# Pulsar Detection Using Deep Learning Techniques

## **Abstract**

Pulsars, rapidly rotating neutron stars, emit beams of electromagnetic radiation, providing valuable insights into astrophysical phenomena. However, identifying pulsars within massive datasets collected by radio telescopes is challenging due to noise and radio frequency interference (RFI). Traditional methods, while effective, are computationally expensive and often require manual inspection. Deep learning techniques have emerged as a promising approach for automating and improving the efficiency of pulsar detection. This research paper explores the application of various deep learning architectures, including Artificial Neural Networks (ANNs), Convolutional Neural Networks (CNNs), multi-modal networks combining ANNs and CNNs, and the Pulsar Hunters Classification eXtreme (PHCX) pipeline, for pulsar candidate classification. We analyze their performance on benchmark datasets, discuss their limitations, and suggest future research directions.

## 1. Introduction

Pulsars, formed from the remnants of massive stars, are highly magnetized, rapidly rotating neutron stars that emit beams of electromagnetic radiation. These beams, much like a lighthouse, sweep across the sky, and when they intersect with Earth, we observe them as periodic pulses. Detecting these signals is crucial for understanding fundamental physics, stellar evolution, and the dynamics of the interstellar medium. However, the process of identifying pulsars within the vast amounts of data generated by radio telescopes presents a significant computational challenge.

Traditional methods for pulsar detection often rely on algorithms such as the Fast Fourier Transform (FFT) and the Fast Folding Algorithm (FFA) to search for periodicities in the data. While these methods have been successful in discovering numerous pulsars, they are often hampered by the presence of noise and RFI, leading to a high number of false positives that require time-consuming manual inspection by astronomers. This bottleneck hinders the efficiency of large-scale pulsar surveys and potentially leads to missed detections, especially for faint pulsars.

Deep learning, a subfield of machine learning, has revolutionized various fields, including image recognition, natural language processing, and now, astrophysics. Its ability to learn complex patterns and representations from raw data makes it well-suited for the task of pulsar candidate classification. Deep learning models can be trained to differentiate between genuine pulsar signals and noise or RFI, effectively automating the detection process and potentially uncovering faint pulsars that might be missed by traditional methods. This automation not only accelerates the pace of discovery but also allows astronomers to focus on more complex analysis and interpretation of the detected signals.

This paper investigates the application of different deep learning architectures for pulsar detection, each with its own strengths and limitations:

- Artificial Neural Networks (ANNs): ANNs are versatile models capable of learning complex relationships between input features and output classes. They have been successfully used in various classification tasks, including pulsar candidate selection.
- Convolutional Neural Networks (CNNs): CNNs excel at processing image data and have shown promising results in identifying patterns in pulsar signals represented as images.
- Multi-modal Networks: Combining ANNs and CNNs can leverage the strengths of both architectures, potentially improving classification accuracy by integrating different types of information.
- PHCX: The Pulsar Hunters Classification eXtreme (PHCX) project is a citizen science initiative that utilizes a Random Forest classifier to identify pulsar candidates, demonstrating the effectiveness of machine learning in this domain and the potential for collaboration between humans and machines.

This research paper aims to provide a comprehensive overview of deep learning techniques for pulsar detection, compare their performance on benchmark datasets, discuss their limitations, and suggest future research directions. By exploring the capabilities and challenges of these techniques, we aim to contribute to the ongoing efforts to improve the efficiency and accuracy of pulsar detection, ultimately leading to a deeper understanding of these fascinating celestial objects.

# 2. Deep Learning Architectures for Pulsar Detection

# 2.1 Artificial Neural Networks (ANNs)

ANNs, inspired by the structure of the human brain, are a fundamental type of deep learning model. They consist of interconnected nodes (neurons) organized in layers: an input layer, one or more hidden layers, and an output layer. Each connection between neurons has an associated weight, and the network learns by adjusting these weights during the training process to minimize the difference between its predictions and the actual labels of the training data. This learning process allows the network to capture complex relationships and patterns within the data.

In the context of pulsar detection, ANNs can be used to classify candidates based on features extracted from the raw telescope data. These features might include:

- Statistical measures of the signal: Mean, standard deviation, skewness, and kurtosis, which capture the distribution and shape of the pulsar signal.
- Information derived from the DM-SNR curve: This curve characterizes the dispersion of the signal as it travels through the interstellar medium, providing valuable information for identifying pulsars.

One study <sup>1</sup> employed an ANN with three dense layers and ReLU activation functions to classify pulsar candidates from the HTRU2 dataset. The input layer consisted of 8 nodes, corresponding to the 8 features in the dataset. The first hidden layer had 12 nodes, the second hidden layer had 8 nodes, and the output layer had 1 node with a sigmoid activation function for binary

classification. This model achieved high accuracy, demonstrating the potential of ANNs for pulsar candidate classification. Another study <sup>2</sup> utilized an ANN-based pipeline to suppress RFI and correct for interstellar dispersion, significantly improving the efficiency of pulsar searches by reducing the computational burden and the number of false positives.

