Simulation of CXRS spectra using Raysect and CHERAB frameworks

Aleksei Shabashov a.shabashov@iterrf.ru

Contents

1	Pur	pose	2
2	Sco	pe	2
3	Defi	initions	2
4	Intr	roduction	3
	4.1	CXRS Diagnostics	3
		4.1.1 CXRS for Plasma Measurements	3
		4.1.2 CXRS in ITER	3
	4.2	Development of the New Simulation Code	5
		4.2.1 Existing Code	5
		4.2.2 Motivation	5
	4.3	Raysect and CHERAB	6
5	Sim	ulation of CXRS Spectra	6
6	Con	nclusion	6
$\mathbf{A}_{\mathbf{J}}$	ppen	dices	7
\mathbf{A}	For	Users	7
	A.1	Installation	7
	A.2	Preparations for the first run	8
	A.3	Usage	8
		A.3.1 Performing a simulation	8
		A.3.2 Print plasma profiles	9
		A.3.3 Print plasma composition	9
		A.3.4 Configuration	9
		A.3.5 Emission parameters	9
		A.3.6 Plasma, beam, fibre, optics, camera, scanner and spectrometer parameters	9
		A.3.7 Emission lines	9
		A.3.8 DNB	10
		A.3.9 Wavelength ranges and geometry	10
В	For	Developers	10
	B.1	Project Structure	11
	B.2	imas module – Interaction with IMAS database	13
		B.2.1 Supplementary Functions	13
		B.2.2 equilibrium IDS	14

	B.2.3	${ t core_profiles \ IDS \ . \ . \ . \ . \ . \ . \ . \ . \ . \ $	14
	B.2.4	edge_profiles IDS	18
	B.2.5	$\verb charge_exchange IDS $	21
	B.2.6	nbi IDS	21
В.3	Setting	g the Wall	22
B.4	Observ	vers	23
	B.4.1	Base Class	23
	B.4.2	Sightlines	24
	B.4.3	Optics	25
	B.4.4	Fibres	26
	B.4.5	Camera	27
	B.4.6	Other Observers	28
B.5	Popula	ating CHERAB's Atomic Database	30
B.6	Utility	Functions	31
	B.6.1	Getting Information on Plasma Parameters	31
	B.6.2	Parsing XML Configuration File	32
	B.6.3	Setting Emission Parameters	33
	B.6.4	Math Functions	33
	B.6.5	Fitting Routine	33
	B 6 6	Others	33

1 Purpose

empty

2 Scope

This document is applicable to the 55.EC CXRS Edge diagnostic. The CXRS Edge system is an optical diagnostic that collects the light emitted by the plasma upon interaction with the Diagnostic Neutral Beam (DNB) and analyses this light to extract the ion temperature, plasma rotation velocities and impurity content of the plasma. An overview of the CXRS Edge system is given in the Design Description document (DDD) [1].

3 Definitions

For a complete list of ITER abbreviations see [2]. Below abbreviations used in this document are given.

4 Introduction

4.1 CXRS Diagnostics

4.1.1 CXRS for Plasma Measurements

The Charge Exchange Recombination Spectroscopy (CXRS) diagnostics measures line emissions of several impurity isotopes in the plasma excited by charge exchange reactions with neutral hydrogen atoms injected into the plasma by the diagnostic neutral beam (DNB) (eqs. (1) and (2)). This line emission (table 1) provides essential information for plasma control and physics studies: ion temperature T_i (eq. (3)), toroidal (v_{tor}) and poloidal (v_{pol}) plasma rotation velocity (eq. (4)), helium ash and low Z impurity densities (beryllium, carbon, neon etc.) and the derived quantities such as Z_{eff} (eq. (5)).

$$X^{Z+} + H^0 \to X^{(Z-1)+}(n_2) + H^+$$
 (1)

$$X^{(Z-1)+}(n_2) \to X^{(Z-1)+}(n_1) + h\nu$$
 (2)

$$kT_{\rm ion} = mc^2 \frac{\Delta \lambda_{\rm Dopp}^2}{\lambda_o^2} \tag{3}$$

where k – Boltzmann constant, $\Delta \lambda_{\text{dopp}}$ – line's width due to Doppler broadening, λ_0 – "natural", unshifted wavelength of the line's center.

$$v_{\rm rot} = c \frac{\Delta \lambda_{\rm rot}}{\lambda_0 \cos \alpha} \tag{4}$$

where $\Delta \lambda_{\rm rot}$ – line's shift due to Doppler effect, α – angle between line of sight and toroidal direction.

$$n_{\rm imp} = \frac{4\pi \int I(\lambda) \, d\lambda}{n_{\rm b} Q_{\rm CX}^{\rm eff}(v_{\rm b}) \, dl} \tag{5}$$

where $I(\lambda)$ – intensity of the line, $n_{\rm b}$ – local neutral beam density, $Q_{\rm CX}^{\rm eff}(v_{\rm b})$ – effective rate coefficient due to charge exchange, l – coordinate along a line of sight.

