

Lecture 4:Uncertainty

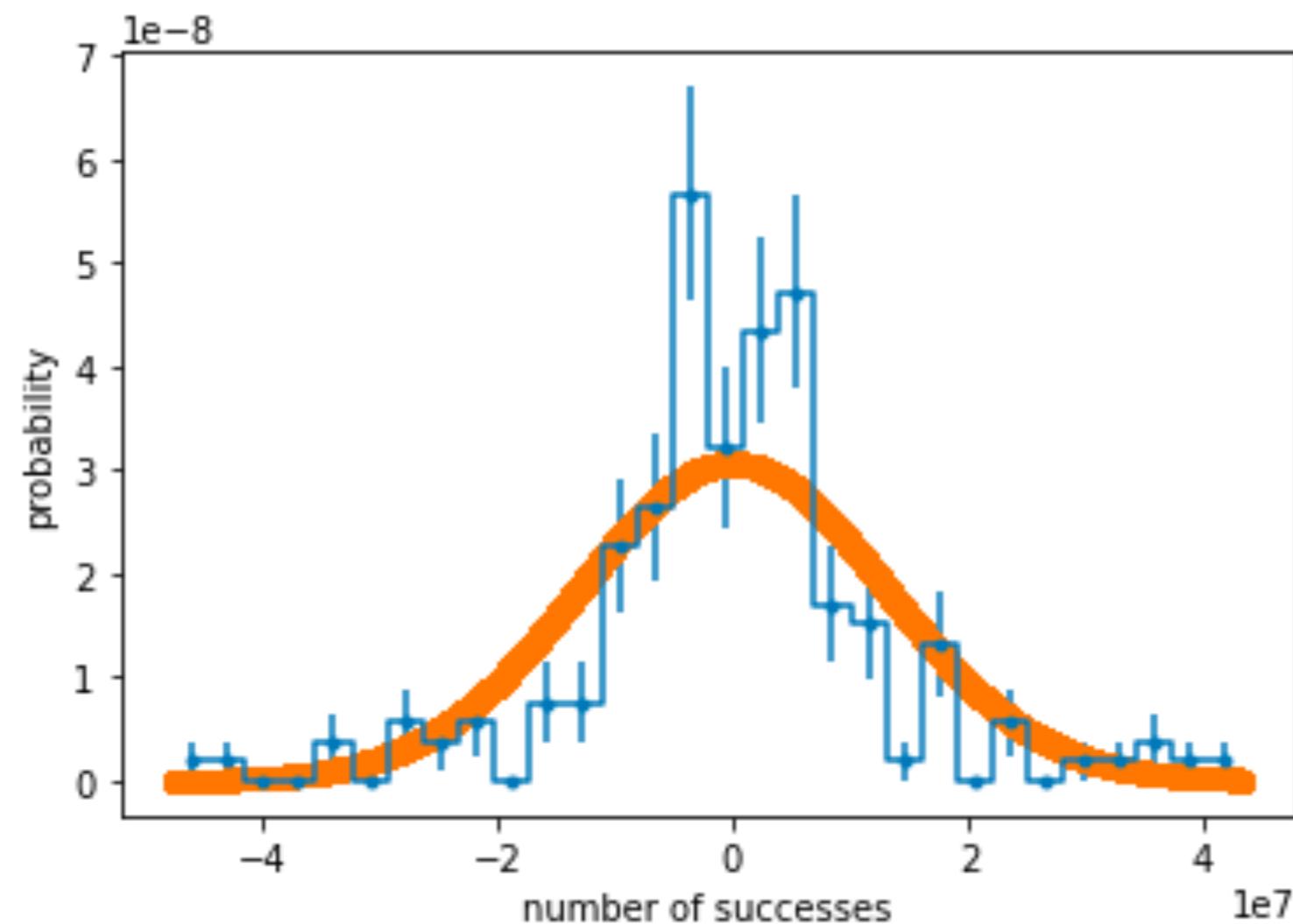
Uncertainty

What are we fitting?

- Lets say a fitted parameter has a distribution
 - We would like to quote the uncertainty on that distribution
 - Traditionally we do this by computing the standard deviation
 - Standard Deviation $\sigma = \sqrt{Var[\theta]}$
 - Standard deviation is just a number
 - It does not reflect what the actual distribution was about

How do we deal with unc?

- Our uncertainties have the same standard deviation
 - You may know there is an implicit assumption its Gaussian



Likelihood

Hikers⁶ in Scotland



A mathematician, a physicist, and an engineer are hiking through Scotland.

The engineer looks out the window, sees a black sheep, and exclaims, "Hey! They've got black sheep in Scotland!"

The physicist looks out the window and corrects the engineer, "Strictly speaking, all we know is that there's at least one black sheep in Scotland."

The mathematician looks out the window and corrects the physicist, " Strictly speaking, all we know is that at least one side of one sheep is black in Scotland."

What is Likelihood?

- Lets take an example:
- 40% of all humans have smelly pee after they eat asparagus
 - Ref: Scientific American Article.
- If you have a room with 56/100 people who have smelly pee
- What is the probability? (binomial)

$$P(56 | p = 40\%) = p^{56}(1 - p)^{44} \frac{100!}{44!56!} \quad 4 \times 10^{-4}$$

- What is the Likelihood of this happening? (given 40%)

$$\mathcal{L}(56 | p = 40\%) = p^{56}(1 - p)^{44} \frac{100!}{44!56!} \quad 4 \times 10^{-4}$$

What is the Max Prob/Like?

- The maximum probability of this distribution occurs at k=40:

$$P(40 | p = 40\%) = p^{40}(1 - p)^{60} \frac{100!}{40!60!} = 0.08$$

- The maximum likelihood occurs at $p=0.56$:

$$\frac{\mathcal{L}(56 | p)}{dp} = 0 = \frac{d}{dp} \left(p^{56}(1 - p)^{44} \frac{100!}{44!56!} \right)$$

$$56(p^{55}(1 - p)^{44}) - 44(p^{56}(1 - p)^{43}) = 0 \rightarrow 56(1 - p) - 44p = 0$$

$$56 - 100p = 0 \rightarrow p = 56/100$$

- Likelihood we vary the underlying parameters to likelihood
- Probability we vary our tests to maximize probability

What is the Log Likelihood?

- The log like of the likelihood:

