

Negative viscosity and stochastic parameterizations of Kinetic Energy Backscatter (KEB) for NEMO 3.6. Implementation details.

Pavel Perezhogin

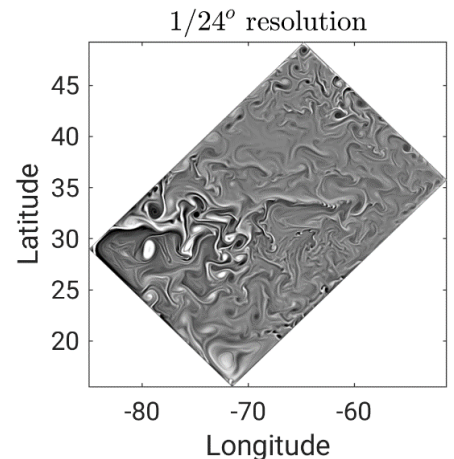
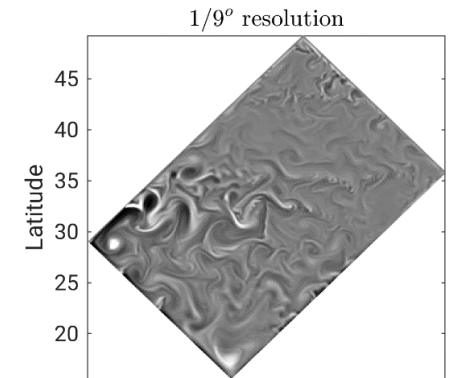
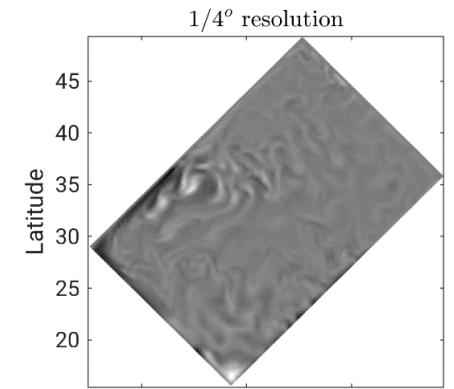
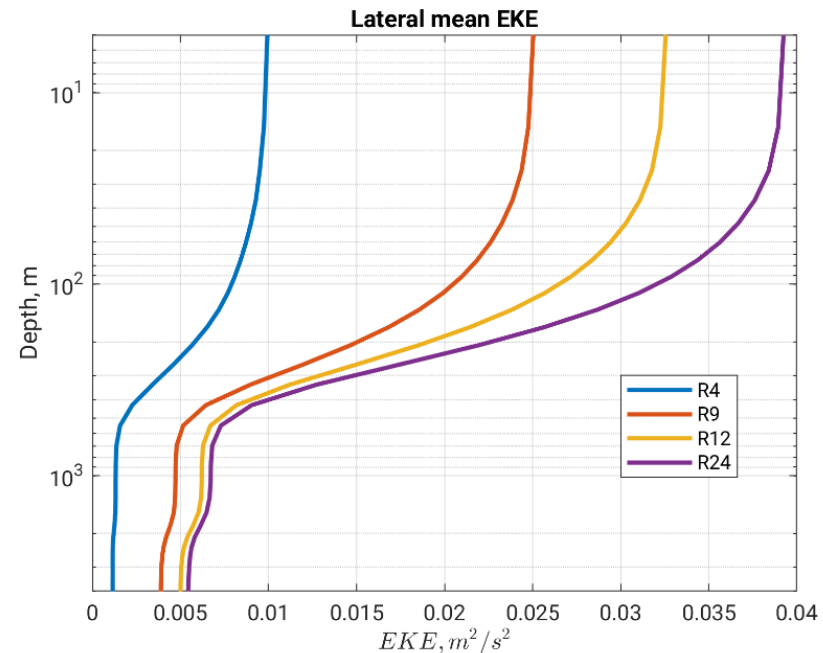
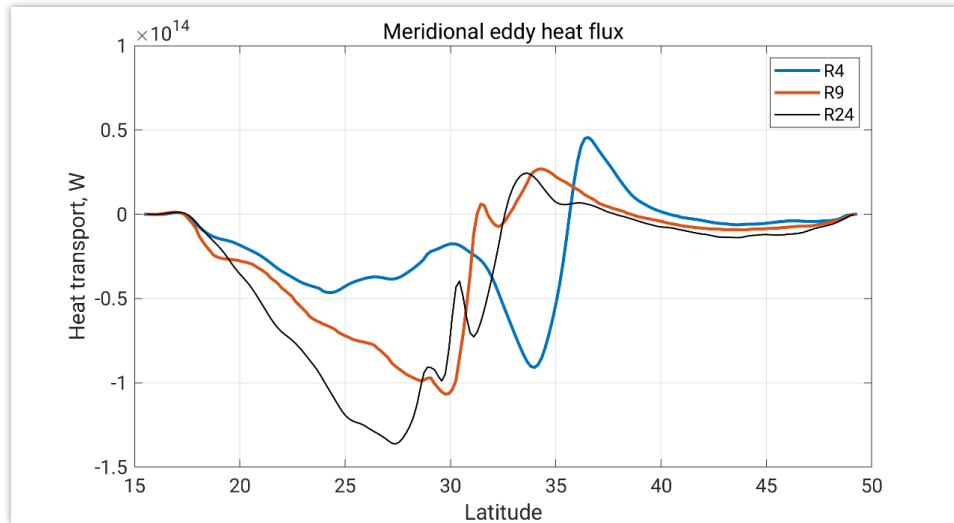
19 Jan 2022, Eddy Closure WG

The work was carried out in Institute of Numerical Mathematics, Moscow

Current affiliation: Courant Institute of Mathematical Sciences, New York

Introduction

- Mesoscale eddies form the most of the kinetic energy in ocean
- Mesoscale eddies only weakly resolved in eddy-permitting simulations
- Eddy kinetic energy and eddy heat fluxes strongly depend on the grid resolution



Eddy-permitting models weakly resolve energy cycle

- The energy of mesoscale eddies is given by two energy pathways:
 - APE->KE conversion
 - Inverse energy cascade
- Both pathways are weakly resolved in eddy-permitting simulations

Fox-Kemper, B., & Marsland, S. (Eds.). (2020). Sources and sinks of ocean mesoscale eddy energy.

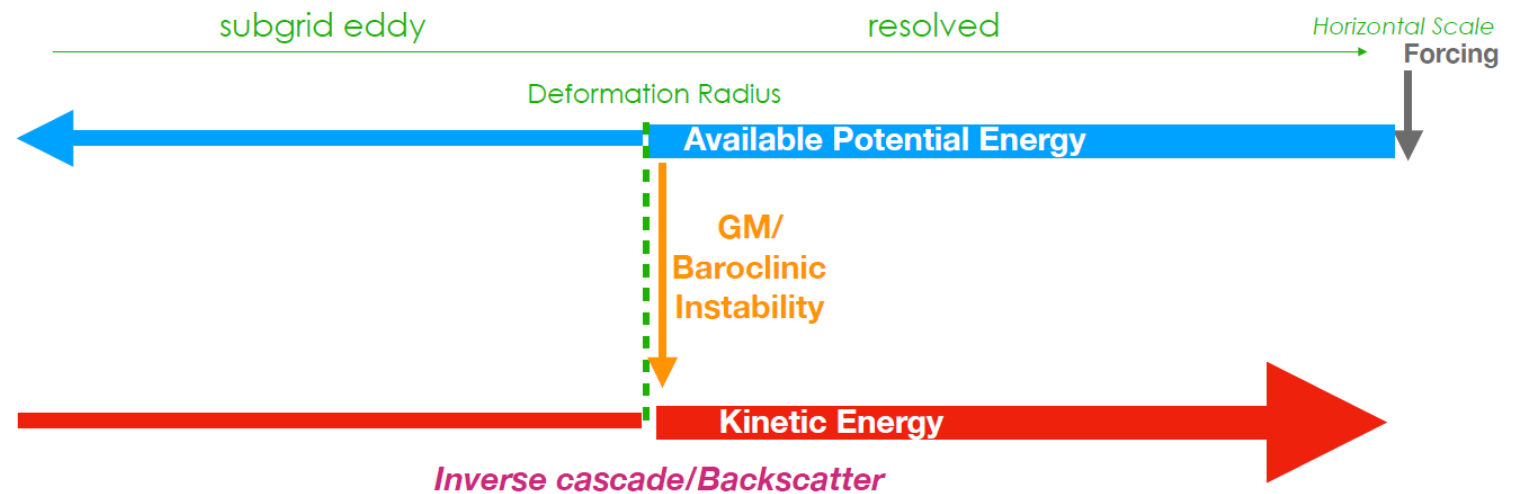


Figure 1. Schematic of energy transfer between reservoirs — potential (blue) and kinetic (red) energy — and horizontal scales (green).