## 2.2 Convolutional Neural Networks (CNNs)

CNNs are specialized neural networks designed for processing grid-like data, such as images. They are particularly effective at identifying spatial hierarchies and patterns within the data. CNNs utilize convolutional filters, small matrices that slide across the input data, extracting features like edges, corners, and textures. These features are then passed through pooling layers, which reduce the dimensionality of the data while preserving important information. This hierarchical feature extraction process allows CNNs to learn complex representations of the input data.

In pulsar detection, CNNs can be applied to images generated from the time-frequency representation of the pulsar signal. These images, often referred to as dynamic spectra, capture the characteristic shape of the pulsar's pulse profile, which can be used to distinguish it from noise and RFI. The convolutional filters in the CNN learn to identify specific patterns and features within these dynamic spectra that are indicative of pulsar signals.

Research has shown that CNNs can achieve high accuracy in pulsar candidate classification <sup>3</sup>. One study <sup>4</sup> proposed a CNN architecture called AR\_Net, which incorporates an attention mechanism to focus on the most important features of the input data. This attention mechanism allows the network to selectively weigh different parts of the input, improving its ability to identify relevant patterns. AR\_Net achieved state-of-the-art performance on the PMPS-26k dataset, demonstrating the effectiveness of CNNs for this task. Another study <sup>4</sup> used a CNN with residual connections, which allow the network to learn more complex features by skipping connections between layers. This model also achieved high accuracy on a pulsar candidate dataset.

#### 2.3 Multi-modal Networks

Multi-modal networks combine different types of deep learning models to leverage their respective strengths. In the context of pulsar detection, a multi-modal approach might involve combining ANNs and CNNs. The ANN could process statistical features extracted from the signal, while the CNN could analyze the time-frequency image representation. By integrating information from different modalities, these networks can potentially achieve higher accuracy and robustness compared to single-model approaches.

A study <sup>5</sup> explored a multi-modal fusion-based pulsar identification model (MFPIM) that uses multiple CNNs to extract features from different diagnostic images of pulsar candidates. These images represent different aspects of the pulsar signal, and by combining information from these different sources, the model can learn a more comprehensive representation of the candidate. This model achieved high accuracy on both the FAST and HTRU datasets, outperforming single-model ANNs and CNNs. The study demonstrated that the multi-modal approach effectively captures complementary information from different modalities, leading to

improved classification performance.

#### **2.4 PHCX**

The Pulsar Hunters Classification eXtreme (PHCX) project is a citizen science initiative that engages volunteers in classifying pulsar candidates. It utilizes a Random Forest classifier, a machine learning algorithm that combines multiple decision trees to improve accuracy and robustness. Each decision tree in the forest learns a different set of rules from the data, and the final classification is made by aggregating the predictions of all the trees. This ensemble approach often leads to more accurate and stable predictions compared to using a single decision tree.

PHCX has been successful in identifying new pulsars, demonstrating the effectiveness of machine learning in this domain. The project highlights the potential of combining human expertise with automated classification tools to accelerate pulsar discovery. Volunteers in PHCX visually inspect candidate plots and provide classifications, which are then used to train and refine the Random Forest classifier. This human-in-the-loop approach leverages the pattern recognition capabilities of both humans and machines, leading to improved results.

# 3. Datasets for Pulsar Detection

Several datasets are commonly used for training and evaluating deep learning models for pulsar detection. These datasets provide diverse sources of pulsar candidate data with varying features and class distributions, allowing researchers to train and evaluate deep learning models under different conditions and assess their generalization capabilities. Here are some of the key datasets used in this field:

Dataset	Source	Samples	Features	Description
HTRU2	High Time Resolution Universe Survey <sup>6</sup>	17,898 <sup>6</sup>	8	A widely used benchmark dataset containing pulsar candidates described by eight features derived from the integrated pulse profile and the DM-SNR curve.

Dataset	Source	Samples	Features	Description
Predicting Pulsar Star	High Time Resolution Universe Survey <sup>7</sup>	Not available	Not available	Another dataset from the High Time Resolution Universe Survey, providing an additional source of pulsar candidate data.
Public List of LAT-Detected Gamma-Ray Pulsars	Fermi Large Area Telescope <sup>8</sup>	Not available	Not available	A dataset focused on gamma-ray pulsars detected by the Fermi Large Area Telescope, providing information on their properties.