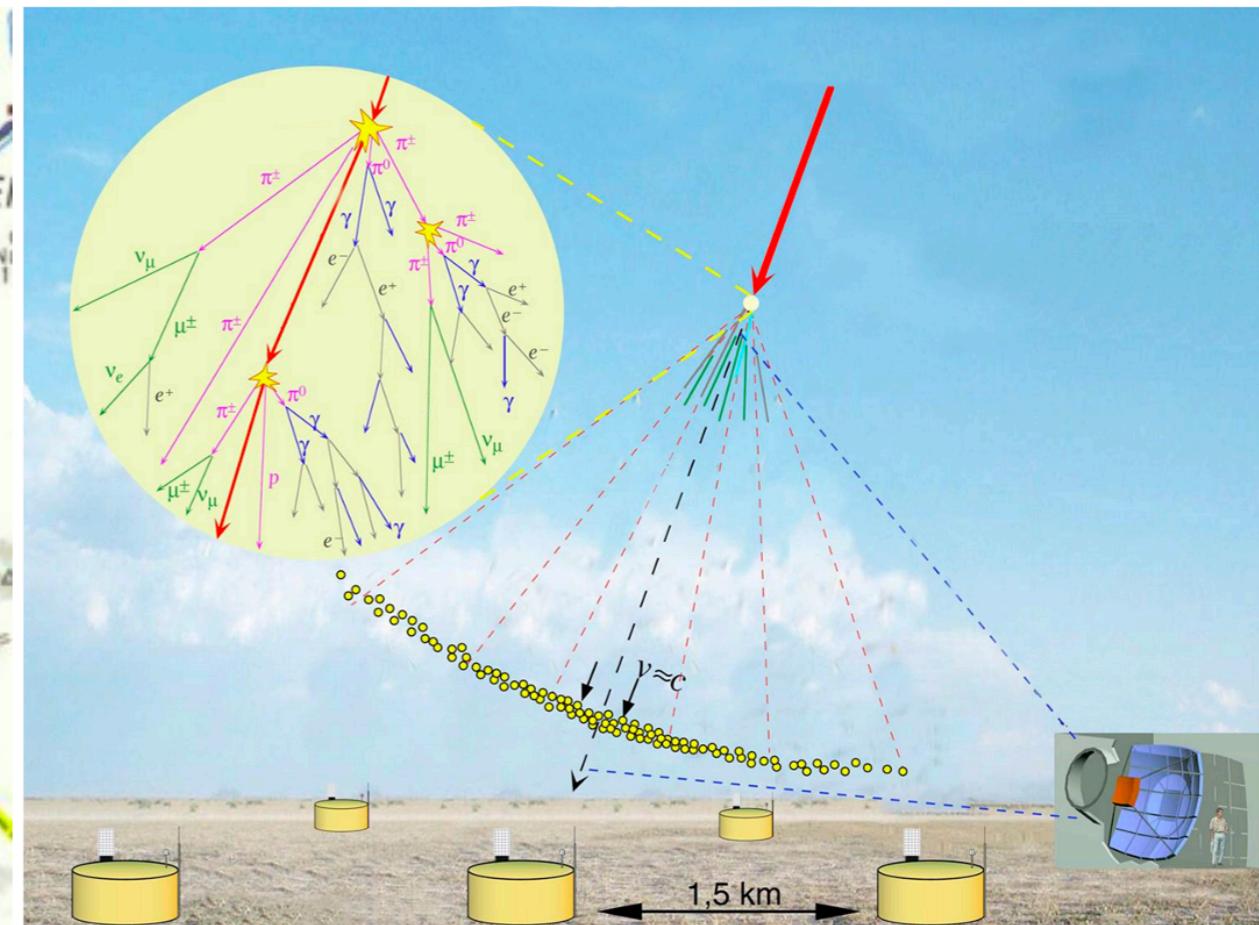
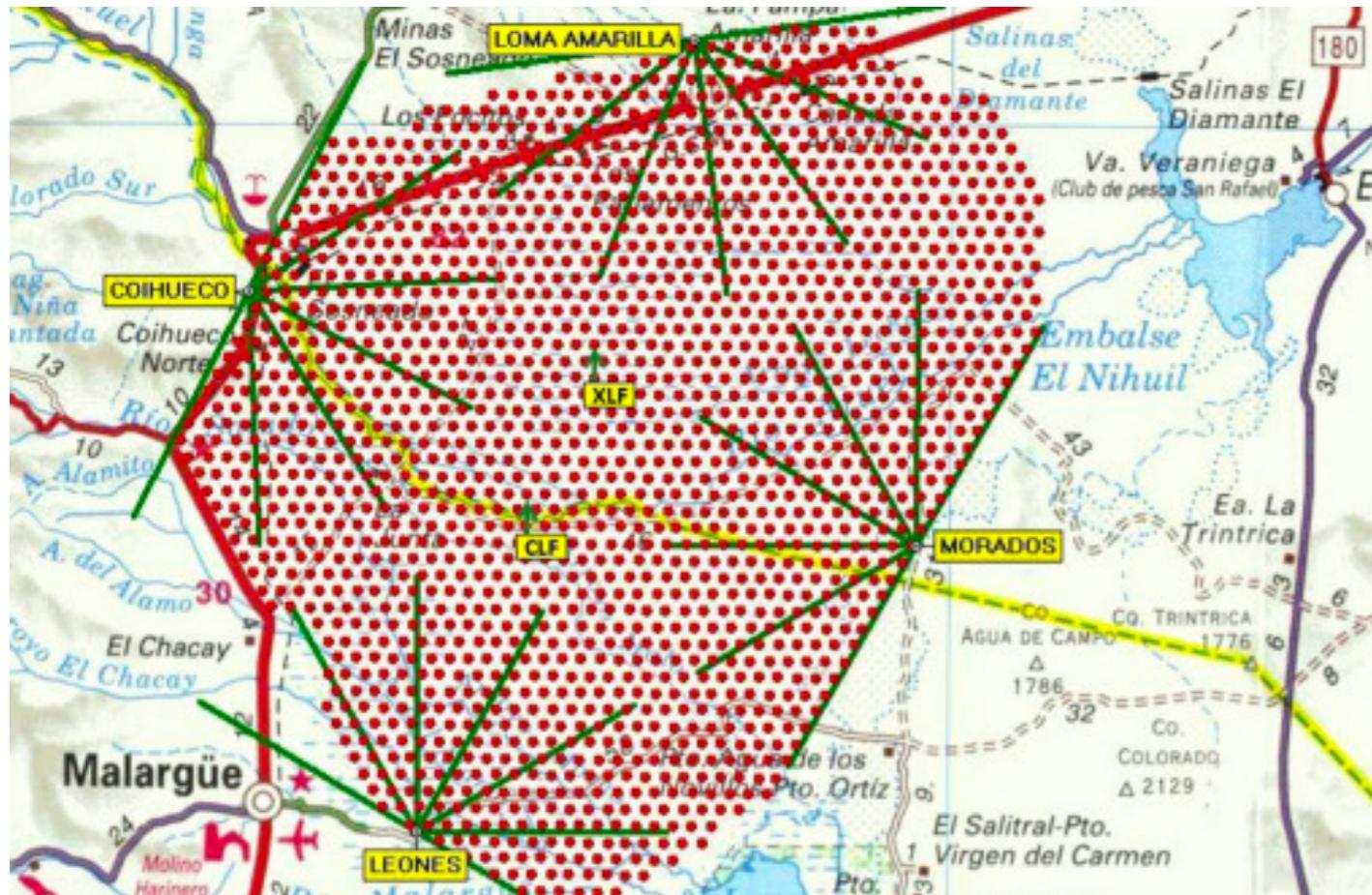
$$\log(\mathcal{L}(56 | p = 40\%)) = 56 \log(p) + 44 \log(1 - p) + \log\left(\frac{100!}{44!56!}\right)$$

- Sometimes the likelihood can get very large
 - That's why we take the log
 - Log is also a 1 to 1 mapping so

