## Function GET\_MW - single-thread version

Calling syntax:

Function parameters:

- 0. Lparms 11-element long integer array of dimensions and global (for all voxels) integer parameters (see below).
- 1. Rparms 5-element double array of global (for all voxels) real parameters (see below).
- 2. Parms array of LOS parameters,  $24 \times Nz$  elements, double. Parms[\*, i] represents the parameters for *i*th voxel (see below).
- 3.  $E_{arr}$  array of energies  $E_i$  where the electron distribution function is specified, NE elements, double, in MeV. The values must be monotonically increasing.
- 4. mu\_arr array of pitch-angle cosines  $\mu_j = \cos \alpha_j$  where the electron distribution function is specified, Nmu elements, double. The values must be monotonically increasing and should cover the entire range of possible values from –1 to +1.
- 5. f\_arr array of electron distribution functions  $f_{ijk} = f_{ij}^{(k)}(E_i, \mu_j)$ , NE × Nmu × Nz elements, double, in cm<sup>-3</sup> MeV<sup>-1</sup>. f\_arr[\*, \*, k] represents the distribution function for kth voxel. The distribution function in each voxel is assumed to satisfy the normalization condition

$$2\pi \int_{E_{\rm min}}^{E_{\rm max}} dE \int_{\mu_{\rm min}}^{\mu_{\rm max}} f(E,\mu) d\mu = n_{\rm b},$$

where  $n_b$  is the local concentration of energetic electrons (in cm<sup>-3</sup>), and the energy E is in MeV.

6. RL – input/output array,  $7 \times Nf$  elements, double. RL[\*, i] corresponds to *i*th frequency (see below).

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NE, Nmu, Nnodes, \$

match\_key, Qopt\_key, arr\_key, log\_key, PK\_key, spline\_key]

- 0. Lparms[0] = Nz number of voxels along the LOS.
- 1. Lparms[1] = Nf number of frequencies in the spectrum.
- 2. Lparms[2] = NE number of energies in the E\_arr array; must be  $\geq 3$  otherwise the array-defined electron distribution function is ignored.
- 3. Lparms[3] = Nmu number of pitch-angle nodes in the mu\_arr array; must be  $\geq 3$  otherwise the array-defined electron distribution function is ignored.

- 4. Lparms[4] = Nnodes number of energy nodes used for integration over energy in the continuous gyrosynchrotron code.
  - a. Minimum value: 16; if  $0 \le \text{Nnodes} < 16$ , 16 nodes are used instead.
  - b. If Nnodes < 0, an adaptive integration grid with the target relative accuracy of  $10^{-5}$  is used.
- 5. Lparms[5] = match\_key controls the behaviour of the hybrid gyrosynchrotron code at the boundary frequencies  $f^{C}$  and  $f^{WH}$  (see Fleishman & Kuznetsov 2010):
  - a. 0: additional re-normalization of the spectrum is performed to remove possible jumps at the boundary frequencies;
  - b.  $\neq 0$ : re-normalization is not performed.
- 6. Lparms[6] = Qopt\_key controls the *Q*-optimization of the continuous gyrosynchrotron code (see Fleishman & Kuznetsov 2010):
  - a. 0: *Q*-optimization is on, which improves accuracy;
  - b.  $\neq 0$ : *Q*-optimization is off, which improves speed.
- 7. Lparms[7] = arr\_key global key specifying whether the contribution of the array-defined electron distribution function is considered:
  - a. 0: the array-defined electron distribution function is enabled and can be used in some voxels, depending on the local on/off keys;
  - b.  $\neq 0$ : the array-defined electron distribution function is disabled for all voxels, regardless on the local on/off keys.
- 8. Lparms[8] = log\_key controls the assumptions about the energy grid for the array-defined electron distribution:
  - a. 0: the nodes are assumed to be logarithmically-spaced  $(E_{i+1}/E_i = const)$ ;
  - b.  $\neq 0$ : the nodes are assumed to be equidistant ( $E_{i+1}$ – $E_i$  = const). Note: if neither of above is applied to your energy grid, choose the option that fits the actual energy spacing better this can improve the calculation accuracy greatly.
- 9. Lparms[9] = PK\_key specifies how the pitch-angle dependence of the array-defined electron distribution function is treated:
  - a. 0: the exact (possibly anisotropic) electron distribution is used (default option);
  - b. 1: the electron distribution at each energy is replaced by an isotropic (pitch-angle-averaged) one;
  - c. 2: same as 1, and the continuous gyrosynchrotron code uses the fast approximation by Petrosian (1981) and Klein (1987).
- 10. Lparms[10] = spline\_key controls the 2D interpolation method for the array-defined electron distribution function:
  - a. 0: spline interpolation is used (usually provides higher speed and accuracy);
  - b.  $\neq 0$ : local linear-quadratic interpolation over 2-3 adjacent nodes is used (sometimes works better for the distributions with very sharp gradients).