Subgrid stress and backscatter

- The role of subgrid eddies in resolved dynamics is guided by the subgrid stress tensor:

$$\tau_{ij} = \overline{u_i u_j} - \bar{u}_i \bar{u}_j$$

$\overline{(\cdot)}$ – spatial filter

- Pointwise energy exchange with subgrid eddies can be of either sign:

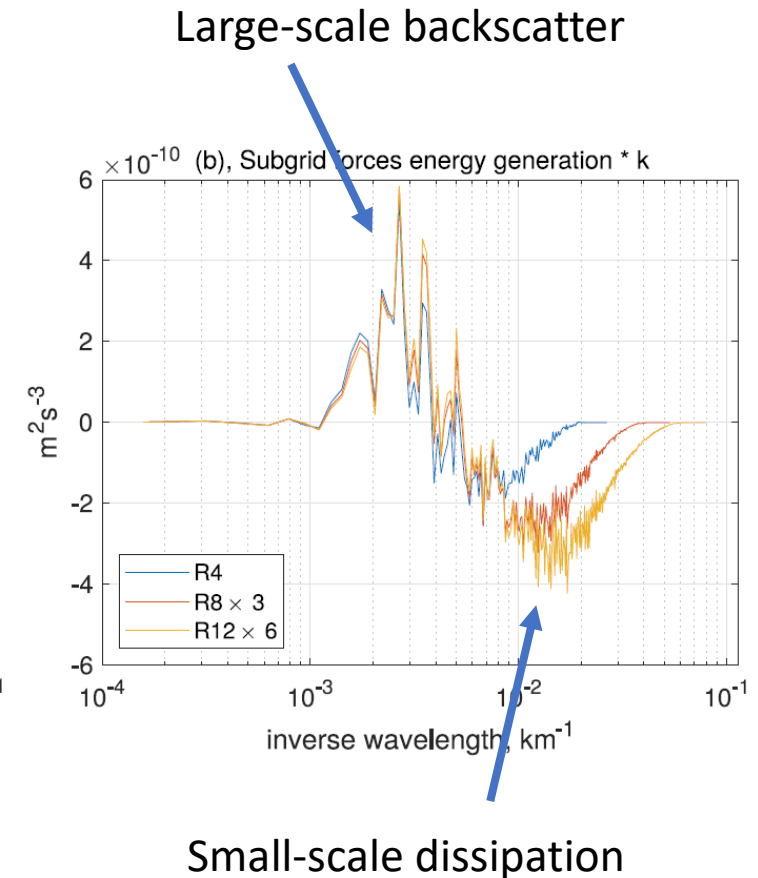
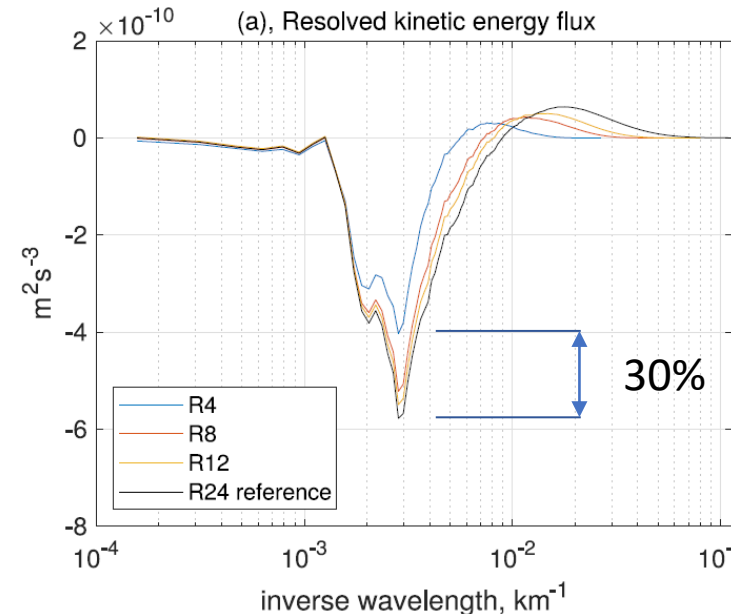
$$-\tau_{ij} S_{ij} > 0 - \text{dissipation}$$

$$-\tau_{ij} S_{ij} < 0 - \textbf{backscatter}$$

Definition for 3D turbulence. Meneveau 2000

Backscatter enhances inverse energy cascade

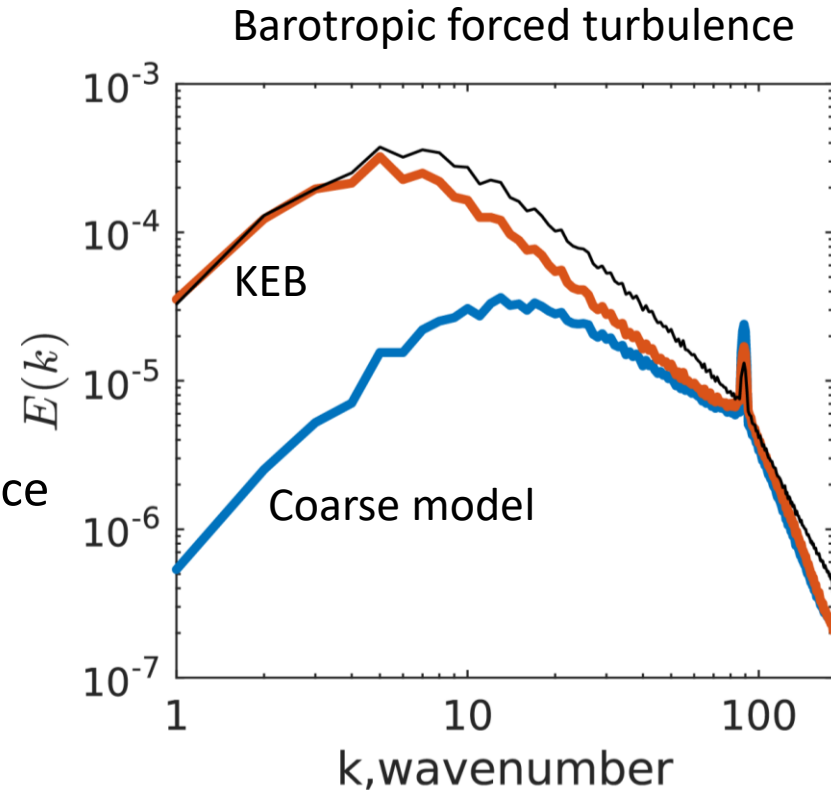
- *Hi-res* simulations show:
 - Resolved part of energy redistribution toward large scales (*due to resolved advection*) decreases with coarsening the resolution
 - Subgrid eddies compensate for unresolved inverse cascade
 - **Backscatter:** Subgrid eddies inject kinetic energy into large scales



