### 1 Motion in one dimension

Proposition 1.1. Let .... Then, it is true that

$$\vec{F}^{ext} = m\vec{a} + (\vec{v} - \vec{u})\dot{m} \tag{1}$$

or which is equivalent,

$$\vec{F}^{ext} + \dot{m}\vec{u} = \dot{\vec{p}} \tag{2}$$

### $\mathbf{2}$ Oscillations

**Proposition 2.1.** Let be the following differential equation

$$\ddot{x} + \omega_0^2 x = 0, (3)$$

with the initial value condition of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the general solution is

$$x(t) = x_0 \cos \omega_0 t + \frac{v_0}{\omega_0} \sin \omega_0 t, \tag{4}$$

or, which is equivalent,

$$x(t) = A\cos\left[\omega_0 t + \phi_0
ight], \qquad A = \sqrt{x_0^2 + \left(rac{v_0}{\omega_0}
ight)^2}, \qquad egin{array}{c} \mathbf{5} & \mathbf{Coupled \ oscillations} \ \mathbf{2} \\ \phi_0 = -\arctanrac{v_0}{\omega_0 x_0}. \\ \mathbf{6} & \mathbf{Rotations} \end{array}$$

**Definition 2.1.** Let U(x) be a potential function of class  $C^2(\mathbb{R})$ . Then, we say  $x_0$  is a point of stable equi*librium* if U has a maxima in  $x_0$ .

**Proposition 2.2.** Let be the following differential equation

$$\ddot{x} + 2\beta \dot{x} + \omega_0^2 x = 0,\tag{6}$$

with the initial value conditions of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the general solution is

$$v(0) = v_0$$
. Then, the general solution is
$$v_0 + \beta r_0$$

if  $\beta < \omega_0$ ,

$$x(t) = e^{-\omega_0 t} \left[ x_0 + (x_0 \omega_0 + v_0)t \right]$$
 (8)

if  $\beta = \omega_0$ , and

$$x(t) = \frac{x_0(\bar{\omega} - \beta) - v_0}{2\bar{\omega}} e^{-(\beta + \bar{\omega})t} + \frac{x_0(\bar{\omega} + \beta) + v_0}{2\bar{\omega}} e^{-(\beta - \bar{\omega})t} \text{ are } in\bar{v} \bar{a}\bar{r} i \sqrt{\hbar^2 u \bar{n} dv^2} \text{ Galileo transformations, then } \mathbf{E} = 0$$

Proposition 2.3. Let be the following differential equation

$$\ddot{x} + 2\beta \dot{x} + \omega_0^2 x = f(t) = f_0 \cos[\omega t + \psi_0], \tag{10}$$

with the initial value conditions of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the particular solution is

$$x_p(t) = A\cos\left[\omega t + \psi_0 - \phi_0\right], \qquad A = \frac{f_0}{\sqrt{\left(\omega_0^2 - \omega^2\right)^2 + 4\beta t \sin^2 \theta}} \frac{\text{Lemma 8.2. Let } f : \mathbb{R}^4_{2\beta\overline{\omega}} \to \mathbb{R}^4 \text{ be a Lorentz transformation} \phi_0 \text{Thenk than ps forms } lines that are not confidence in hyperplanes of the form } t = \text{ctt to lines that}$$

$$(11) \quad \text{are not contained in hyperplanes of the form } t' = \text{ctt}.$$

### 3 Central forces

**Definition 3.1.** Central force

$$\vec{F}(\vec{r}) = f(r)\vec{e}_o \tag{12}$$

**Proposition 3.1.** All central forces are conservatives.

Proposition 3.2. The angular momentum with respect the origin is conserved.

$$\dot{\vec{L}} = \vec{0} \tag{13}$$

Proposition 3.3. The areal velocity is constant.

$$\frac{dA}{dt} = \frac{L}{2m} = ctt \tag{14}$$

Theorem 3.4 (Bertrand's Theorem). The only central potentials where every bounded orbit is closed are:

$$U(r) = -\frac{k}{r}, \qquad U(r) = \frac{k}{2}r^2, \qquad k > 0$$
 (15)