Array of global real parameters Rparms:

Rparms = 
$$[S, f_0, \Delta f, f^C, f^{WH}]$$

- 0. Rparms[0] = S visible source area, in cm<sup>2</sup>.
- 1. Rparms[1] =  $f_0$  starting frequency of the spectrum, in Hz:
  - a. is used, only if  $f_0 > 0$ ;
  - b. if  $f_0 \le 0$ , the frequencies are taken from the RL[0, \*] array.
- 2. Rparms[2] =  $\Delta f$  logarithmic frequency step used to produce the spectrum,  $f_{i+1}/f_i = 10^{\Delta f}$  (is used only if  $f_0 > 0$ ).
- 3. Rparms[3] =  $f^c$  boundary frequency of the hybrid gyrosynchrotron code (Fleishman & Kuznetsov 2010), expressed in units of the local electron gyrofrequency.
  - If the emission frequency  $f < f^c$ , the exact code with summation over cyclotron harmonics is used.
  - If  $f > f^c$ , the continuous code is used.
  - If  $f^{\mathbb{C}} < 0$ , the code is purely continuous with additional re-normalization using the exact parameters computed at  $f = f^{\mathbb{WH}}$ .
- 4. Rparms[4] =  $f^{WH}$  boundary frequency for the exact/approximated expressions for the Bessel functions in the exact gyrosynchrotron code, expressed in units of the local electron gyrofrequency.
  - If  $f < f^{WH}$ , the exact gyrosynchrotron code (at  $f < f^{C}$ ) uses the exact expressions for the Bessel functions.
  - If  $f > f^{WH}$ , the exact gyrosynchrotron code (at  $f < f^{C}$ ) uses the approximate expressions for the Bessel functions by Wild & Hill (1971).

Array of parameters Parms (for a single voxel, 24 parameters):

- 0. Parms[0] =  $\Delta z$  voxel length, in cm.
- 1. Parms[1] =  $T_0$  plasma temperature, in K.
- 2. Parms[2] =  $n_0$  either thermal electron concentration or total atomic concentration (depending on other parameters, see the separate diagram), in cm<sup>-3</sup>.
  - 3. Parms[3] = B magnetic field strength, in G.
  - 4. Parms[4] =  $\theta$  viewing angle, in degrees.
- 5. Parms[5] emission mechanism flag (rounded down to the nearest integer):
  - a. 0: all emission mechanisms (gyrosynchrotron + e-ions + e-neutrals) are included;
  - b. 1: gyrosynchrotron is off;
  - c. 2: e-ions is off;
  - d. 4: e-neutrals is off.

Several flags can be combined by usual or bitwise summation: e.g., Parms[5] = 2 + 4 turns off both e-ions and e-neutrals, etc.

6. Parms[6] – specifies the chosen analytical electron distribution over energy (index of the model distribution function, see the separate document); non-integer values are rounded down to the nearest integer. *Default option: 0.* 

Note: if the kappa-distribution (Parms[6] = 6) is selected, the e-ions contribution is also computed using the formulae for the kappa-distribution (Fleishman & Kuznetsov 2014); in all other cases, the Maxwellian thermal distribution is assumed.

- 7. Parms[7] =  $n_b$  concentration of nonthermal electrons in the analytical electron distributions, in cm<sup>-3</sup>.
- 8. Parms[8] =  $\varepsilon$  or  $\kappa$  either the matching parameter  $\varepsilon$  in the thermal/nonthermal electron distributions or the parameter  $\kappa$  in the kappadistribution.
- 9. Parms[9] =  $E_{min}$  the low-energy cutoff in the analytical electron distributions (when relevant), in MeV.
- 10. Parms[10] =  $E_{max}$  the high-energy cutoff in the analytical electron distributions (when relevant), in MeV.
- 11. Parms[11] =  $E_{break}$  the break energy in the double-power-law analytical electron distributions, in MeV.
- 12. Parms[12] =  $\delta_1$  the power-law index in the single-power-law analytical electron distributions or the low-energy power-law index in the double-power-law analytical electron distributions.
- 13. Parms[13] =  $\delta_2$  the high-energy power-law index in the double-power-law analytical electron distributions.
- 14. Parms[14] specifies the chosen analytical electron distribution over pitch-angle (index of the model distribution function, see the separate document); non-integer values are rounded down to the nearest integer. *Default option: 0.*
- 15. Parms[15] =  $\alpha_c$  or  $\alpha_0$  either the loss-cone boundary  $\alpha_c$  in the loss-cone analytical electron distributions or the beam direction  $\alpha_0$  in the beam-like analytical electron distributions, in degrees.
- 16. Parms[16] =  $\Delta\mu$  either the loss-cone boundary width or the beam angular width in the loss-cone or beam-like analytical electron distributions, respectively.
- 17. Parms[17] =  $a_4$  the coefficient  $a_4$  in the supergaussian beam-like analytical electron distribution.
- 18. Parms[18] =  $n_p$  proton concentration, in cm<sup>-3</sup>; is used only as a switch (see the separate diagram).
- 19. Parms[19] =  $n_{\rm HI}$  neutral hydrogen concentration, in cm<sup>-3</sup> (see the separate diagram).
- 20. Parms[20] =  $n_{Hel}$  neutral helium concentration, in cm<sup>-3</sup> (see the separate diagram).
- 21. Parms[21] = arr\_key\_local local key specifying whether the contribution of the array-defined electron distribution function is considered in this voxel:

