

# **Part 2 of PSD**

## **The Spectral Leakage of the PSD**

**Prof L. Gelman**

# The Spectral Leakage

- Consider the power spectral density of a **finite-length signal**  $x_L[n]$ , where  $L$  is the length of a signal
- It is frequently useful to interpret  $x_L[n]$  as the result of multiplying an infinite signal,  $x[n]$ , by a finite-length rectangular window,  $w_R[n]$  :

$$x_L[n] = x[n] \cdot w_R[n]$$

# The Spectral Leakage

- The effect of the window and the spectral leakage are best understood for sinusoidal data
- Suppose that  $x[n]$  is composed of a sum of  **$M$  complex sinusoids**:

$$x[n] = \sum_{k=1}^M A_k e^{j\omega_k n}$$

where  $A_k$  and  $\omega_k = 2\pi f_k$  are sinusoid amplitude and circular frequency respectively

# The Spectral Leakage

□ Its Fourier transform (for an infinite signal) is

$$X(f) = f_s \sum_{k=1}^M A_k \delta(f - f_k)$$

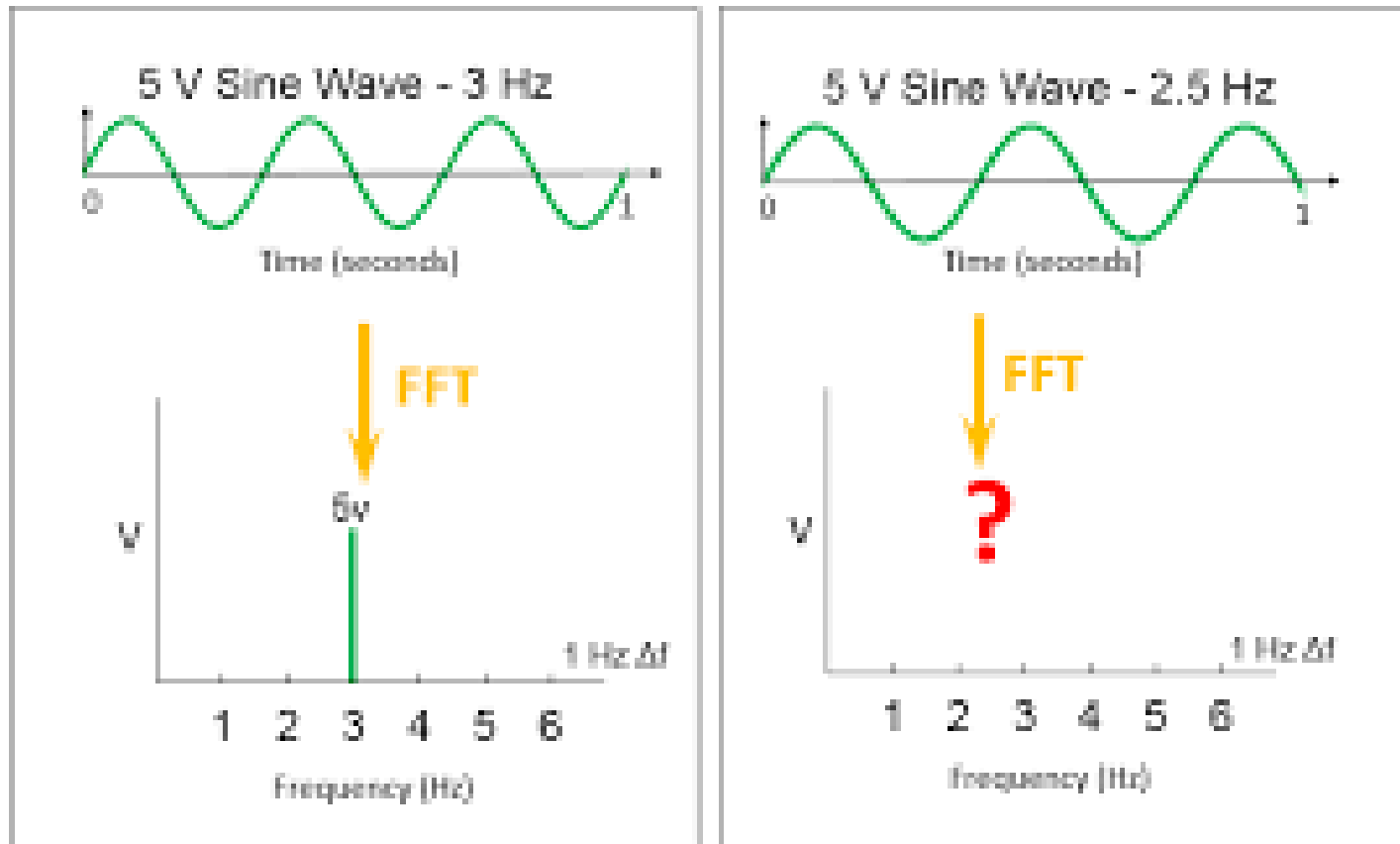
which for a finite-length signal becomes

$$X_L(f) = \int_{-f_s/2}^{f_s/2} \left( \sum_{k=1}^M A_k \delta(\rho - f_k) \right) W_R(f - \rho) d\rho = \sum_{k=1}^M A_k W_R(f - f_k)$$

