Applycations of laser-wakefield scaling laws

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Abstract

We apply analytical scaling laws to the parameters of the FLAME laser system, in order to estimate the maximum obtainable electron energy and total charge in a laser-wakefield acceleration scenario. We also roughly predict the expected betatron radiation spectrum, emitted by the accelerated electrons. Finally, we apply the same analysis to the case of the envisioned GBS parameters.

1. FLAME parameters

The FLAME Laser System (FLAME) has a wavelength $\lambda_L=800$ nanometers. The laser operating pulse energy can ramge from 0.03 J up to 3 J, the pulse duration can range from 30 fs up to 300 fs. The spot at waist is of $30\mu m$. We can take the minimum pulse duration to be $\tau_L=30$ fs at FWHM in intensity.

2. Electron acceleration estimate

For optimal acceleration, we impose the condition that the dephasing length should be equal to the pump depletion length $L_{\rm dephasing} = L_{\rm depletion}$ [1].

$$L_{
m dephasing} = rac{4}{3}igg(rac{\omega_L}{\omega_p}igg)^2rac{\sqrt{a_0}}{k_p} = L_{
m depletion} = igg(rac{\omega_L}{\omega_p}igg)^2c au_L$$

In practical units, we have $\lambda_p[\mu {
m m}] = 3.34 imes 10^{10} (n_p [{
m cm}^{-3}])^{-1/2}$, therefore $k_p[\mu {
m m}^{-1}] = 2\pi/\lambda_p[\mu {
m m}] = 1.88 imes 10^{-10} (n_p [{
m cm}^{-3}])^{1/2}$ and $\omega_p [{
m fs}^{-1}] = 5.64 imes 10^{-11} (n_p [{
m cm}^{-3}])^{1/2}$. This allows us to obtain the ratio a_0/n_p from above.

$$rac{a_0}{n_p [{
m cm}^{-3}]} = 1.79 imes 10^{-21} (au_L [{
m fs}])^2$$

Plasma density and normalized vector potential

If we want to use any external guiding (eg via capilaries), then we also need to impose the condition for self-guided propagation of the laser:

$$a_0 \geq \left(rac{n_c}{n_p}
ight)^{1/5}$$

with the critical density $n_c[{
m cm^{-3}}]=1.12\times 10^{21}/(\lambda_L[\mu{
m m}])^2$, which is $1.75\times 10^{21}{
m cm^{-3}}$ in our case. Substituting above and taking the lower limit, we get

$$n_p [{
m cm}^{-3}] = rac{1.12 imes 10^{21}}{a_0^5 (\lambda_L [\mu {
m m}])^2}$$

Substituting n_p in the ratio a_0/n_p from above, we finally get:

$$a_0pprox 2^{1/6}igg(rac{ au_L[ext{fs}]}{\lambda_L[\mu ext{m}]}igg)^{1/3}$$

$$n_p [{
m cm}^{-3}] = 6.29 imes 10^{20} rac{1}{(\lambda_L [\mu {
m m}])^{1/3} (au_L [{
m fs}])^{5/3}}$$

For FLAME parameters, we get $a_0\approx 3.7$ and $n_p\approx 2.3\times 10^{18}{\rm cm}^{-3}=13\times 10^{-4}n_c$. For $a_0\geq 4-5$ we also get self-injection from pure Helium. Helium has the ionization energies 24.59 eV (He⁺) and 54.42 (He²⁺), corresponding to laser intensities 1.4×10^{15} , respectively 8.8×10^{15} W/cm² (Gibbon, "Short pulse laser interactions with matter", p. 22), and will therefore be easily ionized by the laser prepulse. Knowing the density, we get $\omega_p=8.63\times 10^{-2}$ fs⁻¹, and more importantly, $\omega_p^{-1}=11.59$ fs, which is the same order of magnitude as the pulse duration $\tau_L=30$ fs. For $\tau_L\gg \omega_p^{-1}$ (eg. $\tau_L=600$ fs), one would either be in the SMLWFA or DLA regime, depending on the value of a_0 .

Beam waist

We also need to match the beam waist w_0 of the focused Gaussian laser pulse to the plasma, giving the condition $w_0k_p=2\sqrt{a_0}$, so

$$w_0[\mu{
m m}] = 1.06 imes 10^{10} igg(rac{a_0}{n_n [{
m cm}^{-3}]}igg)^{1/2} = 44.84 rac{ au_L [{
m fs}]}{100}$$

For FLAME, we get $w_0 pprox 13 \mu {
m m}$ at $1/e^2$ intensity, so FWHM (1/2 intensity) = $w_0 \sqrt{2 \ln 2} pprox 18 \mu {
m m}$.