4.1.2 CXRS in ITER

The CXRS Edge diagnostic is a distributed system with components throughout the ITER tokamak complex. The primary viewing components (front-end optics) are installed in Equatorial port 3 (EP3), viewing the Diagnostic Neutral Beam (DNB) that enters the plasma in the

Table 1. Spectroscopic lines used in CXRS.

Ion	Transition	Wavelength
BeIV	$6 \rightarrow 5$	465.8 nm
BeIV	$8 \rightarrow 6$	468.5 nm
HeII	$4 \rightarrow 3$	$468.5~\mathrm{nm}$
ArXVIII	$16 \rightarrow 15$	522.5 nm
NeX	$11 \rightarrow 10$	524.9 nm
CVI	$8 \rightarrow 7$	529.1 nm
$H\alpha$		656.3 nm
MSE		659.1 nm

neighboring section 4. Front-end optics consists of two light collecting systems: Upper and Lower. The scope of the 55.EC CXRS Edge diagnostic for Upper system ends at the image plane in the port interspace (11-L1-C03) where the collected signal is coupled into optical fibre bundles. For Lower system, however, 55.EC CXRS Edge diagnostic is responsible for light collection system, light transportation and detection. The collected signal is transported through optical fibre bundles to the Tritium building (Building 14 – Level 2 – Room 4), where the detection systems are located.

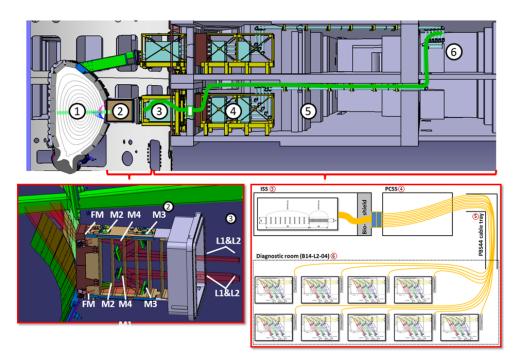


Figure 1. An overview of the CXRS Edge diagnostic design.

4.2 Development of the New Simulation Code

4.2.1 Existing Code

Simulation of Spectra (SOS) code by M. G. von Hellermann [3] Features:

- Simulation takes into account many physical effects (halo effect, crossection effect, plume effect and others);
- Written in Matlab;
- Has Graphical User Interface (fig. 2).

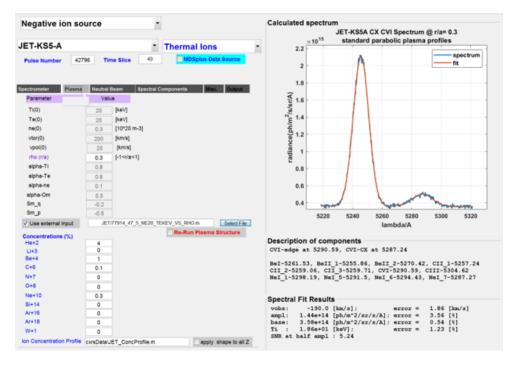


Figure 2. SOS interface.

4.2.2 Motivation

Existing code (Simulation of Spectra – SOS) lacks some features:

- Simplified plasma, tokamak and diagnostic geometry (e.g. elliptical plasma, point emission and others);
- Does not take reflections into account;
- Cannot use data from IMAS directly;
- Requires Matlab license, hard to extend by new developers.

The goal was to create an open and extensible simulation code using Python. Sub goals:

- Implement interaction with IMAS database (read and write);
- Use IMAS data to create a plasma and diagnostic beam with spatial distributions;

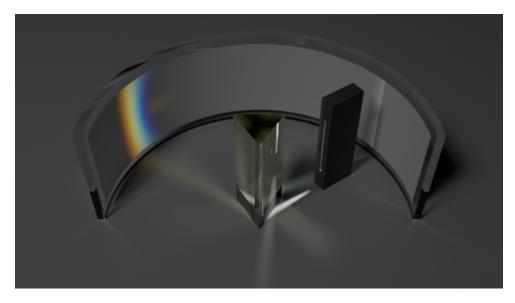


Figure 3. Demonstration of Raysect features.

- Use a ray-tracing engine to simulate spectra, this includes how reflections affect simulated spectra;
- Ensure that emission models include all physics already captured by SOS.

4.3 Raysect and CHERAB

Raysect [4] is a ray-tracing framework for Python designed for scientific purposes.

- Supports scientific ray-tracing of spectra from physical light sources such as plasmas.
- Easily extensible, written with user customisation of materials and emissive sources in mind.
- Different observer types supported such as pinhole cameras and optical fibres.

CHERAB [5] is a Python library for forward modelling diagnostics based on spectroscopic plasma emission which provides physical models for Raysect. Provided models for Raysect:

- Tools for plasma and diagnostic beam simulations;
- Physical emission models (active charge exchange, bremsstrahlung and more).

