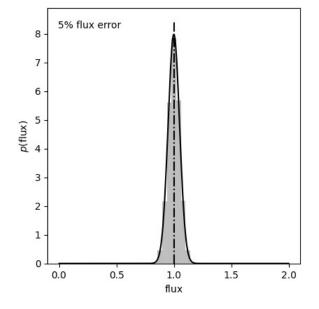
Assignment-1

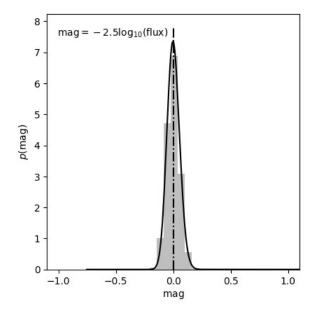
Pradeep Mundlik, AI21BTECH11022

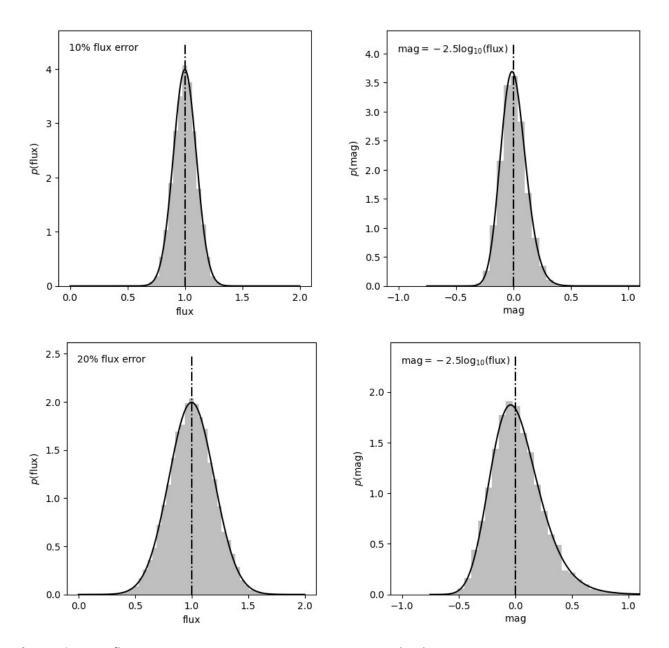
```
import numpy as np
from matplotlib import pyplot as plt
from scipy import stats
from scipy.stats import norm, cauchy
from astropy.stats import median_absolute_deviation
from astroML.stats import sigmaG
import pandas as pd
```

```
## Note: Code for this question is taken from book
def plot (err):
        # Create our data
        # generate 10000 normally distributed points
        np.random.seed(1)
        dist = norm(loc=1, scale=err)
        flux = dist.rvs(10000)
        flux fit = np.linspace(0.001, 2, 1000)
        pdf flux fit = dist.pdf(flux fit)
        # transform this distribution into magnitude space
        mag = -2.5 * np.log10(flux)
        mag fit = -2.5 * np.log10(flux fit)
        pdf mag fit = pdf_flux_fit.copy()
        pdf mag fit[1:] /= abs(mag fit[1:] - mag fit[:-1])
        pdf mag fit /= np.dot(pdf mag fit[1:], abs(mag fit[1:] -
mag fit[:-1]))
        #-----
        # Plot the result
        fig = plt.figure(figsize=(10, 5))
        fig.subplots_adjust(bottom=0.17, top=0.9, left=0.12,
right=0.95, wspace=0.3)
        ax = fig.add subplot(121)
        ax.hist(flux, bins=np.linspace(0, 2, 50),
histtype='stepfilled', fc='gray', alpha=0.5, density=True)
        ax.plot(flux_fit, pdf_flux_fit, '-k')
        ax.plot([1, 1], [0, max(pdf_flux_fit)+0.5], '-.',
color='#000')
        ax.set_xlabel(r'${\rm flux}$')
        ax.set ylabel(r'$p({\rm flux})$')
        ax.text(0.04, 0.95, f'{int(err*100)}% flux error', ha='left',
```

```
va='top', transform=ax.transAxes)
        # plt.grid()
        # next plot the magnitude distribution
        ax = fig.add subplot(122)
        ax.hist(mag, bins=np.linspace(-1, 2, 50),
histtype='stepfilled', fc='gray', alpha=0.5, density=True)
        ax.plot(mag_fit, pdf_mag_fit, '-k')
        ax.plot([0, 0], [0, max(pdf_mag_fit)+0.5], '-.', color='#000')
        ax.set_xlim(-1.1, 1.1)
        # ax.set ylim(0, 5)
        ax.set_xlabel(r'${\rm mag}$')
        ax.set_ylabel(r'$p({\rm mag})$')
        ax.text(0.04, 0.95, r'${\rm mag} = -2.5\log {10}({\rm flux})
$', ha='left', va='top', transform=ax.transAxes)
        # plt.grid()
        plt.show()
plot (0.05)
plot (0.10)
plot_{0.20}
```