Laser energy, power and intensity

We now estimate the required laser energy ε_L . The peak laser intensity in the focal plane is

$$I_L[10^{18}\mathrm{W/cm}^2]pprox 6 imes 10^4rac{arepsilon_L[\mathrm{J}]}{ au_L[\mathrm{fs}](w_0[\mu\mathrm{m}])^2}$$

so for our parameters $I_L\approx 2.96\times 10^{19}~{
m W/cm}^2$. For a relative scale, the atomic Coulomb field is on the order of $10^{14}~{
m W/cm}^2$ and relativistic effects become important for laser intensities above $10^{17}~{
m W/cm}^2$ ($a_0\geq 1$), while QED effects such as radiation reaction only become important for intensities beyond $\sim 2\times 10^{21}~{
m W/cm}^2$. We know that $a_0=0.855\lambda_L[\mu{
m m}](I_L[10^{18}{
m W/cm}^2])^{1/2}$, so

$$arepsilon_L[{
m J}] = 2.28 imes 10^{-5} a_0^2 (w_0[\mu{
m m}])^2 rac{ au_L[{
m fs}]}{(\lambda_L[\mu{
m m}])^2} = 5.68 imes 10^{-6} rac{(au_L[{
m fs}])^{11/3}}{(\lambda_L[\mu{
m m}])^{8/3}}$$

Therefore, for FLAME we need $arepsilon_Lpprox 2.7$ J on target. The laser power on target is

$$P[ext{TW}] = 2.41 imes 10^{18} rac{a_0^3}{n_p [ext{cm}^{-3}]} rac{1}{(\lambda_L [\mu ext{m}])^2}$$

$$P[\mathrm{TW}] = 2 imes 10^3 \sqrt{rac{\ln 2}{\pi}} rac{1}{ au_L[\mathrm{fs}]} arepsilon_L[\mathrm{J}] = 5.34 imes 10^{-3} igg(rac{ au_L[\mathrm{fs}]}{\lambda_L[\mu\mathrm{m}]}igg)^{8/3}$$

so approximately 84 TW in our case.

Acceleration distance

The optimum propagation distance is equal to $L = L_{\rm dephasing} (= L_{\rm depletion})$.

$$L[{
m mm}] = 3.34 imes 10^{17} rac{ au_L[{
m fs}]}{(\lambda_L[\mu{
m m}])^2} rac{1}{n_p[{
m cm}^{-3}]} = 5.31 imes 10^{-4} rac{(au_L[{
m fs}])^{8/3}}{(\lambda_L[\mu{
m m}])^{5/3}}$$

which for FLAME amounts to $L \approx 7$ mm.

Maximum electron energy and charge

The maximum electron energy that can be reached under these circumstances is

$$egin{aligned} \gamma &= rac{\Delta arepsilon}{mc^2} = rac{2}{3} a_0 rac{\omega_L^2}{\omega_p^2} \ \ \gamma &= 7.44 imes 10^{20} rac{a_0}{n_p [ext{cm}^{-3}]} rac{1}{(\lambda_L [\mu ext{m}])^2} = 1.33 igg(rac{ au_L [ext{fs}]}{\lambda_L [\mu ext{m}]}igg)^2 \ \ \Delta arepsilon [ext{GeV}] &= 5.11 imes 10^{-4} \gamma = 6.8 imes 10^{-4} igg(rac{ au_L [ext{fs}]}{\lambda_L [\mu ext{m}]}igg)^2 \end{aligned}$$

For FLAME, we get $\Delta \varepsilon = 0.95$ GeV and a Lorentz factor $\gamma = 1870$.

We assume the blowout radius $R=w_0=44.84 au_L[{
m fs}]/100=18~\mu{
m m}$. The number of accelerated electrons $N_e\propto R^3n_p$ is equal to the ionic charge of the bubble, so (Lu et. al, 2007)

$$N_epprox 3.13 imes 10^8 \lambda_L [\mu{
m m}] (P[{
m TW}])^{1/2} = 2.28 imes 10^7 rac{(au_L [{
m fs}])^{4/3}}{(\lambda_L [\mu{
m m}])^{1/3}}
onumber \ N_epprox 4.85 imes 10^{17} rac{a_0^{3/2}}{(n_p [{
m cm}^{-3}])^{1/2}}$$

giving a charge

$$Q_e[ext{pC}] = N_e imes e[ext{pC}] pprox 3.65 rac{(au_L[ext{fs}])^{4/3}}{(\lambda_L[\mu ext{m}])^{1/3}}$$

For FLAME we get $Q_e \approx 366\,\mathrm{pC} \approx 0.36$ nC and $N_e = 2.29 \times 10^9$ electrons. As a last note, with PW lasers, the higher laser energy can be focused to a larger focal spot matched by a lower plasma density.

External guiding

3J, $a_0=2$, $n_p=5.1\times 10^{17}$ cm $^{-3}$, $w_0=21\,\mu\mathrm{m}$, $au_L=47\,\mathrm{fs}$ - needs external guiding (capilaries) - needs external injection - can reach $\Delta\varepsilon=2.4$ GeV after 52 mm of plasma

3. Betatron emission

After matching the laser spot size, and pulse duration to the plasma density, we can look at betatron emission [2]. The betatron oscillation radius $r_{\beta}=3~\mu{\rm m}$ for a blowout radius $R=21~\mu{\rm m}$. Since the betatron oscillations are limited by the bubble size, it is reasonable to assume a linear dependence on R and therefore take $r_{\beta}=2.6~\mu{\rm m}$ in our case.

