HW #3: Thin-film Flows and Inertia-less Convection

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Note: All the codes used for plotting are available here.

1 Q 1: Adhesive force in a 'squeeze film':

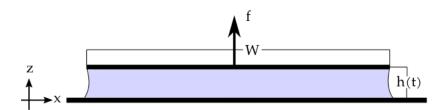


Figure 1: Thin film beneath a knife

Thin-film equations are valid here.

$$\frac{\partial p}{\partial x} = \mu \frac{\partial^2 u}{\partial z^2},
\frac{\partial p}{\partial z} = \mu \frac{\partial^2 w}{\partial z^2},
\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(1)

Let $x \sim W, z \sim h_0, u \sim U, p \sim P$. The continuity equation demands $\frac{\partial u}{\partial x}$ and $\frac{\partial w}{\partial z}$ to balance each other, hence $U/W \sim \tilde{W}/h$, giving a scale for w in terms of U, i.e., $\tilde{W} \sim U h_0/W = U \epsilon$.

The x-momentum equation gives P in terms of U. $\frac{P}{W} \sim \frac{\mu U}{h_0^2}$, giving $P \sim \frac{\mu UW}{h_0^2} = \mu U/(\epsilon^2 W)$.

The dimensionless equations then become:

$$\begin{split} \frac{\partial p}{\partial x} &= \frac{\partial^2 u}{\partial z^2}, \\ \frac{\partial p}{\partial z} &= \epsilon^2 \frac{\partial^2 w}{\partial z^2}, \\ \frac{\partial u}{\partial x} &+ \frac{\partial w}{\partial z} &= 0, \end{split} \tag{2}$$

where all the terms are now dimensionless. The boundary conditions (BCs) are:

$$u = w = 0 \quad \text{at } z = 0,$$

$$u = 0 \quad \text{at } z = h,$$

$$w = \partial_t h + u \partial_x h \quad \text{at } z = h$$

$$w = \partial_t h \quad \text{using } u = 0 \quad \text{at } z = h,$$

$$p = p_0 \quad \text{at } x = 0, 1.$$

$$(3)$$

The leading order z-momentum equation $(\partial_z p = 0)$ tells us that p is not a function of z, i.e., $p \equiv p(x,t)$. Integrating the x-momentum equation wrt z, obtain

$$\frac{\partial u}{\partial z} = \frac{\partial p}{\partial x} \int_{0}^{z} dz$$

$$\frac{\partial u}{\partial z} = \frac{\partial p}{\partial x} z + c_{1}(x, t)$$

$$\Rightarrow u = \frac{\partial p}{\partial x} \frac{z^{2}}{2} + c_{1}(x, t)z + c_{2}(x, t)$$

$$u = 0 \text{ at } z = 0, h,$$

$$c_{2} = 0$$

$$c_{1} = -\frac{\partial p}{\partial x} \frac{h}{2}$$

$$u = \frac{1}{2} \frac{\partial p}{\partial x} \left[z^{2} - hz\right].$$
(4)

Integrating the continuity equation across the domain wrt z:

$$\int_{z=0}^{h(x)} [\partial_x u + \partial_z w_z = 0] dz,$$

$$w \Big|_0^h + \int_{z=0}^{h(x)} \partial_x u dz = 0,$$

$$\partial_t h + u \Big|_h \partial_x h - 0 + \int_{z=0}^{h(x)} (\partial_x u) dz = 0 \quad \dots \text{ using BCs for } w,$$

$$\partial_t h + u \Big|_h \partial_x h + \partial_x \int_{z=0}^{h(x)} u dz - u \Big|_h \partial_x h = 0 \quad \dots \text{ Leibniz rule,}$$

$$\partial_t h + \partial_x \left[\int_{z=0}^{h(x)} u dz \right] = 0$$

$$\frac{dh}{dt} + \partial_x \left[\frac{1}{2} \frac{\partial p}{\partial x} \left[z^3 / 3 - h z^2 / 2 \right]_0^h \right] = 0$$

$$\frac{dh}{dt} - \partial_x \left[\frac{1}{2} \frac{\partial p}{\partial x} \frac{h^3}{6} \right] = 0$$

$$\frac{dh}{dt} - \frac{h^3}{12} \frac{\partial^2 p}{\partial x^2} = 0 \quad \dots h \equiv h(t) \text{ only.}$$

Integrating twice wrt x, we get p.

$$p = \frac{12}{h^3} \frac{dh}{dt} \frac{x^2}{2} + c_1 x + c_2$$

$$p = p_0 \quad \text{at } x = 0, 1,$$

$$c_2 = p_0$$

$$c_1 = -\frac{6}{h^3} \frac{dh}{dt}$$

$$p - p_0 = \frac{6}{h^3} \frac{dh}{dt} (x^2 - x)$$
(6)