$$\max(\log(\mathcal{L})) = \max \mathcal{L}$$

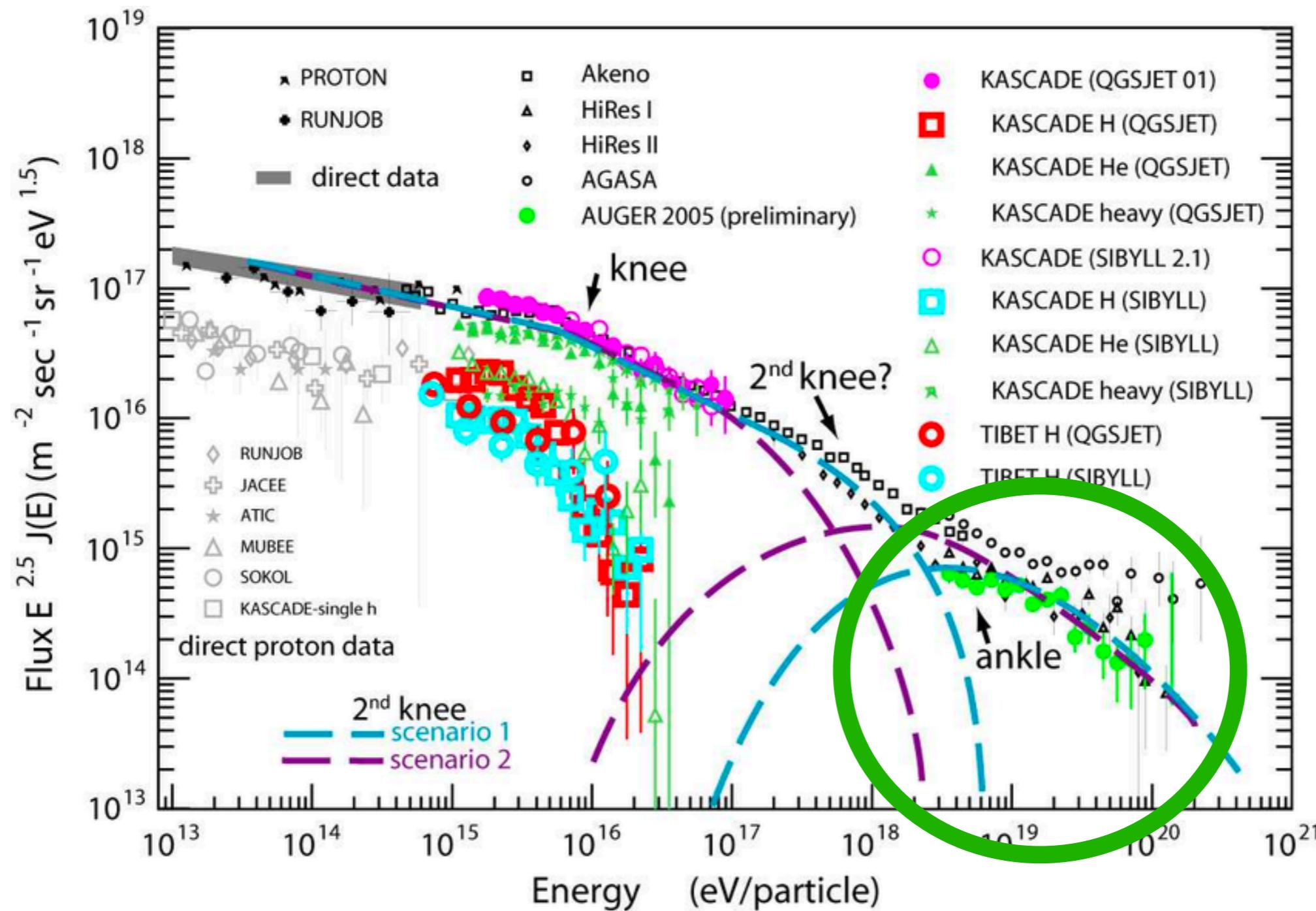
Auger Experiment

Auger Experiment



- Experiment is a giant array of detectors
- These detectors reconstruction particle showers
 - Particles are the highest energy particles in the world
- In the end you get a position and energy of each particle

What we learn w/Auger?



- We probe the highest energy particles in the world w/Auger
- Nobody knows where they come from!

Auger Data

- Data can be found here :
 - <https://www.auger.org/index.php/science/data>
- We are going to analyze data from here :
 - [https://www.auger.org/index.php/document-centre/
viewdownload/115-data/4643-ra-data-event-lists-flux-map](https://www.auger.org/index.php/document-centre/viewdownload/115-data/4643-ra-data-event-lists-flux-map)
- Here is the paper on it:
 - Science 357 (2017) 1266-1270
 - <https://arxiv.org/abs/1709.07321>

Auger Data

- Data format:

Two Datasets

```
## events_a4 : Particles with energy from 4–8 EeV
## events_a8 : Particles with energy > 8 EeV
## Columns:
#
```

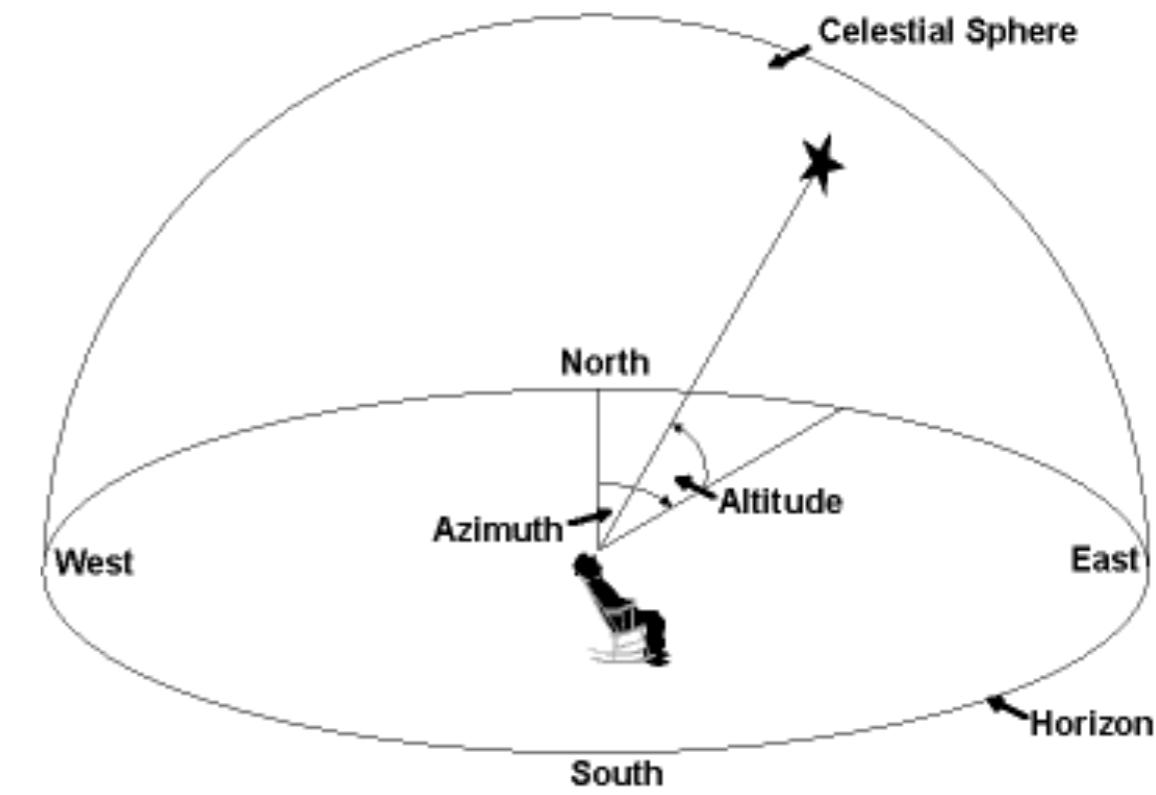
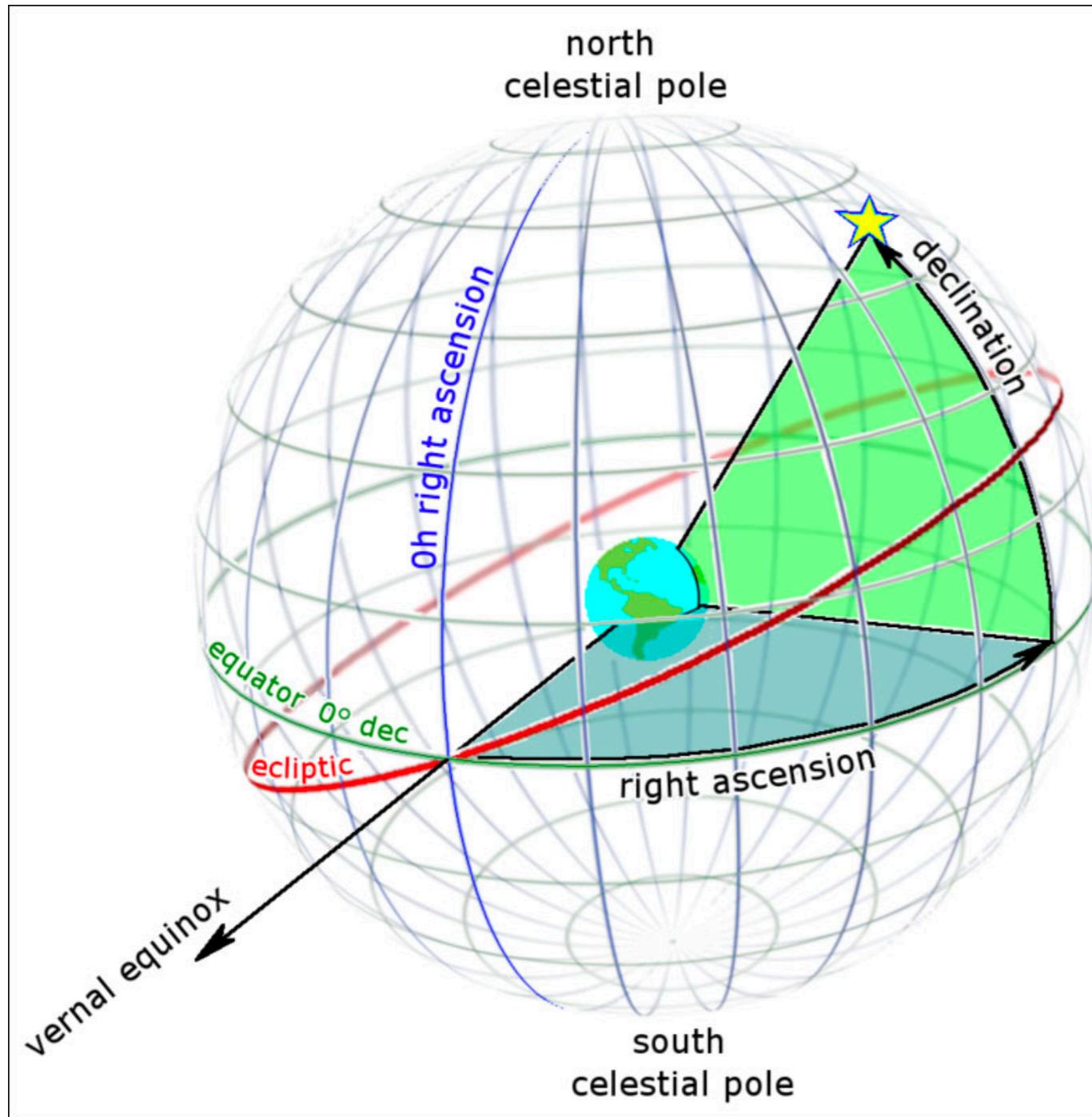
```
# (1) year
# (2) day
# (3) declination (deg.)
# (4) right ascension (deg.)
# (5) local azimuth (deg.)
# (6) weight
```