We now need to work out the five relevant parameters for the wiggler: the relativistic electron factor γ , the number of electrons N_e , the strength parameter K, the period of the wiggler λ_u and the number of (betatron) periods N. We already know the first two parameters. Once we have the parameters, we can calculate the critical radiation energy $\hbar\omega_c$, the opening angle of the radiation θ_r and the number of emitted photons per electron and per betatron period, N_γ . We can then get the total number of emitted photons per laser shot, $N_\gamma \times N_e \times N$.

$$egin{aligned} heta_r[\mathrm{mrad}] &= 10^3 imes rac{K}{\gamma} \ N_\gamma &= 3.31 imes 10^{-2} K \ \hbar \omega_c[\mathrm{keV}] &= 1.86 imes 10^{-3} rac{K \gamma^2}{\lambda_u[\mu \mathrm{m}]} \end{aligned}$$

Wiggler parameters

In the case of betatron oscillations, the strength parameter is given by

$$K = \sqrt{rac{\gamma}{2}} \left(k_p [\mu {
m m}^{-1}]
ight) (r_eta [\mu {
m m}]) = 1.33 imes 10^{-10} r_eta [\mu {
m m}] (\gamma \, n_p [{
m cm}^{-3}])^{1/2}$$

so $K \approx 26$ in our case, which puts us in the wiggler regime $K \gg 1$.

$$\lambda_u[\mu{
m m}] = \sqrt{2} \gamma^{1/2} (\lambda_p[\mu{
m m}]) = 4.72 imes 10^{10} igg(rac{\gamma}{n_p[{
m cm}^{-3}]}igg)^{1/2}$$

so $\lambda_u=1334\,\mu\mathrm{m}=1.33$ mm.

Radiation properties

We now easily obtain $\theta_r=14$ mrad and $N_\gamma=0.87$ photons / (electron x betatron period). While the electrons radiate throughout the whole acceleration process, theefficiency main contribution comes from the part of the trajectory where their energy is maximum, ie, after the distance L. If the electron is around its maximum energy for about $N\sim 3$ betatron periods, the total number of emitted photons per shot will be $N_\gamma^{\rm shot}=N_\gamma\times N_e\times N=0.87\times 2.28\times 10^9\times 3\approx 6\times 10^9$ photons. Finally, the critical energy is in our case $\hbar\omega_c=128$ keV.

Note: The betatron amplitude $r_{eta} \propto \sqrt{a_0/n_p}$ and the numer of betatron oscillations $N \propto 1/\sqrt{n_p}$.

4. Quantum effects

We are now ready to estimate the size of the quantum corrections. Quantum effects become important when the electron energy loss due to photon emission becomes comparable to the electron energy.

Nonlinear quantum parameter

We first look at the so-called nonlinear quantum parameter χ_0 defined in (Di Piazza et al., 2012) for the case of an ultrarelativistic electron counter-propagating with respect to a plane wave:

$$\chi_0 = 5.9 imes 10^{-3} \Delta arepsilon [{
m GeV}] (I_L [10^{18} {
m W/cm^2}])^{1/2}$$

giving $\chi_0 \approx 3 \times 10^{-2}$. χ_0 can be interpreted as the aplitude of the plane wave's electric field in the electron rest frame (in units of the critical QED field $m^2/e=1.3\times 10^{16}$ V/cm $=1.3\times 10^6$ TV/m) and controld quantum effects such as photon recoil and spin. Since $\chi_0 \ll 1$, these effects play a minor role here. In fact our electric field is just $E_L[{\rm TV/m}]=3.21a_0/\lambda_L[\mu{\rm m}]=0.151$ TV/cm. Inside the plasma, the maximum accelerating field achievable at a given plasma density is $\approx a_0^{1/2}E_0$ (for $a_0 \geq 2$), where E_0 is the cold nonrelativistic wave-breaking field, $E_0[{\rm V/cm}]\approx 0.96(n_p[{\rm cm}^{-3}])^{1/2}$. In our case $a_0^{1/2}E_0=2.84\times 10^9$ V/cm.

Interaction duration

The other way to estimate quantum effects is by following (Corde et al., 2013), which explains that radiation reaction can be neglected when the electron - laser interaction duration is much smaller than the electron energy loss rate, or equivalently, the number of oscillations $N \ll N_{\rm RR}$, with

$$N_{
m RR} = 2.7 imes 10^7 rac{\lambda_u [\mu {
m m}]}{\gamma K^2}$$

so for FLAME** parameters $N_{
m RR}\sim 2.8 imes 10^4\gg N.$

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