- a. 0: contribution the array-defined electron distribution function is considered (provided that  $NE \ge 3$  and  $Nmu \ge 3$  and this contribution is enabled by the global key);
- b. ≠0: the array-defined electron distribution function in this voxel is ignored even if it is specified.
- 22. Parms[22] element abundance model (used to compute the e-ions contribution):
  - a. -1: "classical" formulae from Dulk (1985) are used;
  - b. 0: solar coronal abundance (by Feldman 1992) is used (*default option*);
  - c. 1: solar photospheric abundance (by Scott et al. 2015) is used.
  - 23. Parms[23] currently unused.

Input/output array RL:

0. First row (RL[0, \*]) – emission frequencies, in GHz. On input, this array is used if  $f_0$  = Rparms[1]  $\leq$  0; otherwise, the frequencies are computed using the  $f_0$  and  $\Delta f$  parameters:  $f_1 = f_0 10^{\Delta f}$ ,  $f_2 = f_1 10^{\Delta f}$ , etc. On output, this array contains the computed or pre-defined emission frequencies.

Other rows – emission intensities, as observed from the Earth, in sfu:

- 1. RL[1, \*] left polarization, weak mode coupling;
- 2. RL[2, \*] right polarization, weak mode coupling;
- 3. RL[3, \*] left polarization, strong mode coupling;
- 4. RL[4, \*] right polarization, strong mode coupling;
- 5. RL[5, \*] left polarization, exact mode coupling.
- 6. RL[6, \*] right polarization, exact mode coupling.

On input, these arrays specify the emission intensities at the start of the line-of-sight; on output, they contain the emission intensities at the end of the line-of-sight.

## Return value:

- 0: no errors:
- -1: error (insufficient number of parameters);
- 1: error (incorrect parameters of the analytical electron distribution function);
- 2: error (incorrect parameters of the array-defined electron distribution function).

In case of any errors, the input/output array RL remains unchanged. *Note:* the parameter checking has not been fully implemented yet, so that some invalid parameter combinations can pass without notice.

## Function GET\_MW\_SLICE - multi-thread version

Calling syntax:

Function parameters:

- 0. Lparms\_M 12-element long integer array of dimensions and global (for all voxels and LOSs) integer parameters (see below).
- 1. Rparms\_M array of real parameters common for all voxels within each LOS, 5 × Npix elements, double (see below).
- 2. Parms\_M array of voxel parameters, 24 × Nz × Npix elements, double (see below).
- 3. E\_arr array of energies where the electron distribution function is specified, NE elements, double, in MeV. This parameter is the same as in the GET\_MW function.
- 4. mu\_arr array of pitch-angle cosines where the electron distribution function is specified, Nmu elements, double. This parameter is the same as in the GET MW function.
- 5.  $f_{arr_M} array of electron distribution functions, NE × Nmu × Nz × Npix elements, double, in cm<sup>-3</sup> MeV<sup>-1</sup> (see below).$ 
  - 6. RL\_M input/output array,  $7 \times Nf \times Npix$  elements, double (see below).

Array of dimensions and global integer parameters Lparms\_M:

Lparms\_M = [Npix, Nz, Nf, NE, Nmu, Nnodes, \$

match\_key, Qopt\_key, arr\_key, log\_key, PK\_key, spline\_key]

0. Lparms\_M[0] = Npix - number of LOSs.

Other elements (1st to 11th) are respectively the same as the 0th to 10th elements of the Lparms array in the GET\_MW function. In particular:

- all LOSs have the same number of voxels Nz;
- the number of frequencies Nf is the same for all LOSs (although the frequency grids can be different);
- the energy and pitch-angle grids (including their dimensions NE and Nmu) are the same in all voxels of all LOSs;
- all other global parameters and keys (Nnodes, match\_key, Qopt\_key, arr\_key, log\_key, PK\_key, spline\_key) are applied to all voxels of all LOSs.

Other parameters: sub-arrays Rparms\_M[\*, i], Parms\_M[\*, \*, i], f\_arr\_M[\*, \*, \*, i] and RL\_M[\*, \*, i] correspond respectively to the parameters Rparms, Parms, f\_arr and RL of the single-thread GET\_MW function, for *i*th LOS.

## Return value:

- 0: no errors;
- -1: error (insufficient number of parameters); the input/output array RL\_M remains unchanged;
- 1: error (incorrect parameters of an electron distribution function in, at least, one of the LOSs); the elements of the input/output array RL\_M corresponding to those incorrect LOSs remain unchanged.