5 Simulation of CXRS Spectra

Main part

6 Conclusion

Conclusion

References

- [1] Zvonkov A., Serov S., and Tugarinov S. System Design Description (DDD) 55.EC CXRS Edge. Version 4.1. Apr. 2020.
- [2] ITER Abbreviations. Version 1.17. Mar. 2018.
- [3] Manfred von Hellermann et al. "Simulation of Spectra Code (SOS) for ITER Active Beam Spectroscopy". In: *Atoms* 7.1 (Mar. 2019), p. 30. DOI: 10.3390/atoms7010030.
- [4] Dr Alex Meakins and Matthew Carr. raysect/source: v0.5.2 Release. Version v0.5.2. Aug. 2018. DOI: 10.5281/zenodo.1341376. URL: https://doi.org/10.5281/zenodo.1341376.
- [5] Dr Carine Giroud et al. CHERAB Spectroscopy Modelling Framework. Version v0.1.0. Mar. 2018. DOI: 10.5281/zenodo.1206142. URL: https://doi.org/10.5281/zenodo. 1206142.

Appendices

A For Users

A.1 Installation

Clone this repository then go to its root folder:

```
git clone ssh://git@git.iter.org/diag/cxrs.git
cd cxrs
```

First of all you have to load all required modules:

```
source env.sh
```

This will purge all loaded modules and load ones that necessary to install and use cxrs. To install cxrs on your computer run

```
pip install --user .
```

while in the the cxrs root directory.

Note: you can omit python -m part later by adding \$HOME/.local/bin to your PATH. To do so, locate file .bashrc in your home directory and add the following line at the end of the file:

```
export PATH="$PATH:/home/ITER/<username>/.local/bin"
```

where <username> is your username at ITER GPC. You can look at it by executing next command via command line: echo \$USER.

A.2 Preparations for the first run

Step 1: Create environment file

In order to create environment file, run

This will create the default environment file, that loads all needed modules. To do so, run

It will purge all loaded modules and will load modules necessary to use cxrs.

Note: you have to do this for every new session on ITER GPC.

Step 2: Create configuration file

cxrs behavior controlled by configuration file which is necessary to run the code.

To create default configuration file, run

This will create a configuration file with a name config.xml in current directory.

Step 3: Populate local atomic database

cxrs uses a local atomic database. To create one, run

Atomic data will be copied to the \$HOME/.cherab directory.

A.3 Usage

A.3.1 Performing a simulation

Main part of cxrs is simulation routine, to use it for pulse with <shot> shot number and <run> run number for time slice <time> use:

```
python -m cxrs simulate -s <shot> -r <run> -t <time> -c <config>
```

where **<config>** is a path to configuration file.

It will perform a simulation and will store result in newly created folder cxrs_output.

Note: time can be omitted. In that case time slice in the middle of the time range will be used. This is particularly useful for time slices with a lot of digits after after decimal separator. For example:

```
python -m cxrs simulate -s 134000 -r 30 -c config.xml
```

For additional information use help:

```
python -m cxrs simulate --help
```

A.3.2 Print plasma profiles

cxrs info subprogram can be used to inspect different distributions of plasma and DNB parameters. To use it, run

```
python -m cxrs info -s <shot> -r <run>
```

It will produce variety of plots and tables and place it in cxrs_output folder. *Note:* 1D profiles represent values at the DNB axis.

A.3.3 Print plasma composition

You can quickly look at plasma composition for requested pulse by using

```
python -m cxrs composition -s <shot> -r <run>
```

Omit arguments to print composition for all pulses available by scenario_summary program.

A.3.4 Configuration

cxrs accepts user's configurations as xml-file.

To change option's value, change value attribute in double quotes corresponding to the option. For more information on the meaning of certain parameter, look at its description.

Note: type attribute is actually important. For string values use **str**, for floating-point values use **float** and **int** for integers.

Warning: Do not change tags, like <user_options> or others, this will stop the program from working.

A.3.5 Emission parameters

Section emission_parameters controls which emission types are used during a simulation. Each of them can be turned on or off separately.

A.3.6 Plasma, beam, fibre, optics, camera, scanner and spectrometer parameters

Each of those sections controls appropriate aspect of simulation. For information on each parameter look at its description in configuration file.

A.3.7 Emission lines

List of emission lines suited for observations is placed in <emission_parameters> section:

```
<emission_lines description="CXRS elements to observe.">
    <HI>
        <name
            description="Name of the element."
            unit="n.a."
            type="str"
            value="hydrogen"/>
        <charge
            description="Charge of the ion."
            unit="n.a."
            type="int"
            value="0"/>
        <transition_levels</pre>
            description="Transition levels [upper, lower]."
            unit="n.a."
            type="int"
            value="[3, 2]"/>
    </HI>
    ... more lines omitted
</emission_lines>
```

You can change and add new entries in this section following given template. Added emission lines will be simulated if atomic data exists for them.