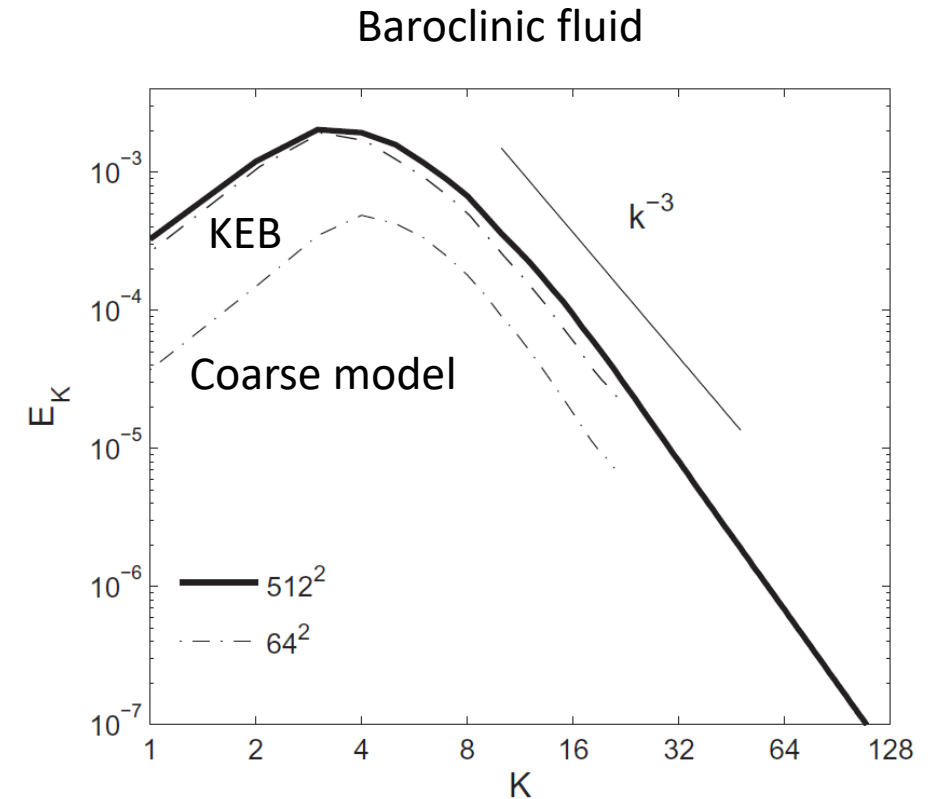
Backscatter is more important in baroclinic fluids

Barotropic fluid:
Improved kinetic energy in large scales

Baroclinic fluid:
Energizing eddies, we enhance energy extraction from the mean flow



Perezhogin et. al. 2019



Jansen Held 2014

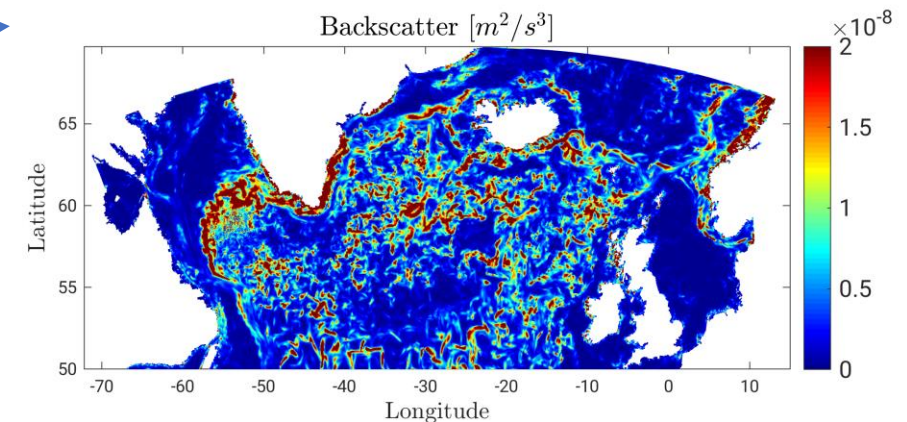
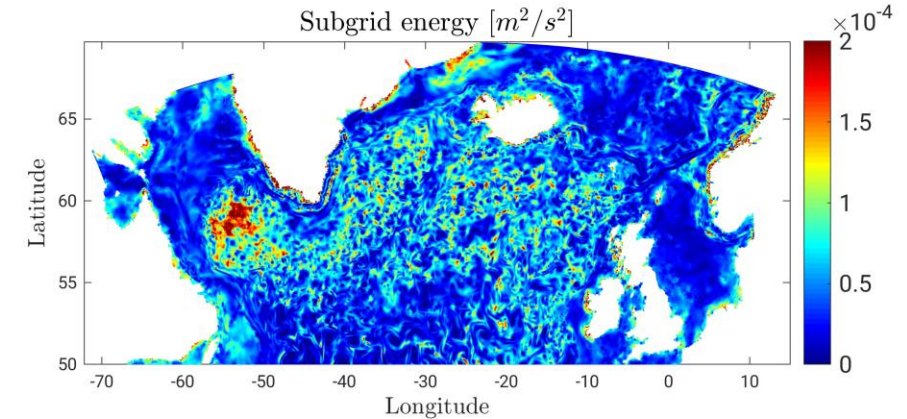
To conclude the introduction

- There are two popular approaches to simulate KEB:
 - Stochastic: Berner 2009, Jansen & Held 2014, Grooms & Majda 2015
 - Negative viscosity: Jansen & Held 2014, Juricke et.al. 2019, Bachman 2019
- Some elements of these approaches are implemented in NEMO 3.6

History of development

- Early 2019
 - KEB in cartesian coordinates in **NEMO 3.6**
 - Perezhugin, P. "Deterministic and stochastic parameterizations of kinetic energy backscatter in the NEMO ocean model in Double-Gyre configuration." IOP Conference Series. 2019.
 - Perezhugin, Pavel A. "Testing of kinetic energy backscatter parameterizations in the NEMO ocean model." RJNMM. 2020.
- Late 2019
 - Curvilinear coordinates
 - Cooperation with *Polina Verezhenskaya* from Shirshov Institute of Oceanology, Moscow
 - 1 model run in North Atlantic **NNATL12 in NEMO 4**
- August 2021
 - Cooperation with *Colin Jones* from Metoffice
 - Large update of the code for **NEMO 3.6**
 - <https://github.com/Pperezhugin/Kinetic-energy-backscatter-for-NEMO>

Two years old picture



Code overview

$$\dot{E}_{diss} = \nu_4 (\Delta \mathbf{u})^2$$

dynldf_bilap.F90

$$\dot{E}_{source} = F(c_{diss} \dot{E}_{diss})$$

KEB_apply()

$$\frac{de}{dt} = \dot{E}_{source} - \dot{E}_{back}$$

KEB_negvisc_tendency()

$$\psi(x, y, z) \sim \sqrt{\dot{E}_{source}} \phi(x, y)$$

KEB_AR1_tendency()

Statistics is collected
KEB_statistics()

stp	99.5%	98.156s
tra_adv	19.2%	18.899s
zdf_tke	15.1%	14.882s
keb_apply	12.7%	12.544s
keb_ar1_tendency	10.2%	10.047s
filter_laplace_t3d_ntimes	2.1%	2.061s
loop in keb_apply at KEB_module.f90:2	0.3%	0.296s
loop in keb_apply at KEB_module.f90:2	0.1%	0.128s
mpp_ink_3d	0.0%	0.012s
dyn_zdf	7.9%	7.830s
dyn_jdf	7.3%	7.184s

(b) EXP_AR1

stp	99.5%	101.104s
tra_adv	18.5%	18.809s
keb_apply	15.5%	15.714s
keb_negvisc_tendency	10.2%	10.354s
klower_cdiss	2.8%	2.870s
filter_laplace_t3d_ntimes	2.1%	2.143s
loop in keb_apply at KEB_module.f90:2	0.3%	0.327s
keb_statistics	0.0%	0.020s
zdf_tke	14.8%	15.005s
dyn_jdf	6.9%	6.990s
dyn_zdf	6.6%	6.742s

(c) EXP_negvisc

Design of the filters

- Filter is based on the diffusion operator:

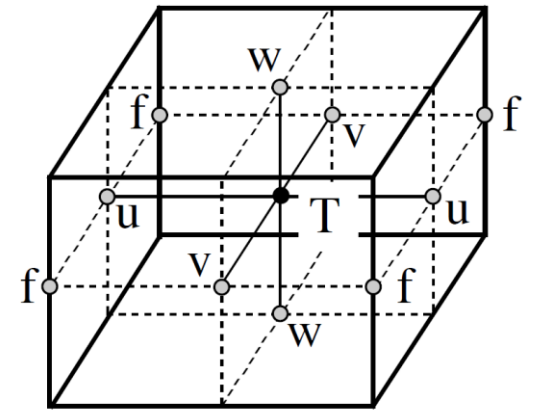
$$F(\psi(x, y, \sigma)) = 1 + \nabla_{\sigma} \left(\begin{bmatrix} \frac{dx^2}{8} & 0 \\ 0 & \frac{dy^2}{8} \end{bmatrix} \nabla_{\sigma} \psi \right)$$

∇_{σ} – Nabla operator along vertical coordinate isosurface $\sigma = \text{const}$

- In Cartesian coordinates:

- $F(\psi_{i,j}) = \frac{1}{8} (4\psi_{i,j} + \psi_{i+1,j} + \psi_{i-1,j} + \psi_{i,j+1} + \psi_{i,j-1})$