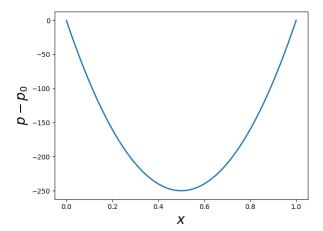


Figure 2: Gauge pressure $p-p_0$ in the film

Force per unit length (into the paper) exerted by the fluid on the knife is

$$f_{1} = \int_{0}^{1} (p - p_{0}) dx,$$

$$f_{1} = \frac{6}{h^{3}} \frac{dh}{dt} \int_{0}^{1} (x^{2} - x) dx$$

$$f_{1} = \frac{6}{h^{3}} \frac{dh}{dt} \left[\frac{x^{3}}{3} - \frac{x^{2}}{2} \right]_{0}^{1}$$

$$f_{1} = -\frac{1}{h^{3}} \frac{dh}{dt}$$
(7)

Therefore, the force (per unit length into the plane of paper) needed to pull the knife upward is $f=-f_1=\frac{1}{h^3}\frac{dh}{dt}$, which is huge if dimensionless $h\sim\epsilon$ is small.

2 Q 2: Static shape of a pendant droplet with uniform surface tension and gravity:

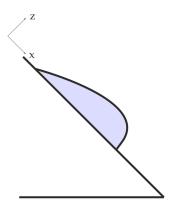


Figure 3: Pendant droplet on an incline

The dimensional governing equations can be written as:

$$\frac{\partial p}{\partial x} = \mu \frac{\partial^2 u}{\partial z^2} + \rho g \sin \alpha$$

$$\frac{\partial p}{\partial z} = \mu \frac{\partial^2 w}{\partial z^2} - \rho g \cos \alpha$$

$$\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(8)

and the dimensional boundary conditions (BCs) are:

u = w = 0 at z = 0,

$$\text{kinematic BC:} \quad \frac{D(z-h)}{Dt} = 0 \Rightarrow \quad w = \partial_t h + u \partial_x h \quad \text{at} \quad z = h(x),$$

Dynamic BC (tangential): $t_i \sigma_{ij} n_j - \underline{t_i} \sigma_{\alpha ij} n_j = 0$ at z = h(x),

Dynamic BC (normal):
$$n_i \sigma_{ij} n_j - \underbrace{n_i \sigma_{aij} n_j}_{p_0} \stackrel{p_0}{=} \gamma K$$
 at $z = h(x)$, (9)

Let us non-dimensionalize the governing equations and BCs.

The scalings used are as follows:

$$x \sim L, z \sim h, u \sim U, w \sim W, p \sim P$$
 (10)

From the continuity equation $O(\partial_x u) \sim O(\partial_z w)$ for balancing each other. This immediately yields the scaling for W in terms of U, i.e.

$$\frac{U}{L} \sim \frac{W}{h_0},$$

$$\Rightarrow W \sim \frac{Uh_0}{L} = \epsilon U.$$
(11)

where $\epsilon = h_0/L \ll 1$ is the thin-film approximation. We now turn to the x-momentum equation to obtain the scale for pressure in terms of U. Balancing the pressure gradient and the viscous terms, we obtain

$$\frac{P}{L} \sim \frac{\mu U}{h^2},$$

$$P \sim \frac{\mu U L}{h^2} = \frac{\mu U}{\epsilon^2 L}.$$
(12)

We now turn to the normal stress boundary condition at z=h(x). As shown in class, the normal stress boundary condition at the leading order reduces to the so called Young-Laplace equation $p-p_0=-\epsilon^3\bar{c}^{-1}\partial_x^2h$, where all the variables are dimensionless and $\bar{c}=\frac{\mu U}{\gamma}$ is the capillary number. In order to retain the effects of surface tension at the leading order, we demand $\epsilon^3\bar{c}^{-1}=O(1)$. Specifically, re-scaling $\epsilon^{-3}\bar{c}=C$, where C=O(1) is the new capillary number. This yields the scaling for u,

$$C = \bar{c}/\epsilon^{3},$$

$$= \mu U/\gamma \epsilon^{3},$$

$$U \sim \gamma \epsilon^{3}/\mu.$$
(13)

Using these scales, the dimensionless governing equations and boundary conditions become:

$$\frac{\partial p}{\partial x} = \frac{\partial^2 u}{\partial z^2} + (\rho g L^2 \sin \alpha / \gamma \epsilon)$$

$$\frac{\partial p}{\partial z} = \epsilon^2 \frac{\partial^2 w}{\partial z^2} - (\rho g L^2 \cos \alpha / \gamma)$$

$$\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(14)