A.3.8 DNB

DNB section of the configuration file contains parameters for Diagnostic Neutral Beam. Structure of this section replicates structure of nbi IDS.

A.3.9 Wavelength ranges and geometry

These two sections contain all information for CXRS diagnostics.

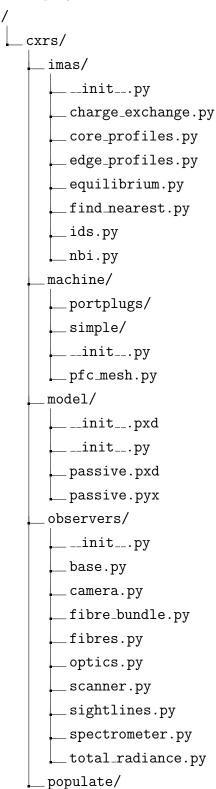
B For Developers

This section is aimed to help future developers in their work on cxrs. Most of the functions implemented in this project contain docstrings written in "classic" Python style using reStructuredText syntax. Each docstring contains a short description of a function, list of arguments and their types and list of exceptions which can be risen during execution of the function, but description for some of the arguments can be missing.

Also this section tries to be a proper documentation for a software project. But, really, it is more of a reasoning for why this thing is done like that and not like this.

B.1 Project Structure

cxrs project is divided in several modules, each in its respective folder:



```
__init__.py
       _create.py
       _{-} openadas.py
    subprograms/
       _{-} init_..py
       _composition.py
       _idslist.py
       _{
m linfo.py}
       _local_copy.py
       _populate.py
       _read_ids.py
       _search.py
       _simulate.py
       _write_ids.py
    utility/
       __init__.py
        info/
          _{-} init...py
          _beam.py
          _equilibrium.py
          _{
m passive.py}
          _{
m plasma.py}
         \_\texttt{profiles.py}
       annotation.py
        data.py
        emission.py
       _fit.py
       math.py
       _timer.py
       _xml.py
     __init__.py
    \_{\tt main}\_.{\tt py}
    _create_beam.py
    env.sh
   \_\mathtt{matplotlibrc}
_tests/
_{-}\,.\mathtt{gitignore}
\_ config.xml
\_env.sh
```

```
MANIFEST.in
matplotlibrc
README.md
requirements.txt
setup.cfg
setup.py
```

B.2 imas module – Interaction with IMAS database

Module responsible for reading data from IMAS and creating appropriate plasma and DNB models is stored in cxrs/imas directory.

B.2.1 Supplementary Functions

charge_exchange_ids = ids_get(

ids_get File ids.py contains only one function ids_get which is used to check and load requested IDS. For example

```
name="charge_exchange",
         shot = 134000,
         run=30,
         user="public",
         database="iter",
         version="3",
         occurrence=0
    )
name is a name of requested IDS,
shot is an IMAS database shot ID,
run is an IMAS data run ID,
user is an IMAS database user ID,
database is an IMAS database ID,
version is an IMAS major version number,
occurrence is an IDS occurrence number. Each IDS can store several occurrences (refer to
     the description of an IDS). For example in charge_exchange IDS occurrences used to store
     data related to different diagnostics.
```

find_nearest File find_nearest.py contains only one function find_nearest. It is used for locating an index of the nearest time slice to the time requested by user.

Example:

B.2.2 equilibrium IDS

File equilibrium.py contains EquilibriumIDS class which is used to read data from equilibrium IDS and create CHERAB's EFITEquilibrium object via time method. It is later used to build core plasma model.

Load an equilibrium IDS:

```
equilibrium_ids = EquilibriumIDS(
        shot=134000,
        run=30,
        user="public",
        database="iter",
        version="3",
    )
shot is an IMAS database shot ID,
run is an IMAS data run ID,
user is an IMAS database user ID,
database is an IMAS database ID,
version is an IMAS major version number,
   Create an EFITEquilibrium object:
    equilibrium = equilibrium.time(time=-1)
time is a requested time. Here -1 for time is used to acquire a time slice in the middle of time
     range.
```

Other methods:

ids() returns the equilibrium IDS object;

psi_1d() returns one-dimensional Ψ profile stored in ids.time_slice[i].profiles_1d.psi. It is implemented in case if core_profiles (section B.2.3) IDS does not contain its own profile.

