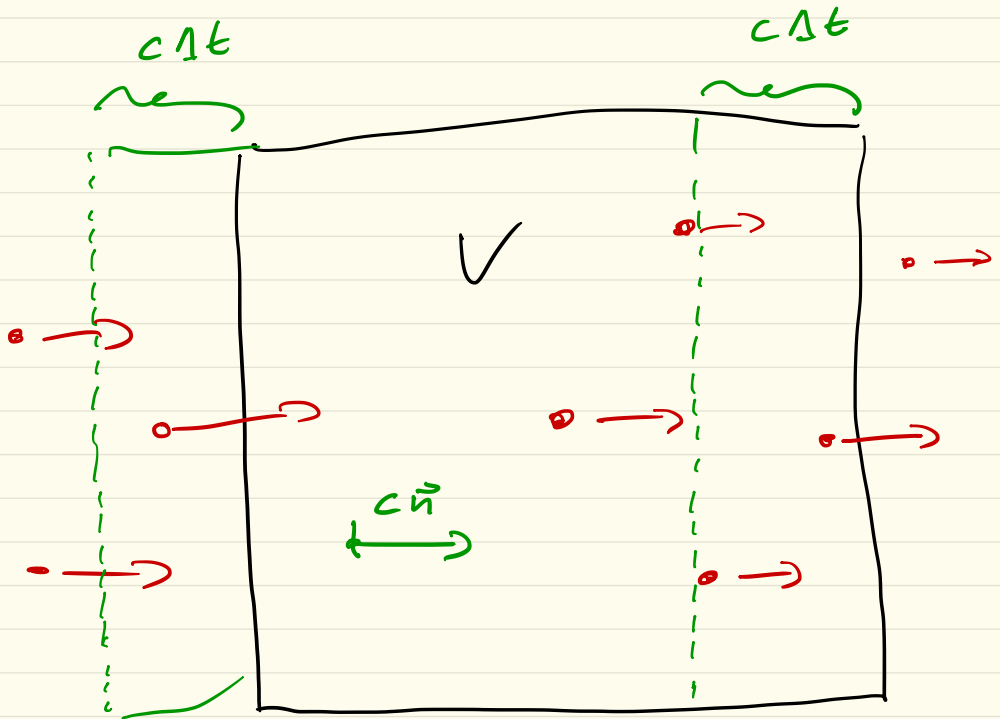


Reynolds transport theorem



photons traveling with velocity $c \vec{n}$

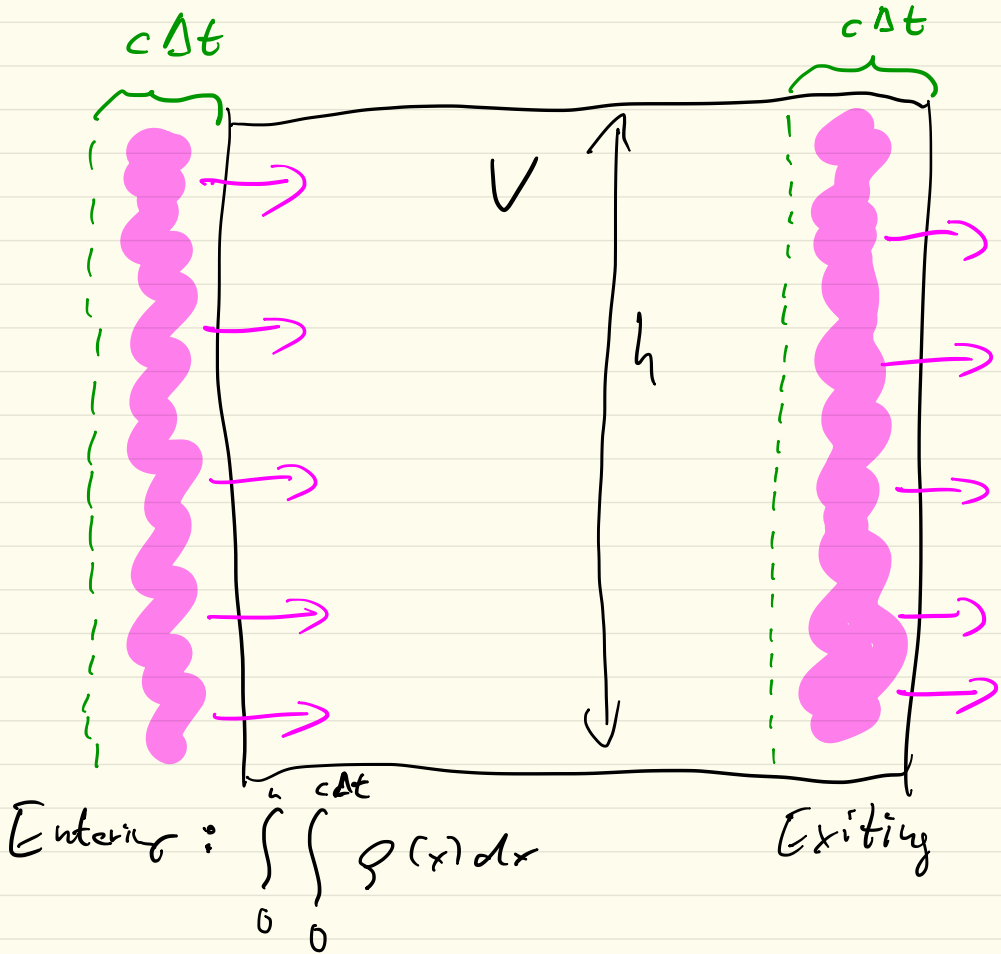
Change of number of photons in V
overtime interval $[t, t + \Delta t]$

