

# 1 Motion in one dimension

**Proposition 1.1.** *Let .... Then, it is true that*

$$\vec{F}^{ext} = m\vec{a} + (\vec{v} - \vec{u})\dot{m} \quad (1)$$

or which is equivalent,

$$\vec{F}^{ext} + \dot{m}\vec{u} = \dot{\vec{p}} \quad (2)$$

# 2 Oscillations

**Proposition 2.1.** *Let be the following differential equation*

$$\ddot{x} + \omega_0^2 x = 0, \quad (3)$$

with the initial value condition of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the general solution is

$$x(t) = x_0 \cos \omega_0 t + \frac{v_0}{\omega_0} \sin \omega_0 t, \quad (4)$$

or, which is equivalent,

$$x(t) = A \cos [\omega_0 t + \phi_0], \quad A = \sqrt{x_0^2 + \left(\frac{v_0}{\omega_0}\right)^2}, \quad \phi_0 = -\arctan \frac{v_0}{\omega_0 x_0}. \quad (5)$$

**Definition 2.1.** Let  $U(x)$  be a potential function of class  $C^2(\mathbb{R})$ . Then, we say  $x_0$  is a point of stable equilibrium if  $U$  has a maxima in  $x_0$ .

**Proposition 2.2.** *Let be the following differential equation*

$$\ddot{x} + 2\beta\dot{x} + \omega_0^2 x = 0, \quad (6)$$

with the initial value conditions of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the general solution is

$$x(t) = e^{-\beta t} \left[ x_0 \cos \tilde{\omega} t + \frac{v_0 + \beta x_0}{\tilde{\omega}} \sin \tilde{\omega} t \right], \quad \tilde{\omega} = \sqrt{\omega_0^2 - \beta^2} \quad (7)$$

if  $\beta < \omega_0$ ,

$$x(t) = e^{-\omega_0 t} [x_0 + (x_0 \omega_0 + v_0)t] \quad (8)$$

if  $\beta = \omega_0$ , and

$$x(t) = \frac{x_0(\tilde{\omega} - \beta) - v_0}{2\tilde{\omega}} e^{-(\beta+\tilde{\omega})t} + \frac{x_0(\tilde{\omega} + \beta) + v_0}{2\tilde{\omega}} e^{-(\beta-\tilde{\omega})t} \quad (9)$$

if  $\beta > \omega_0$ .

**Proposition 2.3.** *Let be the following differential equation*

$$\ddot{x} + 2\beta\dot{x} + \omega_0^2 x = f(t) = f_0 \cos [\omega t + \psi_0], \quad (10)$$

with the initial value conditions of  $x(0) = x_0$  and  $v(0) = v_0$ . Then, the particular solution is

$$x_p(t) = A \cos [\omega t + \psi_0 - \phi_0], \quad A = \frac{f_0}{\sqrt{(\omega_0^2 - \omega^2)^2 + 4\beta^2\omega^2}}, \quad \phi_0 = \arctan \frac{\omega(\omega_0^2 - \omega^2)}{2\beta\omega^2} \quad (11)$$

# 3 Central forces

**Definition 3.1.** Central force

$$\vec{F}(\vec{r}) = f(r)\vec{e}_\rho \quad (12)$$

**Proposition 3.1.** *All central forces are conservatives.*

**Proposition 3.2.** *The angular momentum with respect the origin is conserved.*

$$\dot{\vec{L}} = \vec{0} \quad (13)$$

**Proposition 3.3.** *The areal velocity is constant.*

$$\frac{dA}{dt} = \frac{L}{2m} = ctt \quad (14)$$

**Theorem 3.4** (Bertrand's Theorem). *The only central potentials where every bounded orbit is closed are:*

$$U(r) = -\frac{k}{r}, \quad U(r) = \frac{k}{2}r^2, \quad k > 0 \quad (15)$$

# 4 Coupled oscillations 1

# 5 Coupled oscillations 2

# 6 Rotations

# 7 Dynamics of rigid body

**Proposition 7.1.** *The vector  $\Omega$  is independent on the origin of the system  $S$ .*

**Proposition 7.2.** *The energy of the rigid body is an invariant scalar under change of basis.*

# 8 Special relativity

**Theorem 8.1.** *If Maxwell's equations*

$$\langle \nabla, \mathbf{E} \rangle_I = \frac{\rho}{\epsilon_0}, \quad (16)$$

$$\langle \nabla, \mathbf{B} \rangle_I = 0, \quad (17)$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}, \quad (18)$$

$$\nabla \times \mathbf{B} = \mu_0 \left( \mathbf{J} + \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right). \quad (19)$$

are invariant under Galileo transformations, then  $\mathbf{E} = \mathbf{B} = \mathbf{0}$ .