B.2.3 core_profiles IDS

File core_profiles.py contains CoreProfilesIDS class which is used to read data from core_profiles IDS and create CHERAB's Plasma object via create_plasma method. It is

```
poses as core plasma model.
   Load a core_profiles IDS:
    core_profiles_ids = CoreProfilesIDS(
         shot=134000,
         run=30,
        user="public",
         database="iter",
         version="3",
    )
shot is an IMAS database shot ID,
run is an IMAS data run ID,
user is an IMAS database user ID,
database is an IMAS database ID,
version is an IMAS major version number,
   Create a Plasma object:
    core_plasma = core_profiles_ids.create_plasma(
         equilibrium=equilibrium,
        psi_1d=equilibrium.psi_1d(),
         integration_step=0.001,
         integration_samples=5,
        parent=world,
         transform=None,
        name="Core Plasma"
    )
equilibrium is an EFITEquilibrium object (section B.2.2);
psi_1d is one-dimensional \Psi profile (section B.2.2). It is used if one stored in core_profiles
     IDS is missing;
integration_step is step of volumetric integration in meters;
integration_samples is a number of integration samples;
parent is a parent node;
transform is transformation matrix;
name is a name of this plasma.
Note: for more information on integration_step and integration_samples refer to CHERAB's
documentation on Plasma. For more information on parent, transform and name refer to Ray-
sect's documentation on Node.
```

Note that create_plasma method does not require time value since it uses one that stored in equilibrium.

One required argument of create_plasma is equilibrium. It provides Ψ_{norm} distribution which is used to map density, temperature and bulk velocity distributions of plasma. Functions distribution_density, distribution_temperature and distribution_velocity are doing exactly that. *Note:* distribution_temperature tries to use average ion temperature or electron temperature if species' own temperature profile is absent in the IDS.

detect_species Function detect_species is used to recognize ion and neutral species from label given in IDS. Since there was no convention on the labeling at the time this appeared to be a huge problem. detect_species uses regular expressions to math species label along some other tricks. It returns CHERAB's Element object and species charge as a number (0 for neutrals).

species, charge = detect_species(

```
structure=core_profiles_ids.ids.profiles_1d[0].ion[0],
         ion=True
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
ion set to True for ion species or to False for neutrals. It changes regular expression pattern
     and sets charge to 0 for neutrals.
distribution_density
    n_d = distribution_density(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
```

This function checks if density profile stored in the IDS is correct: not empty, greater than zero, not equal to zero and not equal to 1.0. If it is not correct than message is produced and this species is not included in the plasma model.

```
distribution_temperature
```

```
t_d = distribution_temperature(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium,
         t_average=core_profiles_ids.ids.profiles_1d[i].t_i_average,
         t_electrons=core_profiles_ids.ids.profiles_1d[i].electrons.temperature
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
t_average is an average ion temperature profile stored is the IDS. It is used in case if ion
     temperature profile for requested species is absent.
t_electrons is an electron temperature profile. It is used in case if ion temperature profile for
     requested species is absent.
distribution_velocity
    v_d = distribution_velocity(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium,
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
```

B.2.4 edge_profiles IDS

sect's documentation on Node.

File edge_profiles.py contains EdgeProfilesIDS class which is used to read data from edge_profiles IDS and create CHERAB's Plasma object via create_plasma method. It is poses as edge plasma model.

```
Load a edge_profiles IDS:
    edge_profiles_ids = EdgeProfilesIDS(
        shot=134000,
        run=30,
        user="public",
        database="iter",
        version="3",
    )
shot is an IMAS database shot ID,
run is an IMAS data run ID,
user is an IMAS database user ID,
database is an IMAS database ID,
version is an IMAS major version number,
   Create a Plasma object:
    edge_plasma = edge_profiles_ids.create_plasma(
        time=0.0,
        equilibrium=equilibrium,
        integration_step=0.001,
        integration_samples=5,
        parent=world,
        transform=None,
        name="Edge Plasma"
    )
time is a requested time;
equilibrium is an EFITEquilibrium object (section B.2.2);
integration_step is step of volumetric integration in meters;
integration_samples is a number of integration samples;
parent is a parent node;
transform is transformation matrix;
name is a name of this plasma.
Note: for more information on integration_step and integration_samples refer to CHERAB's
documentation on Plasma. For more information on parent, transform and name refer to Ray-
```

```
distribution_density
    n_d = distribution_density(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium,
         interpolator=interp_options[interp],
         interpolation_data=interp_data[interp],
        mesh_lookup=te_mesh_lookup
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
interpolator – select an interpolator based on where values are defined: faces or nodes.
interpolation_data - select a function to process data based on where values are defined: faces
     or nodes.
mesh_lookup empty
   All interpolator, interpolation_data and mesh_lookup are defined in the body of create_plasma
method.
   This function checks if density profile stored in the IDS is correct: not empty, greater than
zero, not equal to zero and not equal to 1.0. If it is not correct than message is produced and
this species is not included in the plasma model.
distribution_temperature
    t_d = distribution_temperature(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium,
         interpolator=interp_options[interp],
         interpolation_data=interp_data[interp],
         mesh_lookup=te_mesh_lookup
```

