

# **Part 4 of the PSD**

## **Indirect Method of PSD Estimation**

**Prof. L. Gelman**

# Indirect Method of PSD Estimation

- For a wide-sense stationary (WSS) process, its autocorrelation function is

$$K(\tau) = E(x^*(t)x(t + \tau))$$

where  $E$  denotes the **statistical average**, the superscript asterisk denotes complex conjugation

- Then, via the **Wiener-Khinchine theorem**, the power spectral density of the WSS process is the **Fourier transform of the autocorrelation function**:

$$S_{xx}(f) = \int_{-\infty}^{\infty} K_{xx}(\tau) e^{-i2\pi f\tau} d\tau$$

# The Indirect Method of PSD Estimation

- This approach is called the *indirect method* because it requires two steps
- First, the **autocorrelation function** is computed
- Second, the Fourier transform of this function is computed (using the FFT) to obtain the power spectral density

# The Wiener-Khinchine Theorem

The Wiener-Khinchine theorem has two sides:

- the power spectral density of the **wide-sense stationary signal** is the Fourier transform of the autocorrelation function:

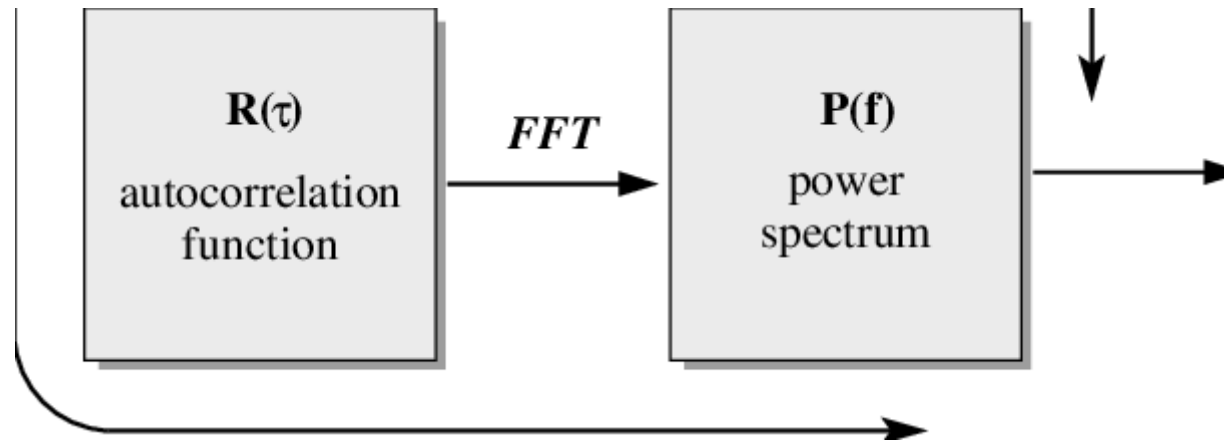
$$S_{xx}(f) = \int_{-\infty}^{\infty} K_{xx}(\tau) e^{-i2\pi f\tau} d\tau$$

- the **autocorrelation function** of the WSS signal is the **inverse Fourier transform** of the power spectral density

# The Wiener-Khinchine Theorem

- ❑ The Wiener-Khinchine theorem is a very important algorithm
- ❑ It means that the autocorrelation function and the power spectral density contain the *same information* about the signal
- ❑ Since, neither of these contains any *phase information*, it is impossible to reconstruct the signal from the autocorrelation function or from the power spectral density