**Definition 8.1** (Reference system). We define a reference system  $S$  as a set of three axis and one origin over which we have determined an orientation. We will suppose we have selected a unit of length and that in each point  $a$  in the immobile space with respect the axis there is a clock  $q_a$  such that the clocks  $q_a$  and  $q_b$  corresponding to two different points  $a$  and  $b$  immobile with respect these axis are synchronized

**Lemma 8.2.** *Let  $f : \mathbb{R}^4_{2\beta\omega} \rightarrow \mathbb{R}^4$  be a Lorentz transformation. Then, for any world lines that are not contained in hyperplanes of the form  $t = ctt$  to lines that are not contained in hyperplanes of the form  $t' = ctt$ .*

**Lemma 8.3.** Let  $f : \mathbb{R}^4 \rightarrow \mathbb{R}^4$  be a Lorentz transformation. Let  $f : \mathbb{R}^4 \rightarrow \mathbb{R}^4$  be a Lorentz transformation. Then,  $f$  transforms planes that are not contained in hyperplanes of the form  $t = \text{ctt}$  to planes that are not contained in hyperplanes of the form  $t' = \text{ctt}$ .

**Lemma 8.4.** Let  $f : \mathbb{R}^4 \rightarrow \mathbb{R}^4$  be a Lorentz transformation. Then,  $f$  transforms hyperplanes that are not of the form  $t = \text{ctt}$  to hyperplanes that are not of the form  $t' = \text{ctt}$ .

**Theorem 8.5.** Let  $f : \mathbb{R}^4 \rightarrow \mathbb{R}^4$  be a Lorentz transformation. Then,  $f$  is an affine transformation.

**Theorem 8.6.** Let  $S, S'$  be two inertial reference systems. We can make orthogonal changes (isometries) of axis to  $S$  and  $S'$  and a change of origin of time such that the Lorentz transformation has the form of the equation ??.

**Theorem 8.7.** (Lorentz transformation) Let  $S, S'$  be two reference systems with the same origin such that  $S'$  moves with a constant velocity  $\mathbf{v} = v\mathbf{e}_x$ . Then,

$$P_{s'} = \Lambda P_s \Leftrightarrow P_{s'}^\nu = \Lambda_\mu^\nu P_s^\mu, \quad \Lambda = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad (20)$$

**Proposition 8.8.** Let  $\mathbf{R} \in L$  be a vector and  $a \in \mathbb{R}$  a scalar. Then,  $\|\mathbf{R}\|_m = |a|\|\mathbf{R}\|_m$ .

**Proposition 8.9.** Every subspace  $W$  of  $L$  is either timelike, spacelike, or lightlike. Besides,

1.  $S$  is timelike  $\Leftrightarrow W^\perp$  is spacelike.
2.  $S$  is spacelike  $\Leftrightarrow W^\perp$  is timelike.
3.  $W$  is lightlike  $\Leftrightarrow W^\perp$  is lightlike.

**Proposition 8.10.** Two orthogonal vectors different from zero and non spacelike are necessarily lightlike and collinear. In particular, there is not a subspace of dimension 2 where  $\langle, \rangle$  is null.

**Proposition 8.11.** Let  $\mathbf{R}_1, \mathbf{R}_2 \in T$  be two timelike vectors. Then, the following statements are true.

1.  $|\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m| \geq \|\mathbf{R}_1\|_m \|\mathbf{R}_2\|_m$ , and the equality is equivalent to both vectors being collinear.
2.  $\mathbf{R}_1, \mathbf{R}_2$  are in the same time cone ( $C_+$  or  $C_-$ ) if and only if  $\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m < 0$ . In this case,
  - (a) There is a unique  $\varphi \in \mathbb{R}$  such that

$$\cosh \varphi = -\frac{\langle \mathbf{R}_1, \mathbf{R}_2 \rangle_m}{\|\mathbf{R}_1\|_m \|\mathbf{R}_2\|_m}. \quad (21)$$

We call this  $\varphi$  the hyperbolic angle.