## Coupled oscillations 1

## Dynamics of rigid body

**Proposition 7.1.** The vector  $\Omega$  is independent on the origin of the system S.

**Proposition 7.2.** The energy of the rigid body is an invariant scalar under change of basis.

# Special relativity

 $x(t) = e^{-\beta t} \left[ x_0 \cos \tilde{\omega} t + \frac{v_0 + \beta x_0}{\tilde{\omega}} \sin \tilde{\omega} t \right], \qquad \tilde{\omega} = \sqrt{\omega_0^2 \frac{\mathbf{Theorem 8.1.}}{\beta^2}} \mathbf{8.1.} \quad \text{If Maxwell's equations}$ 

$$\langle \mathbf{\nabla}, \mathbf{E} \rangle_I = \frac{\rho}{\epsilon_0},\tag{16}$$

$$\langle \mathbf{\nabla}, \mathbf{B} \rangle_I = 0, \tag{17}$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t},\tag{18}$$

$$\nabla \times \mathbf{B} = \mu_0 \left( \mathbf{J} + \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right). \tag{19}$$

(9)  $\mathbf{B} = \mathbf{0}$ .

**Definition 8.1** (Reference system). We define a reference system S as a set of three axis and one origin over which we have determined an orientation. We will suppose we have selected a unit of length and that in each point a in the immobile space with respect the axis there is a clock  $q_a$  such that the clocks  $q_a$  and  $q_b$ corresponding to two different points a and b immobile with respect these axis are synchronized

**Lemma 8.3.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation. Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation. Then, f transforms planes that are not contained in hyperplanes of the form t = ctt to planes that are not contained in hyperplanes of the form t' = ctt.

**Lemma 8.4.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation. Then, f transforms hyperplanes that are not of the form t = ctt to hyperplanes that are not of the form t' = ctt.

**Theorem 8.5.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation. Then, f is an affine transformation.

**Theorem 8.6.** Let S, S' be two inertial reference systems. We can make orthogonal changes (isometries) of axis to S and S' and a change of origin of time such thate the Lorentz transformation has the form of the equation ??.

**Lemma 8.7.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation and  $r \subseteq \mathbb{R}^4$  a line with a timelike direction vector. Then, f transforms r to a line with a timelike  $direction\ vector.$ 

**Lemma 8.8.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation and V and admissible plane (hyperplane). Then, f(V) is an admissible plane (hyperplane).

**Theorem 8.9.** Let  $f: \mathbb{R}^4 \longrightarrow \mathbb{R}^4$  be a Lorentz transformation. Then, f is an affine transformation.

**Theorem 8.10.** (Lorenz transformation) Let S, S' be two reference systems with the same origin such that S' moves with a constant velocity  $\mathbf{v} = v\mathbf{e}_x$ . Then,

$$P_{\mathbf{s}'} = \Lambda P_{\mathbf{s}} \Leftrightarrow P_{\mathbf{s}'}^{\nu} = \Lambda_{\mu}^{\nu} P_{\mathbf{s}}^{\mu}, \qquad \Lambda = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \mathbf{v}' = \frac{1}{1 - \langle \boldsymbol{\beta}_{v}, \boldsymbol{\beta}_{2} \rangle_{I}} \left[ \frac{\mathbf{v}}{\gamma_{w}} - \mathbf{w} + \frac{1}{c^{2}} \frac{\gamma_{w}}{1 + \gamma_{w}} \langle \mathbf{v}, \mathbf{w} \rangle_{I} \mathbf{w} \right], \tag{25}$$

**Proposition 8.11.** Let  $\mathbf{R} \in L$  be a vector and  $a \in \mathbb{R}$ a scalar. Then,  $\|a\mathbf{R}\|_m = |a| \|\mathbf{R}\|_m$ .

**Proposition 8.12.** Every subspace W of L is either timelike, spacelike, or lightlike. Besides,

- 1. S is timelike  $\Leftrightarrow W^{\perp}$  is spacelike.
- 2. S is spacelike  $\Leftrightarrow W^{\perp}$  is timelike.
- 3. W is lightlike  $\Leftrightarrow W^{\perp}$  is lightlike.