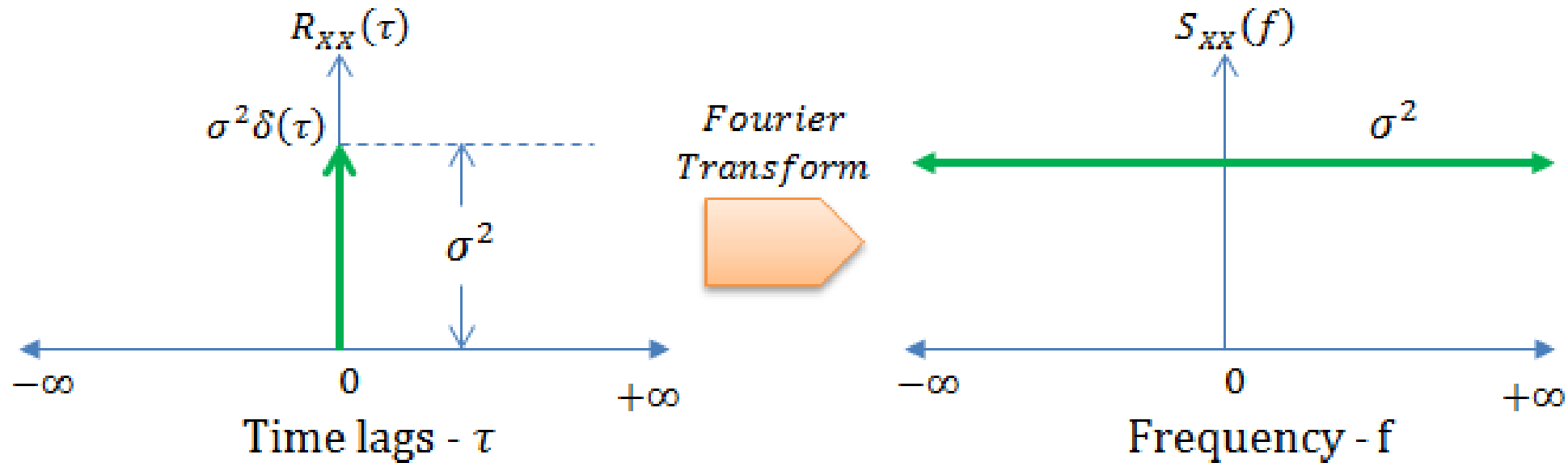
# The Indirect Method of PSD Estimation



# The Autocorrelation Functions and the PSDs

Signal	The autocorrelation function	The power spectral density (one-sided)
Sine wave	$\frac{A^2}{2} \cos 2\pi f_0 \tau$	$\frac{A^2}{2} \delta(f - f_0)$
White noise	$\frac{a}{2} \delta(\tau)$	$a$
Low-pass white noise	$aB \left( \frac{\sin 2\pi B \tau}{2\pi B \tau} \right)$	$a$ if $0 \leq f \leq B$
Band-pass white noise	$aB \left( \frac{\sin \pi B \tau}{\pi B \tau} \right) \cos 2\pi f_0 \tau$	$a$ if $f_0 - (B/2) \leq f \leq f_0 + (B/2)$

# The Autocorrelation Function and the PSD: White Noise

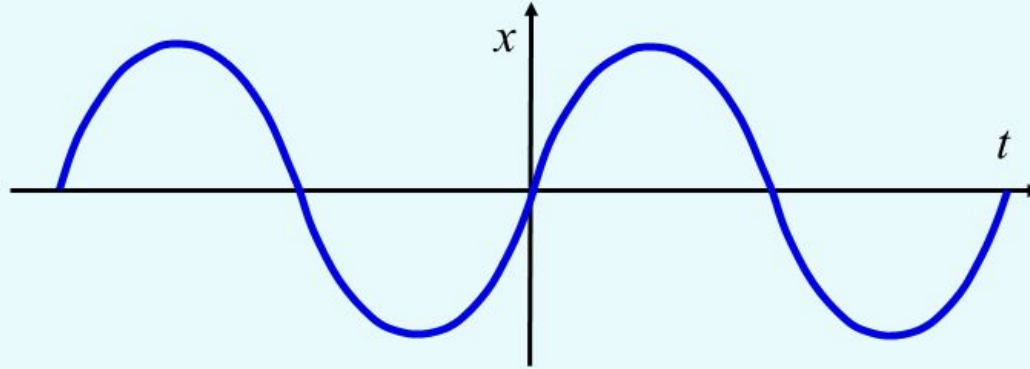




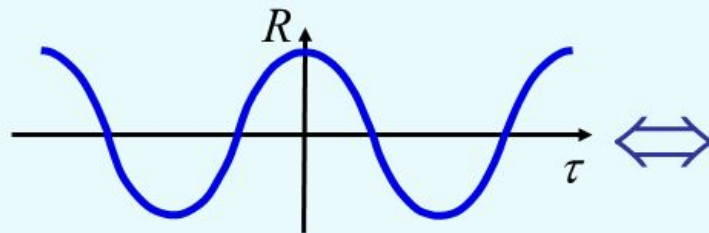
# The Autocorrelation Function and the PSD: Random Sine Wave

**Examples** of correlation functions & power spectra:

Random sine wave  $\tilde{x} = \tilde{A} \sin(\omega_0 t)$

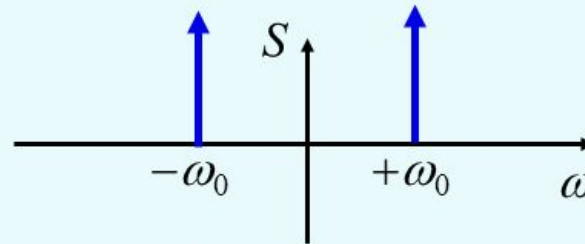


Autocorrelation:



$$R(\tau) = \frac{A^2}{2} \cos(\omega_0 \tau)$$

Power Spectrum:



$$S(\omega) = \frac{\pi A^2}{2} [\delta(\omega - \omega_0) + \delta(\omega + \omega_0)]$$