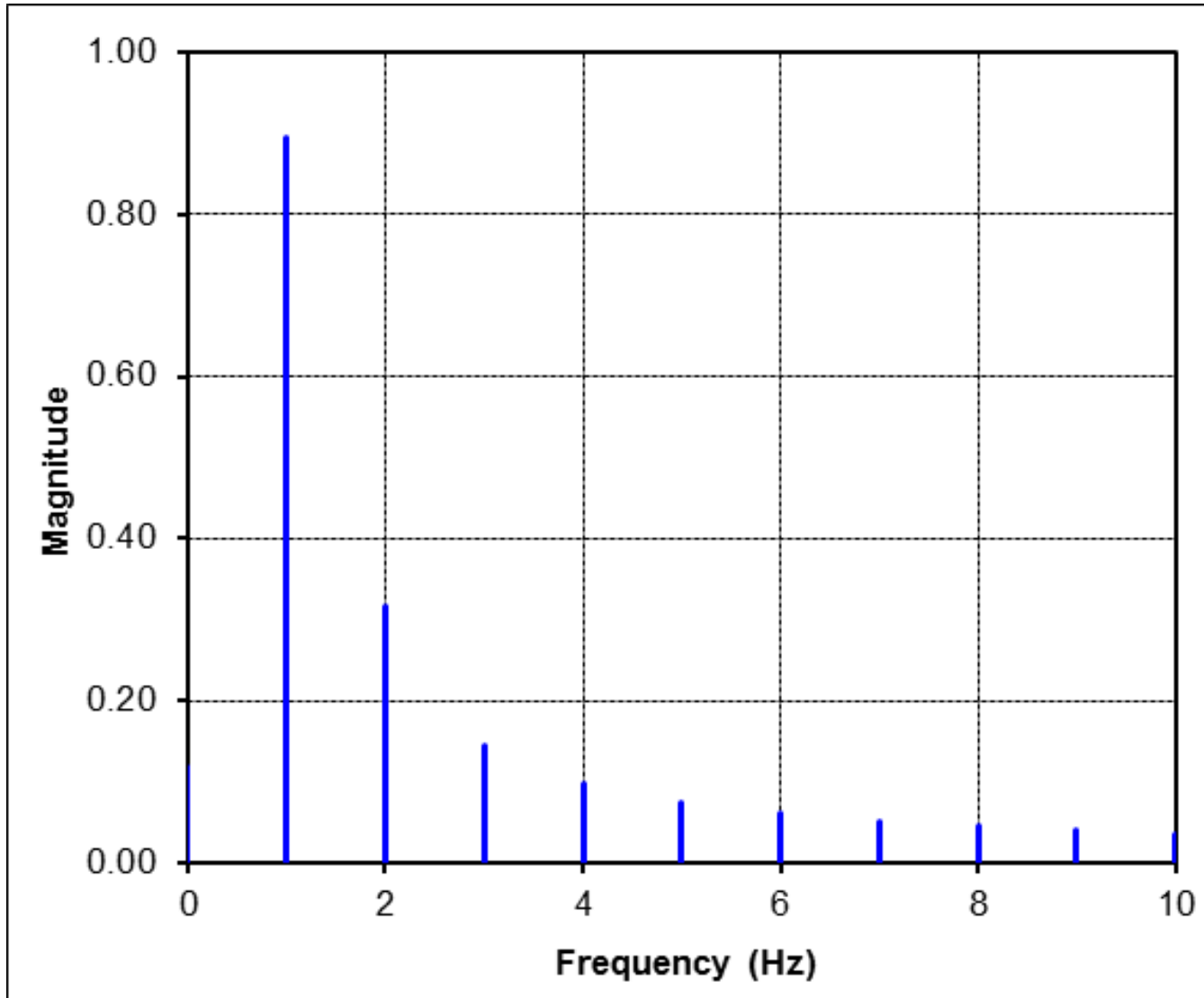
where  $f_s$  is sampling frequency,  $\delta(t)$  is the Dirac function