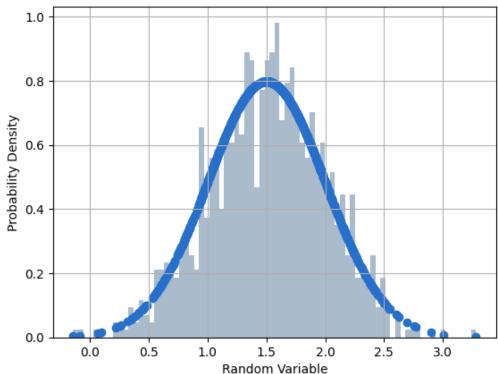
Obervation: As flux error increses, assymetricity in magnitude plot increases.

```
dist = norm(1.5, 0.5)
rv = dist.rvs(1000)
p = dist.pdf(rv)

plt.hist(rv, bins=80, density=True, histtype='barstacked', fc='#ABC')
plt.scatter(rv, p, color='#266FCB')
plt.grid()
plt.title("PDF of Normal Distribution of 1000 draws of mean 1.5 and standard deviation 0.5")
plt.xlabel("Random Variable")
```

```
plt.ylabel("Probability Density")
plt.show()
```

PDF of Normal Distribution of 1000 draws of mean 1.5 and standard deviation 0.5



```
print('Sample mean is: ', np.mean(rv))
print('Sample variance is: ', np.var(rv))
print('Skewnness is: ', stats.skew(rv))
print('Kurtosis is: ', stats.kurtosis(rv))
print('MAD is: ', median_absolute_deviation(rv))
print('$\sigma_G$ is: ', sigmaG(rv))

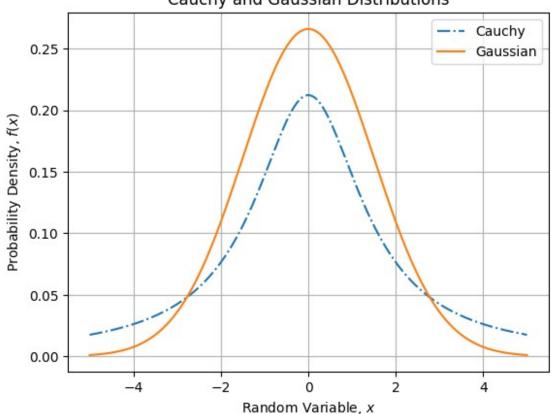
Sample mean is: 1.5106087998628224
Sample variance is: 0.2451947534991927
Skewnness is: -0.1153698344677762
Kurtosis is: -0.014897295036694125
MAD is: 0.3385080545162056
$\sigma_G$ is: 0.5063075940485616
```

```
dist1 = cauchy(0, 1.5)
dist2 = norm(0, 1.5)

x = np.arange(-5, 5, 1e-3)
plt.plot(x, dist1.pdf(x), linestyle = '-.')
plt.plot(x, dist2.pdf(x))
```

```
plt.legend(['Cauchy','Gaussian'])
plt.title("Cauchy and Gaussian Distributions")
plt.xlabel("Random Variable, $x$")
plt.ylabel("Probability Density, $f(x)$")
plt.grid()
plt.show()
```

