \lim_{\epsilon \rightarrow 0^+} \int_a^b \frac{\epsilon/\pi}{x^2 + \epsilon^2} f(x) dx + \lim_{\epsilon \rightarrow 0^+} \int_a^b \frac{x^2}{x^2 + \epsilon^2} \frac{f(x)}{x} dx,$$

where $a < 0 < b$, and we have the **Plemelj formula**

$$\lim_{\epsilon \rightarrow 0^+} \frac{1}{x \pm i\epsilon} = \mp i\pi \delta(x) + \text{P} \frac{1}{x}. \quad (5)$$

This is a well known result in complex analysis, which has been widely used in QFT (expected, because it shows the analytic structure of the propagator of a particle). Now, formally, we have

$$\int dx e^{ikx} \text{sgn}(x) = \int_0^\infty dx e^{ikx} - \int_{-\infty}^0 dx e^{ikx}.$$

The two integrals on the RHS does not really make sense because they diverge. However, note that the sign function is usually multiplied to function quickly decaying as $|x| \rightarrow \infty$, and we can “spare” a part of the decaying factor to the sign function, and we therefore introduce an imaginary part into the frequency, i.e.

$$\begin{aligned} \int dx e^{ikx} \text{sgn}(x) &= \lim_{\epsilon \rightarrow 0^+} \left(\int_0^\infty dx e^{i(k+i\epsilon)x} - \int_{-\infty}^0 dx e^{i(k-i\epsilon)x} \right) \\ &= \lim_{\epsilon \rightarrow 0^+} \left(\frac{i}{k+i\epsilon} + \frac{i}{k-i\epsilon} \right) \\ &= \lim_{\epsilon \rightarrow 0^+} \left(\frac{k^2}{k^2 + \epsilon^2} \frac{2i}{k} \right) = \frac{2i}{k}. \end{aligned}$$

The last line gives the Fourier transformation of the sign function:

$$\int dx e^{ikx} \text{sgn}(x) = \frac{2i}{k}. \quad (6)$$

Substituting this equation into (4), we have

$$\chi_e(\tau) = \int \frac{d\nu}{2\pi} \chi_e(\nu) \times \frac{2i}{\omega - \nu},$$

or in other words

$$\text{Re } \chi_e(\omega) = -\frac{1}{\pi} \text{P} \int_{-\infty}^\infty \frac{\text{Im } \chi_e(\nu)}{\omega - \nu} d\nu, \quad \text{Im } \chi_e(\omega) = \frac{1}{\pi} \text{P} \int_{-\infty}^\infty \frac{\text{Re } \chi_e(\nu)}{\omega - \nu} d\nu. \quad (7)$$

These are called the **Kramers–Kronig relations**.

It is obvious that the proof of (7) does not involve anything other than that $\chi_e(\tau)$ satisfies the law of causality, and that it is analytic enough. The Kramers–Kronig relations, therefore, hold for every reasonable definition of response functions.

In optics people usually do not like negative frequencies. Note that for $\chi_e(\omega)$ we have

$$\chi_e(-\omega) = \chi_e^*(\omega), \quad (8)$$

or

$$\text{Im } \chi_e(\omega) = -\text{Im } \chi_e(-\omega),$$

and therefore we have

$$\begin{aligned} \text{Re } \chi_e(\omega) &= \frac{1}{\pi} \text{P} \int_{-\infty}^\infty \left(\frac{\nu \text{Im } \chi_e(\nu)}{\nu^2 - \omega^2} + \frac{\omega \text{Im } \chi_e(\nu)}{\nu^2 - \omega^2} \right) d\nu \\ &= \frac{1}{\pi} \text{P} \int_{-\infty}^\infty \frac{\nu \text{Im } \chi_e(\nu)}{\nu^2 - \omega^2} d\nu, \end{aligned}$$

so now we have a more widely used equation concerning only the positive frequencies:

$$\text{Re } \chi_e(\omega) = \frac{2}{\pi} \text{P} \int_0^\infty \frac{\nu}{\nu^2 - \omega^2} \text{Im } \chi_e(\nu) d\nu. \quad (9)$$