# The Fourier Transform for an Infinite Sine Signal

## Sine Wave FFT Analysis

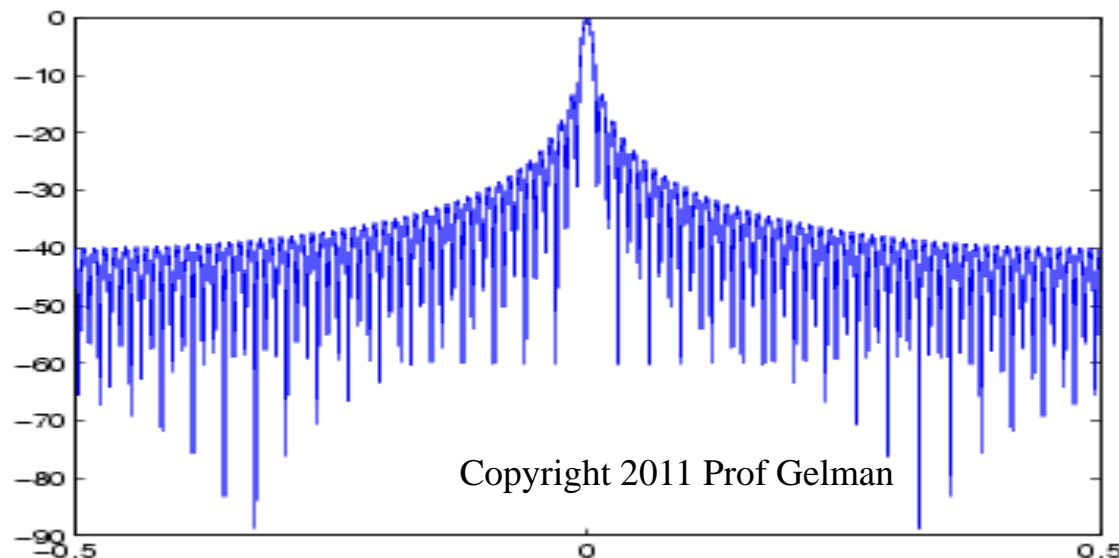


# The Fourier Transform for an Infinite Sine Signal

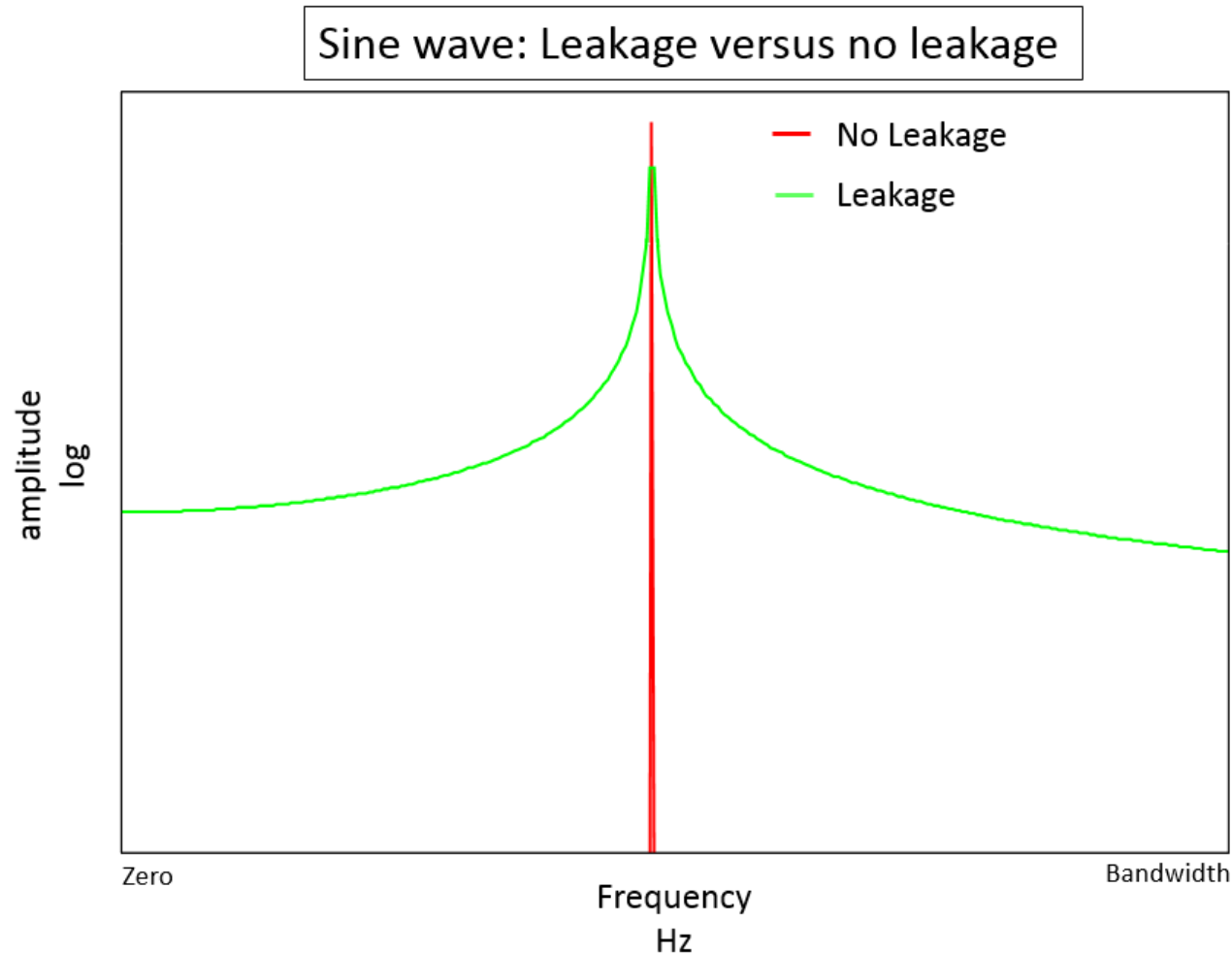


# The Spectral Leakage

- ❑ So, in the Fourier transform of the finite-length signal, the Dirac functions *have been replaced* by terms of the form  $W_R(f - f_k)$ , which corresponds to the Fourier transform of a rectangular window
- ❑ The frequency response of a rectangular window has the shape of a **sinc signal**, as shown below

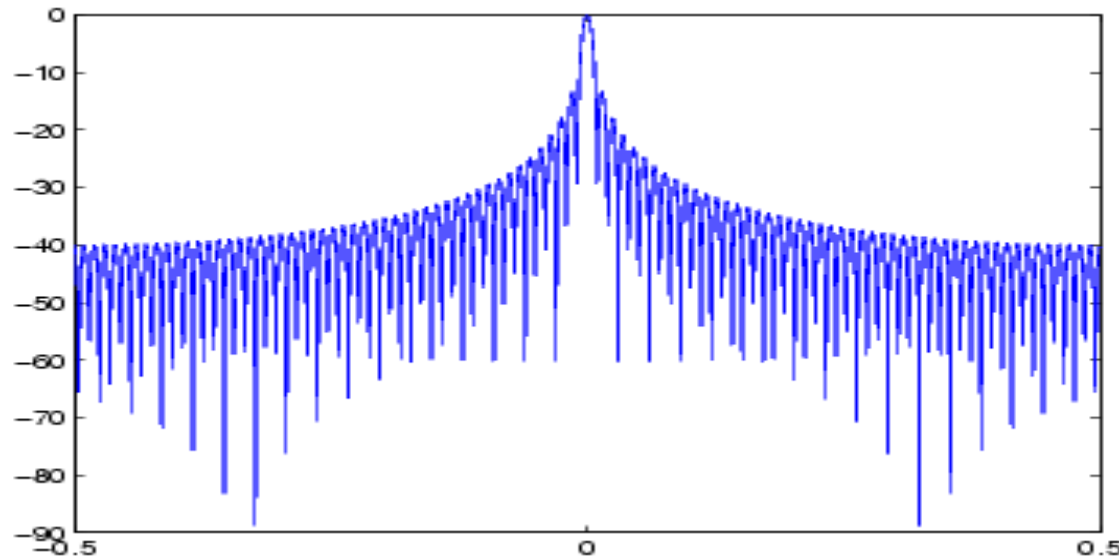


# The Fourier Transform for an Infinite Sine Signal





# The Spectral Leakage



- The plot displays the main lobe and several side lobes, the largest of which is 13.5dB below the main lobe peak
- These lobes account for the effect known as the *spectral leakage* (the Gibbs phenomenon).

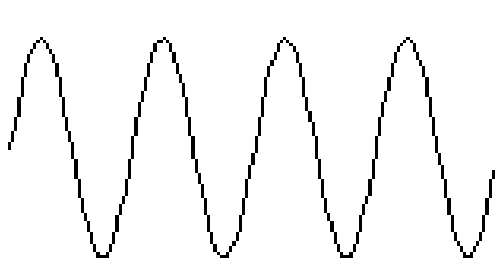
# The Spectral Leakage

- ❑ While the *infinite-length signal* has its power spectral density concentrated *exactly* at the discrete frequency points, the **windowed (or truncated) signal** has a continuum of power "leaked" around the discrete frequency points
- ❑ Let us consider conditions for the leakage appearance

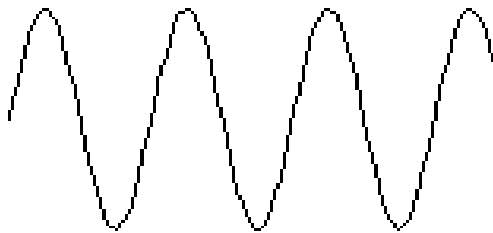
# The Spectral Leakage

- If a signal is periodic, **two main cases arise**:
- an integral number of periods fit into the total duration of a signal
- an integral number of periods does **not** fit into the total duration of a signal; therefore, signal has little “glitch”
- The “glitch” is a short signal