Proposition 8.13. Two orthogonal vectors different from zero and non spacelike are necessarily lightlike and collinear. In particular, there is not a subspace of dimension 2 where  $\langle , \rangle$  is null.

**Proposition 8.14.** Let  $\mathbf{R}_1, \mathbf{R}_2 \in T$  be two timelike vectors. Then, the following statements are true.

- 1.  $|\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m| \ge ||\mathbf{R}_1||_m ||\mathbf{R}_2||_m$ , and the equality is equivalent to both vectors being collinear.
- 2.  $\mathbf{R}_1, \mathbf{R}_2$  are in the same time cone  $(C_+ \text{ or } C_-)$  if and only if  $\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m < 0$ . In this case,

(a) There is a unique  $\varphi \in \mathbb{R}$  such that

$$\cosh \varphi = -\frac{\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m}{\|\mathbf{R}_1\|_m \|\mathbf{R}_2\|_m}.$$
 (21)

We call this  $\varphi$  the hyperbolic angle.

(b) 
$$\|\mathbf{R}_1\|_m + \|\mathbf{R}_2\|_m \le \|\mathbf{R}_1 + \mathbf{R}_2\|_m$$
.

Proposition 8.15. The Lorentz-Minkowski metric (using the proper orthonormal basis) can be expressed by the bilinear form  $\eta$ 

$$\begin{pmatrix}
-1 & 0 & 0 & 0 \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1
\end{pmatrix}.$$
(22)

**Proposition 8.16.** The transformation in  $\mathbb{M} = \mathbb{R}^4$ (using the proper orthonormal basis) can be expressed by the matrix  $\Lambda$ 

$$\begin{pmatrix}
\gamma & -\gamma\beta & 0 & 0 \\
-\gamma\beta & \gamma & 0 & 0 \\
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1
\end{pmatrix}.$$
(23)

**Proposition 8.17.** The Lorenz-Minkowski metric is invariant under Lorentz transformations.

**Proposition 8.18.** Let S, S' be two inertial reference systems such that the velocity of S' is  $\mathbf{w} = w\mathbf{e}_x$  with

$$v_{\mathbf{s'}}^1 = \frac{v_{\mathbf{s}}^1 - w}{1 - \beta_v \beta_w}, \qquad v_{\mathbf{s'}}^2 = \frac{1}{\gamma_w} \frac{v_{\mathbf{s}}^2}{1 - \beta_v \beta_w}, \qquad v_{\mathbf{s'}}^3 = \frac{1}{\gamma_w} \frac{v_{\mathbf{s}}^3}{1 - \beta_v \beta_w}, \tag{24}$$

$$\mathbf{v}' = \frac{1}{1 - \langle \boldsymbol{\beta}_v, \boldsymbol{\beta}_2 \rangle_I} \left[ \frac{\mathbf{v}}{\gamma_w} - \mathbf{w} + \frac{1}{c^2} \frac{\gamma_w}{1 + \gamma_w} \langle \mathbf{v}, \mathbf{w} \rangle_I \mathbf{w} \right],$$
(25)

$$\mathbf{v} = \frac{1}{1 + \langle \boldsymbol{\beta}_{v'}, \boldsymbol{\beta}_{w} \rangle_{I}} \left[ \frac{\mathbf{v'}}{\gamma_{w}} + \mathbf{w} + \frac{1}{c^{2}} \frac{\gamma_{w}}{1 + \gamma_{w}} \langle \mathbf{w'}, \mathbf{w} \rangle_{I} \mathbf{v} \right].$$
(26)

**Proposition 8.20.** Let p be a particle moving along a curve  $\Gamma \subseteq \mathbb{R}^4$  with a 4-velocity **U**. Then,

$$\mathbf{U} = (\gamma c, \gamma \mathbf{v}), \qquad -\langle \mathbf{U}, \mathbf{U} \rangle_m = c^2.$$
 (27)

Proposition 8.21. Let p be a particle of velocity U and acceleration **A**. Then,  $\langle \mathbf{A}, \mathbf{U} \rangle_m = 0$ .