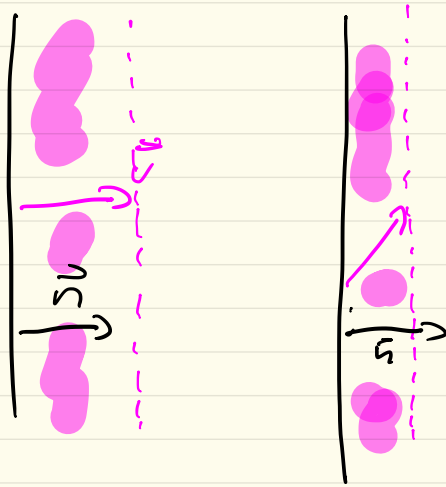
= number of photons entering
- number of photons leaving

Transition to density ρ

- Number of photons in volume V :

$$\# = \int_V \rho(x) dx$$





Travel speed : c

Travel direction : \vec{d}

Outer normal to the domain : \vec{n}

Width of the strip : $c \Delta t \vec{n} \cdot \vec{d}$

- integrate in vertical direction

- Limit for $\Delta t \rightarrow 0$

- [density and domain are suff. smooth]

$$\frac{d}{dt} \int_V \rho(x) dx = - \int_{\partial V} \rho(x) c \vec{n} \cdot \vec{d} ds$$

Physical principle

$$\frac{d}{dt} \int_V \rho(x) dx = - \int_{\partial V} \rho(s) \vec{d} \cdot \vec{n} ds$$

Gauss-Theorem

$$\begin{aligned} \vec{v} \cdot \vec{u} &= \partial_x u_x \\ &\quad + \partial_y u_y \end{aligned}$$

$$\int_V \vec{v} \cdot \vec{u} = \int_{\partial V} \vec{u} \cdot \vec{n} ds$$

$$\begin{aligned} \int_V \frac{\partial}{\partial t} \rho(x) dx &= - \int_{\partial V} \rho(s) \vec{d} \cdot \vec{n} ds \\ &= - \int_V \nabla \cdot (\rho(x) \vec{d}) dx \end{aligned}$$

Differential equation

$$\frac{1}{c} \partial_t \rho + \nabla \cdot (\rho \vec{d}) = 0$$

Different Forms of the flux term
conservative form

$$\vec{V} \cdot (\varrho \vec{d}) = \partial_x (\varrho d_x) + \partial_y (\varrho d_y)$$

$$= \partial_x \varrho d_x + \partial_y \varrho d_y \\ + \varrho \partial_x d_x + \varrho \partial_y d_y$$

$$= (\underbrace{\vec{d} \cdot \nabla}_{\text{directional derivative}}) \varrho + \varrho \underbrace{\nabla \cdot \vec{d}}_{\text{radiation: } \vec{d} \equiv \text{const}}$$

$$= (\vec{d} \cdot \nabla) \varrho \quad \text{advective form}$$

Transport of a density function ϱ by a
(divergence free) vector field \vec{v}

$$\partial_t \varrho + (\vec{v} \cdot \nabla) \varrho = 0$$

Application to radiation

1. Consider density in phase space

$$\mathcal{S}(\vec{x}, \vec{n}_\Omega, \nu)$$

2. Apply transport theorem to all directions

$$\frac{1}{c} \partial_t \mathcal{S}(\vec{x}, \vec{n}_\Omega, \nu) + (\vec{n}_\Omega \cdot \nabla) \mathcal{S}(\vec{x}, \vec{n}_\Omega, \nu) = 0$$

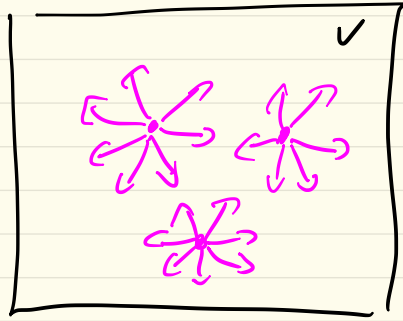
3. Convert photon density to energy density
= specific intensity

$$I_\nu(\vec{x}, \vec{n}_\Omega) = h\nu \mathcal{S}(\vec{x}, \vec{n}_\Omega, \nu)$$

Radiation transport equation in vacuum

$$\frac{1}{c} \partial_t I_\nu + (\vec{n}_\Omega \cdot \nabla) I_\nu = 0$$

Emission



- Let there be a source of photons releasing particles in V at a certain rate $\eta/\Delta t$

Then, # photons in V at $t + \Delta t$ - # photons in V at $t = \eta$

- Represent this source by a source density in space time

$$q_s(\vec{x}, t)$$

Then, with no movement

$$\int_V \rho(x + \Delta t) dx = \int_V \rho(x) dx + \int_V \int_t^{t+\Delta t} q_s(\vec{x}, s) ds dx$$

- Since this applies to every volume V , we derive the pointwise eqn.

$$\rho(x + \Delta t) = \rho(x) + \int_t^{t+\Delta t} q_s(\vec{x}, s) ds$$

Finally, we convert this to differential form

$$\partial_t g(\vec{x}) = q_g(\vec{x})$$

Application to radiation

- rescaling

$$g(\vec{x}, \vec{n}_x, \nu) \rightarrow I_\nu(\vec{x}, \vec{n}_x)$$

$$q_g(\vec{x}, \vec{n}_x, \nu) \rightarrow q_\nu(\vec{x}, \vec{n}_x, \nu)$$

$$\frac{1}{c} \partial_t I_\nu = q_\nu$$

Radiation transport in an emitting atmosphere
(unphysical)

$$\frac{1}{c} \partial_t I_\nu + \vec{n}_x \cdot \vec{\nabla} I_\nu = q_\nu$$