This equation holds if the conditions for the Kramers–Kronig relations hold and we have (8). The relation (9) is used frequently in experiments, where the absorption spectrum can be measured easily, and from which we can get the real part of the response. Similarly we have

$$\text{Im } \chi_e(\omega) = -\frac{2}{\pi} \text{P} \int_0^\infty \frac{\omega}{\nu^2 - \omega^2} \text{Re } \chi_e(\nu) d\nu. \quad (10)$$

Note that this time the remaining term is $\omega \text{Re } \chi_e(\nu)$ instead of $\nu \text{Re } \chi_e(\nu)$, because we have

$$\text{Re } \chi_e(\omega) = \text{Re } \chi_e(-\omega),$$

and therefore $\nu \text{Re } \chi_e(\omega)$ terms cancel with each other.

3 The Kramers–Kronig relations from analyticity

It is a well known fact that a reasonable response function’s frequency domain form has no singularities on the upper complex plane, because we have

$$\chi(\tau) = \frac{1}{2\pi} \int d\tau e^{-i\omega\tau} \chi(\omega),$$

and if $\chi(\omega)$ has a pole ω_0 on the upper plane, the pole contributes a term into $\chi(\tau)$ which is like

$$\chi(\tau) \sim e^{-i(\text{Re } \omega_0 + i \text{Im } \omega_0)\tau} \propto e^{\text{Im } \omega_0 \tau},$$

so the response *increases* as time goes by, which is ridiculous.

What we want to do in this section is to show that the Kramers–Kronig relations can also be derived from this analytic condition. This, actually, is much easier than the proof in Section 2.

Since there is no singularity on the upper plane, we have

$$\oint d\nu \frac{\chi(\nu)}{\omega - \nu} = 0,$$

where we set the contour to be

4 The analytic structure of electromagnetic response functions

It should be noticed that the Kramers–Kronig relations shown in the last two sections may *fail* for some materials. We often include the ohmic currents into χ_e , which introduces a divergent imaginary part when $\omega \rightarrow 0$. Eq. (7.18) in [the optics note](#) tells us that the current introduces a $i\sigma/\omega$ term in ϵ , and since we have

$$1 + \chi_e = \frac{\epsilon}{\epsilon_0}, \quad (11)$$

a $i\sigma/\epsilon_0\omega$ term is introduced in χ_e . We see that $\chi_e(\omega)$ for metals have an additional pole at the origin, which is not considered when deriving the Kramers–Kronig, where we simply assume there is no singularity on the real axis. In this case, we have a modified version of (10):

$$\text{Im } \chi_e(\omega) = \frac{\sigma}{\epsilon_0\omega} - \frac{2}{\pi} \text{P} \int_0^\infty \frac{\omega}{\nu^2 - \omega^2} \text{Re } \chi_e(\nu) d\nu. \quad (12)$$

This formula can be proved by design a contour walking around the origin. Another way to think of the problem is to consider χ_e *without* conductivity, i.e. we analyze the structure of $\chi_e - i\sigma/\epsilon_0\omega$. It can be easily verified that (9) does not need modification.

From the analytic properties of the electromagnetic response functions we can also have some insight about how much light can a material film absorbs. First we introduce the **superconvergence theorem**. Suppose we have

$$g(y) = \text{P} \int_0^\infty \frac{f(x)}{y^2 - x^2} dx \quad (13)$$

and $f(x)$ decays at least as fast as $(x \log x)^{-1}$, then we have

$$\int_0^\infty f(x) dx = \lim_{y \rightarrow \infty} (y^2 g(y)). \quad (14)$$

Applying this theorem to χ_e . By (9), we can let

$$f(\nu) = \nu \text{Im } \chi_e(\nu)$$

and

$$g(\omega) = -\frac{\pi}{2} \text{Re } \chi_e(\omega),$$

so we have

$$\int_0^\infty \nu \operatorname{Im} \chi_e(\nu) d\nu = \lim_{\omega \rightarrow \infty} \left(-\frac{\pi}{2} \omega^2 \operatorname{Re} \chi_e(\omega) \right). \quad (15)$$