#### **Feature Explanation:**

The **integrated pulse profile** refers to the average shape of the pulsar's pulse over many rotations. It is a characteristic feature that can be used to distinguish pulsars from other sources. The **DM-SNR curve** represents the signal-to-noise ratio of the pulsar signal as a function of dispersion measure (DM). DM is a measure of the electron density along the line of sight to the pulsar, and it affects the way the pulsar signal is dispersed in time. These features are extracted from the raw telescope data and used as input to the deep learning models<sup>9</sup>.

# 4. Data Preprocessing

Before being fed into deep learning models, the raw pulsar data often undergoes several preprocessing steps to improve the performance and efficiency of the models. These steps include:

• **Data scaling:** This involves scaling the features to a similar range, which can prevent features with larger values from dominating the learning process and improve the

- convergence speed of the models.
- Handling imbalanced datasets: Pulsar candidates are rare compared to noise and RFI, leading to imbalanced datasets where the number of non-pulsar examples significantly outweighs the number of pulsar examples. This imbalance can bias the models towards classifying candidates as non-pulsars. Techniques like oversampling the minority class, undersampling the majority class, or using weighted loss functions can help address this issue and improve the models' sensitivity to real pulsars<sup>10</sup>.

These preprocessing techniques play a crucial role in preparing the data for deep learning models and ensuring that the models can effectively learn the underlying patterns and relationships within the data.

# 5. Code Implementations

Several code implementations of deep learning models for pulsar detection are available online, providing valuable resources for researchers and practitioners interested in exploring and applying these techniques. These implementations offer a starting point for developing and experimenting with different architectures and training strategies. Here are some notable examples:

- Pulsar Candidate Classification using Artificial Neural Networks: This GitHub repository
   provides code for an ANN-based pulsar candidate classifier using the HTRU2 dataset.
- **DeepPulsarNet:** This GitHub repository <sup>3</sup> hosts the code for DeepPulsarNet, a CNN-based pipeline for RFI suppression, dispersion correction, and pulsar candidate classification. This pipeline offers an end-to-end solution for pulsar detection, integrating multiple stages of the process into a single deep learning model.
- **Finding Pulsar:** This repository <sup>11</sup> contains code for a hybrid model combining CNN and Random Forest for pulsar detection. This approach explores the potential of combining different machine learning techniques to improve classification accuracy.
- Pulsar Candidate Recognition Using Deep Neural Network Model: This study <sup>4</sup> describes the implementation of AR\_Net, a CNN architecture with an attention mechanism for pulsar candidate recognition.

These code implementations provide valuable resources for researchers and practitioners interested in exploring and applying deep learning techniques for pulsar detection.

# 6. Performance Comparison

Comparing the performance of different deep learning models for pulsar detection requires careful consideration of the evaluation metrics and datasets used. Common metrics include:

Metric	Description
Accuracy	The overall proportion of correctly classified candidates.
Precision	The proportion of correctly classified pulsar candidates among all candidates classified as pulsars.
Recall	The proportion of correctly classified pulsar candidates among all actual pulsar candidates.
F1-score	A harmonic mean of precision and recall, providing a balanced measure of the model's performance.
Matthews correlation coefficient	A measure of the correlation between the predicted and actual classifications, taking into account true and false positives and negatives.

While a direct comparison of all the models mentioned in this paper is challenging due to variations in datasets and evaluation procedures, some general observations can be made:

- CNNs generally outperform ANNs: CNNs' ability to process time-frequency images of
  pulsar signals often leads to higher accuracy compared to ANNs that rely on statistical
  features. This suggests that the spatial information captured in the images is crucial for
  accurate classification.
- Multi-modal networks can further improve performance: Combining ANNs and CNNs can
  leverage the strengths of both architectures, potentially leading to better classification results.
  By integrating information from different modalities, these networks can capture a more
  comprehensive representation of the pulsar candidates.
- PHCX demonstrates the effectiveness of machine learning: The success of this citizen science project highlights the potential of machine learning for pulsar detection, even with simpler models like Random Forest. This demonstrates that machine learning can be a valuable tool for both professional astronomers and citizen scientists involved in pulsar searches.

Studies have demonstrated that CNNs can achieve high accuracy in pulsar candidate classification <sup>3</sup>. For example, one study <sup>4</sup> found that a CNN model performed well even on noisy data, indicating its robustness to challenging conditions.

It is important to note that the performance of deep learning models can be significantly influenced by factors such as the dataset used, the specific architecture and hyperparameters of

the model, and the preprocessing steps applied to the data. Therefore, careful consideration of these factors is crucial when comparing and evaluating different models.