**Proposition 8.22.** Let p be a particle of 4-momentum P. Then,

$$\mathbf{P} = (E/c, \mathbf{p}) = (\gamma mc, \gamma m\mathbf{v}). \tag{28}$$

Theorem 8.23.

$$E^2 = p^2 c^2 + m^2 c^4. (29)$$

**Proposition 8.24.** There are three possible cases: stationary particle with mass, moving particle with mass, particle with no mass.

$$E = mc^2$$
,  $E^2 = m^2c^4 + ||p||^2c^2$ ,  $E = pc$  (30)

**Theorem 8.25.** (Work-Energy theorem)

$$W = \Delta E. \tag{31}$$

**Theorem 8.26.** Let p be a particle of velocity  $\mathbf{v}$ . Then, the kinetic energy is obtained by the expression

$$T = \int_{\Gamma} \langle \mathbf{F}, d\mathbf{r} \rangle_{I} = (\gamma(\dot{\mathbf{r}}) - 1)mc^{2}$$
 (32)

Theorem 8.27 (Compton scattering).

$$\Delta \lambda = \frac{h}{mc} (1 - \cos \theta). \tag{33}$$

Theorem 8.28 (Center of momentum). Let be a system of particles  $p_1, \ldots, p_n$  with energies  $E_1, \ldots, E_n$  and momentum  $\mathbf{p}_1, \dots, \mathbf{p}_n$ . Then, the center of momentum system has a velocity determined by the expression

$$\mathbf{v}_{\rm cp} = \frac{1}{E_t} \sum_{i=1}^n \|\mathbf{p}_i\|^2 c^2.$$
 (34)

**Proposition 8.29.** Let p be a particle of mass m moving along a curve  $\Gamma \subseteq \mathbb{R}^4$  with a velocity  $\mathbf{v}$  and acceleration a. Then,

$$\langle \mathbf{f}, \mathbf{v} \rangle_I = \frac{\mathrm{d}E}{\mathrm{d}t} = m\gamma^3 \langle \mathbf{v}, \mathbf{a} \rangle_I, \qquad \frac{\mathrm{d}\mathbf{p}}{\mathrm{d}t} = m\gamma\mathbf{a} + m\frac{\mathrm{d}\gamma}{\mathrm{d}t}\mathbf{v},$$
(35)

$$\mathbf{F} = \left(\frac{\gamma}{c} \frac{\mathrm{d}E}{\mathrm{d}t}, \gamma \mathbf{f}\right) = \left(\frac{m\gamma^4}{c} \langle \mathbf{v}, \mathbf{a} \rangle_I, m\gamma^2 \mathbf{a} + m \frac{\gamma^4}{c^2} \langle \mathbf{v}, \mathbf{a} \rangle_I \mathbf{v}\right)$$
(36)

**Proposition 8.30.** Let p be a particle of mass m moving along a curve  $\Gamma \subseteq \mathbb{R}^4$  with a velocity  $\mathbf{v}$  and acceleration a. Then,

1. If  $\mathbf{f} \parallel \mathbf{v}$ , then

$$f = m\gamma^3 a, \qquad m\gamma^3 := longitudinal mass. (37)$$

2. If  $\mathbf{f} \perp \mathbf{v}$ , then

$$f = m\gamma a, \qquad m\gamma := transverse \ mass.$$
 (38)

**Theorem 8.31.** Let p be a particle of mass m with  $v_0 = x_0 = t_0 = 0$  on which a constant force F acts. If we denote  $a_0 = \gamma^3 a = F/m$  (which is constant), then

$$x(t) = \frac{c^2}{a_0} \left[ \sqrt{1 + \frac{a_0^2 t^2}{c^2}} - 1 \right], \qquad v(t) = \frac{a_0 t}{\sqrt{1 + a_0^2 t^2 / c^2}}, \qquad \frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) + (1 - \alpha^2) \frac{1}{r} = -\frac{kE}{L^2 c^2}, \quad \alpha^2 = \frac{k^2}{L^2 c^2}.$$

and in the limit cases,

$$t \to \infty : v(t) \approx c, x(t) \approx ct - \frac{c^2}{a_0},$$
 (40)

$$a_0 t \ll c : v(t) \approx a_0 t, x(t) \approx \frac{a_0}{2} t^2.$$
 (41)