We have already shown in Eq. (6.29) in [the solid state physics note](#) that

$$1 + \chi_e = \epsilon_r = 1 - \frac{\omega_p^2}{\omega^2}$$

when ω is large, where $\omega_p \propto n_e$ according to Eq. (6.28) in [the solid state physics note](#). Therefore, from (15) we have

$$\int_0^\infty \nu \operatorname{Im} \chi_e(\nu) d\nu = \lim_{\omega \rightarrow \infty} \left(-\frac{\pi}{2} \omega^2 \operatorname{Re} \chi_e(\omega) \right) = \lim_{\omega \rightarrow \infty} \left(-\frac{\pi}{2} \omega^2 \times \frac{\omega_p^2}{\omega^2} \right) = \frac{\pi}{2} \omega_p^2,$$

so thus we have

$$\omega_p^2 = \frac{2}{\pi} \int_0^\infty \nu \operatorname{Im} \epsilon_r(\nu) d\nu. \quad (16)$$

This is called the **Thomas-Reiche-Kuhn sum rule (TRK sum rule)**. It is the electrodynamics version of the f-sum rule. Since ω_p is proportion to the number of free electrons, and the RHS of (16) is proportion to the absorption in the material, we conclude that the absorption is bounded roughly by the free electron number.

Repeating this procedure for (10) or more generally for (12), this time setting

$$f(\nu) = \operatorname{Re} \chi_e(\nu)$$

and

$$g(\omega) = \frac{\pi}{2} \left(\frac{1}{\omega} \operatorname{Im} \chi_e(\omega) - \frac{\sigma}{\epsilon_0 \omega^2} \right),$$

we have

$$\int_0^\infty \operatorname{Re} \chi_e(\nu) d\nu = \frac{\pi}{2} \lim_{\omega \rightarrow \infty} \left(\omega \operatorname{Im} \chi_e(\omega) - \frac{\sigma}{\epsilon_0} \right) = -\frac{\pi \sigma}{2 \epsilon_0},$$

where the second equation holds because the response decays quickly enough as $\omega \rightarrow \infty$. The result is

$$\int_0^\infty \operatorname{Re} \chi_e(\nu) d\nu = \int_0^\infty \operatorname{Re}(\epsilon_r(\nu) - 1) d\nu = -\frac{\pi \sigma}{2 \epsilon_0}. \quad (17)$$

If the material is insulator, $\sigma = 0$, and therefore we have

$$\overline{\operatorname{Re} \epsilon_r} = 1. \quad (18)$$

Sum rules (16) and (18) are both substantial constraints on electromagnetic responses. The Kramers–Kronig relations can be simply seen as the redundancy of the frequency domain response functions, but by just adding some information like “ $\omega_p \propto n_e$ ” or “ χ_e for metals have additional poles at the origin”, the consequences of the Kramers–Kronig relations impose real and substantial constraints on the response functions.

5 Poor man’s linear response theory

A thorough introduction of the linear response theory in condensed matters can be found in Section 5.4.2 in [the solid state physics note](#). Here to make things easier, we use a time-dependent orbital-free DFT approach (formulated in a fluid dynamics way) to demonstrate basic ideas of linear response, and how momentum dependence of electromagnetic response functions arises.