Defining $G = \rho g L^2 / \gamma$ to be the "gravity" number, we obtain

$$\frac{\partial p}{\partial x} = \frac{\partial^2 u}{\partial z^2} + \frac{G \sin \alpha}{\epsilon}$$

$$\frac{\partial p}{\partial z} = \epsilon^2 \frac{\partial^2 w}{\partial z^2} - G \cos \alpha$$

$$\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(15)

And the BCs become:

$$u = w = 0$$
 at $z = 0$,
 $w = uh_x$ at $z = h(x)$,
 $u_z = 0$ at $z = h(x)$, (16)
 $p - p_0 = -C^{-1}h_{xx}$ at $z = h(x)$,
 $h = 0$, at $x = 0, 1$.

At the leading order, the z-momentum equation becomes $p_z = -G\cos\alpha$, which is just hydrostatic balance. Integrating wrt z, we get:

$$p = -[G\cos\alpha]z + \tilde{p}(x),\tag{17}$$

where $\tilde{p}(x)$ is a constant of integration. Applying the normal-stress BC (the Young-Laplace condition), we obtain, $p|_h = p_0 - C^{-1}h_{xx}$.

$$p_0 - h_{xx} = -[G\cos\alpha]h + \tilde{p}(x),$$

$$\Rightarrow \tilde{p}(x) = p_0 + [G\cos\alpha]h - C^{-1}h_{xx}$$

$$\Rightarrow p = p_0 + [G\cos\alpha](h - z) - C^{-1}h_{xx}.$$
(18)

Therefore,
$$p = p_0 + [G\cos\alpha](h-z) - C^{-1}h_{xx}$$

Substituting in the x-momentum equation and integrating wrt z twice, obtain u:

$$\int_{z=0}^{z} \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} = u_{zz} + \frac{G\sin\alpha}{\epsilon} \right] dz,$$

$$u_{z} = \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] z + \tilde{u}_{z}(x),$$

$$\therefore u_{z} = 0 \quad \text{at} \quad z = h(x),$$

$$\tilde{u}_{z}(x) = -\left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] h,$$

$$u_{z} = \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] (z - h),$$

$$\Rightarrow u = \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] \left(\frac{z^{2}}{2} - hz \right) + \tilde{u}(x),$$

$$\therefore u = 0 \quad \text{at} \quad z = 0,$$

$$\tilde{u} = 0,$$

$$\Rightarrow u = \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] \left(\frac{z^{2}}{2} - hz \right).$$

$$(19)$$

Now, integrating continuity equation across the domain wrt z, we obtain:

$$\int_{z=0}^{h(x)} [\partial_x u + \partial_z w_z = 0] dz,$$

$$w\Big|_0^h + \int_{z=0}^{h(x)} \partial_x u dz = 0,$$

$$\partial_t h + u\Big|_h \partial_x h - 0 + \int_{z=0}^{h(x)} (\partial_x u) dz = 0 \quad \dots \text{ using BCs for } w, \qquad (20)$$

$$\underline{u}\Big|_h \partial_x h + \partial_x \int_{z=0}^{h(x)} u dz - \underline{u}\Big|_h \partial_x h = 0 \quad \dots \text{ Leibniz rule,}$$

$$\int_{z=0}^{h(x)} u dz = c$$

However, $Q = \int_{z=0}^{h(x)} u dz$ corresponds to the volume flux and there is no volume flux here. So c = 0. We then get the equation for h.

$$\int_{0}^{h} \left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] \left(\frac{z^{2}}{2} - hz \right) dz = 0$$

$$\left[(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} \right] \left(\frac{z^{3}}{6} - h\frac{z^{2}}{2} \right) \Big|_{0}^{h} = 0$$

$$(G\cos\alpha)h_{x} - C^{-1}h_{xxx} - \frac{G\sin\alpha}{\epsilon} = 0.$$
(21)

For gravity to do anything, it must have an O(1) effect in the x-direction. Redefining $\tilde{G} = G/\epsilon$ and demanding $\tilde{G} \sim O(1)$, we get

$$(\epsilon \tilde{G}\cos\alpha)h_x - C^{-1}h_{xxx} - \tilde{G}\sin\alpha = 0.$$
 (22)