- (b)  $\|\mathbf{R}_1\|_m + \|\mathbf{R}_2\|_m \leq \|\mathbf{R}_1 + \mathbf{R}_2\|_m$ .

**Proposition 8.12.** The Lorenz-Minkowski metric is invariant under Lorentz transformations.

**Proposition 8.13.** Let  $S, S'$  be two inertial reference systems such that the velocity of  $S'$  is  $\mathbf{w} = w\mathbf{e}_x$  with respect to  $S$ . Then,

$$v_{s'}^1 = \frac{v_s^1 - w}{1 - \beta_v \beta_w}, \quad v_{s'}^2 = \frac{1}{\gamma_w} \frac{v_s^2}{1 - \beta_v \beta_w}, \quad v_{s'}^3 = \frac{1}{\gamma_w} \frac{v_s^3}{1 - \beta_v \beta_w}, \quad (22)$$

**Proposition 8.14.** If the system has a general velocity  $\mathbf{w}$ , then

$$\mathbf{v}' = \frac{1}{1 - \langle \beta_v, \beta_w \rangle_I} \left[ \frac{\mathbf{v}}{\gamma_w} - \mathbf{w} + \frac{1}{c^2} \frac{\gamma_w}{1 + \gamma_w} \langle \mathbf{v}, \mathbf{w} \rangle_I \mathbf{w} \right], \quad (23)$$

$$\mathbf{v} = \frac{1}{1 + \langle \beta_{v'}, \beta_w \rangle_I} \left[ \frac{\mathbf{v}'}{\gamma_w} + \mathbf{w} + \frac{1}{c^2} \frac{\gamma_w}{1 + \gamma_w} \langle \mathbf{v}', \mathbf{w} \rangle_I \mathbf{v} \right]. \quad (24)$$

**Proposition 8.15.** Let  $S, S'$  two inertial reference systems. Then,

$$\mathbf{U}_{s'} = \Lambda \mathbf{U}_s \Leftrightarrow U_{s'}^\nu = \Lambda_\mu^\nu U_s^\mu. \quad (25)$$

**Theorem 8.16.**

$$E^1 = p^2 c^2 + m^2 c^4. \quad (26)$$

**Theorem 8.17.** (Work-Energy theorem)

$$W = \Delta E. \quad (27)$$

## 9 Generalized coordinates

**Definition 9.1.** Let  $S$  be a system of particles  $p_1, \dots, p_n$  with masses  $m_1, \dots, m_n$ . Then, we say the system has *non stationary holonomic constraints* or *rheonomic constraints* if and only if there is a function  $\mathbf{f} : \mathbb{R}^{3n} \times \mathbb{R} \rightarrow \mathbb{R}^k$  such that

$$\mathbf{f}(\mathbf{r}_1, \dots, \mathbf{r}_n, t) = \mathbf{0}. \quad (28)$$

**Proposition 9.1.** Let  $S$  be a system of  $n$  particles with a constraint  $\mathbf{f} : \mathbb{R}^{3n} \times \mathbb{R} \rightarrow \mathbb{R}^k$  and  $V$  the set of possible velocities at an instant  $t$ . If  $\dot{\mathbf{x}} = (\mathbf{v}_1, \dots, \mathbf{v}_n)$ , then

$$\sum_{i=1}^n \langle \nabla_i f_j, \mathbf{v}_i \rangle_I + \frac{\partial f_j}{\partial t} = 0, \quad j = 1, \dots, k. \quad (29)$$

**Theorem 9.2** (D'Alembert's principle). Let  $S$  be a system of particles. Then,

$$\sum_{i=1}^n \langle \mathbf{F}_i - m\mathbf{a}_i, \delta \mathbf{r}_i \rangle_I = 0, \quad \forall \delta \mathbf{r}_i. \quad (30)$$

**Theorem 9.3.** Let  $S$  be a system of  $n$  particles with generalized coordinates  $q^1, \dots, q^r$ . Then,

$$\sum_{\mu=1}^n \left\langle \mathbf{F}_\mu, \frac{\partial \mathbf{p}_\mu}{\partial q^\nu} \right\rangle_I = \frac{d}{dt} \frac{\partial T}{\partial \dot{q}^\nu} - \frac{\partial T}{\partial q^\nu}, \quad \nu = 1, \dots, r. \quad (31)$$