- Removes check-board noise
- Neumann or Dirichlet B.C. are possible
- Conservative



Source of subgrid energy (\dot{E}_{source})

- Dissipation by biharmonic eddy viscosity has sign-definite form:

$$\dot{E}_{diss}(x, y, \sigma) = |\nu_4|(\Delta_\sigma \mathbf{u})^2$$

- Source of subgrid energy:

$$\dot{E}_{source}(x, y, \sigma) = F^n(c_{diss}\dot{E}_{diss})$$

- Parameter $c_{diss} \leq 1$ may be constant or given by Klöwer et.al. 2018:

$$c_{diss}(x, y, \sigma) = \left(1 + \frac{R_o}{R_{diss}}\right)^{-1}, R_o = \frac{|D|}{f}$$

Only biharmonic viscosity is provided in the code. Easily extended to the Smagorinsky biharmonic viscosity

```
ndiss          = 2  
dirichlet_filter = .true.
```

```
constant_cdiss = .false.  
R_diss=0.5
```

Negative viscosity KEB. Operator for returning energy

- Vector-invariant form of Laplacian operator:

$$\frac{\partial \mathbf{u}_h}{\partial t} = -\nabla_\sigma \times (v_2 F^n(\omega) \mathbf{k})$$

- B.C.: zero vorticity ω on the boundary
- Divergence-free
- Sign-definite formula for returning energy:

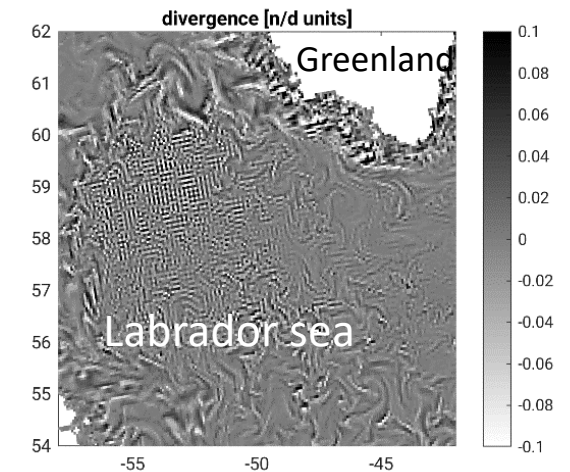
$$\dot{E}_{back}(x, y, \sigma, t) = |v_2| \left(F^{\frac{n}{2}}(\omega) \right)^2$$

note $n/2$ as filter is self-adjoint

- Formula for negative viscosity:

$$v_2(x, y, \sigma, t) = c_{back} \sqrt{dxdy} \sqrt{\max(2e, 0)}$$

nback = 4



c_back = 0.1

Negative viscosity KEB. Subgrid energy equation

- Time evolution for 3D subgrid energy:

$$\frac{\partial e(x, y, \sigma, t)}{\partial t} + \nabla(\mathbf{u}e) = \dot{E}_{source} - \dot{E}_{back} + \nu_e \Delta_\sigma e$$

```
tke_adv = .true.  
dirichlet_TKE = .true.  
nu_TKE = 300.
```

- Due to sign-definite conversion terms \dot{E}_{source} and \dot{E}_{back} , energy is positive:

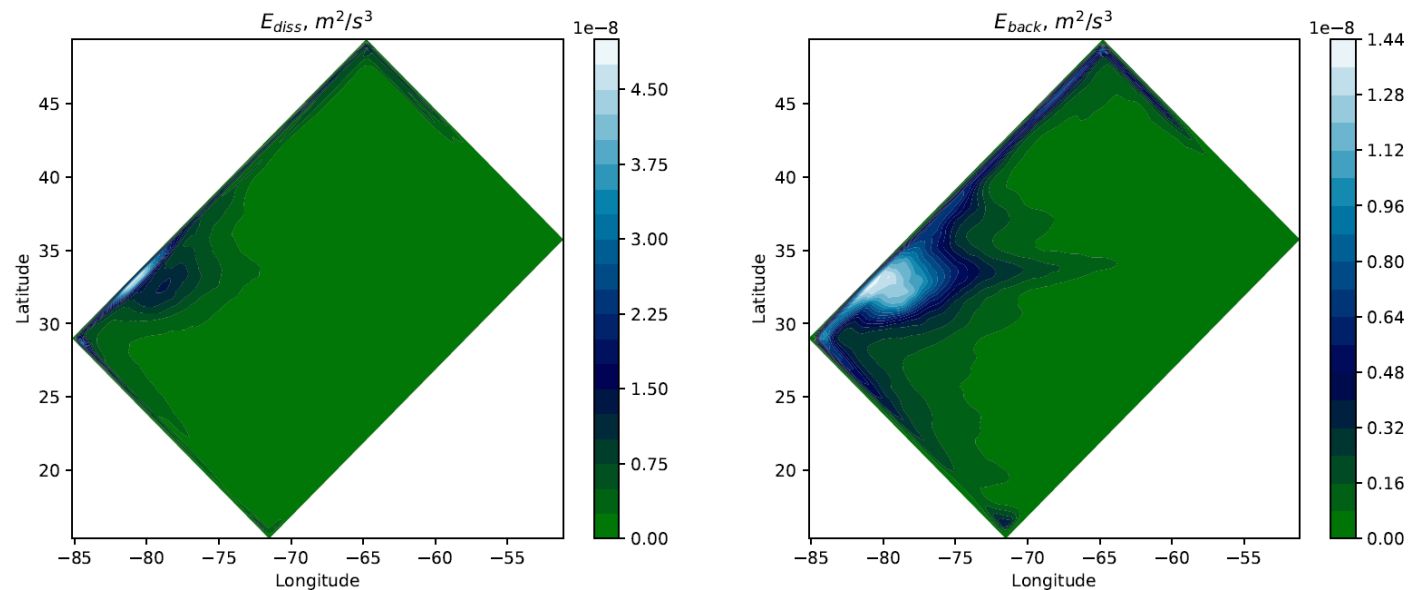
$$e \geq 0$$

- Advantages of positive e :
 - Fast check for bugs in new configurations
 - Negative energy can accumulate along boundaries

Is advection of subgrid energy important?

$$\frac{\partial e}{\partial t} + \nabla(\mathbf{u}e) = \dot{E}_{source} - \dot{E}_{back} + \nu_e \Delta_\sigma e$$

- There are evidences that subgrid energy can be transported on large distances [Grooms 2013]
- Large part of transport need not to be investigated – it is advection of subgrid energy. It makes KEB to be Galilean invariant:
 - Backscattered energy is injected downstream the mean flow



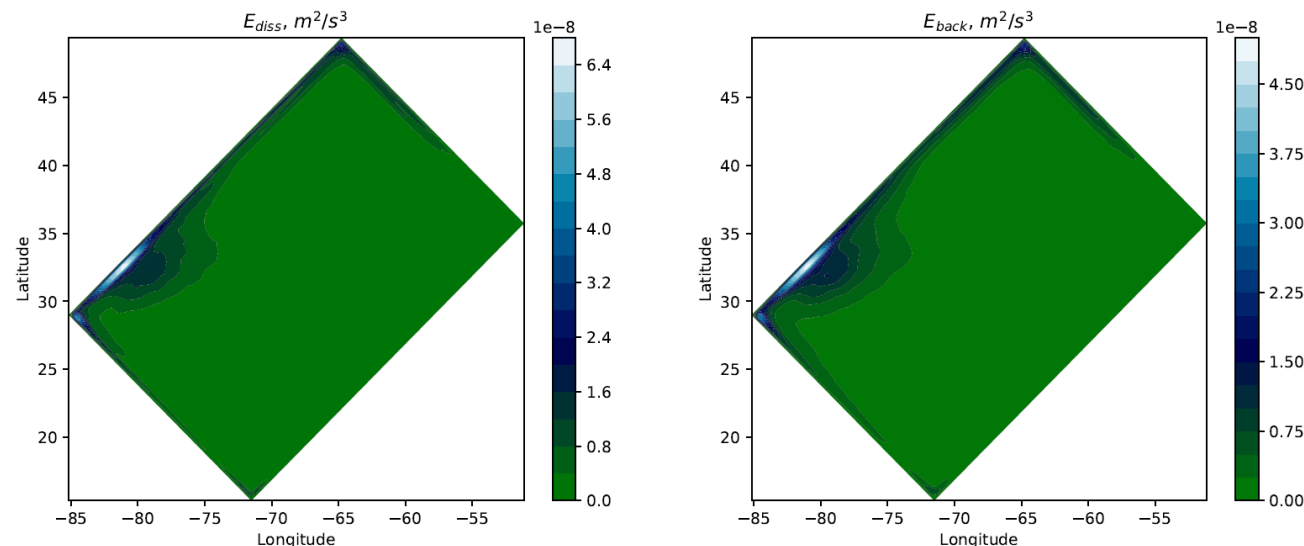
How to choose c_{back} ?