	#	year	day	dec	RA	azimuth	weight
2004	1	-11.8	208.9	164.7	1.0040		
2004	2	-70.4	36.6	-111.6	1.0001		
2004	3	-47.8	334.7	-143.8	0.9973		
2004	3	-22.1	87.2	-162.8	1.0080		
2004	7	-5.8	178.6	162.7	1.0046		
2004	9	3.4	60.5	82.9	1.0072		
2004	12	2.6	242.2	126.2	1.0007		

Earth Coordinates in Equatorial form

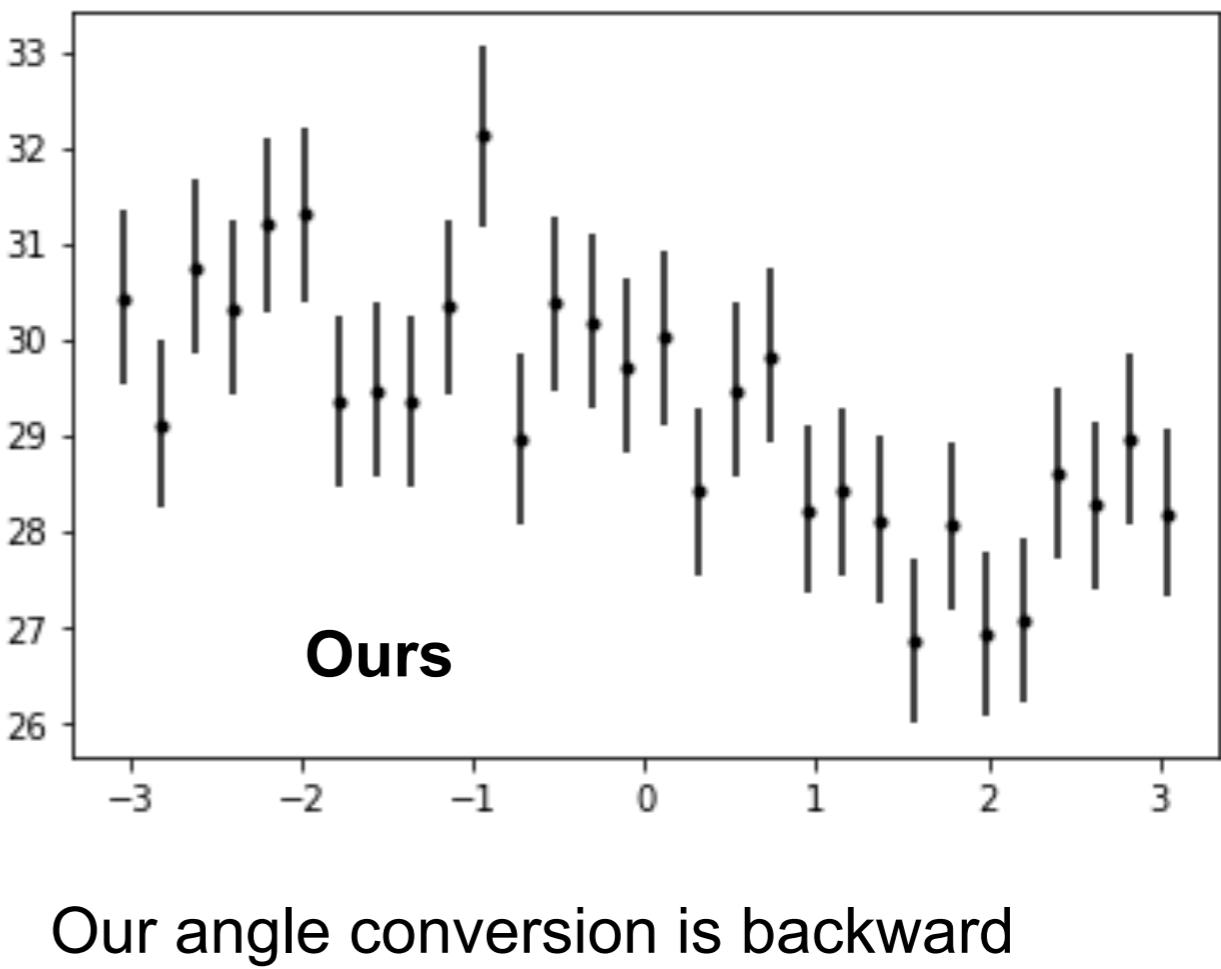
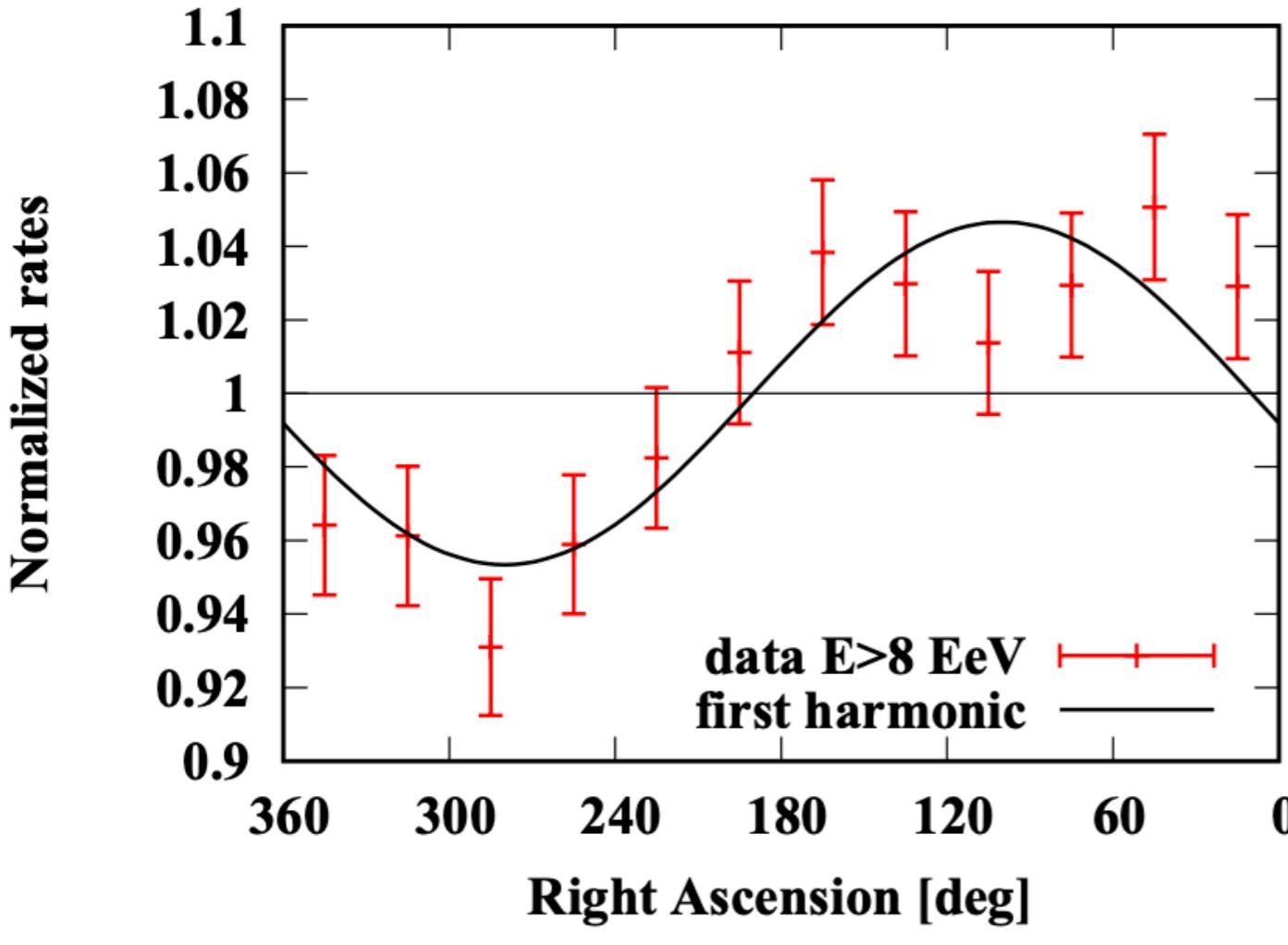
Uncertainty correction for data
its close to 1 we will ignore it