t_average=core_profiles_ids.ids.profiles_1d[i].t_i_average,

```
t_electrons=core_profiles_ids.ids.profiles_1d[i].electrons.temperature
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
interpolator – select an interpolator based on where values are defined: faces or nodes.
interpolation_data - select a function to process data based on where values are defined: faces
     or nodes.
mesh\_lookup \ empty
t_average is an average ion temperature profile stored is the IDS. It is used in case if ion
     temperature profile for requested species is absent.
t_electrons is an electron temperature profile. It is used in case if ion temperature profile for
     requested species is absent.
distribution_velocity
    v_d = distribution_velocity(
         structure=core_profiles_ids.ids.profiles_1d[i].ion[j],
         symbol="D",
         charge=1,
         psi_normalised=psi_normalised,
         equilibrium=equilibrium,
         interpolator=interp_options[interp],
         interpolation_data=interp_data[interp],
         mesh_lookup=te_mesh_lookup
    )
structure is an IDS structure containing information on species (ids.profiles_1d[i].ion[j]
     or ids.profiles_1d[i].neutral[j]).
symbol is a species symbol. It is used for messages.
charge is a species charge. It is used for messages.
psi_normalised is a \Psi_{norm} profile. It is used to map density values.
equilibrium is an EFITEquilibrium object (section B.2.2).
interpolator – select an interpolator based on where values are defined: faces or nodes.
interpolation_data - select a function to process data based on where values are defined: faces
     or nodes.
mesh_lookup empty
```

detect_species is a copy of a function described in section B.2.3.

B.2.5 charge_exchange IDS

File charge_exchange.py contains ChargeExchangeIDS class which is used to read data from charge_exchange IDS and create different types of observers.

Load an charge_exchange IDS:

B.2.6 nbi IDS

File nbi.py contains NBIIDS class which is used to read data from nbi IDS and create CHERAB's Beam object via create_beam method. It is poses as DNB model.

Load an nbi IDS:

nbi_ids = NBIIDS(

```
shot=134000,
run=30,
user="public",
database="iter",
version="3",
)
shot is an IMAS database shot ID,
run is an IMAS data run ID,
user is an IMAS database user ID,
database is an IMAS database ID,
version is an IMAS major version number,
```

```
beam = nbi_ids.create_beam(
        time=0.0,
        plasma=core_plasma,
        atomic_data=adas,
        attenuation_instructions=attenuation_instructions,
        emission_instructions=emission_instructions,
        length=4.0,
        integration_step=0.001,
        integration_samples=5,
        parent=world
        transform=None,
        name="DNB"
    )
time is a requested time;
plasma is CHERAB's Plasma object;
atomic_data is modified CHERAB's OpenADAS object;
attenuation_instructions is a dictionary with attenuation instructions;
emission_instructions is a dictionary with emission instructions (see section B.6.3);
integration_step is step of volumetric integration in meters;
integration_samples is a number of integration samples;
parent is a parent node;
transform is transformation matrix;
name is a name of this plasma.
```

Note: for more information on integration_step and integration_samples refer to CHERAB's documentation on Beam. For more information on parent, transform and name refer to Raysect's documentation on Node.

At the time Beam supports model with only one beamlet and create_beam is designed with this in mind.

B.3 Setting the Wall

File cxrs/machine/pfc_mesh.py contains function load_pfc_mesh that is used to create a reactor wall model. All meshes are stored in machine/portplugs and machine/simple subfolders.

```
wall = load_pfc_mesh(
    path=os.path.join(os.path.dirname(__file__)),
    reflections=True,
    roughness={"Be": 0.26, "W": 0.29, "Ss": 0.13},
    detailed=False,
```

```
parent=world,
transform=None,
name="Wall",
)

path is a path to .rsm mesh files;
reflections sets reflections on or off;
roughness is a roughness dictionary for PFC materials ("Be", "W", "Ss");
detailed — Load the detailed wall model instead of the simple one;
parent is a parent node;
transform is a transformation matrix;
name is a name of this plasma.
```

Here reflections assigns material properties to appropriate wall segments if set to True and sets the wall as perfect absorber if set to False. It is used to effectively turn reflections on and off. roughness argument sets *roughness* of the materials assigned to the wall segments (beryllium, tungsten and stainless steel). For more information on parent, transform and name refer to Raysect's documentation on Node.

B.4 Observers

Module cxrs/observers contains several files defining different observer classes built upon Raysect's observers.

B.4.1 Base Class

File observers/base contains ObserverGroup class that represents a group of observers, for example CXRS sightlines. The class has observe method that is used to perform an observation by all observers in a group one by one and store the results. Methods display and savetxt are used to show registered spectra as an image or plot and save results in the text format respectively. They should be implemented in subclasses.

Example:

```
ObserverGroup(
    charge_exchange_ids=charge_exchange_ids,
    config="config.xml",
    wavelength_range=1,
    relative_error=0.05,
    scenario=scenario,
    parent=world,
    transform=None,
```

```
name="CXRS Edge Sightlines"
)
charge_exchange_ids is a ChargeExchangeIDS object (section B.2.5),
```

config is a path to configuration file,

wavelength_range is a number representing a wavelength range for observation (it is defined in a configuration file),

relative_error is a minimal value of a desired relative error (if achieved relative error is higher than this value, number of pixel samples will be increased and observation will be performed again until desired relative error is achieved),

scenario is a dictionary containing all simulation labels: shot number, run number, time, used emission types, etc.,

parent is a parent node in a scenegraph,

transform is a transformation matrix,

name is a name for this group of observers.