# Spectral Leakage: Two Conditions



If the period fits the time,  
the spectrum is correct



If the period does not fit  
the time, spurious  
spectral lines result



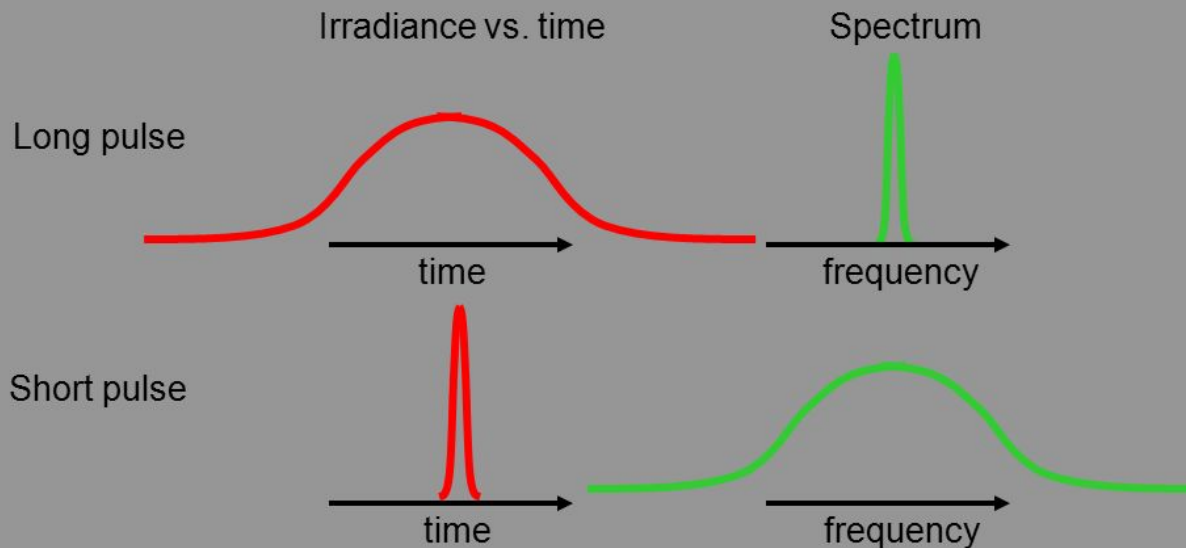
# The Spectral Leakage

- **There is a direct relation between a signal's duration in time and the width of its power spectral density:**
  - **short signals have a broad power spectral density**
  - **long signals have a narrow power spectral density**
- **So, short glitches have a broad power spectral density.**
- **This broadening is superimposed on the power spectral density of the actual signal**

# Long vs. Short Pulses

## Long vs. short pulses of light

The uncertainty principle says that the product of the temporal and spectral pulse widths is greater than  $\sim 1$ .



# The Spectral Leakage: Two Conditions

- if an **integral number of periods** exactly fits a signal duration, the power spectral density is represented by one single line
- if an integral number of periods does **not** match a signal duration, the power spectral density is broadened

# The Spectral Leakage

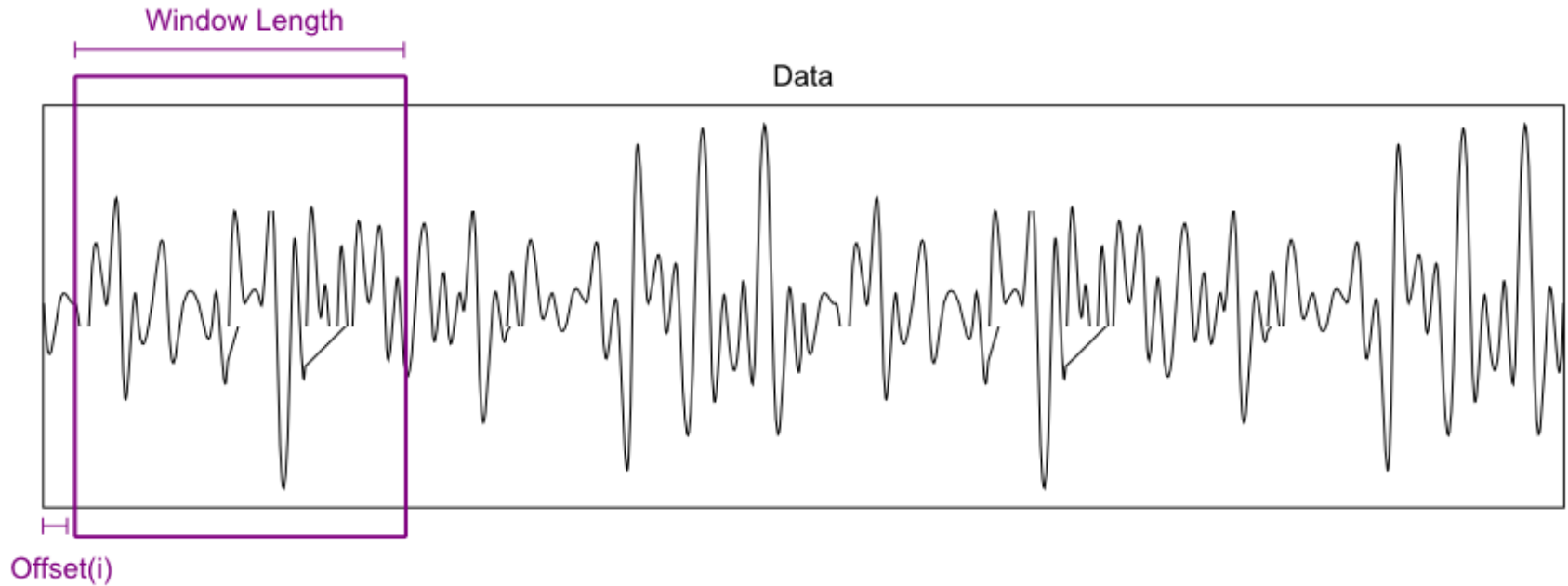
- ❑ For example, a sine wave should have a power spectral density which consists of one single line
- ❑ But in practice, if measured by a spectrum analyzer, the power spectral density will be a broad line - with the side lobes flapping up and down
- ❑ *When we see a **perfect (single line) power spectral density**, this has in fact been obtained by **tuning** the signal frequency carefully so that an integral number of periods exactly fits the measurement time and the power spectral density is the best obtainable*