Cosmic Rays



To understand Auger Data
We need to understand
Spherical Coordinates

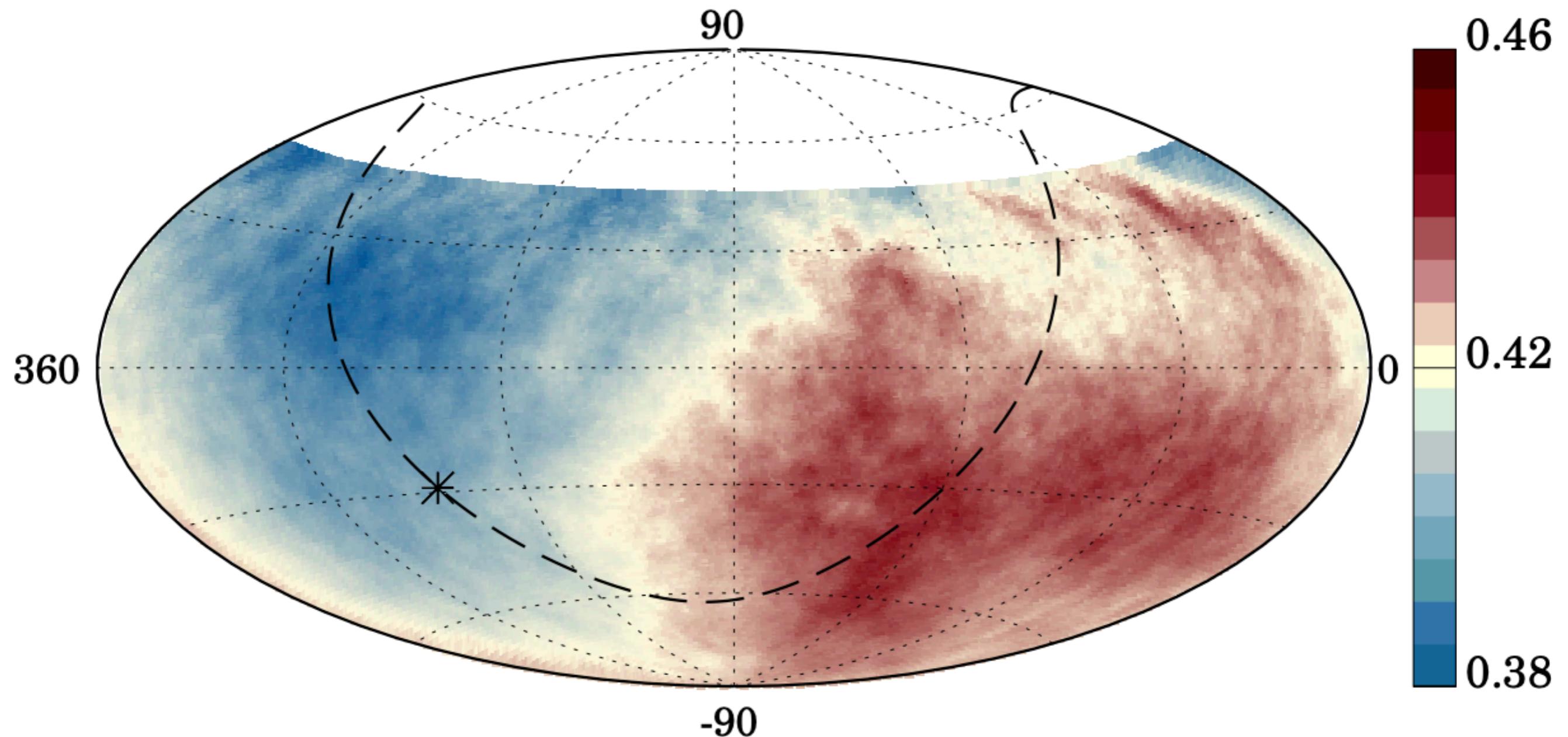
From the paper we find



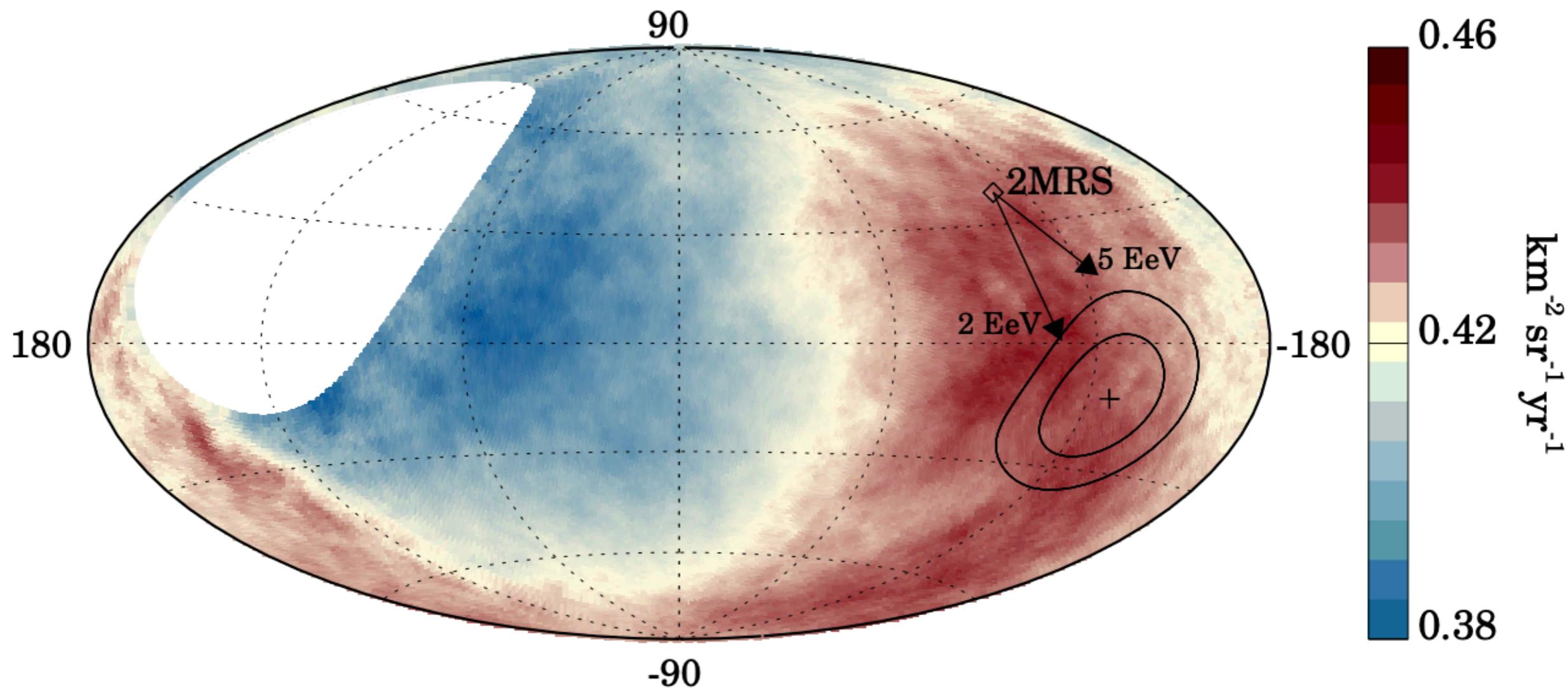
- An asymmetry in right ascension
- Our x-value is flipped and we didn't normalize (same trend)

We also find

To get this you need to know details about acceptance
I have a guess in the notebook