$$v_2(x, y, \sigma, t) = c_{back} \sqrt{dx dy} \sqrt{\max(2e, 0)}$$

- Parameter c_{back} controls amount of subgrid energy [Jansen 2015]
- Also we may want to optimize subgrid energy lifetime:

$$T = \frac{\langle e \rangle}{\langle \dot{E}_{back} \rangle} \sim \langle e \rangle$$

- For lifetime of interest ($T \approx 30days$), when advection on large distances is possible, the level of subgrid energy is too large
- For realistic level of subgrid energy lifetime is only $T \approx 2days$



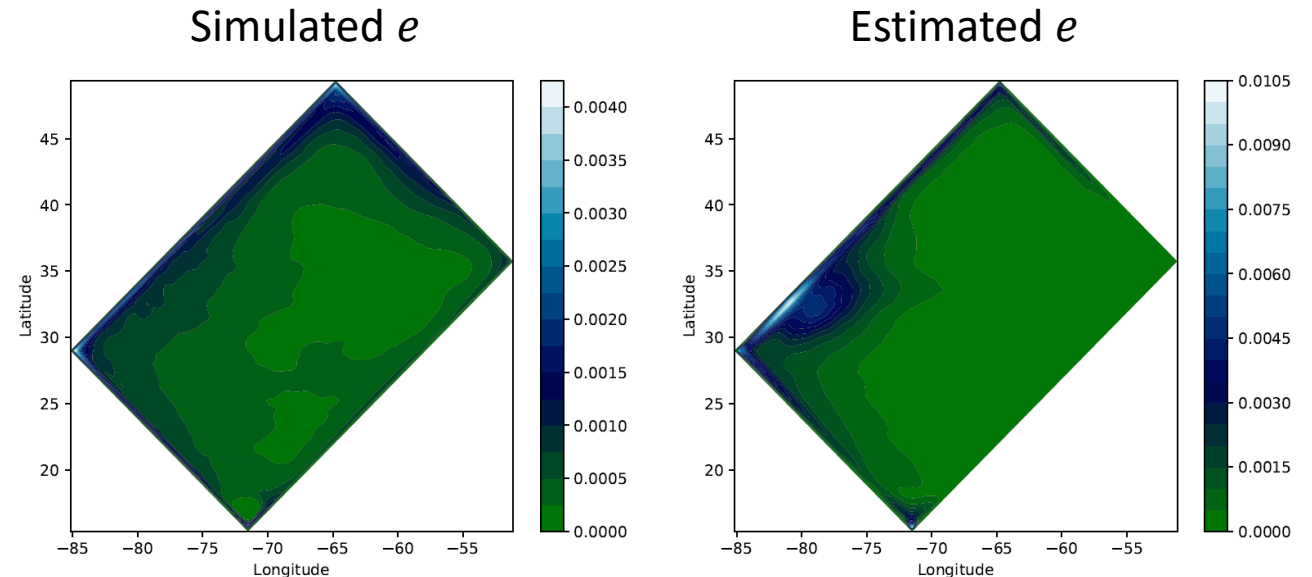
The current model does not allow to simulate both long subgrid energy propagation and realistic subgrid energy level

How to calibrate parameters of subgrid energy equation?

- Compare subgrid energy with estimation [Pomraning&Rutland2002]

$$e \equiv \frac{1}{2} tr(\tau_{ij}) \approx \frac{1}{2} tr(\tau_{ij}^{VGM}) = \frac{\Delta_f^2}{48} (|S|^2 + \omega^2)$$

- Too much subgrid energy near boundaries?
 - Apply Dirichlet B.C. for subgrid energy and filters, increase R_{diss} in Klöwer formula
- Too much subgrid energy?
 - Increase c_{back}



Stochastic KEB

- Return energy as follows:

$$\frac{\partial \mathbf{u}_h}{\partial t} = \begin{bmatrix} f_x(x, y, \sigma, t) \\ f_y(x, y, \sigma, t) \end{bmatrix}$$

f_x, f_y – stochastic fields

- Amount of returning energy:

$$\tau(\langle f_x^2 \rangle + \langle f_y^2 \rangle)$$

T_decorr=86400.

where τ – correlation time

- Desired properties:
 - Divergence-free
 - Correlated in space
 - Prescribed energy input: \dot{E}_{source}

Stochastic KEB. Generation of noise with desired properties

- The proposed form for stochastic field:

$$\begin{bmatrix} f_x \\ f_y \end{bmatrix} = \nabla_\sigma \times (\psi \mathbf{k})$$

- Where streamfunction is given by:

$$\psi(x, y, \sigma) = F^n(A(x, y, \sigma)\phi(x, y))$$

nstoch=6

$\phi(x, y)$ – grid white noise $N(0,1)$

- According to O’Neil “A new stochastic backscatter model for large-eddy simulation of neutral atmospheric flows” 2016:

$$\langle f_x^2 \rangle + \langle f_y^2 \rangle \approx A^2 \cdot 2\rho_0(1 - \rho_1) \left(\frac{1}{dx^2} + \frac{1}{dy^2} \right)$$

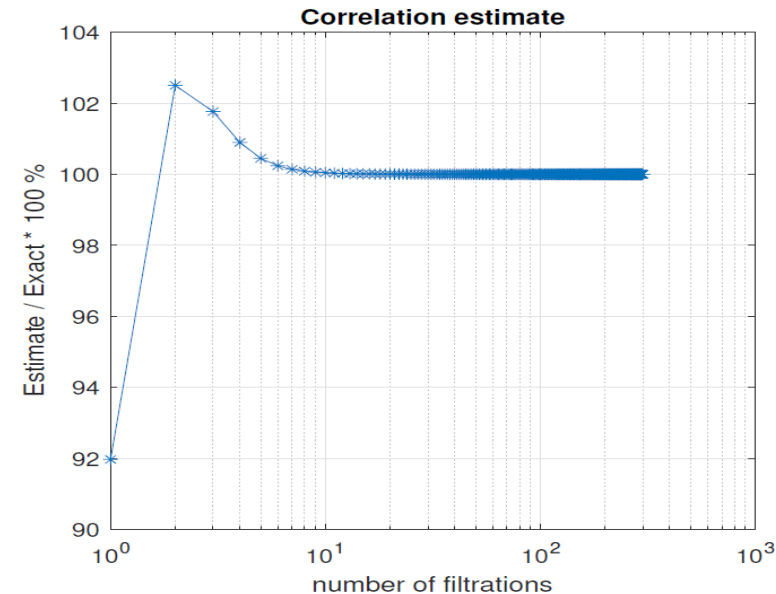
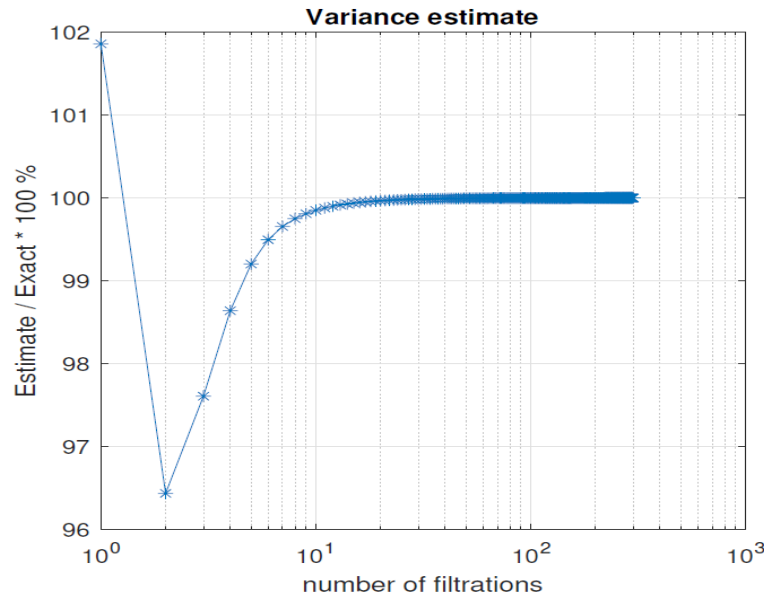
- Final formula for A :

$$A(x, y, \sigma) = \sqrt{\dot{E}_{source}} \left(2\tau\rho_0(1 - \rho_1) \left(\frac{1}{dx^2} + \frac{1}{dy^2} \right) \right)^{-1/2}$$