For more information on parent, transform and name refer to Raysect's documentation on Node.

B.4.2 Sightlines

```
sightlines = SightlineGroup(
   ids=charge_exchange_ids,
   config="config.xml",
   wavelength_range=1,
   relative_error=0.05,
   scenario=scenario,
   parent=world,
   transform=None,
   name="CXRS Sightlines",
)
```

 $\verb|charge_exchange_ids| is a \verb|ChargeExchangeIDS| object (section $B.2.5$),$

config is a path to configuration file,

wavelength_range is a number representing a wavelength range for observation (it is defined in a configuration file),

relative_error is a minimal value of a desired relative error (if achieved relative error is higher than this value, number of pixel samples will be increased and observation will be performed again until desired relative error is achieved),

scenario is a dictionary containing all simulation labels: shot number, run number, time, used emission types, etc.,

parent is a parent node in a scenegraph,

```
transform is a transformation matrix,
name is a name for this group of observers.
   Methods:
display used to show and save produced image(s):
     SightlineGroup.display(
         show=True,
         save=True,
         filename="cxrs_sightlines",
         dirname="output"
     )
     show - show produced images in separate window;
     save – save images to a disk;
     filename - name of the file to save to. Image is saved in .png format;
     dirname - name of the directory to save to.
savetxt used to save simulation results in text format:
     SightlineGroup.savetxt(
         filename="cxrs_sightlines",
         dirname="output"
     )
     filename - name of the file to save to. Text is saved in .txt format;
     dirname - name of the directory to save to.
draw_scheme used to draw a simple scheme of a diagnostic:
     SightlineGroup.draw_scheme(
         plane="xy",
         save=True,
         filename="cxrs_sightlines",
         dirname="output"
     )
     plane – produce a scheme in x-y ("xy") or r-z ("rz") plane;
     save - save images to a disk;
     filename – name of the file to save to. Image is saved in .png format;
     dirname - name of the directory to save to.
B.4.3 Optics
    optics = OpticsAssembly(
         target_distance=4.0,
         aperture_radius=0.05,
```

```
magnification=0.1,
    refractive_index=1.52,
    parent=None,
    transform=None,
    name="Lens",
)

target_distance - distance to a target in meters;
aperture_radius - radius of the aperture in meters;
magnification - lens' magnification;
refractive_index - refractive index of the lens' material;
parent is a parent node in a scenegraph,
transform is a transformation matrix,
name is a name for this group of observers.
```

For more information on parent, transform and name refer to Raysect's documentation on Node.

B.4.4 Fibres

```
fibres = FibreGroup(
    ids=charge_exchange_ids,
    config="config.xml",
    wavelength_range=1,
    relative_error=0.05,
    scenario=scenario,
    parent=world,
    transform=None,
    name="CXRS Fibres",
)
```

 $\verb|charge_exchange_ids| is a \verb|ChargeExchangeIDS| object (section B.2.5),$

config is a path to configuration file,

wavelength_range is a number representing a wavelength range for observation (it is defined in a configuration file),

relative_error is a minimal value of a desired relative error (if achieved relative error is higher than this value, number of pixel samples will be increased and observation will be performed again until desired relative error is achieved),

scenario is a dictionary containing all simulation labels: shot number, run number, time, used emission types, etc.,

parent is a parent node in a scenegraph,

```
transform is a transformation matrix,
name is a name for this group of observers.
   Methods:
display used to show and save produced image(s):
     FibreGroup.display(
          show=True,
          save=True,
          filename="cxrs_sightlines",
          dirname="output"
     )
     show - show produced images in separate window;
     save – save images to a disk;
     filename - name of the file to save to. Image is saved in .png format;
     dirname - name of the directory to save to.
savetxt used to save simulation results in text format:
     FibreGroup.savetxt(
          filename="cxrs_sightlines",
          dirname="output"
     )
     filename - name of the file to save to. Text is saved in .txt format;
     dirname - name of the directory to save to.
draw_scheme used to draw a simple scheme of a diagnostic:
     FibreGroup.draw_scheme(
         plane="xy",
          save=True,
          filename="cxrs_sightlines",
          dirname="output"
     )
     plane – produce a scheme in x-y ("xy") or r-z ("rz") plane;
     save - save images to a disk;
     filename – name of the file to save to. Image is saved in .png format;
     dirname - name of the directory to save to.
B.4.5 Camera
    ccd_camera = CCDCamera(
         config="config.xml",
```