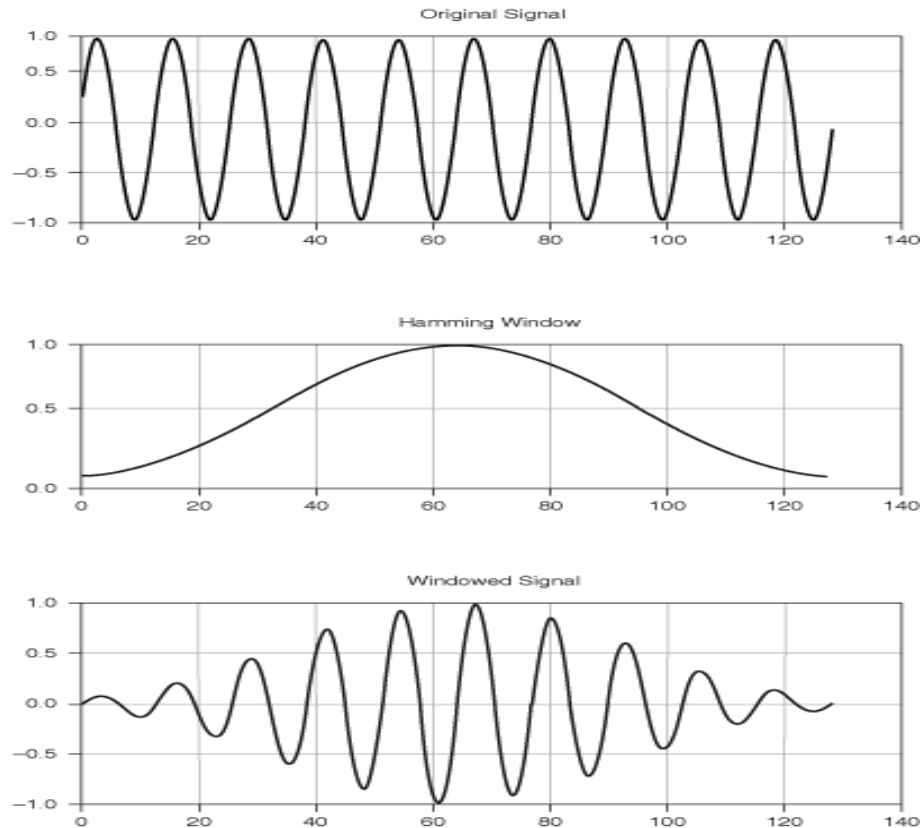
# Time Windowing for Decrease of the Spectral Leakage

- Non-rectangular signal window decreases the spectral leakage
- The choice of the *analysis window* is important, since it directly affects the spectral leakage (i.e. **side-lobe attenuation**)
- To understand the effect of the window, let us consider its effect on a complex sinusoidal signal
- It is well-known that the Fourier transform of a windowed sinusoid is the Fourier transform of the window function shifted to be centred at the frequency of the sinusoid

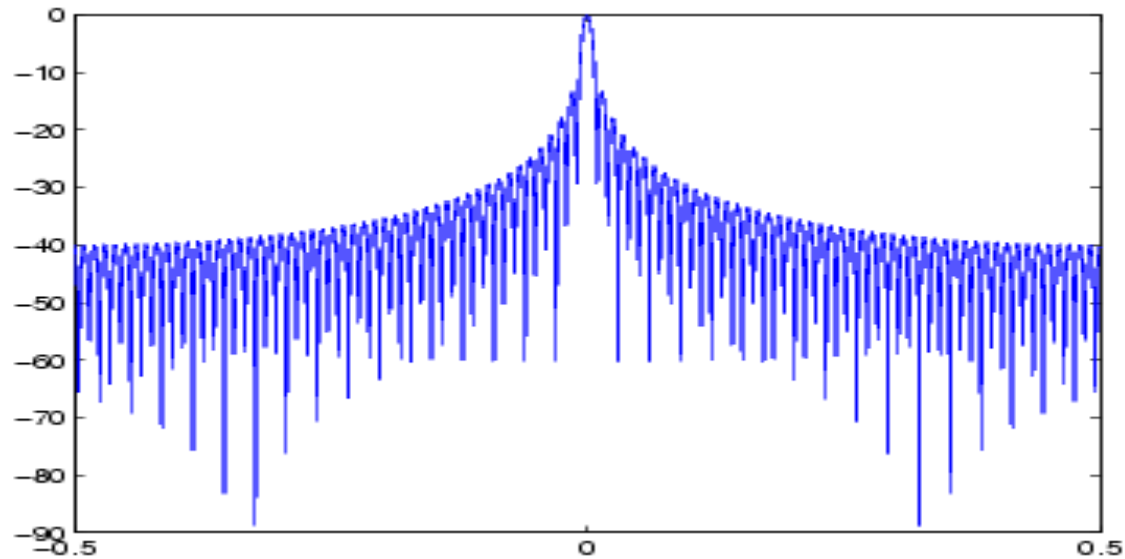
# Time Windowing for Decrease of the Spectral Leakage