Moments of the filter

- For filter F^n define the moments:

$$\rho_0 = \sum_{i,j} w_{i,j}^2 \rightarrow \frac{1}{n\pi}$$
$$\rho_1 = \frac{\sum_{i,j} w_{i,j} w_{i+1,j}}{\rho_0} \rightarrow \exp\left(-\frac{1}{n}\right)$$



Correlation in time

- Numerical integration with modified Leap-Frog scheme
- Introduce time correlation without changing variance [Schumann1995]:

$$\hat{\psi}^{n+1} = \left(1 - \frac{dt}{\tau}\right) \hat{\psi}^n + \left(\frac{dt}{\tau} \left(2 - \frac{dt}{\tau}\right)\right)^{1/2} \psi$$

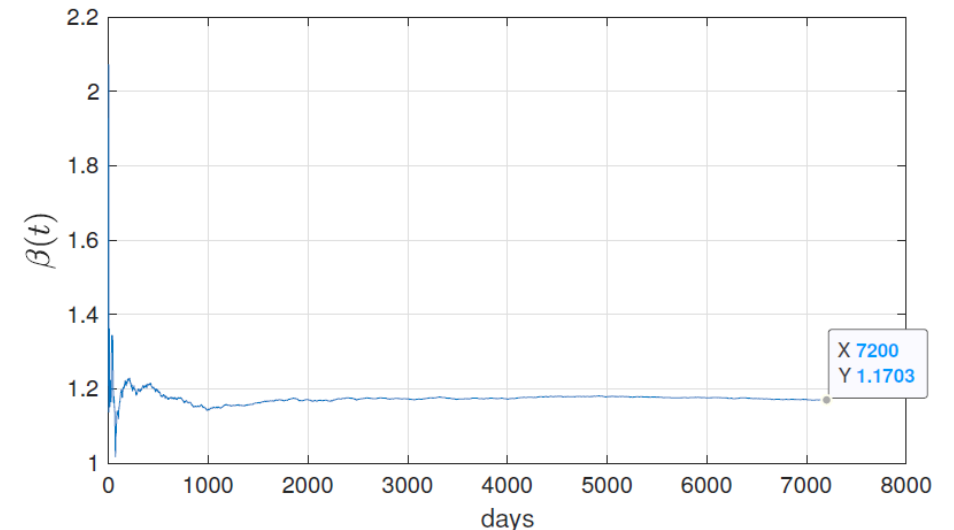
- A posteriori correction:

$$\beta(t) = \frac{\langle \dot{E}_{source} \rangle_{space \times time}}{\langle \hat{\mathbf{f}} \cdot \mathbf{u}_h \rangle_{space \times time}}$$

- Final tendency:

$$\frac{\partial \mathbf{u}_h}{\partial t} = \beta(t) \begin{bmatrix} \hat{f}_x \\ \hat{f}_y \end{bmatrix}$$

Correction increases input +37%



How to be sure that in complex configuration the code works as expected? KEB_test

- Reference values for some integrals provided

```
! ----- diffusion ----- !
conservation:
no-flux (1e-20 ok): -1.064180091878790E-021
Dirichlet (1e-4 ok): -1.252268632356181E-004

positiveness:
min ok: 1.850456085375350E-007 1.856420384649364E-007
max ok: 8.217884550423646E-002 8.206124773202066E-002

! ----- z-filter ----- !
positiveness:
min ok: 1.850456085375350E-007 1.850865339105703E-007
max ok: 8.217884550423646E-002 8.066954583786658E-002
conservation (1e-20 ok): -7.289300561377718E-020

! ----- local c_diss ----- !
min value for c_diss (0 ok): 0.220658658338855
max value for c_diss (1 ok): 0.999355240226306

! ----- laplace filter ----- !
filter_laplace_T3D_ntimes, no-flux:
conservation (1e-20 ok): -9.063963607859050E-022
min ok: 9.955149248888630E-013 8.190910945193417E-012
max ok: 1.660738531266926E-006 9.642326610140910E-007
```

Results.

$$\frac{\partial T}{\partial t} + \mathbf{U} \cdot \nabla T = F_T \frac{\partial S}{\partial t} + \mathbf{U} \cdot \nabla S = F_S$$

$$\rho = \rho_0(1 - a(T - T_0) + b(S - S_0))$$

$$\frac{\partial p}{\partial z} = -\rho g$$

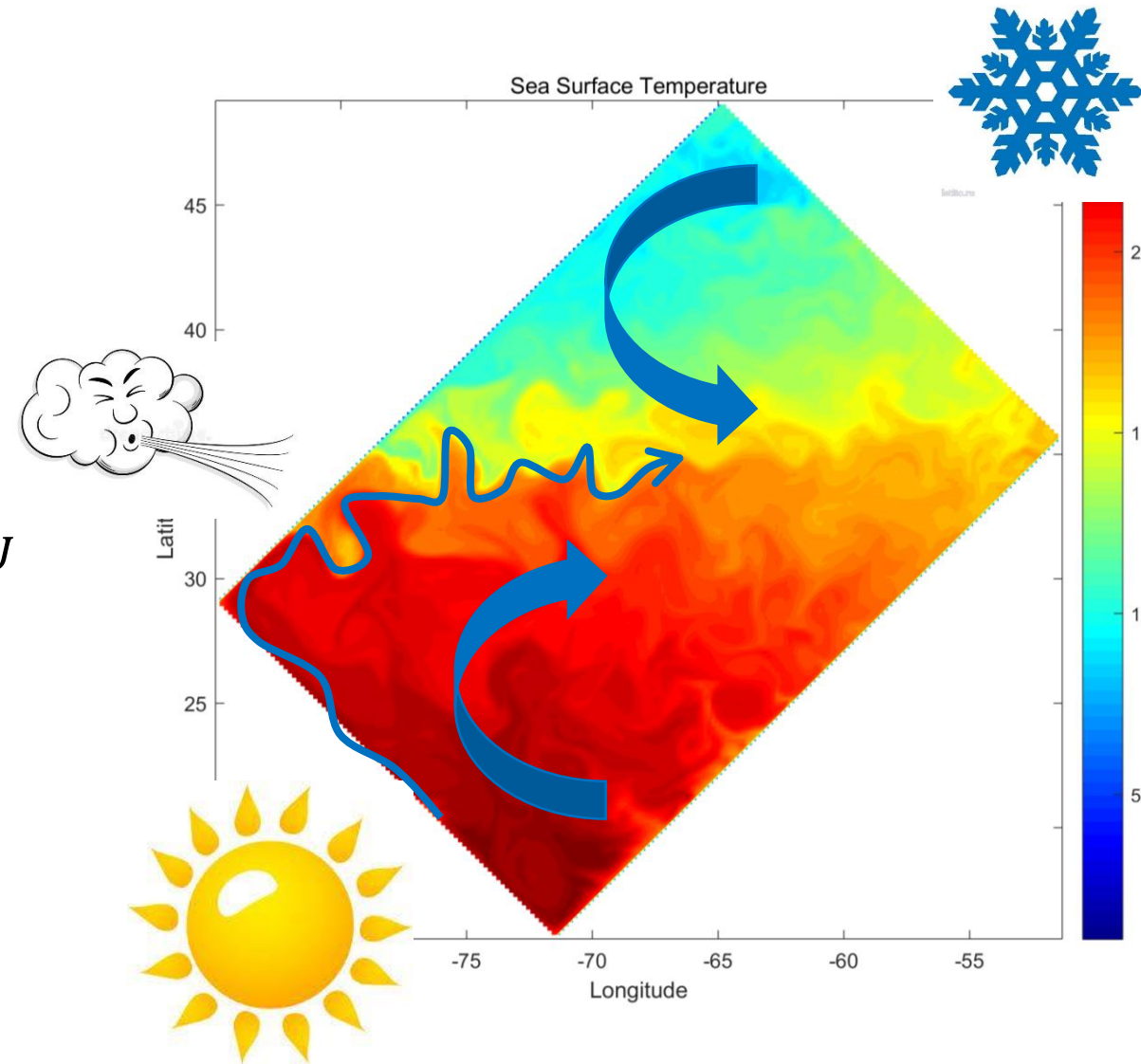
$$\frac{\partial \mathbf{U}_h}{\partial t} + \mathbf{U} \cdot \nabla \mathbf{U}_h + \mathbf{cor}_h = -\frac{1}{\rho_0} \nabla_h p + F_U$$