Neglecting the $O(\epsilon)$ term at the leading order, we obtain

$$C^{-1}h_{xxx} + \tilde{G}\sin\alpha = 0 \tag{23}$$

Finally defining the Bond number to be $B=\tilde{G}C=\rho gL^2C/\epsilon\gamma$, and integrating thrice in x, we get:

$$h_{xxx} = (-B\sin\alpha) \tag{24}$$

$$h_{xx} = (-B\sin\alpha)x + c_1\tag{25}$$

$$h_x = (-B\sin\alpha)\frac{x^2}{2} + c_1x + c_2 \tag{26}$$

$$h = (-B\sin\alpha)\frac{x^3}{6} + c_1\frac{x^2}{2} + c_2x + c_3 \tag{27}$$

Since h=0 at $x=0,1,\ c_3=0$ and $c_1/2+c_2=B\sin\alpha/6$. Hence, $h=(-B\sin\alpha)\frac{x^3}{6}+c_1\frac{x^2}{2}+c_2x$.

Also, the volume V_0 is preserved. In dimensionless terms $V_0 = \int_0^1 h dx$. Therefore, we get,

$$V_0 = -B\sin\alpha/24 + c_1/6 + c_2/2 \tag{28}$$

Solving $c_1/2+c_2=B\sin\alpha/6$ and Eqn.(28) simultaneously, we obtain: $c_1=\frac{-24V_0+B\sin\alpha}{2}$ and $c_2=\frac{72V_0-B\sin\alpha}{12}$

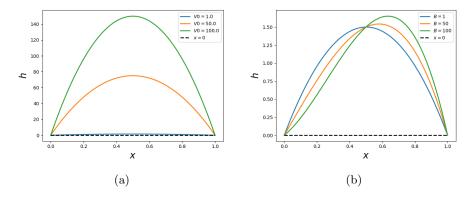


Figure 4: h(x) vs x for $\alpha=\pi/4$. (a) Fix B=1, vary V_0 and (b) Fix $V_0=1$, vary B.

3 Q 3: Linear stability of a liquid film with non-uniform surface tension and destabilizing gravity.

Governing Equations and BCs:

The governing equations, as before, can be written as

$$\frac{\partial p}{\partial x} = \frac{\partial^2 u}{\partial z^2}
\frac{\partial p}{\partial z} = \epsilon^2 \frac{\partial^2 w}{\partial z^2} - G
\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(29)

These can be easily retrieved from Eqns.(15) by setting $\alpha=0$. Here, $G=\rho g L^2/\gamma_0$ as defined in the previous problem.

The tangential stress boundary condition can be written as follows:

$$t_{i}\sigma_{ij}n_{j} - t_{i}\sigma_{xij}\overset{0}{n_{j}} = \partial_{s}\gamma \quad \text{sis the arclength along the surface,}$$

$$\hat{t} = \frac{\hat{e}_{x} + \partial_{x}h\hat{e}^{z}}{\sqrt{1 + (\partial_{x}h)^{2}}}, \quad \hat{n} = \frac{-\partial_{x}h\hat{e}_{x} + \hat{e}^{z}}{\sqrt{1 + (\partial_{x}h)^{2}}}$$

$$\therefore t_{1}\sigma_{11}n_{1} + t_{1}\sigma_{12}n_{2} + t_{1}\sigma_{12}n_{2} + t_{2}\sigma_{22}n_{2} = \partial_{s}\gamma$$

$$\mu(\partial_{z}u + \partial_{x}w)(1 - (\partial_{x}h)^{2}) - 4\mu\partial_{x}h(\partial_{x}u - \partial_{z}w) = \partial_{s}\gamma.$$

$$(30)$$

But $ds \approx \sqrt{dx^2 + dy^2} = dx(\sqrt{1 + (\partial_x h)^2})$ at y = h. This gives $\partial_s = \frac{\partial_x}{\sqrt{1 + (\partial_x h)^2}}$. In dimensionless terms, using the scaling for u from Eqn. (13), $U \sim \gamma_0 \epsilon^3/\mu$ we get:

$$\frac{1}{1+\epsilon^{2}(\partial_{x}h)^{2}} \left[\mu(\partial_{z}u+\partial_{x}w)(1-\epsilon^{2}(\partial_{x}h)^{2})U/(L\epsilon)-4\mu\partial_{x}h(\partial_{x}u-\partial_{z}w)\epsilon U/L\right]$$

$$=(\gamma_{0}/L)\frac{\partial_{x}\gamma}{\sqrt{1+(\partial_{x}h)^{2}}},$$

$$\frac{1}{\sqrt{1+\epsilon^{2}(\partial_{x}h)^{2}}} \left[\mu(\partial_{z}u+\partial_{x}w)(1-\epsilon^{2}(\partial_{x}h)^{2})(\gamma_{0}\epsilon^{3}/\mu)/(\epsilon)-4\mu\partial_{x}h(\partial_{x}u-\partial_{z}w)\epsilon(\gamma_{0}\epsilon^{3}/\mu)\right]$$

$$=\gamma_{0}\partial_{x}\gamma,$$

$$\boxed{\epsilon^{2}\partial_{z}u=\partial_{x}\gamma} \quad \text{at the leading order.}$$
(31)