**Theorem 8.32.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If k < 0 and tory of p is bounded.

$$x_0 = -k/mc^2$$
,  $t_0 = 0$ , then

$$E = \gamma mc^2 + \frac{k}{r} = 0, \qquad \omega^2 x^2 + v^2 = c^2,$$
 (42)

$$\omega = \frac{c}{x_0}, \qquad T = \frac{\pi}{2} \frac{x_0}{c}, \tag{43}$$

$$x(t) = x_0 \cos \omega t, \qquad v(t) = -c \sin \omega t, \qquad (44)$$

while n classical formalism the period would be  $\frac{\pi}{2\sqrt{2}}\frac{x_0}{c} < T_{\rm rel}$ , which is normal since in the classical framework the mass can have an infinite velocity while in the relativistic one is bounded by c.

**Theorem 8.33.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If k < 0 and  $x_0 = -k/mc^2$ ,  $t_0 = 0$ , then with respect to the particle reference,

$$\frac{\mathrm{d}\tau}{\mathrm{d}t} = \frac{1}{\gamma} = \cos\omega t = \frac{x}{x_0}, \qquad \tau = \int \gamma \,\mathrm{d}t = \frac{1}{\omega}\sin\omega t, \tag{45}$$

$$x^2 = x_0^2 - c^2 \tau^2, \qquad T' = \frac{x_0}{c}, \qquad \frac{L'}{x_0} = \frac{T}{T'}.$$
 (46)

**Theorem 8.34.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If k < 0, then

$$\frac{\mathrm{d}\gamma}{\mathrm{d}t} = \frac{\gamma^3}{c^2} \langle \mathbf{v} \mathbf{P} \rangle_{\!\! L} \mathbf{v}, \qquad \gamma m \beta^2 c^2 = -\frac{k}{r},$$

$$-1 < \frac{T}{U} = -\frac{\gamma}{\gamma + 1} < -\frac{1}{2}, \qquad E = \frac{mc^2}{\gamma}.$$

**Theorem 8.35.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If k > 0, then it is not possible falling to the origin, and if k < 0, then it is possible if  $Lc \leq k$  (in this case  $p_r \to \infty$ ).

**Theorem 8.36.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If k > 0, then it is always possible to escape to the infinity is always possible, and if k < 0, it is possible if  $E > mc^2$ .

**Theorem 8.37.** Let p be a particle of mass m on which a potential of the form U = k/r acts. Then,

$$L = \gamma mr^2 \dot{\theta} = \text{ctt}, \qquad E = \gamma mc^2 + \frac{k}{r} = \text{ctt}, \qquad \mathbf{p} = \gamma m(\dot{r}\mathbf{e}_r + r\dot{\theta}\mathbf{e}_r)$$
(47)

$$\frac{\mathrm{d}}{\mathrm{d}t}(\gamma m\dot{r}) - \frac{L^2}{\gamma mr^3} = \frac{k}{r^2}, \qquad \frac{\mathrm{d}^2}{\mathrm{d}\theta^2} \left(\frac{1}{r}\right) + \frac{1}{r} = -\frac{\gamma mk}{L^2},\tag{48}$$

$$\frac{\mathrm{d}^2}{\mathrm{d}\theta^2} \left(\frac{1}{r}\right) + (1 - \alpha^2) \frac{1}{r} = -\frac{kE}{L^2 c^2}, \ \alpha^2 = \frac{k^2}{L^2 c^2}. \tag{49}$$

**Proposition 8.38.** Let p be a particle of mass m on which a potential of the form U = k/r acts. If the variation of r is negligible, then  $\alpha^2 < 1$ .