In Galactic Coordinates

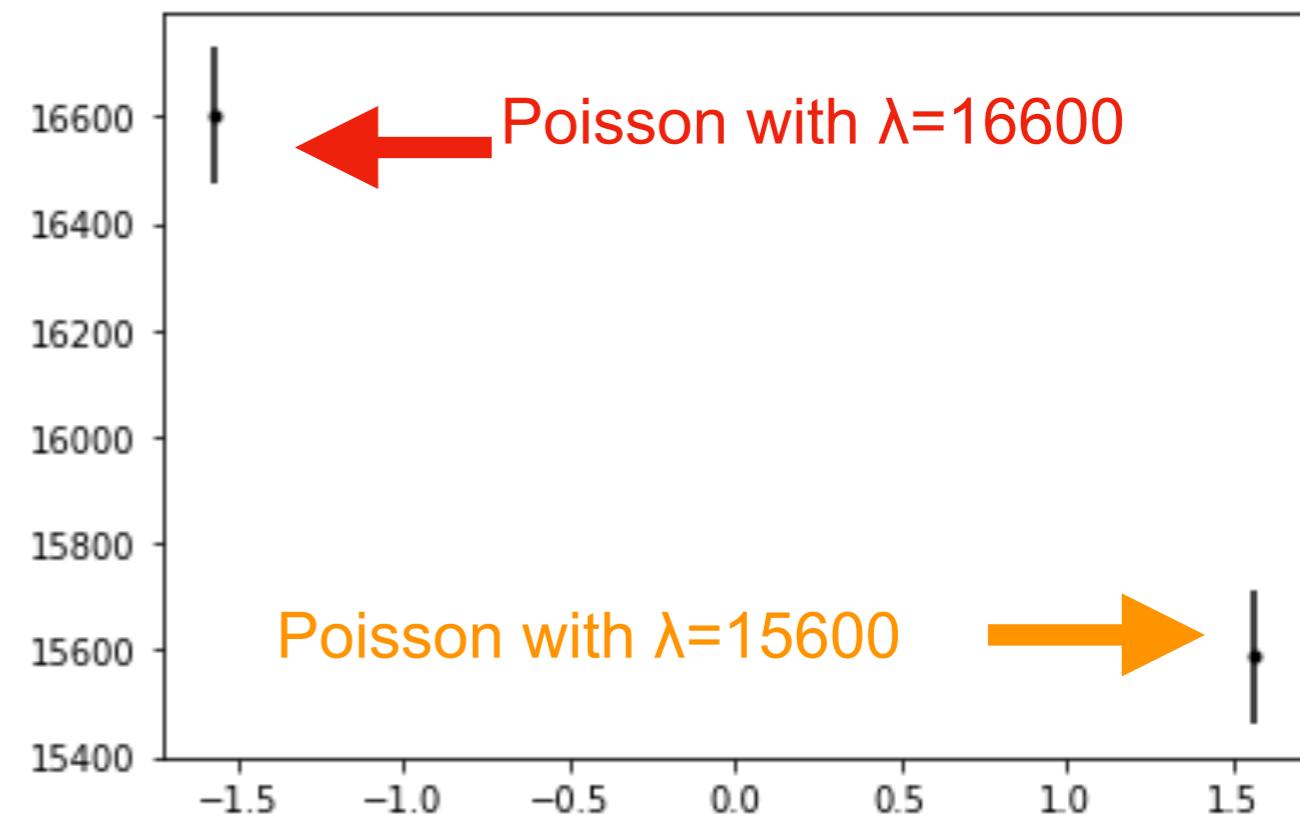
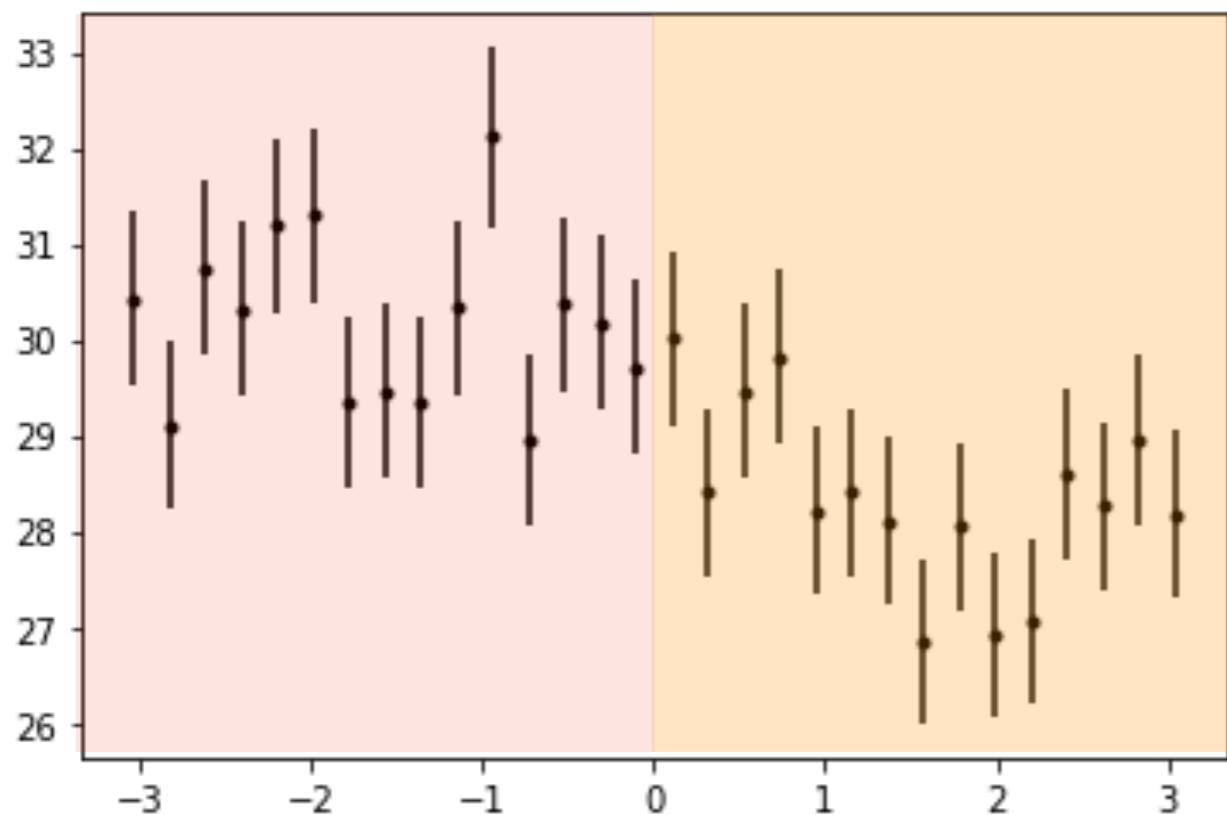


Challenge: Try to make this plot

How do we quantify asymmetry?

Highest Energy > 8 EeV

First lets just make 2 bins



Given an underlying poisson distribution for left and right what is the likelihood of each?