Writing $\gamma = 1 + \epsilon^2 \gamma$ and equating terms of the same order in ϵ , the tangential BC at the leading order reduces to $\partial_z u = \partial_x \gamma_1$. Also, the dynamic boundary condition (in the normal direction), in dimensional terms is the Young-Lapace equation $p - p_0 = -\gamma \partial_x^2$. In dimensionless terms, we remember that γ is no longer a constant. The dimensionless version will read $p - p_0 = C^{-1} \gamma \partial_x^2 h$, where all the quantities are now dimensionless.

Hence the BCs become:

$$u = w = 0 \quad \text{at} \quad z = 0,$$

$$w = \partial_t h + u \partial_x x \quad \text{at} \quad z = h(x),$$

$$u_z = (\partial_x \gamma_1) \quad \text{at} \quad z = h(x),$$

$$p - p_0 = -C^{-1} (1 + \epsilon^2 \gamma_1) \partial_x^2 h \quad \text{at} \quad z = h(x),$$

$$h = 0, \quad \text{at} \quad x = 0, 1.$$

$$(32)$$

As before, integrating the z-momentum equation at the leading order is just the hydrostatic balance. Integrating the z-momentum equation in z, we obtain the pressure distribution. This is similar to the previous question and we directly write p, by setting $\alpha = 0$ in Eqn.(18).

$$p = p_0 + G(h - z) - C^{-1}(1 + \epsilon^2 \gamma_1) h_{xx}.$$
 (33)

Let us define $\pi = p - p_0$ to be the gauge pressure. Therefore, $\pi = G(h - z) - C^{-1}(1 + \epsilon^2 \gamma_1)h_{xx}$.

Substituting in the x-momentum equation and integrating twice wrt z, we get:

$$\partial_z^2 u = \partial_x \pi \qquad \dots \qquad \partial_x p = \partial_x \pi$$

$$\Rightarrow \quad \partial_z u = (\partial_x \pi) z + c_1(x, t)$$

$$\partial_z u = (\partial_x \gamma_1) \quad \text{at} \quad z = h(x),$$

$$\Rightarrow \quad u_z = (\partial_x \pi) (z - h) + (\partial_x \gamma_1)$$

$$\Rightarrow \quad u = (\partial_x \pi) \left(\frac{z^2}{2} - hz\right) + (\partial_x \gamma_1) z + c_2(x, t)$$

$$u = 0 \quad \text{at} \quad z = 0 \Rightarrow c_2(x, t) = 0.$$

$$\Rightarrow \quad u = (\partial_x \pi) \left(\frac{z^2}{2} - hz\right) + (\partial_x \gamma_1) z \right].$$
(34)

Substituting into conservation of mass Eqn. (20),

$$\partial_{t}h + \partial_{x} \int_{0}^{h(x)} \partial_{x}\pi \left(\frac{z^{2}}{2} - hz\right) + (\partial_{x}\gamma_{1})zdz = 0$$

$$\partial_{t}h + \partial_{x} \left[\partial_{x}\pi \left(\frac{z^{3}}{6} - \frac{hz^{2}}{2}\right) + (\partial_{x}\gamma_{1})\frac{z^{2}}{2}\right]_{0}^{h} = 0$$

$$\partial_{t}h + \partial_{x} \left[(\partial_{x}\gamma_{1})\frac{h^{2}}{2} - (\partial_{x}\pi)\frac{h^{3}}{3}\right] = 0$$

$$\partial_{t}h + \partial_{x} \left[(\partial_{x}\gamma_{1})\frac{h^{2}}{2} - (G\partial_{x}h - C^{-1}(1 + \epsilon^{2}\gamma_{1})\partial_{x}^{3}h - \epsilon^{2}C^{-1}(\partial_{x}\gamma_{1})(\partial_{x}^{2}h))\frac{h^{3}}{3}\right] = 0.$$
(35)

Rescaling time to T=t/C, old time scale U_0/L changes to $T=t'/(\mu L/\epsilon^3\gamma_0)$, where t' is the dimensional time. Also, defining $CG=C\rho gL^2/\gamma_0$ to be the "Bond number", we obtain:

$$\partial_T h + \partial_x \left[C(\partial_x \gamma_1) \frac{h^2}{2} - \left(B \partial_x h - (1 + \epsilon^2 \gamma_1) \partial_x^3 h - \epsilon^2 (\partial_x \gamma_1) (\partial_x^2 h) \right) \frac{h^3}{3} \right] = 0.$$
 at the leading oredr $O(1)$

$$\boxed{\partial_T h + \partial_x \left[C(\partial_x \gamma_1) \frac{h^2}{2} - \left(B \partial_x h - \partial_x^3 h \right) \frac{h^3}{3} \right] = 0}.$$
(36)