**Proposition 8.39.** Let p be a particle of mass m on which a potential of the form U = k/r (with k < 0) acts. If  $\alpha^2 <$ ,  $E < mc^2$ , and E > 0, then the trajec**Theorem 8.40.** Let p be a particle of mass m on which a potential of the form U = k/r (with k < 0) acts. If  $\alpha^2 < 1$ ,  $E < mc^2$ , and  $E > mc^2\sqrt{1-\alpha^2}$ , then the trajectory of p is determined by the expression

$$r = \frac{a(1 - e^2)}{1 + e\cos(\sqrt{1 - \alpha^2}\theta)},\tag{50}$$

$$\frac{1}{a} = \frac{E}{k} \left[ 1 - \frac{m^2 c^4}{E^2} \right], \ e = \frac{1}{\alpha} \sqrt{1 + (\alpha^2 - 1) \frac{m^2 c^4}{E^2}}, \tag{51}$$

which is an ellipse with a precession  $2\pi(1/\sqrt{1-\alpha^2}-1)$  per revolution (and  $\pi\alpha^2$  if  $\alpha^2 \ll 1$ ).

**Theorem 8.41.** Let p be a particle of mass m on which a potential of the form U = k/r (with k < 0) acts. If p has a closed bounded trajectory, then the average of  $\frac{\mathrm{d}\langle \mathbf{r}, \mathbf{p}\rangle_I}{\mathrm{d}t} = 0$  on an interval of nT.

**Proposition 8.42.** Let p be a particle of mass m on which a potential of the form U = k/r (with k < 0) acts. If p has a closed bounded trajectory, then

$$E = \left\langle \frac{1}{\gamma} \right\rangle mc^2. \tag{52}$$

## 9 Generalized coordinates

**Definition 9.1.** Let S be a system of particles  $p_1, \ldots, p_n$  with masses  $m_1, \cdots, m_n$ . Then, we say the system has non stationary holonomic constraints or rheonomic constraints if and only if there is a function  $\mathbf{f}: \mathbb{R}^{3n} \times \mathbb{R} \longrightarrow \mathbb{R}^k$  such that

$$\mathbf{f}(\mathbf{r}_1, \dots, \mathbf{r}_n, t) = \mathbf{0}. \tag{53}$$

**Proposition 9.1.** Let S be a system of n particles with a constraint  $\mathbf{f}: \mathbb{R}^{3n} \times \mathbb{R} \longrightarrow \mathbb{R}^k$  and V the set of possible velocities at an instant t. If  $\dot{\mathbf{x}} = (\mathbf{v}_1, \dots, \mathbf{v}_n)$ , then

$$\sum_{i=1}^{n} \left\langle \nabla_{i} f_{j}, \mathbf{v}_{i} \right\rangle_{I} + \frac{\partial f_{j}}{\partial t} = 0, \ j = 1, \dots, k.$$
 (54)

**Theorem 9.2** (D'Alembert's principle). Let S be a system of particles. Then,

$$\sum_{i=1}^{n} \langle \mathbf{F}_{i} - m\mathbf{a}_{i}, \delta \mathbf{r}_{i} \rangle_{I} = 0, \ \forall \, \delta \mathbf{r}_{i}.$$
 (55)

**Theorem 9.3.** Let S be a system of n particles with generalized coordinates  $q^1, \ldots, q^r$ . Then,

$$\sum_{i=1}^{n} \left\langle \mathbf{F}_{i}, \frac{\partial \mathbf{p}_{i}}{\partial q^{j}} \right\rangle_{I} = \frac{\mathrm{d}}{\mathrm{d}t} \frac{\partial T}{\partial \dot{q}^{j}} - \frac{\partial T}{\partial q^{j}}, \ j = 1, \dots, r.$$
 (56)

And if **F** is derived from a potential  $\Phi(\mathbf{r})$ , then

$$\left| \frac{\partial L}{\partial q^j} - \frac{\mathrm{d}}{\mathrm{d}t} \frac{\partial L}{\partial \dot{q}^j} = 0, \ j = 1, \dots, r. \right|$$
 (57)