$$\nabla \cdot \mathbf{U} = 0$$

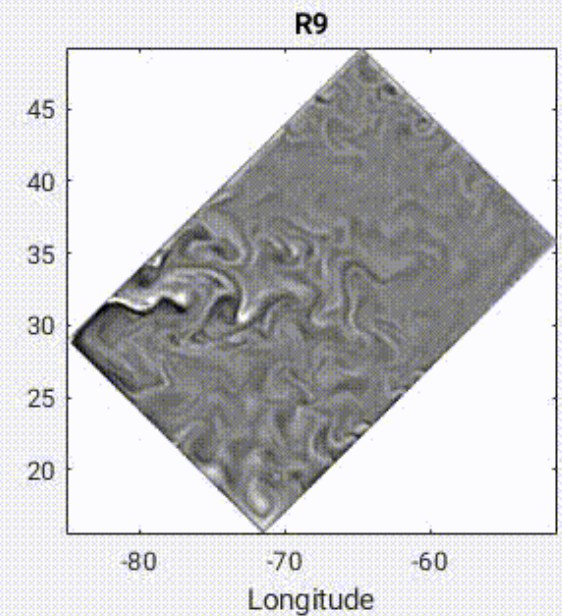
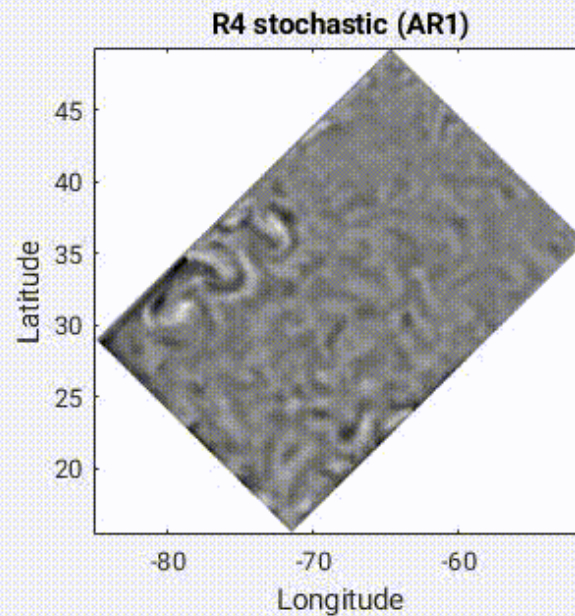
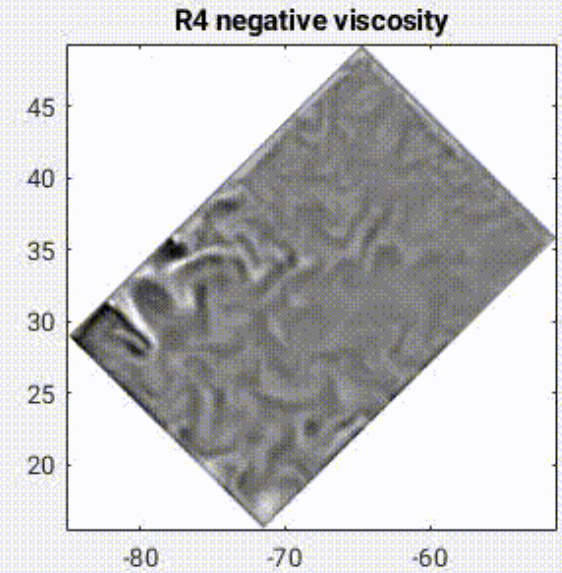
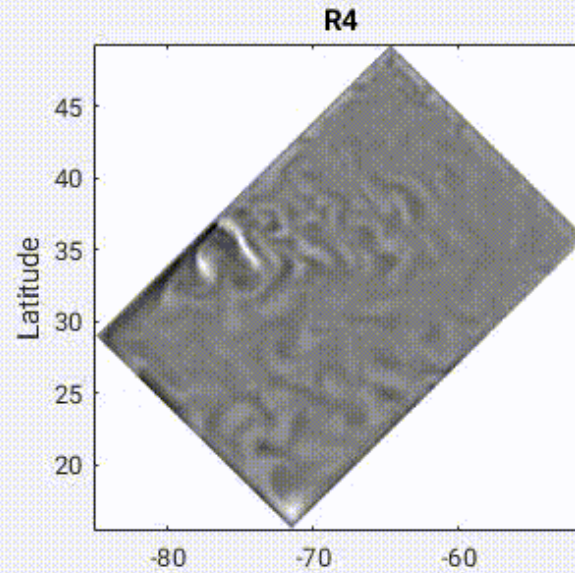
$$\frac{\partial \eta}{\partial t} = -H \nabla_h \cdot (\overline{\mathbf{U}_h})$$

$T, S, \mathbf{U}, \mathbf{U}_h, \overline{\mathbf{U}_h}, p, \eta, \rho, H$ – temperature; salinity; full, horizontal and vertically-averaged velocity; pressure; sea surface height; density; depth.

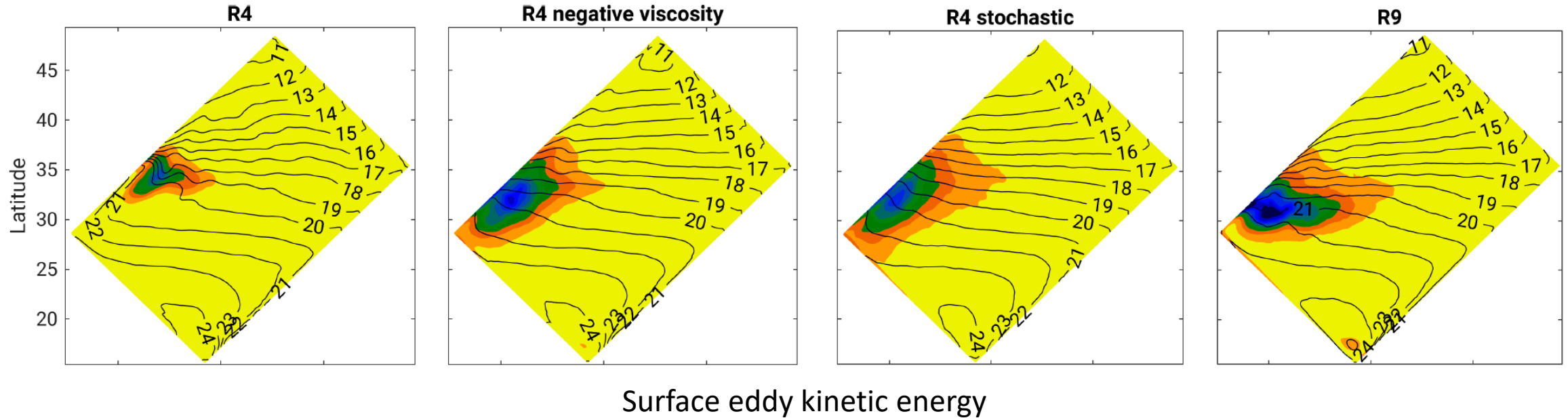
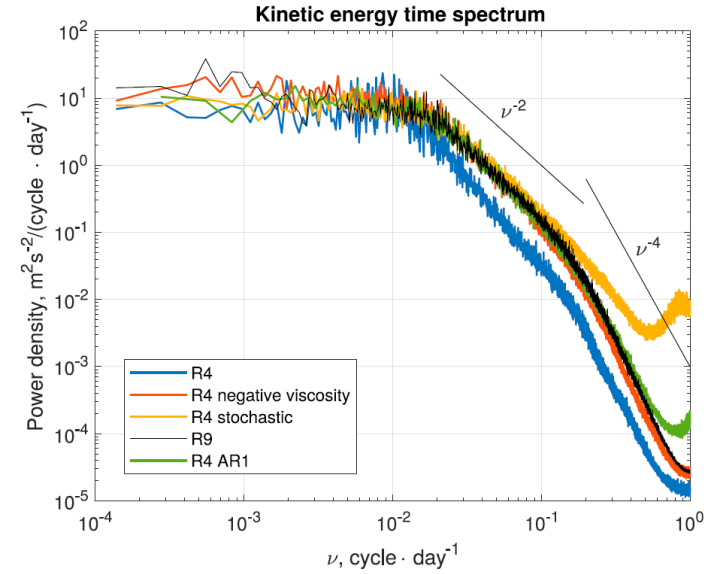
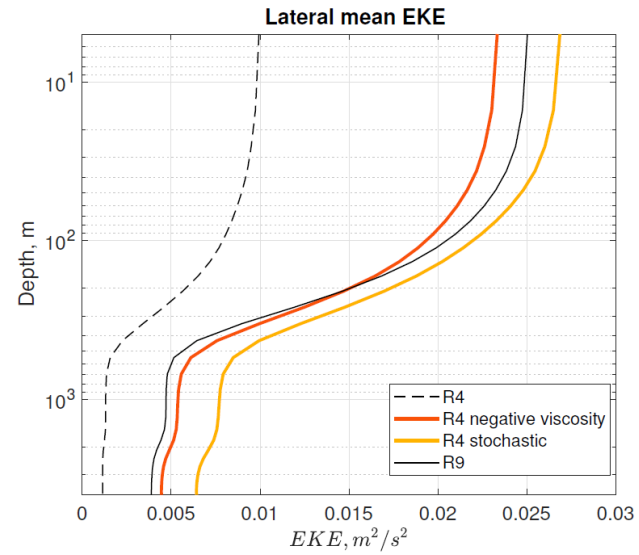
Lateral B.C.: zero heat and salinity flux, free-slip