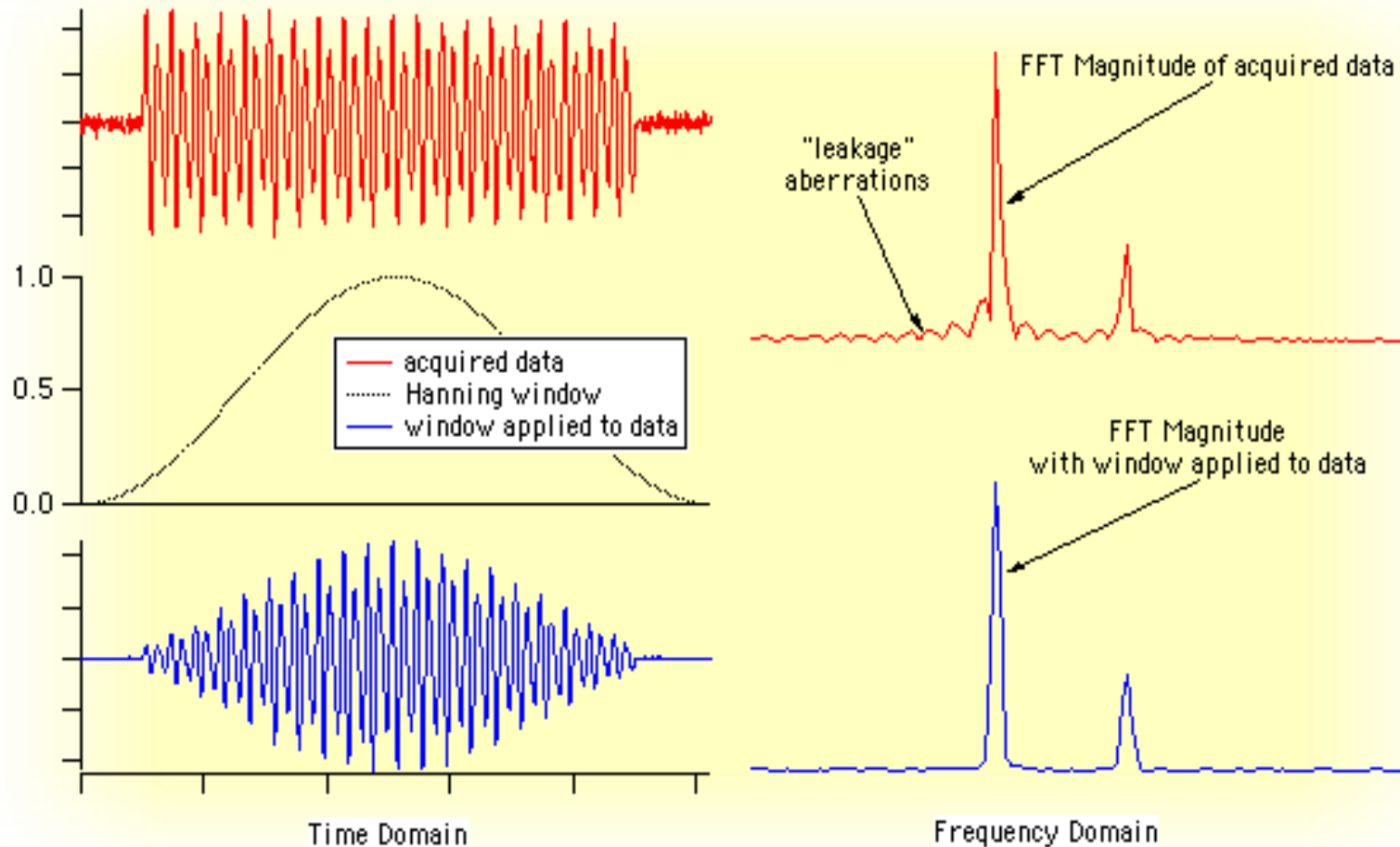
# Time Windowing for Decrease of the Spectral Leakage



# Case Study: the Windowed Sinusoid



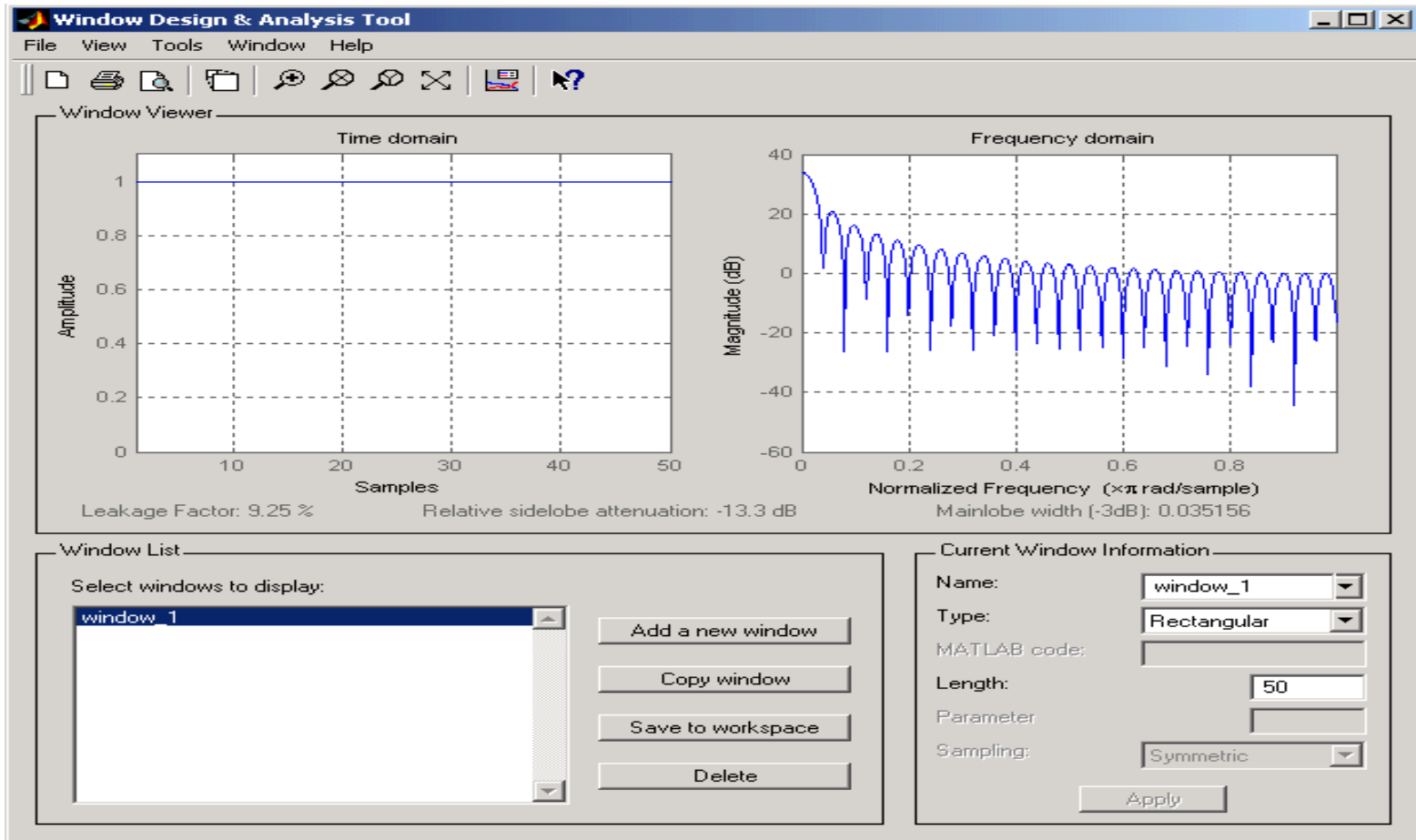
# Time Windowing for Decrease of the Spectral Leakage



# Window Selection

- A good window function must present a magnitude response (i.e. magnitude of the Fourier transform of the window function) characterized by **the ratio of the main-lobe amplitude to the largest side-lobe amplitude**
- This ratio must be **as large as possible**

# The Rectangular Window



The main problem associated with the rectangular window is the relatively low level (13.5dB) of the ratio of the main lobe to the largest side lobe