Linear stability of a uniformly thick film lining the underside of a rigid flat horizontal substrate

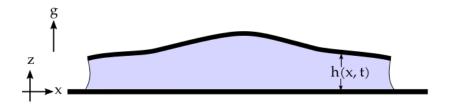


Figure 5: A film lining the underside of a rigid flat horizontal substrate.

We say that a film lining underside of a rigid horizontal substrate is equivalent to the case of regular film on a horizontal substrate, with gravity pointing upwards. So we let $B \to -B$, and $\gamma_1 = \Lambda/h$ ($\Rightarrow \partial_x \gamma_1 = \frac{-\Lambda}{h^2} \partial_x h$), we get:

$$\partial_T h + \partial_x \left[-\frac{C\Lambda}{2} \partial_x h + \left(B \partial_x h + \partial_x^3 h \right) \frac{h^3}{3} \right] = 0 \tag{37}$$

The base state is $h_b = 1$. Introduce a perturbation of the form $h = 1 + \eta$. Substituting in Eqn.(37), obtain:

$$\partial_T \eta + \partial_x \left[-\frac{C\Lambda}{2} \partial_x \eta + \left(B \partial_x \eta + \partial_x^3 \eta \right) \frac{1}{3} \right] = 0,$$

$$\partial_T \eta + \frac{1}{3} \partial_x^4 \eta + \frac{B}{3} \partial_x^2 \eta - \frac{C\Lambda}{2} \partial_x^2 \eta = 0.$$
(38)

Now, we start by "modal analysis", i.e., seek solutions of the form $\eta = Ae^{\sigma t}e^{ikx} +$ c.c. where k is the (known) real wavenumber of the perturbation, σ is the possibly complex growth rate and c.c. denotes the complex conjugate. Substituting into Eqn.(38):

$$\sigma + \frac{k^4}{3} - \left(\frac{B}{3} - \frac{C\Lambda}{2}\right)k^2 = 0 \tag{39}$$

Hence, we get the dispersion relation $\sigma \equiv \sigma(k)$.

$$\sigma = \left(\frac{B}{3} - \frac{C\Lambda}{2}\right)k^2 - \frac{k^4}{3} \tag{40}$$

Unstable when $Re(\sigma) > 0$.

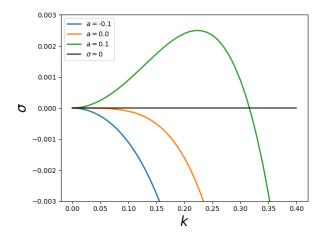


Figure 6: Dispersion relation $(\sigma \equiv \sigma(k))$ for the inear stability of a uniformly thick film lining the underside of a rigid flat horizontal substrate with $a=2B-3C\Lambda$

From Fig.(6), it is clear that when a > 0, a band of modes become unstable. Hence, the condition for instability is $2B - 3C\Lambda \ge 0$ or $B > 3C\Lambda/2$.

4 Q 4: Marangoni convection in the inertia-less limit:

The aim of the analysis is to investigate the possibility that, even in the absence of buoyancy, convection may be possible provided that the temperature-dependence of the surface tension coefficient γ is accounted for.

The dimensional governing equations are the incompressible Stokes equations \mathbf{w}/\mathbf{o} gravity:

$$\partial_x p = \mu(\partial_x^2 u + \partial_z^2 u),
\partial_z p = \mu(\partial_x^2 w + \partial_z^2 w),
0 = \partial_x u + \partial_z w,
\partial_t T + u \partial_x T + w \partial_z T = \kappa(\partial_x^2 T + \partial_z^2 T),$$
(41)

The BCs are:

$$u = w = 0 \quad \text{at } z = 0,$$

$$\mu \partial_z u = \partial_x \gamma \quad \text{at } z = H,$$
where
$$\gamma = \gamma_0 - \Lambda (T - T_0),$$

$$w = 0 \quad \text{at } z = H,$$

$$T = T_0 \quad \text{at } z = 0,$$

$$\partial_z T = -Q_0 \quad \text{at } z = H.$$

$$(42)$$

The surface height H remains constant throughout this analysis.