Let $e = -|e|$ be the electric charge carried by one electron. The basic equations are

$$m_e n \frac{d\mathbf{v}}{dt} = ne(\mathbf{E} + \mathbf{v} \times \mathbf{B}) - n \nabla \left(\frac{\delta G[n]}{\delta n} \right), \quad (19)$$

and

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{v}) = 0. \quad (20)$$

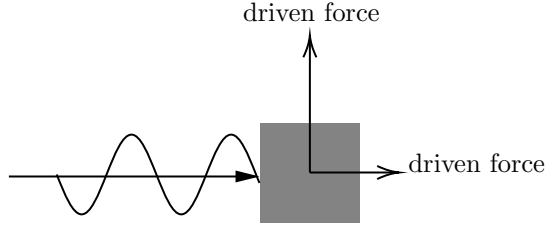


Figure 1: Matter volume in electromagnetic waves

Here $G[n]$ is the density functional of the electron gas with Coulomb interaction, and the total energy functional of the system is

$$H[n] = G[n] + \int d^3\mathbf{r} n \frac{(\mathbf{p} - e\mathbf{A})^2}{2m_e} + e \int d^3\mathbf{r} n \phi, \quad (21)$$

and (19) and (20) can be derived from the total energy.

Theoretically (20) and (19) can be exact for processes of which the length scales are significantly greater than, say, the lattice constant, where the two equations are actually fluid dynamics equations which work well for systems with well defined flowing densities, and the partial derivative of the density functional gives forces acting on a volume of the electronic fluid. This is, nonetheless, practically impossible, for two reasons. (a) The precision of energy functional for orbital-free DFT is unimaginably poor, much less accurate than Kohn-Sham DFT. (b) Some systems simply cannot be described in terms of electron density in a natural way. For example, consider a metal in a strong magnetic field. No one expects the Landau energy levels can be described in a satisfactory way in terms of electron density. Though in theory any expectation can be rewritten as a functional of the electron density, the form of the functional may be strange.

This does not mean that (19) and (20) are useless. At least they can be used to capture qualitatively some physics when the concept of “flowing electronic liquid” still works. Also, since more accurate *ab initio* methods are much more time-consuming, (19) and (20) can be an efficient way to get some first impression of the system under investigation. Another advantage of the time dependent orbital-free DFT method is that it can be used to check the *dynamic* response of materials. Because KSDFT - not to mention field theoretic methods - is quite slow, the way mainstream *ab initio* packages calculate the electromagnetic response functions is to observe the behavior of the material in a static background field. This, for example, is unable to detect ultrafast behaviors, where the time scale of the time evolution of the electrodynamics field is comparable with the time scale of the response of the material. Another example where the background field method is unacceptable is the force acting on a volume of matter. consider Figure 1 on page 5. If both the matter and the light is simulated *ab initio*, then the electromagnetic waves will give the matter volume a rightward force. If, however, the light is seen as a background, the volume can only feel an up-down force. When the characteristic scale of the problem is $\lesssim 1$ nm, the background field method may be a good choice, but when the characteristic scale is between ~ 1 nm and ~ 10 nm, where both the wave nature of the electromagnetic field and the details of the matter are important, the orbital-free DFT and the fluid dynamics formulation are still a useful tool. Standard KSDFT does not work well in this region, and coarse-graining of the matter does not work.

Now we make the linear response approximation, i.e. let

$$n = n_0 + n_1, \quad \mathbf{E} = \mathbf{E}_0 + \mathbf{E}_1, \quad \mathbf{B} = \mathbf{B}_1, \quad (22)$$

where n_0 is a stable solution for (19) and (19) and n_1 is the linear response, \mathbf{E}_0 and \mathbf{B}_0 the stable component of the electric field and the magnetic field, respectively. We ignore the static magnetic field because it creates Landau energy levels, destroying the fluid dynamics picture as we have just mentioned. The ground state density n_0 can be solved from

$$e\mathbf{E}_0 = \nabla \left(\frac{\delta G[n]}{\delta n} \right), \quad (23)$$

which is the stable version of (19). Once after n_0 is found, (19) reads

$$m_e(n_0 + n_1) \frac{d\mathbf{v}}{dt} = e(n_0 + n_1)(\mathbf{E}_1 + \mathbf{v} \times \mathbf{B}_1) - n_1 \left(\nabla \left(\frac{\delta G[n]}{\delta n_0} \right) + \nabla \left(\frac{\delta G[n]}{\delta n_1} \right) \right) - n_0 \nabla \left(\frac{\delta G[n]}{\delta n_1} \right).$$

Since we only consider the linear response,