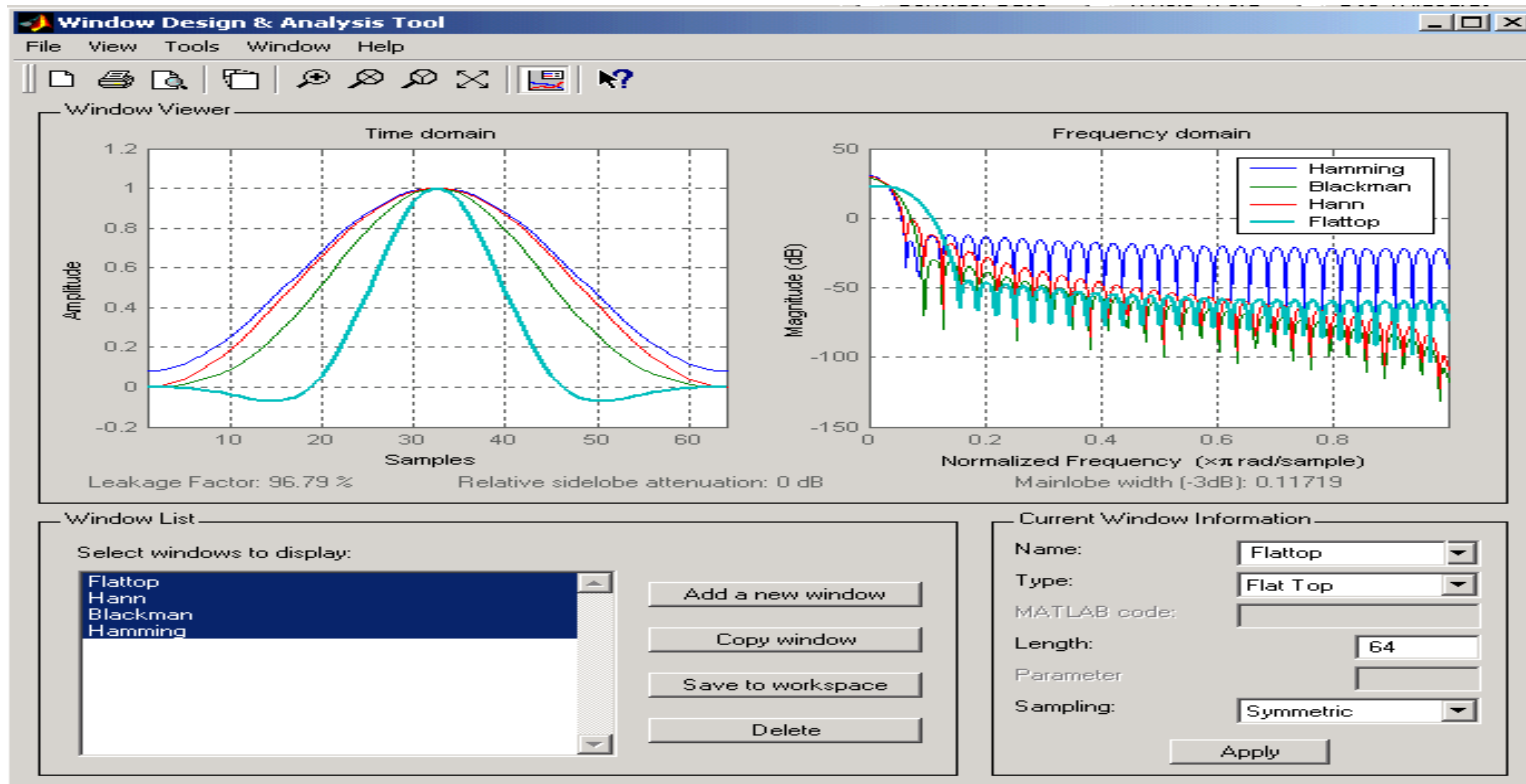
# The Rectangular and Non-Rectangular Windows

- Such a problem is due to the **inherent discontinuity** of the rectangular window in the time domain.
- One way to reduce such a discontinuity is to employ windows, which contain a taper and decays toward zero **gradually**, instead of abruptly, and, therefore, present only small discontinuities near its edges
- Literature lists several window functions that possess desirable magnitude responses: Bartlett (triangular), Blackman, Hamming, Hanning, Flat Top, Kaiser
- All of these functions have significantly increase **ratio of the main-lobe amplitude to the largest side-lobe** compared with the rectangular window



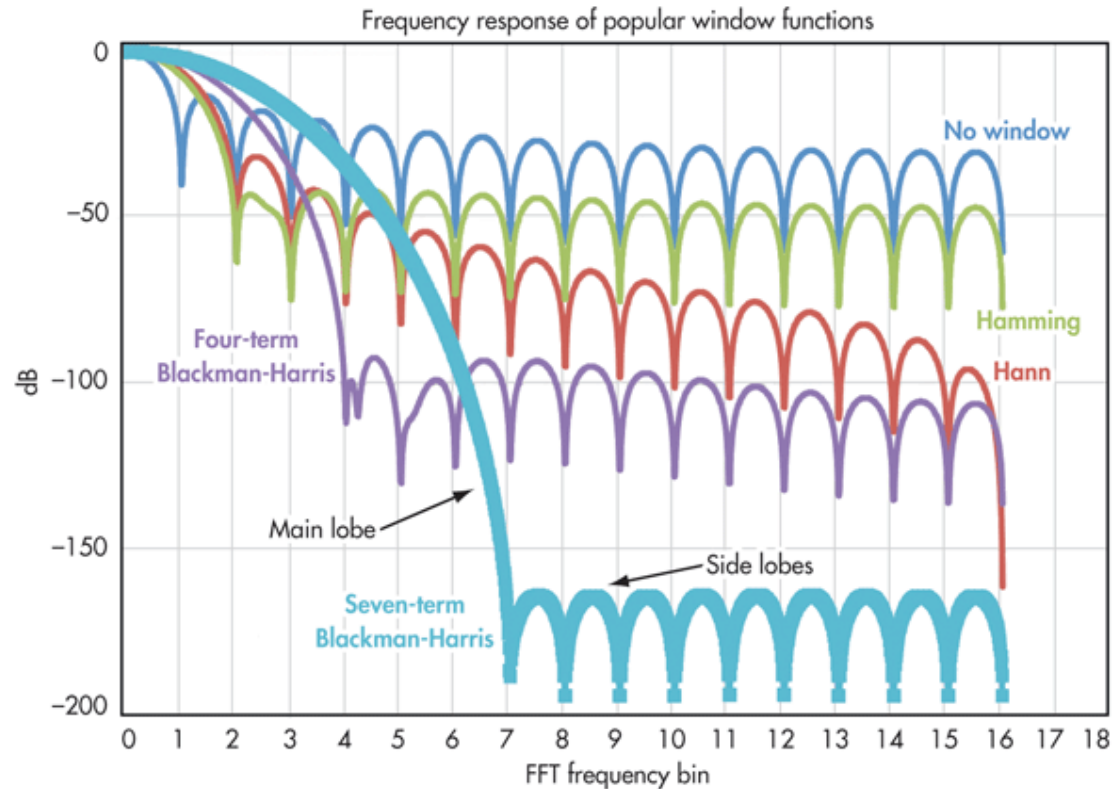
# Non-Rectangular Windows

The Blackman, Flat Top, Hamming, Hann (Hanning), and rectangular windows are all special cases of the *generalized cosine window*