First, we cast the governing equations in terms of the stream function $\psi,$ such that

$$u = \partial_z \psi, \quad w = -\partial_x w.$$
 (43)

The incompressibility condition is then automatically satisfied. Eliminating pressure by taking the curl of the momentum equations:

$$\mu(\partial_x^2 \partial_z u + \partial_z^3 u - \partial_x^3 w - \partial_z^2 \partial_x w) = 0,$$

$$\Rightarrow (\partial_x^4 + 2\partial_x^2 \partial_z^2 + \partial_z^4) \psi = 0,$$

$$\nabla^4 \psi = 0.$$
(44)

The dimensional equations and BCs, in terms of the streamfunction ψ can be written as:

$$\nabla^4 \psi = 0,$$

$$\partial_t T + [u \cdot \nabla] T = \kappa \nabla^2 T.$$
(45)

The BCs become:

$$\partial_z \psi = \partial_x \psi = 0 \quad \text{at } z = 0,$$

$$\mu \partial_{zz} \psi = \partial_x \gamma \quad \text{at } z = H,$$
where $\gamma = \gamma_0 - \Lambda (T - T_0),$

$$\partial_x \psi = 0 \quad \text{at } z = H,$$

$$T = T_0 \quad \text{at } z = 0,$$

$$\partial_z T = -Q_0 \quad \text{at } z = H.$$
(46)

Scaling $x \sim H, y \sim H, u \sim \kappa/H, T \sim Q_0H$, we obtain scalings for time and streamfunction. The scaling for time is obtained from the energy equation, where $\partial_t T$ must balance $\kappa \nabla^2 T$, yielding $t \sim H^2/\kappa$. From the definition of the streamfunction, we get $\psi \sim \kappa$. Using these scales, we obtain the dimensionless equations:

$$\nabla^4 \psi = 0,$$

$$\partial_t T + [u \cdot \nabla] T = \nabla^2 T.$$
(47)

The BCs become:

$$\partial_z \psi = \partial_x \psi = 0 \quad \text{at } z = 0,$$

$$\partial_{zz} \psi = -\tilde{\Lambda} \partial_x T \quad \text{at } z = 1, \quad \text{where } \tilde{\Lambda} = \frac{\Lambda Q_0 H^2}{\kappa \mu},$$

$$\partial_x \psi = 0 \quad \text{at } z = 1,$$

$$T = T_0/(Q_0 H) \quad \text{at } z = 0,$$

$$\partial_z T = -1 \quad \text{at } z = 1.$$

$$(48)$$

All the quantities in the above BCs are now dimensionless. If $\psi = \text{const}$, $\nabla^4 \psi$ is definitely zero and $u_b = \mathbf{0}$ is the base state velocity. Without loss of generality, we take $\psi_b = 0$. We assume a steady conduction base state for the temperature with no x-variation. $\partial_{zz} T_b = 0$, giving $T_b = Az + B$. With $T_b = T_0/(Q_0H)$ at z = 0, we get $B = T_0/(Q_0H)$ and $\partial_z T = -1$ at z = 1 yields A = -1. Therefore, the steady state base temperature profile is $T_b = T_0/(Q_0H) - z$. Perturbing about the base state and substituting $\psi \equiv \psi_b + \psi$ and $T = T_b + \theta$ (noting that $T_{bz} = -1$), into the governing equations and BCs,

$$\nabla^{4}\psi = 0,$$

$$\partial_{t}\theta + \psi_{z}\theta_{x} - \psi_{x}(-1 + \theta_{z}) = \nabla^{2}\theta.$$
Neglecting nonlinear terms
$$\partial_{t}\theta + \psi_{x} = \nabla^{2}\theta.$$
(49)

The BCs become:

Substituting

$$\begin{bmatrix} \theta \\ \psi \end{bmatrix} = \begin{bmatrix} \hat{\theta} \\ \hat{\psi} \end{bmatrix} e^{ikx} e^{\sigma t} + \text{c.c.}, \tag{51}$$

we obtain a linear eigenvalue problem in z.

$$\begin{split} [k^4 - 2k^2D^2 + D^4]\hat{\psi} &= 0 \\ \sigma \hat{\theta} + ik\hat{\psi} &= [-k^2 + D^2]\hat{\theta} \\ \text{combining the above, we obtain,} \\ \\ \hat{\psi} &= \frac{1}{ik}[-k^2 + D^2 - \sigma]\hat{\theta} \\ \\ &= [D^4 - 2k^2D^2 + k^4][D^2 - k^2]\hat{\theta} = \sigma[D^4 - 2k^2D^2 + k^4]\hat{\theta} \\ \end{split}$$
 (52)

where $D \equiv d_z$. Substitute

$$\begin{bmatrix} \hat{\theta} \\ \hat{\psi} \end{bmatrix} = \begin{bmatrix} \theta_0 \\ \psi_0 \end{bmatrix} e^{imz}. \tag{53}$$

We first solve the $\hat{\psi}$ equation, since it is independent. Get $(k^2+m^2)^2=0$, degenerate solutions. Therefore, basis would be formed by $\{e^{kx}, xe^{kx}, e^{-kx}, xe^{-kx}\}$. General solution: $\hat{\psi}=c_1e^{kx}+c_2xe^{kx}+c_3e^{-kx}+c_4xe^{-kx}$.