parent=world,

```
transform=None,
        name="CXRS Fibres",
    )
config is a path to configuration file,
parent is a parent node in a scenegraph,
transform is a transformation matrix,
name is a name for this group of observers.
B.4.6 Other Observers
    fibre_bundle = FibreBundle(
        core_radius=0.5e-3,
        numerical_aperture=0.22,
        spectrum=,
        n_rows=3,
        n_cols=1,
        relative_error=0.05
        parent=world,
        transform=None,
        name="CXRS Fibres",
    )
core_radius - radius of a fibre's core in meters;
numerical_aperture - fibre's numerical aperture;
spectrum - Raysect's Spectrum object;
n_rows - number of rows;
n_cols - number of columns;
relative_error - desired relative error;
parent is a parent node in a scenegraph,
transform is a transformation matrix,
name is a name for this group of observers.
    scanner = Scanner(
        ids=charge_exchange_ids,
        config="config.xml",
        range_vertical=0.2,
        step_vertical=0.01,
        relative_error=0.05,
        scenario=scenario,
        parent=world,
```

```
transform=None,
         name="Scanner",
    )
charge_exchange_ids is a ChargeExchangeIDS object (section B.2.5),
config is a path to configuration file,
range_vertical - vertical range in meters;
step_vertical - size of the scanning step in vertical direction;
relative_error is a minimal value of a desired relative error (if achieved relative error is higher
     than this value, number of pixel samples will be increased and observation will be performed
     again until desired relative error is achieved),
scenario is a dictionary containing all simulation labels: shot number, run number, time, used
     emission types, etc.,
parent is a parent node in a scenegraph,
transform is a transformation matrix,
name is a name for this group of observers.
    total_radiance = TotalRadianceSightlines(
         ids=charge_exchange_ids,
         config="config.xml",
         relative_error=0.05,
         scenario=scenario,
         parent=world,
         transform=None,
         name="Scanner",
    )
charge_exchange_ids is a ChargeExchangeIDS object (section B.2.5),
config is a path to configuration file,
relative_error is a minimal value of a desired relative error (if achieved relative error is higher
     than this value, number of pixel samples will be increased and observation will be performed
     again until desired relative error is achieved),
scenario is a dictionary containing all simulation labels: shot number, run number, time, used
     emission types, etc.,
parent is a parent node in a scenegraph,
transform is a transformation matrix,
name is a name for this group of observers.
    spectrometer = Spectrometer(
         pixel_size=,
```

```
n_pixels=,
    wl_calibration=,
    int_calibration=,
    width=,
    inst_func="rect",
        transmission=1.0,
)

pixel_size ?
n_pixels ?
wl_calibration ?
int_calibration ?
width ?
inst_func ?
transmission ?
    Methods: ?
```

B.5 Populating CHERAB's Atomic Database

```
populate/create
    populate_more(
        download=True,
        repository_path=None,
        adas_path="/work/projects/adas/adas/"
    )
download - attempt to download the ADAS files if missing;
repository_path - alternate path for the OpenADAS repository;
adas_path - alternate path in which to search for ADAS files;
   populate/openadas
    adas = OpenADAS(
        data_path=None,
        permit_extrapolation=False,
        missing_rates_return_null=False
    )
data_path - path to atomic data repository;
permit_extrapolation - if True informs interpolation objects to allow extrapolation beyond
     the limits of the tabulated data;
missing_rates_return_null - if True, allows Null rate objects to be returned when the re-
     quested atomic data is missing.
```

B.6 Utility Functions

B.6.1 Getting Information on Plasma Parameters

```
info/
    \_ \_init\_.py
    _beam.py
    \_ equilibrium.py
    _passive.py
    _{
m plasma.py}
    \_ profiles.py
   beam.py
    beam_info(
        plasma=,
         id_dict=,
        plot_dict=,
        dirname=,
         config=,
    )
plasma
id\_dict
plot_dict
dirname
config
   equilibrium.py
    equilibrium_info(
         id_dict=,
        plot_dict=,
         filename=,
         dirname=,
    )
{\tt id\_dict}
plot_dict
filename
dirname
   passive.py
passive_info(
    plasma=,
```

```
beam=,
    id_dict=,
    plot_dict=,
    dirname=,
)
plasma
beam
id\_dict
plot_dict
dirname
  profiles.py
profiles_info(
    core_plasma=,
    edge_plasma=,
    beam=,
    id_dict=,
    plot_dict=,
    dirname=,
)
core_plasma
edge\_plasma
beam
id\_dict
plot_dict
dirname
      Parsing XML Configuration File
srt2bool(string="on")
bool2str(arg=True)
read_xml_entry(tree_element)
parse_user_options(section, config)
parse_emission_lines(config)
parse_diagnostic_geometry(config)
parse_dnb_parameters(config)
parse_wavelength_ranges(config)
```

B.6.3 Setting Emission Parameters

```
plasma_emission_parameters = plasma_emission(
    plasma=,
    lines=,
    bremsstrahlung=False,
    recombination=False,
    excitation=False,
    passive=False,
)

plasma
lines
bremsstrahlung
recombination
excitation
passive
```

B.6.4 Math Functions

B.6.5 Fitting Routine

B.6.6 Others

utility/annotation utility/data utility/timer utility/fit