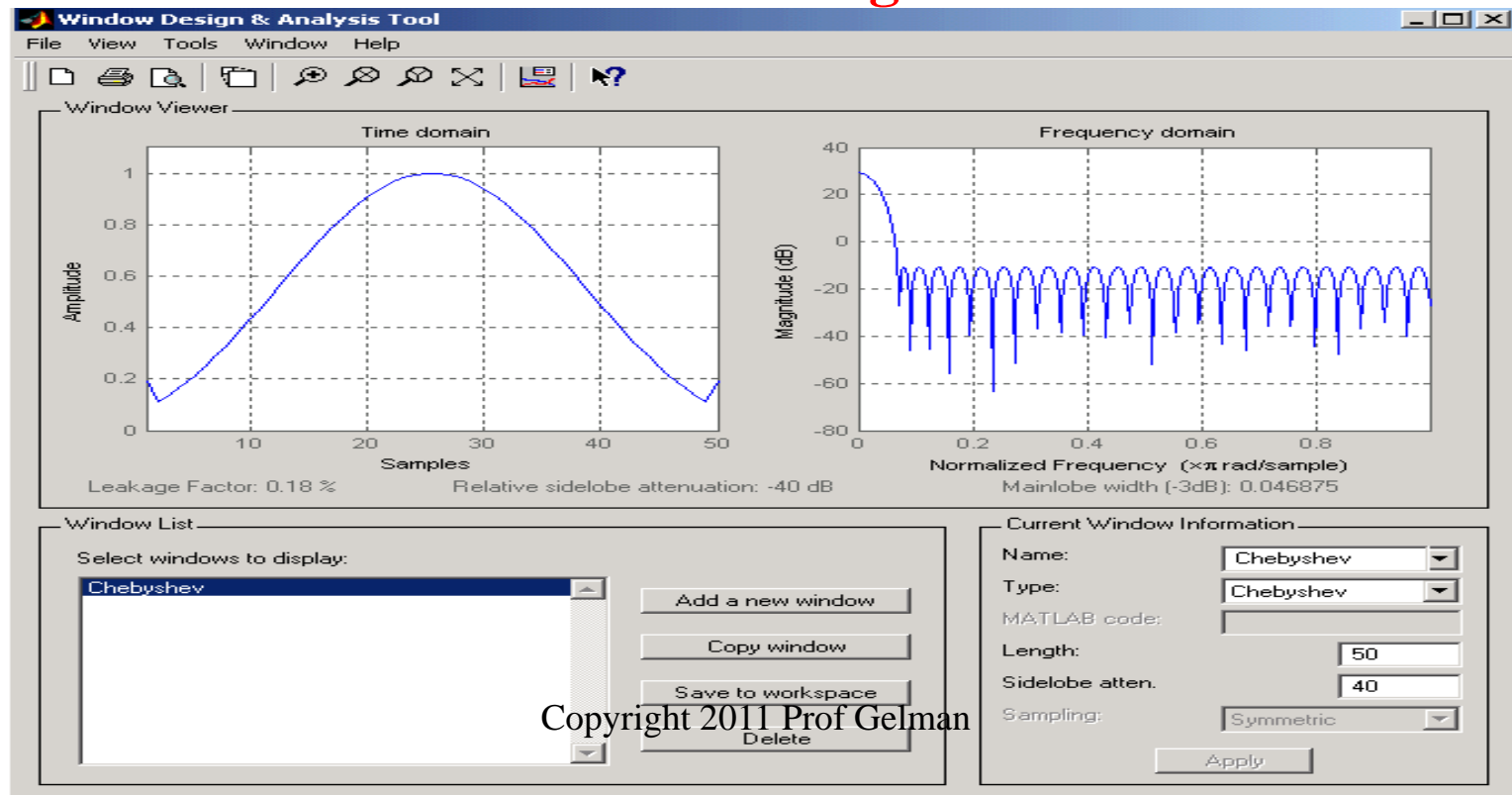
# Non-Rectangular Windows

The Blackman-Harris, Hamming, Hann, and rectangular windows



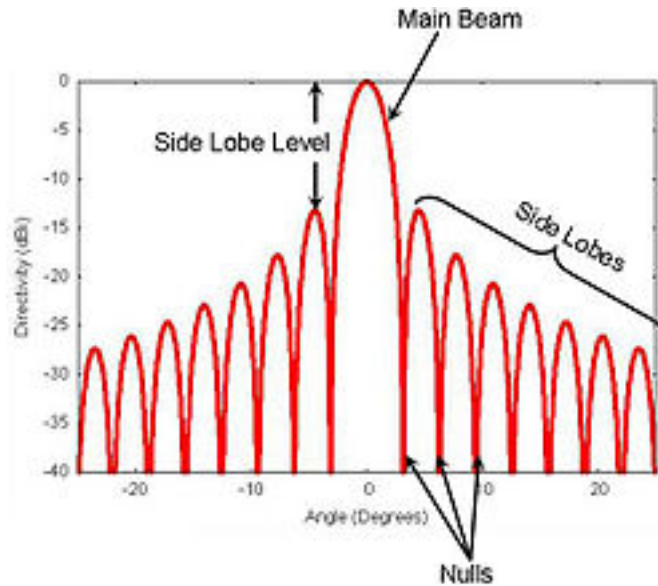
# The Chebyshev Window

- The Chebyshev window also maximizes the main lobe-side lobe ratio; The ratio of the main lobe to the side lobe is relatively high, i. e. 40 dB
- Its all side lobes have the same height



# Window Summary

Table summarizes main lobe-side lobe ratios of the various window functions



Time window	Main lobe/side lobe ratio, dB
Rectangular	13.5
Bartlett	27
Hanning	32
Hamming	43
Blackman	58