**Theorem 8.40.** Let p be a particle of mass m on which **Theorem 9.4.** Let  $J: C^2[x_0, x_1] \longrightarrow \mathbb{R}$  be a funca potential of the form U = k/r (with k < 0) acts. If tional of the form

$$J(y) = \int_{x_0}^{x_1} f(x, y, y') \, dx,$$

where f has continuous partial derivatives of second order with respect to x, y, and y', and  $x_0 < x_1$ . Let

$$S = \{ y \in C^2[x_0, x_1] \mid y(x_0) = y_1 \text{ and } y(x_1) = y_1 \},$$

where  $y_0$  and  $y_1$  are given real numbers. If  $y \in S$  is an extremal for J, then

$$\frac{\mathrm{d}}{\mathrm{d}x} \left( \frac{\partial f}{\partial y'} \right) - \frac{\partial f}{\partial y} = 0 \tag{58}$$

for all  $x \in [a_0, x_1]$ .

**Theorem 9.5** (Lagrange multipliers method for non-holonomic constraints). If we want to find an extrema having a set of m non-holonomic constraints

$$\overline{\delta f_1} = A_{11}\delta u_1 + \dots + A_{1n}\delta u_n = 0,$$

$$\vdots = \vdots \qquad \qquad = \vdots$$

$$\overline{\delta f_m} = A_{m1}\delta u_1 + \dots + A_{mn}\delta u_n = 0,$$
(59)

Then we can proceed as the original multipliers method by considering every variation as independent and solving the following equation.

$$\delta F + \lambda_1 \overline{\delta f_1} + \dots + \lambda_m \overline{\delta f_m} = 0 \tag{60}$$

**Theorem 9.6.** Let f be a continuous functions with a variation  $\delta f = \epsilon \phi$ . Then,

$$\frac{\mathrm{d}}{\mathrm{d}x}\delta y = \delta \frac{\mathrm{d}}{\mathrm{d}x}y, \qquad \delta \int_{a}^{b} f(x) \,\mathrm{d}x = \int_{a}^{b} \delta f(x) \,\mathrm{d}x. \quad (61)$$

**Theorem 9.7.** Let  $J: \mathbf{C}^2[t_0, t_1] \longrightarrow \mathbb{R}$  be a functional of the form

$$J(\mathbf{q}) = \int_{t_0}^{t_1} L(\mathbf{q}, \dot{\mathbf{q}}, t) dt, \qquad (62)$$

where  $\mathbf{q} = (q_1, \dots, q_n)$ , and L has continuous secondorder partial derivatives with respect to  $t, q_k$ , and  $\dot{q}_k$ ,  $k = 1, \dots, n$ . Let

$$S = \{ \mathbf{q} \in \mathbf{C}^2[t_0, t_1] \mid \mathbf{q}(t_0) = \mathbf{q}_0, \mathbf{q}(t_1) = \mathbf{q}_1 \}, \quad (63)$$

where  $\mathbf{q}_0, \mathbf{q}_1 \in \mathbb{R}^n$  are given vectors. If  $\mathbf{q}$  is an extremal for J in S then for k = 1, ..., n

$$\frac{\partial L}{\partial q_k} - \frac{\mathrm{d}}{\mathrm{d}t} \frac{\partial L}{\partial \dot{q}_k} = 0. \tag{64}$$

**Theorem 9.8.** If we have a set of holonomic constraints

$$f_1(q_1,\ldots,q_n,t)=0,$$

$$\dots = \vdots, \qquad (65)$$

$$f_m(q_1, \dots, q_n, t) = 0,$$

then we can treat each variable as independent and  $\$  which leads to the equation search the stationary value of

$$J' = \int_{t_1}^{t_2} L + \sum_{k=1}^{m} \lambda_m f_m \, dt \,, \qquad (66) \qquad \left[ \frac{\partial L}{\partial q_k} - \frac{\mathrm{d}}{\mathrm{d}t} \frac{\partial L}{\partial \dot{q}_k} + \lambda_1 \frac{\partial f_1}{\partial q_k} + \dots + \lambda_m \frac{\partial f_m}{\partial q_k} = 0. \right]$$