Surface relative
vorticity in f
(Coriolis) units

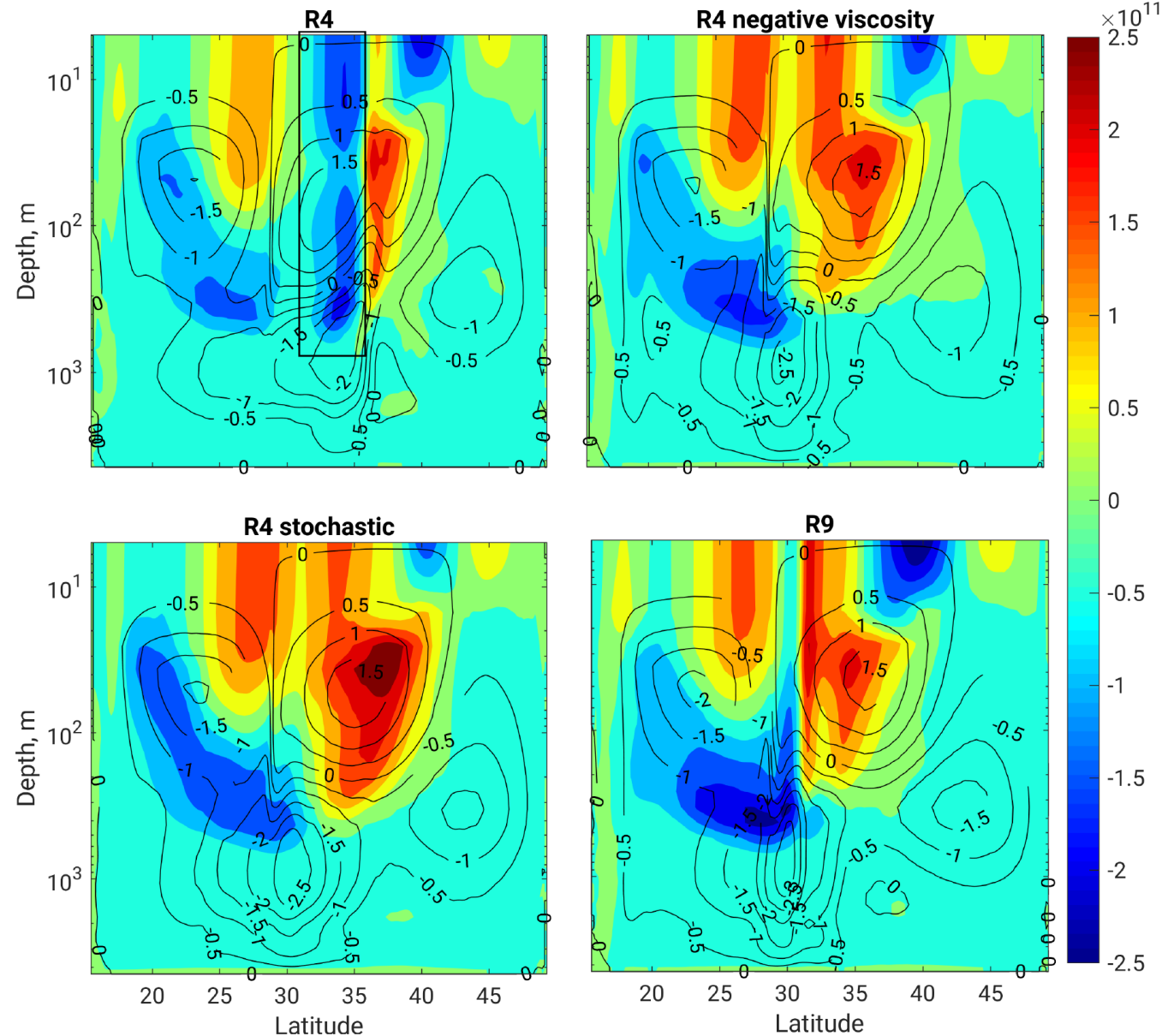
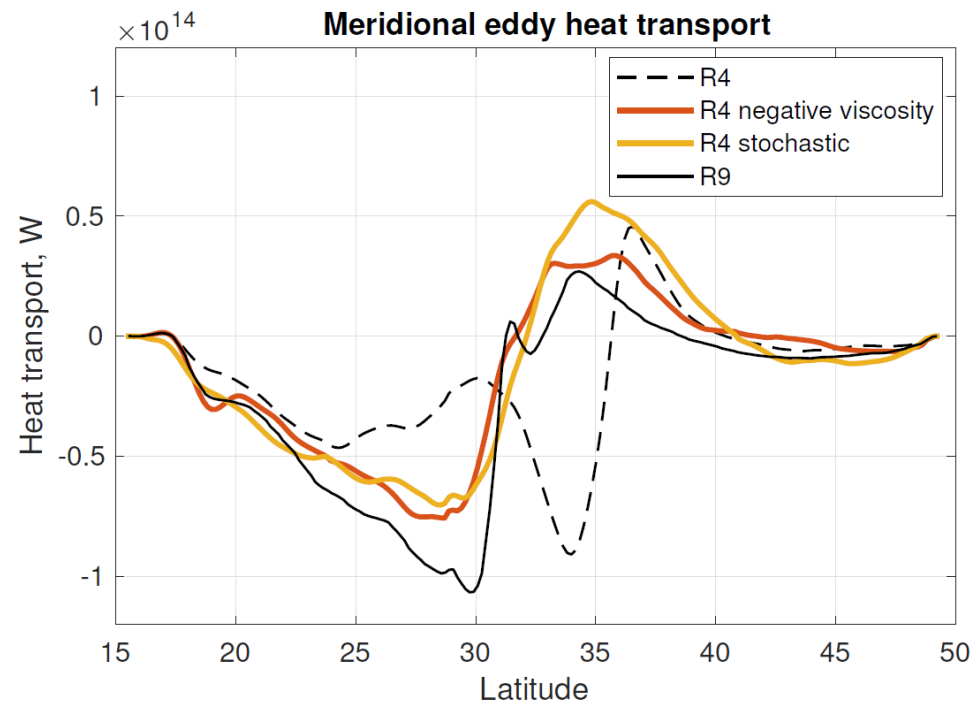


Turbulence statistics



Meridional eddy heat flux and MOC

- The latitude of extreme value shifts southward
- Eddy heat flux at the surface is closer to reference
- Bottom MOC cell shifts southwards



Conclusions

- Two types of kinetic energy backscatter parameterizations are implemented in NEMO 3.6 model
- They work in curvilinear coordinates and generalized vertical coordinate
- They can be used in complex configurations at eddy-permitting resolution
- They can be used together with GM parameterization
- The implementation is the most close to Juricke2019 and Berner2009, and a little less close to Jansen2019 and Bachman2019