BCs on ψ :

$$\frac{d\hat{\psi}}{dz} = \hat{\psi} = 0 \quad \text{at } z = 0,$$

$$\hat{\psi} = 0 \quad \text{at } z = 1,$$

$$\frac{d^2\hat{\psi}}{dz^2} = -\tilde{\Lambda}ik\hat{\theta} \quad \text{at } z = 1.$$
(54)

Can find $\hat{\psi}$ in terms of $\hat{\theta}$ from here and substitute into the $\hat{\theta}$ equation to obtain the dispersion relation $\sigma \equiv \sigma(k)$.

5 Q 5: A lubrication approximation for Darcy flow in semi-saturated porous media:

Consider a 2d shallow-water flow over a porous medium of length L. The lubrication approximation here would be $\epsilon \equiv h/L \ll 1$. We assume incompressibility and use Darcy's law as the momentum equations. The dimensional governing equations become:

$$\mathbf{u} = -\frac{\kappa}{\mu} \nabla (p + \rho g z),$$

$$\nabla \cdot \mathbf{u} = 0.$$
(55)

In component form:

$$u = -\frac{\kappa}{\mu} \frac{\partial p}{\partial x},$$

$$w = -\frac{\kappa}{\mu} \frac{\partial p}{\partial z} - \frac{\kappa \rho g}{\mu},$$

$$\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0.$$
(56)

The boundary conditions are:

$$w(x, z = 0, t) = 0, w(x, z = h(x, t), t) = \partial_t h + u \partial_x h, p(x, z = h(x, t), t) = p_0,$$
(57)

where p_0 is the constant atmospheric pressure impressed on the top of the groundwater layer (and capillary effects are being neglected). Scaling $x \sim L, z \sim h, u \sim U, p \sim P$. The continuity equation implies $\frac{U}{L} \sim \frac{W}{h}$ or $W \sim Uh/L = \epsilon U$. The x-momentum equation implies $U \sim \kappa P/\mu L \Rightarrow P \sim \mu U L/\kappa$. In the z-momentum equation, the relative size of w and $\frac{\delta p}{\mu} \frac{\partial p}{\partial z}$ term can be found to be:

$$|w| / \left| \frac{\kappa}{\mu} \frac{\partial p}{\partial z} \right| \sim \frac{\epsilon U}{UL/h},$$

$$|w| / \left| \frac{\kappa}{\mu} \frac{\partial p}{\partial z} \right| \sim \epsilon^{2}.$$
(58)

Hence, we neglect w at the leading order in the z-momentum equation. At the leading order, the dimensional z-momentum equation reads:

$$\frac{\partial p}{\partial z} = -\rho g. \tag{59}$$

Integrating, we obtain $p = -\rho gz + c(x)$. Using the boundary condition at the top surface z = h, obtain $p - p_0 = \rho g(h - z)$.

Substituting in the x-momentum equation, obtain: $u = -\frac{\kappa \rho g}{\mu} \partial_x h$.

Now, using the depth-averaged version of the continuity equation (see Eqns. (5) and (20))

$$\partial_{t}h + \partial_{x} \left[\int_{0}^{h} u dz \right] = 0,$$

$$\partial_{t}h - \partial_{x} \left[\frac{\kappa \rho g}{\mu} \partial_{x} h[z]_{0}^{h} \right] = 0,$$

$$\partial_{t}h - \partial_{x} \left[\frac{\kappa \rho g}{\mu} h \partial_{x} h \right] = 0,$$

$$M\partial_{t}h = \partial_{x} \left[h \partial_{x} h \right],$$
(60)

where $M=\mu/(\kappa\rho g)$. This is nonlinear diffusion equation for h(x,t). Notice that pressure and h are linearly related in this problem. If there is a Gaussian pressure anomaly localized at x=0 at t=0, it will diffuse as time goes on. The time-scale for this would be governed by the above nonlinear diffusion equation. Namely, $M/t \sim h/L^2$ or $t \sim ML^2/h = ML/\epsilon$. This pressure diffusion is typical of porous media flows.