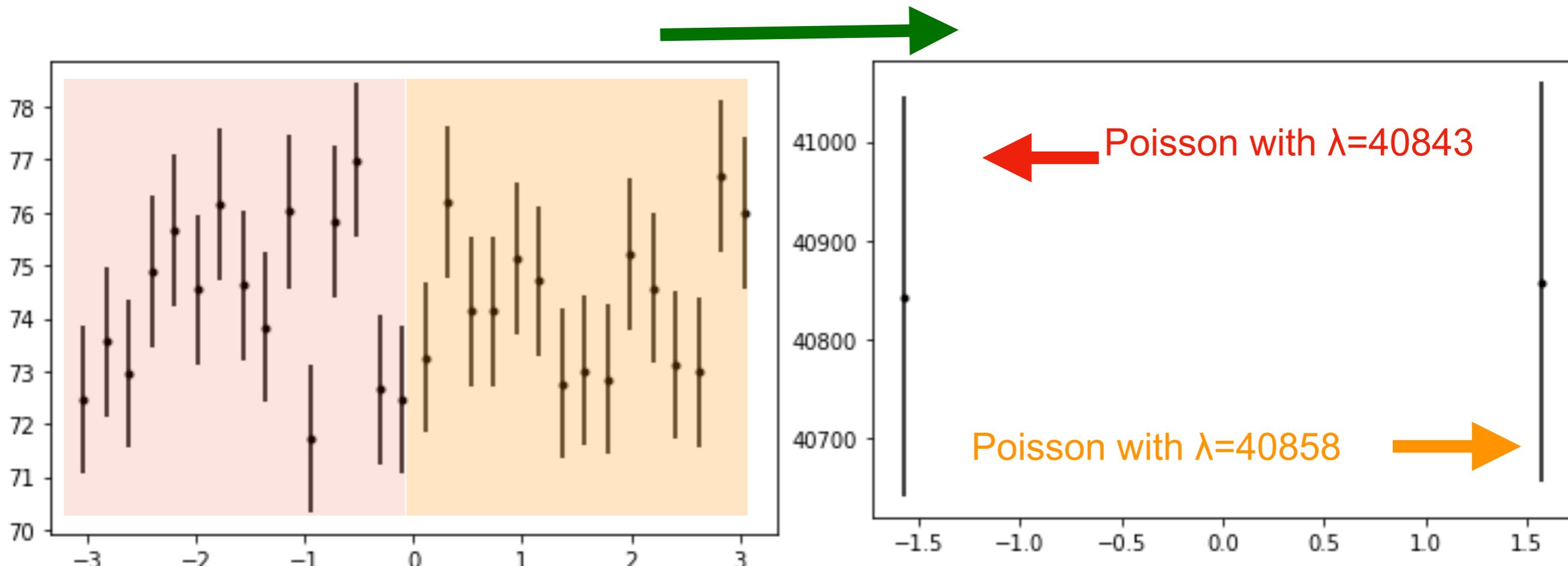
$$p(x = 16600 | \lambda = 16100) / p(x = 16600 | \lambda = 16600) = 10^{-3}$$

1/1000 times unlikely to fluctuate up and down like this

How do we quantify asymmetry?

Lower Energy 4-8 EeV

First lets just make 2 bins



Given an underlying poisson distribution for left and right what is the likelihood of each?

$$p = 0.999$$

Why would it change with energy?

What is the likelihood for a gaussian over N bins?²¹

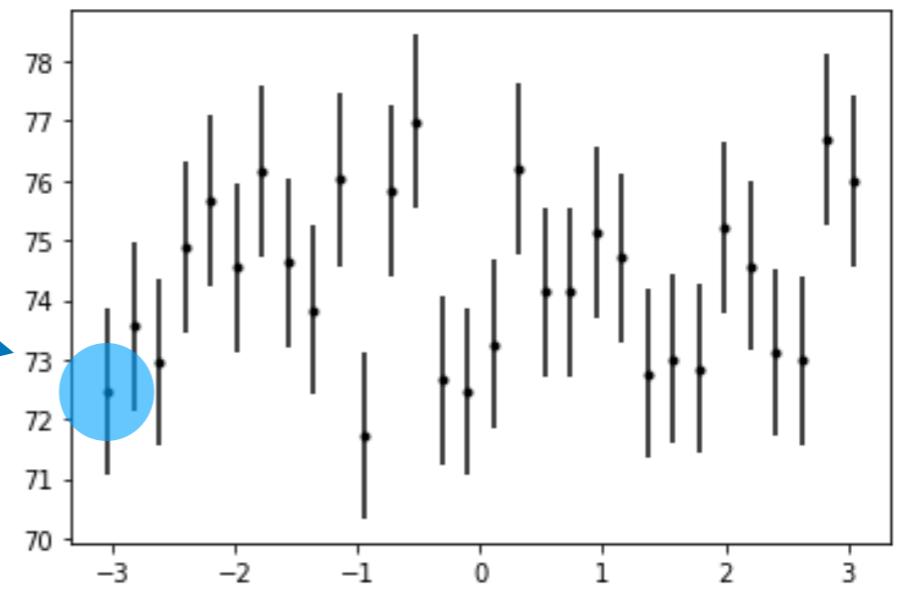
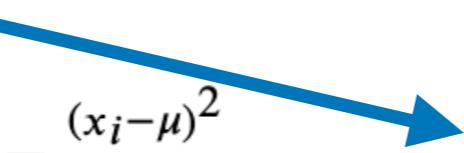
$$L(x|\lambda) = \prod_{i=1}^N p(x_i|\lambda)$$

$$L(x|\lambda) = \prod_{i=1}^N \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(x_i-\mu)^2}{2\sigma^2}}$$

$$-\log(L(x|\lambda)) = \sum_{i=1}^N -\frac{1}{2}\log(2\pi\sigma^2) + \frac{(x_i - \mu)^2}{2\sigma^2}$$

$$L(x|\lambda) = \prod_{i=1}^N \frac{1}{\sqrt{2\pi\lambda}} e^{-\frac{(x_i-\lambda)^2}{2\lambda}}$$

$$-\log(L(x|\lambda)) = \sum_{i=1}^N -\frac{1}{2}\log(2\pi\lambda) + \frac{(x_i - \lambda)^2}{2\lambda}$$



Using the Poisson variance

What is the likelihood for a gaussian over N bins?²²

$$f(x|\lambda) = -\frac{1}{2}\log(2\pi\sigma^2) + \frac{(x_i - \mu)^2}{2\sigma^2}$$

$$E[f(x)] = -\frac{1}{2}\log(2\pi\sigma^2) + E[g(x)]$$

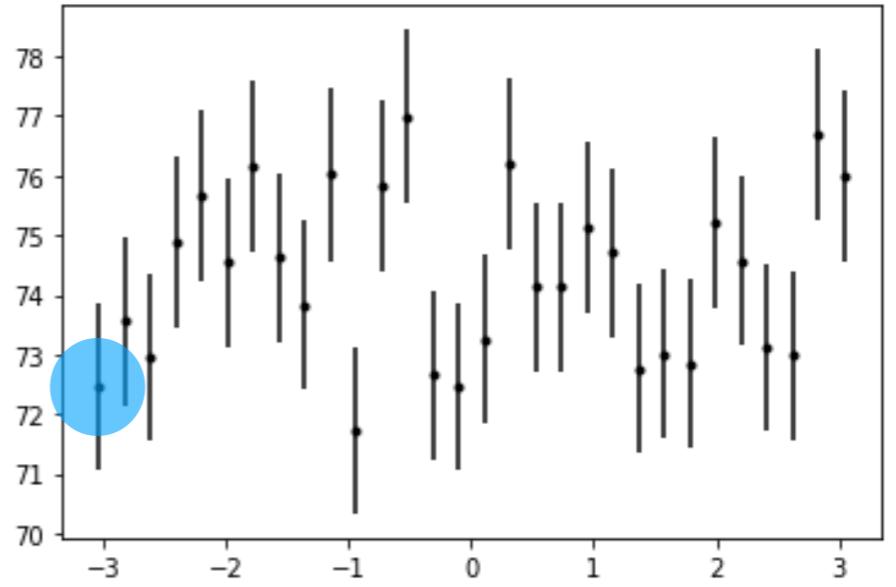
$$g(x) = \frac{(x_i - \mu)^2}{2\sigma^2}$$

$$E[g(x)] = \frac{\text{Var}(x)}{2\sigma^2}$$

$$E[g(x)] = \frac{\sigma^2}{2\sigma^2} = 1/2$$

$$g(x) = \sum_{i=1}^N \frac{(x_i - \mu)^2}{\sigma^2}$$

$$E[g(x)] = N$$



Expectation of log likelihood using a gaussian is $N/2$

N is the number of bins

χ^2 distribution

- χ^2 distribution is the sum of N independent variables X_i
 - Where the distributions X_i are distributed as normal
 - N denotes the number of degrees of freedom