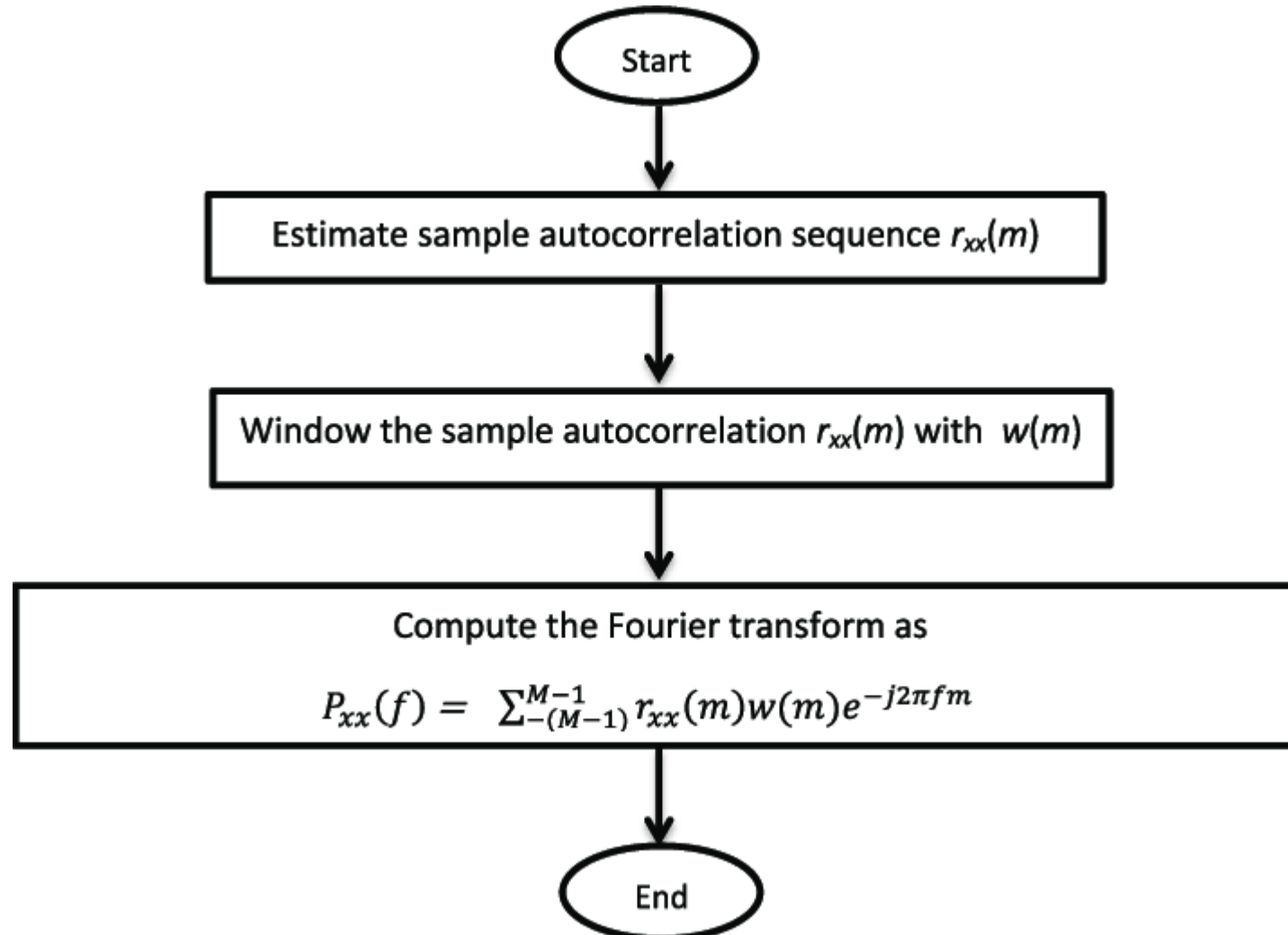
# The Blackman and Tukey Method for PSD Estimation

- Blackman and Tukey proposed the indirect method in which the estimated autocorrelation sequence **is, first, windowed** and then Fourier transformed to yield an estimate of **the power spectral density**
- Thus, the Blackman-Tukey estimate is

$$\hat{S}_{xx}^{BT}(f) = \sum_{m=-(L-1)}^{L-1} K_{xx}(m)W(m)e^{-i2\pi fm}$$

where  $W(m)$  the window function

# The Blackman and Tukey Method for Indirect PSD Estimation



# The Blackman and Tukey Method for Indirect PSD Estimation

It is clear that the effect of windowing the autocorrelation function is to smooth the periodogram estimate, thus, **decreasing the variance of the estimate** at the expense of **poorer frequency resolution**.

# Correlation Estimates via the FFT

The following steps are employed to compute the autocorrelation function via the FFT:

- Compute the **power spectral density** of a signal
- Compute the autocorrelation function using **the inverse FFT**