Cauchy and Gaussian Distributions



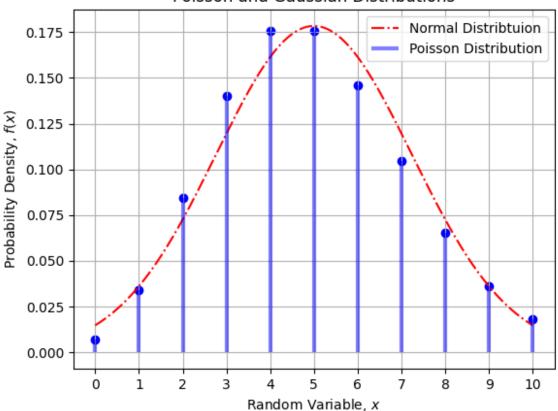
```
x = np.linspace(0 , 10 , 1000)
normal = norm.pdf(x , loc = 5 , scale = 5 ** 0.5)
plt.plot(x , normal , 'r-.', label = 'Normal Distribtuion')

x = np.arange(0, 11)
Poisson = stats.poisson.pmf(x , mu = 5)

plt.vlines(x, 0, Poisson, color='b', lw=3, alpha = 0.5, label='Poisson Distribution')
plt.scatter(x, Poisson, color='b', marker='o')
plt.legend()
plt.xticks(x)
plt.title("Poisson and Gaussian Distributions")
```

```
plt.xlabel("Random Variable, $x$")
plt.ylabel("Probability Density, $f(x)$")
plt.grid()
plt.show()
```

Poisson and Gaussian Distributions



```
data = [\{0.8920, 0.00044\}, \{0.881, 0.009\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.8913, 0.00032\}, \{0.891
{0.9837,0.00048},{0.8958,0.00045}]
df = pd.DataFrame(data, columns=['x', 'error'])
df
                                                            Х
                                                                                                  error
0 0.8920 0.00044
1 0.8810 0.00900
2 0.8913
                                                                                   0.00032
3 0.9837
                                                                                   0.00048
4 0.8958 0.00045
#calculating mean
df['numerator'] = df['x']/ df['error']**2
df['denominator'] = 1/df['error']**2
```

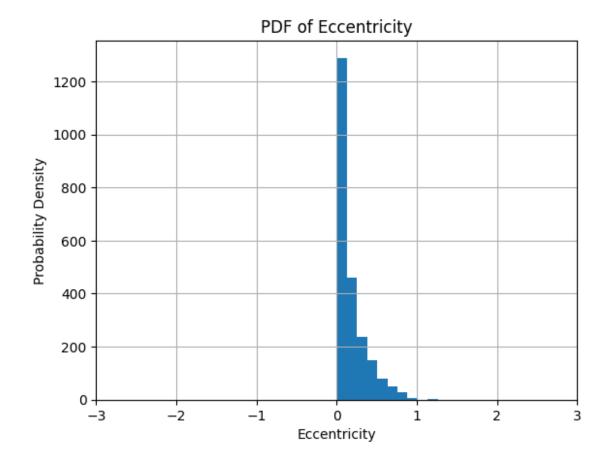
```
#mean
mean = df['numerator'].sum()/df['denominator'].sum()
sd = np.sqrt(1/df['denominator'].sum())

print('Mean is: ', mean)
print('Standard deviation is: ', sd)

Mean is: 0.9089185199574897
Standard deviation is: 0.00020318737026848627
```

```
df = pd.read_csv('exoplanets.csv')
e = df['eccentricity'].dropna()

#histogram of eccentricity
plt.hist(e,bins=25)
plt.xlim(-3,3)
plt.title("PDF of Eccentricity")
plt.xlabel("Eccentricity")
plt.ylabel("Probability Density")
plt.grid()
plt.show()
```



```
#Gaussianizing the data using BoxCox
boxcoxdata = e[e>0]
bc_e, lmbda = stats.boxcox(boxcoxdata)
plt.hist(bc_e,bins=40)
plt.title("PDF of Eccentricity after BoxCox")
plt.xlabel("Eccentricity")
plt.ylabel("Probability Density")
plt.grid()
plt.show()
```