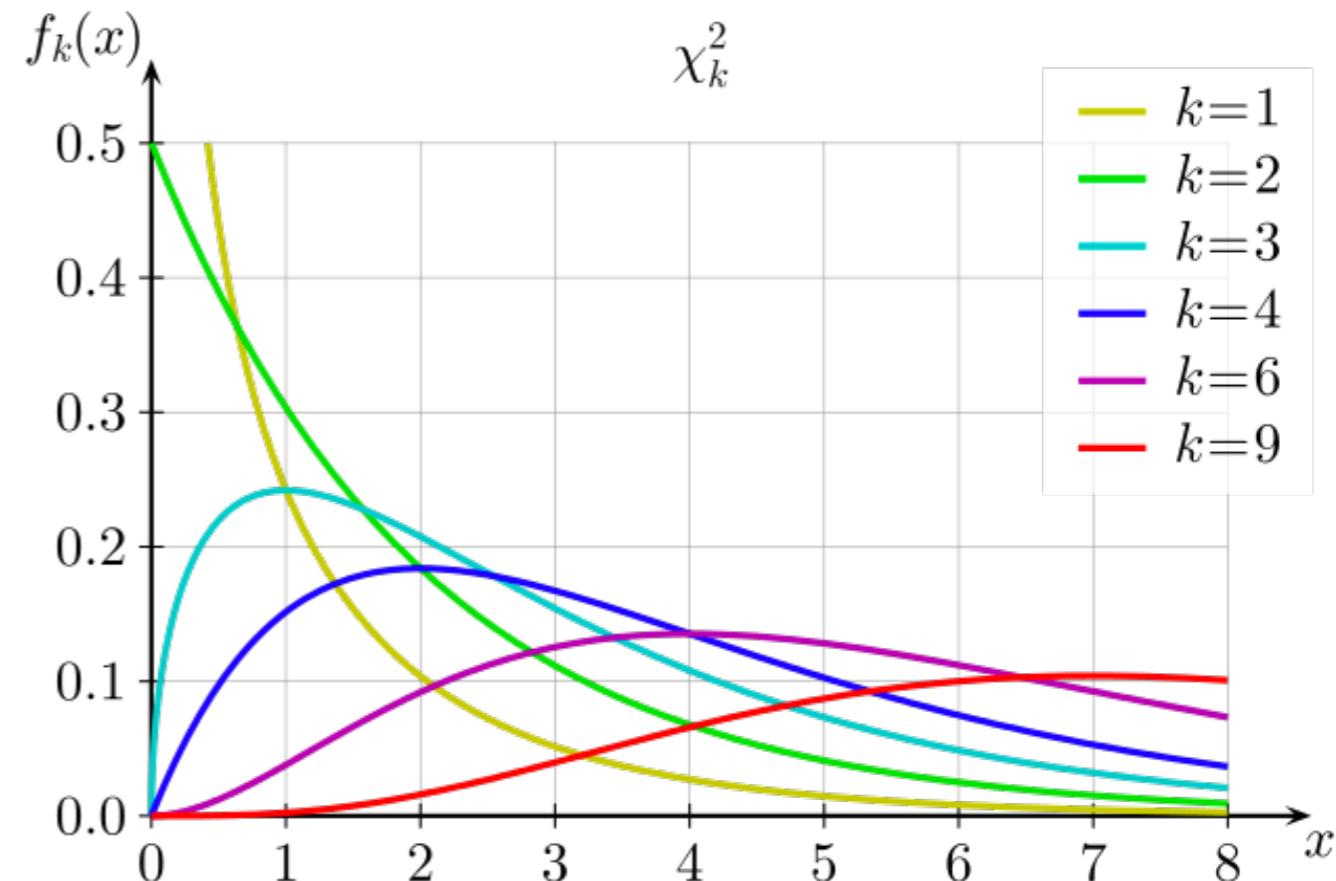
$$\chi^2(x, n) = \sum_{i=1}^N \frac{(x - \mu)^2}{\sigma^2}$$

$$E[\chi^2(x)] \approx N$$

$$E[\chi^2(x)/N] \approx 1$$

$$\text{Var}(\chi^2(x)) = 2N$$

$$\Delta\chi^2(x) = |\chi^2(x) - \chi^2(x \pm \sqrt{2N})|$$



Important Properties

- Taylor expand in our floated parameter (μ in this case)

$$\chi^2(x_i, \mu) = \boxed{\chi^2_{min}(x_i, \mu_0) + \frac{1}{2} \frac{\partial^2}{\partial \mu^2} \chi^2_{min}(x_i, \mu_0) (\mu - \mu_0)^2}$$

$$\frac{1}{2} \frac{d^2 \chi^2}{d \mu^2} \rightarrow \frac{1}{\sigma^2}$$

χ^2 distribution of 1 degree of freedom
 $V[\chi^2(x)] = 1$

$$\Delta \chi^2 = 2 \Delta \log L = 1$$

For one degree of freedom

$$\frac{\partial^2 \chi^2}{\partial \theta^2} = \frac{2}{\sigma_\theta^2}$$

For any floated parameter
uncertainty of that parameter is
given by the 2nd derivative of χ^2

This is known as Wilk's Theorem →

$$\sigma_\theta^2 = \left(\frac{\partial^2 \log L}{\partial \theta^2} \right)^{-1}$$

Important Properties

- Taylor expand in our floated parameter (μ in this case)

$$\chi^2(x_i, \mu) = \boxed{\chi_{min}^2(x_i, \mu_0) + \frac{1}{2} \frac{\partial^2}{\partial \mu^2} \chi_{min}^2(x_i, \mu_0)(\mu - \mu_0)^2}$$

$$\Delta \chi^2 = \frac{1}{2} = \sum_{i=1}^N \frac{(x - \mu_0 + \sigma_\mu)^2}{2\mu_0} - \sum_{i=1}^N \frac{(x - \mu_0)^2}{2\mu_0} \quad \begin{matrix} \chi^2 \text{ distribution of 1 degree of freedom} \\ V[\chi^2(x)] = 1 \end{matrix}$$

$$1 \approx \sum_{i=1}^N \frac{(x - \mu_0 + \sigma_\mu)^2}{\mu_0} - \sum_{i=1}^N \frac{(x - \mu_0)^2}{\mu_0}$$

$$1 = \frac{1}{\mu_0} \sum_{i=1}^N (x - \mu_0 + \sigma_\mu)^2 - (x - \mu_0)^2$$

$$1 = \frac{1}{\mu_0} \sum_{i=1}^N \sigma_\mu^2 + 2\sigma_\mu(x - \mu_0)$$

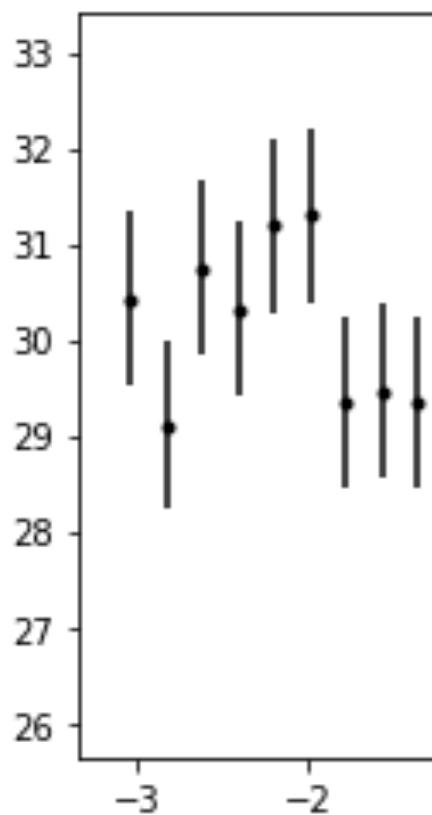
$$1 = \frac{N\sigma_\mu^2}{\mu_0}$$

$$\sigma_\mu^2 = \frac{\mu_0}{N} \quad \begin{matrix} \leftarrow \\ \text{For a poisson distribution we recover the variance per bin} \end{matrix}$$

How do we quantify asymmetry?

First lets just make 2 bins

$$L(x|\lambda) = \prod_{i=1}^N p(x_i|\lambda)$$

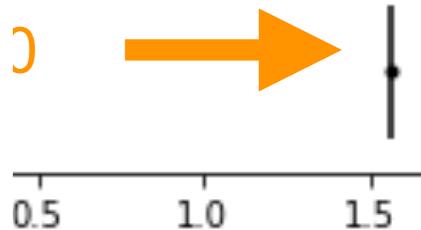


$$L(x|\lambda) = \prod_{i=1}^N \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(x_i-\mu)^2}{2\sigma^2}}$$

$\lambda=16600$

$$-\log(L(x|\lambda)) = \sum_{i=1}^N -\frac{1}{2}\log(2\pi\sigma^2) + \frac{(x_i - \mu)^2}{2\sigma^2}$$

$$L(x|\lambda) = \prod_{i=1}^N \frac{1}{\sqrt{2\pi\lambda}} e^{\frac{-(x_i-\lambda)^2}{2\lambda}}$$



Given an unde

$$p(x = -3) = -\log(L(x|\lambda)) = \sum_{i=1}^N -\frac{1}{2}\log(2\pi\lambda) + \frac{(x_i - \lambda)^2}{2\lambda}$$

od of each?

1/1000 times unlikely to fluctuate up and down like this

We also find