## 7. Limitations and Future Research Directions

Despite the promising results achieved by deep learning in pulsar detection, several limitations and challenges remain:

- Data imbalance: Pulsar candidates are rare compared to noise and RFI, leading to imbalanced datasets that can bias the training of deep learning models. This imbalance can result in models that are overly sensitive to noise and RFI, leading to a high number of false positives, or models that are not sensitive enough to real pulsars, leading to missed detections. Techniques like data augmentation, where synthetic examples are generated to balance the dataset, and weighted loss functions, where the model is penalized more for misclassifying the minority class, can help address this issue.
- Generalization to new data: Deep learning models may struggle to generalize to data from
  different telescopes or with different noise characteristics. This is because the models may
  learn features that are specific to the training data and may not be applicable to new data with
  different properties. Transfer learning, where a model trained on one dataset is fine-tuned on
  another dataset, and domain adaptation, where a model is adapted to a new domain with
  different data distributions, can be explored to improve generalization.
- Interpretability: Understanding why a deep learning model makes a particular classification decision can be challenging. These models often learn complex and non-linear relationships within the data, making it difficult to interpret their decision-making process. Techniques like attention mechanisms, which highlight the parts of the input that the model focuses on, and visualization tools, which allow researchers to visualize the internal representations learned by the model, can provide insights into the model's decision-making process.

Future research directions include:

- Developing more robust and efficient deep learning architectures: Exploring new
  architectures, such as recurrent neural networks (RNNs), which are designed for sequential
  data, and transformers, which have shown promising results in natural language processing,
  could further improve the accuracy and efficiency of pulsar detection.
- Incorporating domain knowledge: Integrating astrophysical knowledge into deep learning
  models can enhance their performance and interpretability. This could involve incorporating
  prior information about pulsar signals, noise characteristics, or the interstellar medium into the
  model architecture or training process.
- Developing unsupervised and semi-supervised learning methods: Reducing the reliance on labeled data can make deep learning more applicable to large-scale pulsar surveys, where obtaining labels for all candidates can be prohibitively expensive. Unsupervised learning methods can learn patterns and structures from unlabeled data, while semi-supervised learning methods can leverage both labeled and unlabeled data to improve performance.
- Applying deep learning to other pulsar-related tasks: Beyond candidate classification, deep learning can be used for tasks like pulsar timing, which involves precisely measuring the arrival times of pulsar pulses, identification of different pulsar types, such as millisecond pulsars or binary pulsars, and searching for new astrophysical phenomena, such as Fast

## 8. Conclusion

Deep learning has emerged as a powerful tool for pulsar detection, offering the potential to automate the classification process, improve efficiency, and uncover faint pulsars that might be missed by traditional methods. This research paper explored various deep learning architectures, including ANNs, CNNs, multi-modal networks, and PHCX, highlighting their strengths and limitations. While challenges remain, ongoing research and development in this field promise to further advance our ability to detect and study these fascinating celestial objects, ultimately leading to new discoveries and a deeper understanding of the universe.

# 9. Synthesis

This research paper provides a comprehensive overview of deep learning techniques for pulsar detection. We have examined various architectures, including ANNs, CNNs, multi-modal networks, and PHCX, and discussed their performance on benchmark datasets. The key findings include:

- Deep learning models, particularly CNNs, have shown promising results in accurately
  classifying pulsar candidates, often outperforming traditional methods based on FFT and
  FFA. This is due to their ability to learn complex patterns and representations from the data,
  effectively distinguishing between pulsar signals and noise or RFI.
- Multi-modal networks that combine ANNs and CNNs can further improve classification accuracy by leveraging the strengths of both architectures. By integrating information from different modalities, such as statistical features and time-frequency images, these networks can capture a more holistic view of the pulsar candidates.
- The PHCX citizen science project demonstrates the effectiveness of machine learning in real-world pulsar detection scenarios and highlights the potential for collaboration between humans and machines. By combining the pattern recognition capabilities of both humans and algorithms, PHCX has successfully identified new pulsars.

Despite the advancements, challenges remain, such as data imbalance, generalization to new data, and interpretability of deep learning models. Data imbalance can bias the models towards classifying candidates as non-pulsars, while generalization issues can limit their applicability to new datasets or telescopes. Interpretability is crucial for understanding the model's decision-making process and ensuring its reliability.

Future research should focus on addressing these limitations and exploring new architectures and learning paradigms to further enhance the accuracy and efficiency of pulsar detection. This includes developing more robust and efficient deep learning architectures, incorporating domain knowledge into the models, developing unsupervised and semi-supervised learning methods to reduce the reliance on labeled data, and applying deep learning to other pulsar-related tasks, such as pulsar timing and the search for new astrophysical phenomena.

The application of deep learning to pulsar detection has the potential to revolutionize our understanding of these fascinating objects. By automating the classification process, improving

efficiency, and uncovering faint pulsars, deep learning can accelerate the pace of discovery and enable astronomers to probe deeper into the mysteries of the universe.

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