And if  $\mathbf{F}$  is derived from a potential  $\Phi(\mathbf{r})$ , then

$$\frac{\partial L}{\partial q^\nu} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}^\nu} = 0, \quad \nu = 1, \dots, r. \quad (32)$$

**Theorem 9.4.** Let  $J : C^2[x_0, x_1] \rightarrow \mathbb{R}$  be a functional of the form

$$J(y) = \int_{x_0}^{x_1} f(x, y, y') dx,$$

where  $f$  has continuous partial derivatives of second order with respect to  $x$ ,  $y$ , and  $y'$ , and  $x_0 < x_1$ . Let

$$S = \{y \in C^2[x_0, x_1] \mid y(x_0) = y_0 \text{ and } y(x_1) = y_1\},$$

where  $y_0$  and  $y_1$  are given real numbers. If  $y \in S$  is an extremal for  $J$ , then

$$\frac{d}{dx} \left( \frac{\partial f}{\partial y'} \right) - \frac{\partial f}{\partial y} = 0 \quad (33)$$

for all  $x \in [a_0, x_1]$ .

**Theorem 9.5** (Lagrange multipliers method for non-holonomic constraints). If we want to find an extrema having a set of  $m$  non-holonomic constraints

$$\begin{aligned} \overline{\delta f_1} &= A_{11}\delta u_1 + \cdots + A_{1n}\delta u_n = 0, \\ \vdots &= \ddots = \vdots \\ \overline{\delta f_m} &= A_{m1}\delta u_1 + \cdots + A_{mn}\delta u_n = 0, \end{aligned} \quad (34)$$

Then we can proceed as the original multipliers method by considering every variation as independent and solving the following equation.

$$\delta F + \lambda_1 \overline{\delta f_1} + \cdots + \lambda_m \overline{\delta f_m} = 0 \quad (35)$$

**Theorem 9.6.** Let  $f$  be a continuous functions with a variation  $\delta f = \epsilon \phi$ . Then,

$$\frac{d}{dx} \delta y = \delta \frac{d}{dx} y, \quad \delta \int_a^b f(x) dx = \int_a^b \delta f(x) dx. \quad (36)$$

**Theorem 9.7.** Let  $J : \mathbf{C}^2[t_0, t_1] \rightarrow \mathbb{R}$  be a functional of the form

$$J(\mathbf{q}) = \int_{t_0}^{t_1} L(\mathbf{q}, \dot{\mathbf{q}}, t) dt, \quad (37)$$

where  $\mathbf{q} = (q_1, \dots, q_n)$ , and  $L$  has continuous second-order partial derivatives with respect to  $t$ ,  $q_k$ , and  $\dot{q}_k$ ,  $k = 1, \dots, n$ . Let

$$S = \{\mathbf{q} \in \mathbf{C}^2[t_0, t_1] \mid \mathbf{q}(t_0) = \mathbf{q}_0, \mathbf{q}(t_1) = \mathbf{q}_1\}, \quad (38)$$

where  $\mathbf{q}_0, \mathbf{q}_1 \in \mathbb{R}^n$  are given vectors. If  $\mathbf{q}$  is an extremal for  $J$  in  $S$  then for  $k = 1, \dots, n$

$$\boxed{\frac{\partial L}{\partial q_k} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_k} = 0.} \quad (39)$$

**Theorem 9.8.** If we have a set of holonomic constraints

$$\begin{aligned} f_1(q_1, \dots, q_n, t) &= 0, \\ &\dots = \vdots, \\ f_m(q_1, \dots, q_n, t) &= 0, \end{aligned} \quad (40)$$

then we can treat each variable as independent and search the stationary value of

$$J' = \int_{t_1}^{t_2} L + \sum_{k=1}^m \lambda_k f_k dt, \quad (41)$$

which leads to the equation

$$\boxed{\frac{\partial L}{\partial q_k} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_k} + \lambda_1 \frac{\partial f_1}{\partial q_k} + \cdots + \lambda_m \frac{\partial f_m}{\partial q_k} = 0.} \quad (42)$$